



# AWI-ICENet1: A convolutional neural network retracker for ice altimetry

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**Abstract.** The Greenlandic and Antarctic Ice Sheet are important indicators of climate change and major contributors to sea level rise. Hence, precise, long-term observations of surface elevation change are required to assess changes and their contribution to sea level rise. Satellite radar altimetry has been used by various missions to measure surface elevation change since 1992. It has been shown that, next to the surface slope and complex topography, one of the most challenging issues is the spatial and temporal variability of radar pulse penetration into the snow pack, especially over the vast East Antarctic plateau. This results in an inaccurate measurement of the true surface elevation and consequently affects surface elevation change (SEC) estimates. To increase the accuracy of surface elevations retrieved by retracking the radar return waveform and thus reduce the uncertainty in SEC, we developed a deep convolutional neural network architecture (AWI-ICENet1). The AWI-ICENet1 is trained using a simulated reference data set with 3.8 million waveforms, taking into account different surface slopes, topography, and attenuation. The successfully trained network is finally applied as AWI-ICENet1-retracker to the full time series of CryoSat-2 Low Resolution Mode (LRM) waveforms over both ice sheets. We compare the AWI-ICENet1 retrieved SEC with estimates of conventional retrackers like TFMRA and ESA ICE1 and ESA ICE2 products. Our results show less uncertainty and a greatly diminished effect of time variable radar penetration, reducing the need to apply corrections based on a close relationship with backscatter- and/or leading edge width, as typically done in SEC processing. This technique provides new opportunities to utilize convolutional neural networks in altimetry, waveform retracking, and processing of satellite altimetry data, which can be applied to historical, recent, and future missions.

## 1 Introduction

Ice sheet mass loss is a major contributor to sea level rise, as already the Greenland Ice Sheet (GrIS) has contributed  $21.0 \pm 1.9$  mm (Otosaka et al., 2023) between 1992-2022 and Antarctica  $7.6 \pm 3.9$  mm 1992-2017 (Shepherd et al., 2018). The non-linearity of mass loss from Antarctica is driven by West Antarctica, where glacier acceleration and retreat has caused an increasing contribution from 2010 onwards. The stability of the West Antarctic Ice Sheet (WAIS) is a major concern, as even a single glacier, Thwaites Glacier, has the potential to contribute 65 cm alone to sea level rise (Scambos et al., 2017). This exemplifies



the need for accurate observation of ice sheet mass loss. Due to the self gravitation effect, local sea level rise along the world's coastlines strongly depends on the spatial distribution of ice sheet mass loss (Larour et al., 2017). This is where the advantage of altimeters comes into play. Altimeter data can provide ice sheet mass loss in a high spatial resolution compared to gravimeters. The accuracy of the altimeter based mass loss products is defined by the accuracy with which the individual elevation measurement is conducted, by the repeat cycle, the spatial interpolation scheme, and the conversion from volume change to mass change including firn densification. With this study, we focus on the improvement on the accuracy of the individual elevation measurement.

Satellite altimeters have been surveying ice sheets since the early 1990's in a substantial coverage, beginning with esa's ERS-1 and ERS-2 (both Ku-band), followed by the ENVISAT 2002-2012 (Ku-band). With the launch of CryoSat-2 (Ku-band) in 2010 the first altimeter dedicated to study the Earth's cryosphere went operational. CryoSat-2 orbits at an unusually high inclination, reaching latitudes of 88°North and South. In addition it came along with major improvements to measure icy surfaces. Next to the conventional Low Resolution Mode (LRM), a Synthetic Aperture Mode (SAR) mode was accomplished to increase the spatial resolution in along track direction. To be able to locate the point of closest approach in sloped terrain in across track direction a second antenna is mounted enabling to measure the phase difference in the so called SAR interferometer (SARIn) mode. This allows to estimate the angle of arrival and thus to relocate the ground return. However, in this study we will focus on CryoSat-2 data in LRM mode. The first spaceborne radar altimeter in Ka-band is SARAL/ALTIKA launched in 2013. The latest radar altimeter is the pair of Sentinel-3 (S3A/S3B) that have been operating since 2016/2018 and both have a Ku-band altimeter onboard that operate in SAR mode. In addition Pseudo-LRM (PLRM) products are provided for Sentinel-3. Next to the radar altimeters, two laser altimeters surveyed polar areas: NASA's ICESat-1 operated from 2003-09 and since 2018 ICESat-2 is operational. The big advantage of laser altimetry is the high precision of a single range measurement as it is not prone to penetration issues over snow and ice surfaces. In addition, due to its small footprint size and high pointing precision the spatial assignment of the observations is high, even in areas of large slope and complex terrain. The disadvantage of laser altimetry is cloud cover, which prevents surface measurements and leads to data gaps. Distorted elevation measurements due to snow drift also affect accuracy. Because we use a three-year measurement period and all six available laser beams, data gaps are not a problem. The advantages outweigh the disadvantages, and therefore, we use elevation change estimates based on ICESat-2 data to compare our results. As dry ice and snow are transparent for radar waves, the penetration of the radar signal complicates the detection of the true surface. In general the returned power over ice sheets consists of surface and volume components: surface scattering at the air-snow interface, scattering from internal layers, and volume scattering from scattering at snow grains. While the onset of surface scattering leads to a sharp rise in power (often called leading edge), volume scattering affects the gentle decline in power over time (the trailing slope (TSL)). In general the waveform shape of a return echo mainly depends on the local surface topography in the area of the radar footprint and the small scale surface roughness, meaning the roughness on scales of the radar wavelength. This leads to distinct differences between waveforms from rough terrains and waveforms from smooth surface topography. On top of this the volume scattering acts as kind of low pass filter, leading to a widening of the waveform while adding more energy to the waveform tail but also enlarging the leading edge width (LEW), especially for conventional LRM waveforms. Volume scattering is caused by scattering of the radar wave that penetrates into



the ice sheet, at snow and firn grains. This depends on the size of the grains, while the absorption loss of the radar wave is mainly governed by temperature. As a result the volume scattering varies widely over ice sheets due to differences in the snow and firn properties. As a consequence, the penetration of the radar signal in Ku-band can lead to a bias in elevation detection in the order of 10-20 cm (Larue et al., 2021) or even more, depending on the retracking method used to measure the range between the antenna mounted on the satellite and the surface. It has been shown by e.g. Davis (1997); Helm et al. (2014); Nilsson et al. (2016) that the choice of the retracker which aims to retrack the range at the lower part of the leading edge, like the Threshold Centre Of Gravity (TCOG) or Threshold First Maxima Retracker Algorithm (TFMRA), can strongly suppress the radar penetration bias. However, a remaining signal contribution is still left, which partly can be corrected using additional waveform parameters such as the LEW or TSL and/or the radar backscatter as proposed by e.g. Flament and Rémy (2012); Simonsen and Sørensen (2017); Schröder et al. (2019). In this study we aim to reduce this penetration bias directly in the waveform retracking by employing a machine learning approach.

Machine learning and in particular Deep Learning (DL) offers a data-driven alternative to traditional physical/statistical approaches for modeling the functional relationship between measurements and target variables. It is particularly successful in cases where the true relationship between input and output is too complex to be approximated by traditional models, which often only capture a part of the full spectrum or make simplifying assumptions to be tractable. Convolutional neural Networks (ConvNets), first proposed for more general computer vision applications such as image classification (LeCun et al., 2015), are a class of DL architectures that have shown tremendous success in a broad variety of Earth observation applications including land cover/use classification (e.g. from multispectral time series (Campos-Taberner et al., 2020)), image processing (e.g. speckle reduction of synthetic aperture radar images (Dalsasso et al., 2022)), and estimation of geo-/bio-physical parameters (e.g. forest height estimation (Lang et al., 2022)).

Machine learning has been applied in various other studies for waveform analysis: Müller et al. (2017) analyses altimetry data in the Arctic to detect open water within sea ice cover using an unsupervised clustering (i.e. k-medoids) of the radar echos to subdivide the waveforms based on different characteristics and subsequently classifies them via k-Nearest-Neighbors. Lee et al. (2016) use Random Forests to detect cracks between ice floes to improve the estimation of sea ice thickness. Random Forests are also used by Shen et al. (2017b, a) to classify sea ice type based on waveform data. These studies focus on the classification of the waveforms to detect different surface types. The regression task to accurately estimate surface elevation has been barely addressed. Fayad et al. (2021) used DL for the detection of surface heights from space-borne laser altimeter data of the GEDI mission (Dubayah et al., 2020).

We use single waveforms of CryoSat-2 and represent them as sequential data to a 1D ConvNet that applies a series of processing layers (in particular convolutions with learned kernels along the time dimension of the waveform) to automatically extract features and agglomerate information. The output of the network is the retracked range that corresponds to the snow/firn surface. In order to engage supervised machine learning for processing of the altimeter waveforms, a large data set with known range is necessary. This cannot be achieved by using airborne or ground-borne sounding with radar altimeters synchronously with laser scanners, as the footprint of the satellite is orders of magnitude larger than those of airborne measurements. Space borne laser altimetry as ICESat-2 to be used as test data set in a DL approach to improve radar derived elevation measurements



is in our opinion also not applicable. Reason is the very different footprint size of the two systems. Whereas the ICESat-2 laser points to areas of less than  $0.02 \text{ km}^2$ , radar altimeter systems illuminate large areas of up to  $10 \text{ km}^2$ , thus both are not spatially assigned and cannot directly compared to each other. Large scale topographic undulation and surface slope involve a slope correction in the post processing to reposition the radar elevation measurement to its point of closest approach. As this correction cannot be extracted from the waveform shape itself a direct comparison between laser and radar in terms of a DL approach is not possible. Instead, we make use of simulated waveforms to create a large reference data set used for training and testing of the new ConvNet retracker. To represent the satellite altimeter waveforms as best as possible, we take the local surface topography over the ice sheet into account. This way, we can create basically an infinite number of training samples to learn a neural network. In a second step we apply the new ConvNet retracker to measured CryoSat-2 waveforms and derive elevation and elevation change estimates, which we compare to ICESat-2 derived elevation change data products. The remainder of this paper is structured as follows: In Section 2 we first present our approach to simulate waveforms, followed by presenting the used ConvNet, provide a brief overview of other retrackers and summarize the used satellite altimeter data and our approach to estimate rates of elevation change. Section 3 shows the performance of the new AWI-ICENet1 on simulated waveforms before we present results on real satellite altimeter data over selected sites and ice sheet wide. In Section 4 we first discuss the results achieved with AWI-ICENet1 on simulated data before we compare it to other retrackers. Finally, the estimates of the rates of elevation change are evaluated and compared with those of ICESat-2.

## 2 Material and Methods

### 2.1 Simulated waveforms

For the success of our approach to develop a DL retracker capable of minimizing the effects of variations in backscatter and radar speckle on range measurements, a good reference dataset is essential. We have defined the following criteria for our reference data set that must be met by the simulated waveforms:

- represent observed CryoSat-2 LRM waveforms as best as possible
- large variability of waveforms shapes
- large number of waveforms for optimal training results

To meet the criteria and to ensure a good training result, several features must be included in the simulation:

- representation of different real topographic situations
- allow different bulk attenuation rates (to mimic volume scattering component)
- add Gaussian noise to each waveform (to mimic radar speckle)
- add noise floor prior the leading edge



- repositioning of waveform within range window to cope tracker gate variations
- include CryoSat-2 system characteristics (Ku-Band frequency (wavelength), antenna gain pattern, range resolution)

We randomly selected 1000 locations within the LRM zone limits of Antarctica as shown in Fig. 1. At each location a  
 125 reference surface echo  $P_{r_s}$  is simulated, using the classical radar equation integrated over the illuminated surface area (Eq. 1, Brown (1977)):

$$P_{r_s}(t) = \frac{P_t \lambda^2}{(4\pi)^3} \int_A \frac{\sigma_S^0(\theta) G(\theta)}{R^4} dA \quad (1)$$

with wavelength  $\lambda = c/f_c$ , speed of light  $c$ , Ku-Band centre frequency  $f_c = 13.5$  GHz, the transmit power  $P_t$ , the backscatter cross-section  $\sigma_S^0(\theta)$  and  $R$  the range from the radar to the surface element  $dA$ .  $\sigma_S^0(\theta)$  is chosen to be homogeneous and the antenna gain pattern  $G(\theta)$  is defined as an elliptical 2D Gaussian function:

$$130 \quad G_{\theta_{al}\theta_{ac}} = G_0 \exp\left(-\left[\frac{\theta_{al}^2}{\beta_{al}^2} + \frac{\theta_{ac}^2}{\beta_{ac}^2}\right]\right) \quad (2)$$

with the look angles  $\theta_{ac} = \tan(R_x(x,y)/R_z(x,y))$  and  $\theta_{al} = \tan(R_y(x,y)/R_z(x,y))$  and the constants  $\beta_{al} = \theta_{al_{3db}}/\sqrt{4\log(2)}$   
 and  $\beta_{ac} = \theta_{ac_{3db}}/\sqrt{4\log(2)}$ . Centered at each location  $(x,y)$ ,  $R_z(x,y) = h_{\text{sat}} - h_s(x,y)$  is estimated by using a mean satellite  
 altitude of  $h_{\text{sat}} = 730$  km and  $h_s(x,y)$  is derived from an high resolution interpolated sub grid of the REMA elevation model  
 (Howat et al., 2019). Each 2D-range pattern  $R(x,y) = \sqrt{R_x(x,y)^2 + R_y(x,y)^2 + R_z(x,y)^2}$  covers the radar footprint and has  
 135 dimensions of  $30 \text{ km} \times 30 \text{ km}$  re-sampled to  $20 \text{ m}$  pixel resolution. For the elliptical Gaussian 2D-antenna pattern we use an  
 antenna beam width of  $\theta_{al_{3db}} = 1.3^\circ$  and  $\theta_{ac_{3db}} = 1.15^\circ$  as given by Wingham et al. (2006). To mimic tracker gate variations  
 and to force the network not to retrack always at the same incremental position we fractional shifted  $P_{r_s}$  randomly within the  
 range gate while updating the reference range epoch.

In the next step we added different volume contributions following the approach of Legrésy and Rémy (1997) by using an  
 140 exponential decay function (Eq. 3)

$$V(z) = \exp(-r(z)L_A/4.342945) \quad (3)$$

and a homogeneous layered model consisting of  $N = 128$  layers with a depth resolution of  $\Delta R_z = 0.468$  m, which correspond  
 to the range bin size and the number of range bins per observed CryoSat-2 LRM waveform. For simplicity the loss factor ( $L_A$ )  
 accounts for loss contributions due to volume scattering, absorption and stratification of the snowpack in a combined manner.  
 This loss factor, called bulk attenuation raten in the fowwlowing, is adjustable to be able to vary the volume contribution. The  
 145 final received power  $P_{r_v}$ , including surface and volume contributions, is given by (Eq. 4)

$$P_{r_v} = \sum_{i=rt_{ref}}^{N-1-rt_{ref}} P_{r_s} V(z_{i-rt_{ref}}). \quad (4)$$

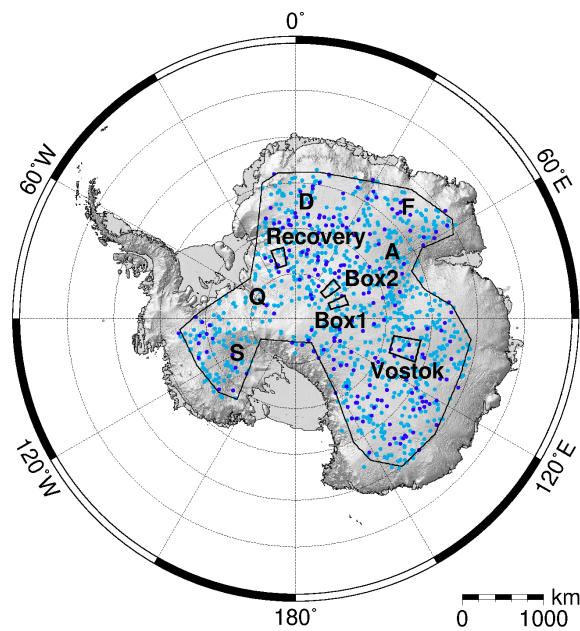


Although the model architecture is simple, it represents the general effect of absorption and scattering leading to an attenuation of the radar wave within snow/ice due to different physical properties like temperature, surface density grain size and layering as explained by Lacroix et al. (2008); Adodo et al. (2018).

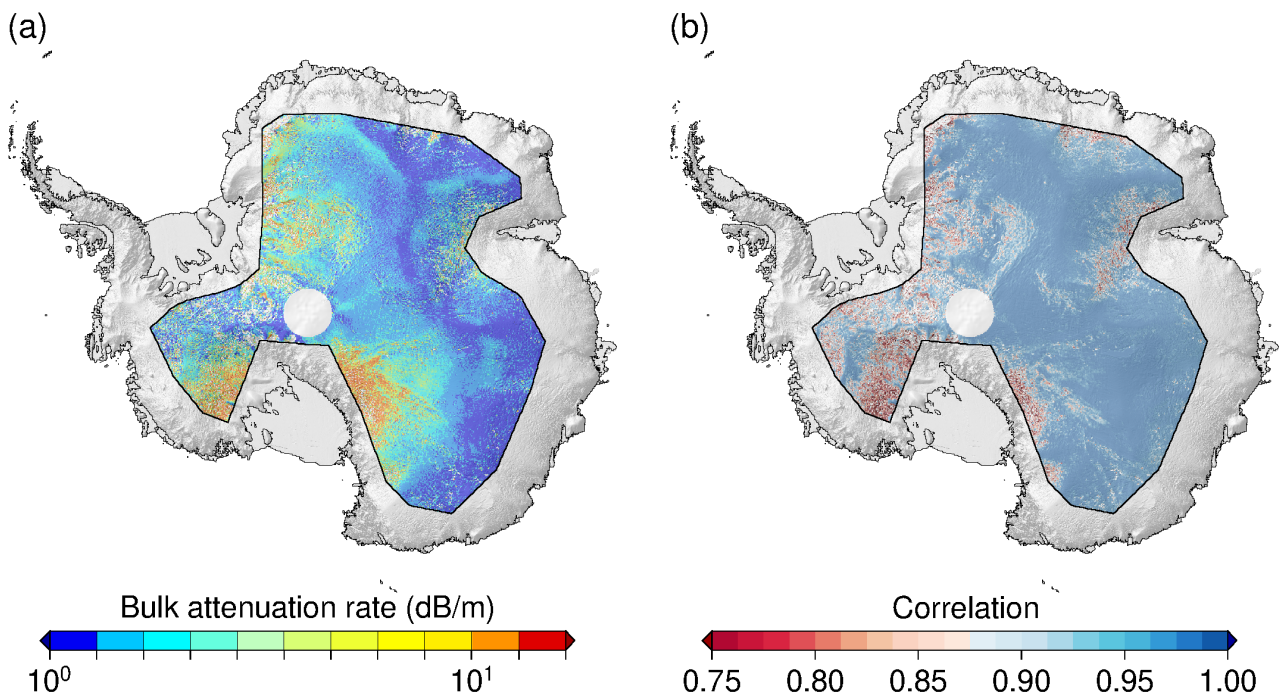
In a simplified model performance test, in which a simulated waveform of a flat surface was fitted to observed CryoSat-2 LRM waveforms only by adjusting  $L_A$  in a minimum least square estimator (MLE), we found very high correlations between measured and fitted waveforms of 0.9 and higher. The test results are shown in Fig. 2 and Fig. A1 for Antarctica and Greenland, where the median of the fitted attenuation rate and the median of the correlation between fitted and observed model are shown on a grid of  $5 \text{ km} \times 5 \text{ km}$  pixel resolution. Based on the findings in Fig. 2(a) we generated 95 reference waveforms at each location using an attenuation rate between 1 dB and 20 dB with a step size of 0.2 dB. Finally, we generated 40 different noisy samples of each of the reference waveforms by adding 40 different randomly generated samples of Gaussian noise. We thereby increased the training data set to 3.8 Mio noisy waveforms ( $m = 1000 \times 95 \times 40$ ).

$$\Psi_i = P_{r_v} * \epsilon_s + \epsilon_f \quad (5)$$

while  $\epsilon_s$  represents radar speckle and  $\epsilon_f$  a Gaussian noise floor. Both were selected to match the noise of CryoSat-2 waveforms. We denote the final noisy waveforms hereafter referred to  $(\Psi_i)_{i \in I}$ , where  $I = \{1, \dots, m\}$  is the index set of the waveforms. Each waveform consists of 128 range bins, matching the data structure of the CryoSat-2 satellite altimetry waveform product. For each of the noisy waveforms the actual true surface or reference range is known and is used as reference for the ConvNet training. A typical example of the different simulation steps is shown in Figure 3, where the observed CryoSat-2 waveform at the same geographical position (Lake Vostok) is overlaid in blue.  $L_A$  was estimated by a MLE to be 1.5 dB/m. More example waveforms simulated in different topographic environments are shown in the Appendix in Figures A2, A3, A4 .

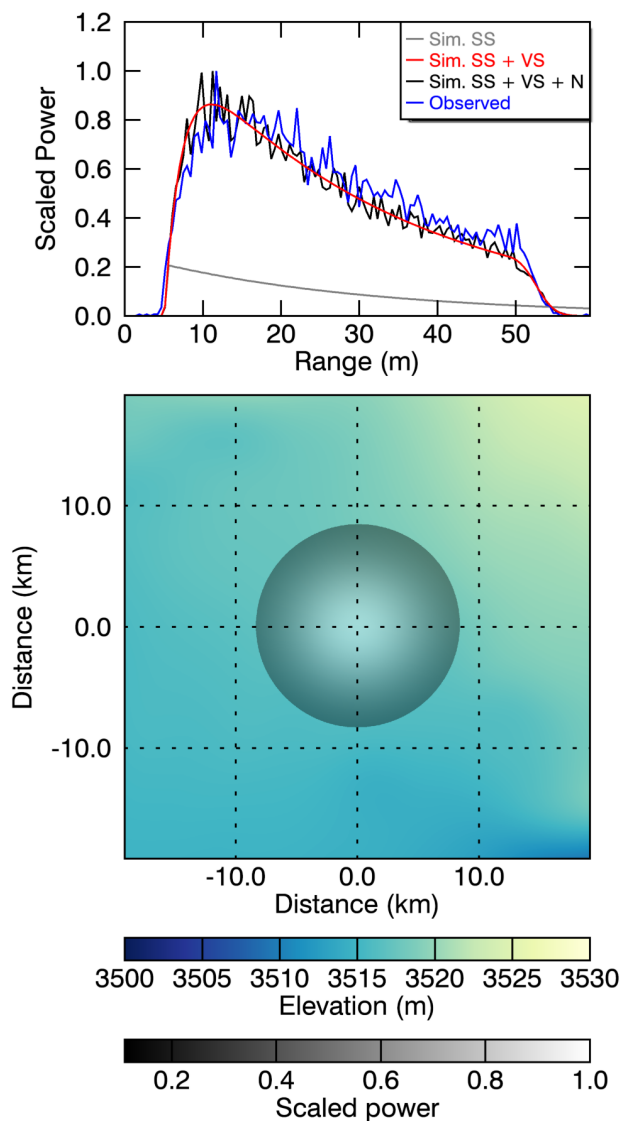


**Figure 1.** Overview of the four selected regions used in the analysis. The entirety of the points show the randomly selected locations where waveforms are simulated. Light blue dots show the 80% used for training and dark blue the 20% used for testing the ConvNet.

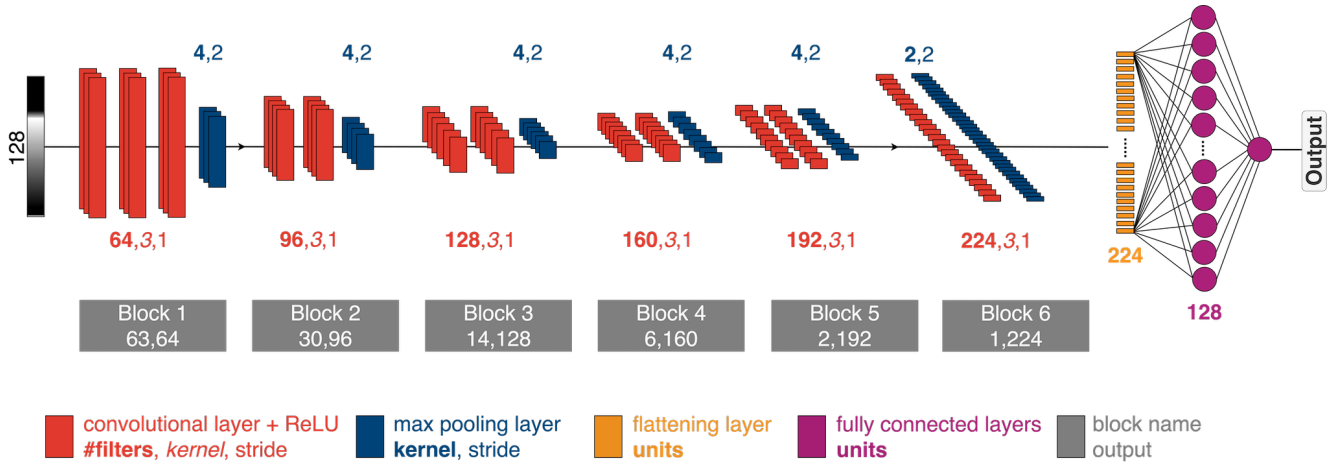


**Figure 2.** Results of the test where a flat surface waveform model is fitted to real CryoSat-2 waveforms by adjusting the attenuation  $L_A$  as the only parameter. (a) Gridded median of the attenuation rate estimated by a MLE fit and (b) median of the correlation between observed and fitted waveforms.





**Figure 3.** Example of the different simulation steps shown for a typical waveform over Lake Vostok. Blue line denotes the observed CryoSat-2 LRM waveform and the gray line the simulated surface waveform (SS) based on Eq. 1. The red line represents the simulated waveform including volume scattering (SS+VS) according to Eq. 3, while the black line depicts the final simulated waveform including surface and volume scattering as well as noise (SS+VS+N) as given in Eq.5. The lower panel represents the 2D elevation model and scaled  $P_{r_s}$ , which is mainly controlled by the Gaussian antenna pattern.



**Figure 4.** A ConvNet is used to map an individual waveform to the corresponding range that correlates with the actual surface. We use standard combinations of convolutions, ReLU, batch-norm, max-pooling, and drop-out layers. The final feature layer is flattened and used as input to a fully-connected layer.

## 2.2 AWI-ICENet1

165 We use a ConvNet  $f$  to model the relationship between a (real or simulated) waveform  $\Psi$  and the corresponding retracked range  $r$ , i.e. the network provides an estimate  $\hat{r} = f(\Psi)$  of the retracked range, which corresponds to the true surface, based on the provided input data. During training, only simulated waveforms are used. The network is then applied to real waveforms during inference.

170 The used ConvNet consists of six blocks of stacked convolutional layers with ReLU activation functions, batch-normalization, maximum pooling and drop-out (with overlapping pooling windows similar to Krizhevsky et al. (2012)). The final feature layer is flattened and processed by a fully connected layer with 128 units. Since the waveforms are 1D input data, all layers and operations are one-dimensional as well. Fig. 4 shows an overview of the network architecture.

The cost function  $\mathcal{L}$  is a standard mean squared error (MSE) loss between the true and estimated range of the simulated waveform, i.e.

$$175 \quad MSE = \frac{1}{|\mathcal{B}|} \sum_{(r, \Psi) \in \mathcal{B}} (r - f(\Psi))^2, \quad (6)$$

where  $\mathcal{B} \subset \mathcal{D}_{Tr}$  is the batch, i.e. a subset of the training set  $\mathcal{D}_{Tr}$ .

The network is implemented in tensorflow-gpu 2.6.2 / keras 2.6.0 and trained for 25 epochs with a batch size of 128 via an Adam optimiser (Kingma and Ba, 2014) with L2 regularization.

180 We split the available training data into five different folds and perform five independent runs with training data  $\mathcal{D}_{Tr}$  consisting of four folds, while the fifth fold serves as test data  $\mathcal{D}_{Te}$ . Performance is averaged over these five runs. Model performance



is assessed via MSE

$$MSE = \frac{1}{|\mathcal{D}_{Te}|} \sum_{(r, \Psi) \in \mathcal{D}_{Te}} (r - f(\Psi))^2, \quad (7)$$

the mean absolute error (MAE)

$$MAE = \frac{1}{|\mathcal{D}_{Te}|} \sum_{(r, \Psi) \in \mathcal{D}_{Te}} |r - f(\Psi)|, \quad (8)$$

185 and the root mean squared error (RMSE)

$$RMSE = \sqrt{\frac{1}{|\mathcal{D}_{Te}|} \sum_{(r, \Psi) \in \mathcal{D}_{Te}} |r - f(\Psi)|^2}. \quad (9)$$

### 2.3 Satellite altimetry data

In this study we use CryoSat-2 Level 1B (including measured waveforms) and Level 2I (including range estimates by ESA retracker) products of the Baseline D/E low resolution mode (LRM) provided by the European Space Agency (ESA). We make use of the ICE1 and ICE2 retracker solutions given in the Level 2I product and apply according to Helm et al. (2014) the TFMRA and the newly developed AWI-ICENet1 retracker to the Level 1B waveform product. ICE1 is based on a threshold centre of gravity retracker (TCOG) (Wingham et al., 1986; Davis, 1997), while ICE2 the UCL land-ice retracker fits a Brown model (Brown, 1977) adapted for CryoSat-2. To analyse the performance of the AWI-ICENet1 in comparison to the other three retracker, we make use of the ATL06.005 ICESat2 data product provided by NASA. We use data from all six beams within the time span 2019/01 to 2021/12. Instead of using the quality flag given in the ATL06 product we filter the data based on the version 2 of the REMA Antarctic elevation model Howat et al. (2022). All data points with a difference larger than  $\pm 50$  m are excluded from further processing.

### 2.4 Elevation change and empirical corrections for time-variable radar-penetration effects

Surface elevation change (SEC) processing as applied by various groups use different strategies to minimize the effect of radar penetration. In most cases the backscatter and/or additional waveform shape parameters such as the leading edge width (LEW) and the trailing edge slope (TES), estimated by the ICE2 retracker introduced by Legresy et al. (2005); Frappart et al. (2016), are used. Those waveform parameters are not provided by other retracker such as the offset centre of gravity (OCOG by Wingham et al. (1986)), a threshold centre of gravity (TCOG by Davis (1997)) or the threshold first maxima retracker algorithm (TFMRA by Helm et al. (2014)). Here, we make use of the LEW and backscatter provided by the ESA LRM-Level 2I ICE2 and ICE1, respectively. The decision to use the backscatter of ICE1 for all retrackers is due to the lower sensitivity of the OCOG amplitude to speckle noise, so that the backscatter of successive waveforms is less noisy, but large-scale and time-dependent fluctuations are preserved. After retracking, the geo-referenced surface elevation is determined for each of the retracking approaches using orbital information such as altitude, latitude and longitude and additional geophysical correction included in the ESA products. In addition, the refined slope correction (Roemer et al., 2007) is applied to relocate the echo to its point of closest approach.



210 This results in a large point cloud of geo-referenced elevation measurements for each of the retracker. The interpolated elevation  
anomaly product and rates of elevation change (dhdt) are generated using a similar approach as described in e.g. McMillan  
et al. (2016); Schröder et al. (2019); Nilsson et al. (2022). We collect all geo-referenced data points within a pixel size of 2 km  
and correct for topographic effects at the pixel scale using the REMA-DEM, rather than fitting any kind of linear or quadratic  
surface as McMillan et al. (2016); Schröder et al. (2019); Nilsson et al. (2022). We then estimate rates of elevation change  
215 using a linear regression for each pixel with sufficient data coverage, again without using additional information such as LEW,  
TeS, backscatter or seasonal components. The residuals are averaged to monthly residuals per pixel. Both gridded products,  
the trend and the monthly residual grids, are finally interpolated using inverse distance weighting with variable radius to form  
the final interpolated grids with 5 km posting. Furthermore, the backscatter and LEW information are processed in the same  
way without any trend or topographic correction. Instead of applying the LEW and/or backscatter correction in the context  
220 of multi parameter fitting, as in e.g. Flament and Rémy (2012); Simonsen and Sørensen (2017); Schröder et al. (2019), we  
apply the correction adapted from Nilsson et al. (2022) to our final interpolated monthly anomalies. We assume, that changes  
of electromagnetic properties of the ice-sheet surface are driven by atmospheric processes that affect the temperature and  
surface density at the kilometer scale and can thus explain the time-varying elevation anomalies (hereafter referred to as  $\Delta h$ )  
as shown by Lacroix et al. (2008)). Our approach of applying the correction to averaged, interpolated products reduces the high  
225 uncertainty of single waveform parameter estimates. This is reflected in high correlations between  $\Delta h$  and anomalies of LEW  
and backscatter as presented in figures A5, A6, A7, A8 for Greenland and Antarctica. Our final product contains four monthly  
elevation estimates for each of the four retracker solutions used to investigate  $\Delta h$  and to derive dhdt. These are compared with  
those derived by ICESat-2 using the same processing strategy.

- $\Delta h$  and dhdt without any correction
- 230 – LEW corrected  $\Delta h$  and dhdt
- Backscatter corrected  $\Delta h$  and dhdt
- LEW and backscatter corrected  $\Delta h$  and dhdt.

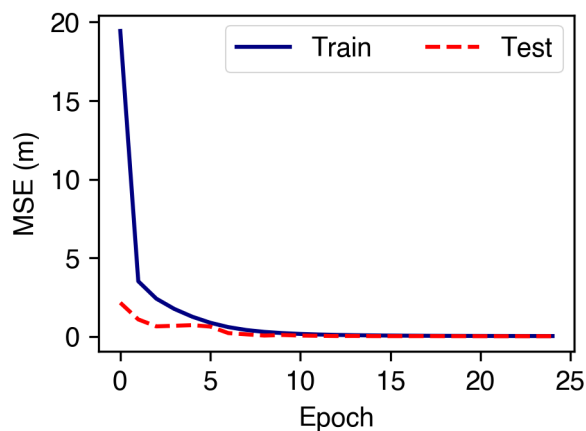


Figure 5. AWI-ICENet learning curve

### 3 Results

To evaluate the performance of the new AWI-ICENet1 retracker, different tests are conducted, which are summarized and structured as follows. First, the AWI-ICENet1 is evaluated using statistical metrics such as the MSE based on the simulated test data set. Learning curve and K-fold cross validation statistics are shown. Second, the simulated data set is retracked using the TFMRA retracker of Helm et al. (2014) and compared to AWI-ICENet1. Third, the new retracker is applied to CryoSat-2 LRM data and a monthly cross point analysis is carried out in four selected regions (ROI) of Antarctica, to assess the retracker's ability to provide reliable elevation estimates in areas with different surface topography and bulk attenuation rate. Fourth, the monthly cross point analysis is carried out across the LRM zone using all four retrackers (AWI-ICENet1, TFMRA and ESA products ICE1 and ICE2). To assess the retracker's performance in terms of its ability to minimize transient penetration effect, the monthly elevation anomaly product is evaluated within the ROIs and spatially over the whole LRM zone of Antarctica and Greenland. Finally, empirical correction to reduce time variant penetration bias are applied to all retrackers and evaluated.

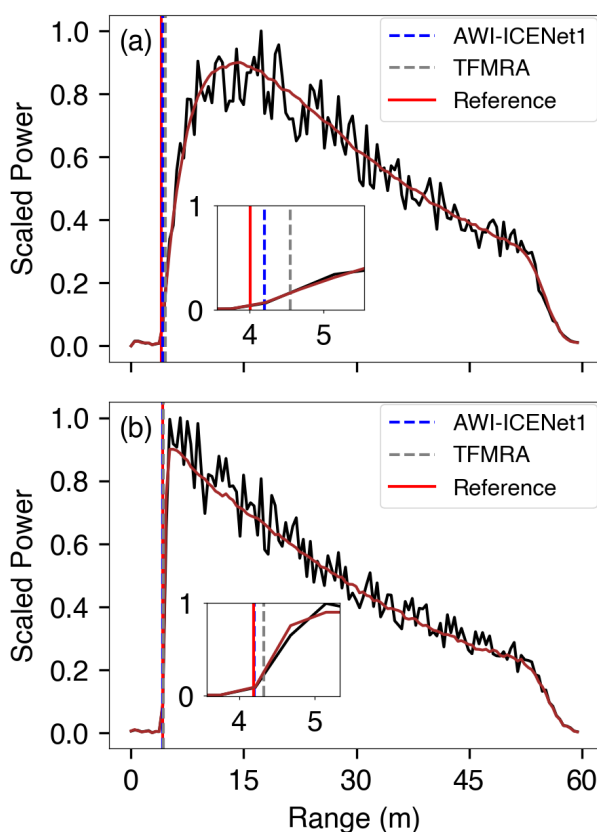
#### 3.1 AWI-ICENet1 results for simulated waveforms

We evaluated the ConvNet performance using MAE, MSE, and RMSE as given in Equations 7, 8 and 9, and by examining the learning curve of the MSE for both the training and test data sets. The learning curve of our final model is shown in Fig. 5. Here, the MSE of training and test data gradually decreases and reaches a constant plateau after approximately 10 epochs. To assess whether the model architecture provides similar metrics for different training and test data, we applied 5-fold cross validation, meaning that the same input data-set is split into five different parts, by using an 80/20 split factor for training/testing. The results of the different folds are listed in Table 1. They show nearly identical values indicating a consistent, repeatable learning



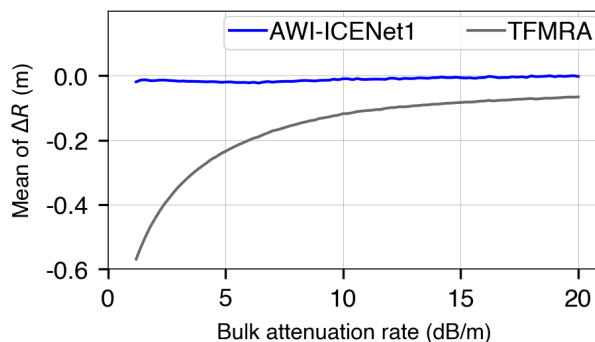
K-fold	MSE (m)	RMSE (m)	MAE (m)
1	0.070	0.056	0.042
2	0.075	0.060	0.047
3	0.070	0.056	0.047
4	0.070	0.056	0.047
5	0.094	0.066	0.047
Mean	0.076	0.059	0.046

**Table 1.** K-fold cross-validation results



**Figure 6.** Examples of simulated waveforms for (a) low  $L_A$  and (b) high  $L_A$ . Initial noise free waveform is shown as red line and with added noise as black line. The reference range is displayed as vertical bar in red color and superimposed is the AWI-ICENet1 and TFMRA retracked range in blue and grey, respectively. The inlets show a zoom into the leading edge where the retracking takes place.

independent of the training data set. Our final model, which was a hold-out with RMSE of 0.056 m and MAE of 0.042 m, was considered for application to real data and further analysis.



**Figure 7.** Comparison of the Mean of  $\Delta R$  with respect to bulk attenuation rate between AWI-ICENet1 and TFMRA.

To evaluate the retracker performance we estimated the difference  $\Delta R = R_{\text{ref}} - R_{\text{RT}}$  between the retracked range  $R_{\text{RT}}$  and the reference range  $R_{\text{ref}}$ . In addition we applied the TFMRA retracker to the same set of simulated waveforms and estimated  $\Delta R$  as well. In Fig. 7 the mean values of  $\Delta R$  for the AWI-ICENet1 and TFMRA are presented. The statistics are calculated for all location across Antarctica subdivided into bins of the bulk attenuation rate. Here, the influence of high volume scatter, due to low  $L_A$ , on range retrieval is clearly shown. For low  $L_A$  we observe only a small offset ( $< 0.03$  m) of the AWI-ICENet1 retracked surface elevation compared to the true surface. In contrast,  $\Delta R$  of TFMRA forms a kind of exponential function, with differences of up to  $-0.5$  m for low  $L_A$  as presented in Fig. 7.



## 260 3.2 AWI-ICENet1 results for observed waveforms

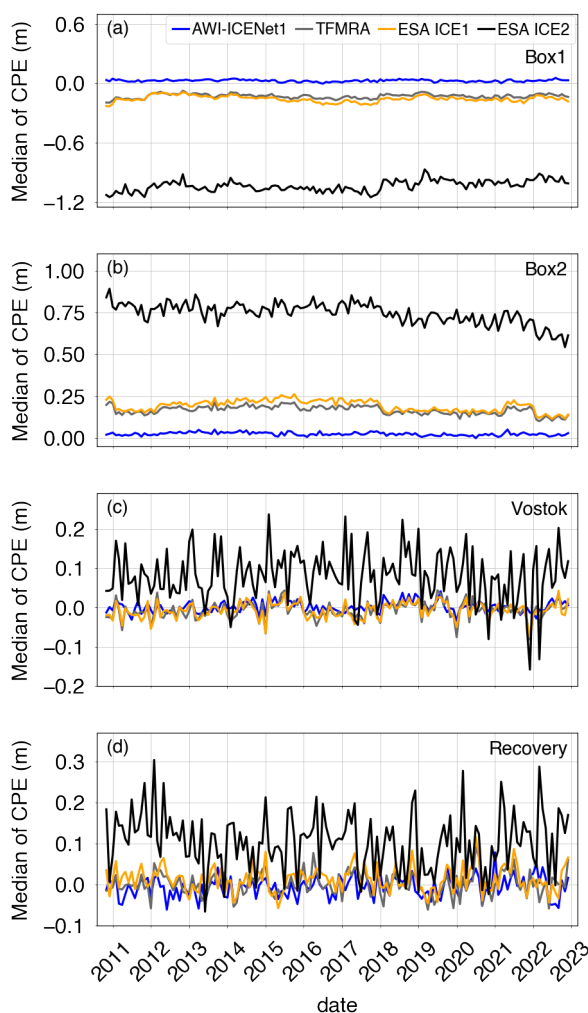
In this section, the final AWI-ICENet1 retracker is applied to the entire CryoSat-2 time series. For each of the ESA Level1B 20 Hz waveforms a range is retracked and combined with the precise orbit information (latitude, longitude and altitude) to form a point cloud of geo-referenced surface elevations. The cross point analysis is carried out on precise orbits. To derive elevation and elevation change products used for further applications, a slope corrected point cloud data set is generated using  
265 the relocation slope correction following Roemer et al. (2007).

### 3.2.1 Cross point error analysis

To avoid time dependent differences, due to changes in radar volume scattering, our accuracy measure of the individual solutions is based on monthly cross point error analysis (CPE) over the whole Antarctic ice sheet. CPE is defined as the elevation difference between ascending and descending track at cross points (CPs). In total more than 3 Mio CPs from 130 months are  
270 used. We filter outliers using the following criteria:  $|CPE| > 10$  m. Finally a grid with 5 km  $x \times 5$  km pixel resolution based on mean and standard deviation (SD) is calculated from all CPEs. Figures 8 and 9 show the time evolution of the median and SD of CPE for four selected regions which are shown in Fig. 1. In the area of Box1 (Fig. 8a) we find a significant negative CPE for TFMRA and both ESA retracker, with the largest CPE for the ESA ICE2 retracker. AWI-ICENet1 has a very low CPE and performs best in this area. In Box2 the picture is similar, but the CPE is in the positive range with again ESA ICE2 showing  
275 the highest CPE (Fig. 8b). Also the SD, displayed in panels (a) and (b) of Figure 9, has exceptional high values for ESA's ICE2 retracker. The other retracker have similar SD's in the order of 20 cm in Box1 and Box2. Next we consider two regions with specific topographic settings (both areas are shown in Fig. 1): the Vostok region is characterized as a very flat area with low surface undulations, whereas the Recovery region is more complex with higher surface slopes of up to  $1^\circ$  and medium scale topographic undulations. The results for median of CPE are presented in Fig. 8 panels (c) and (d). For both areas, the median  
280 of CPE is lower than for the two box areas, with a range of at most 20-30 cm. Again ESA ICE2 shows the highest CPE over the entire time period for both Vostok and Recovery areas. The new AWI-ICENet1, TFMRA and ESA ICE1 have median and SD CPE values at similar order of magnitude with only slight differences. The temporal variability is higher in the Recovery area than in the Vostok region.

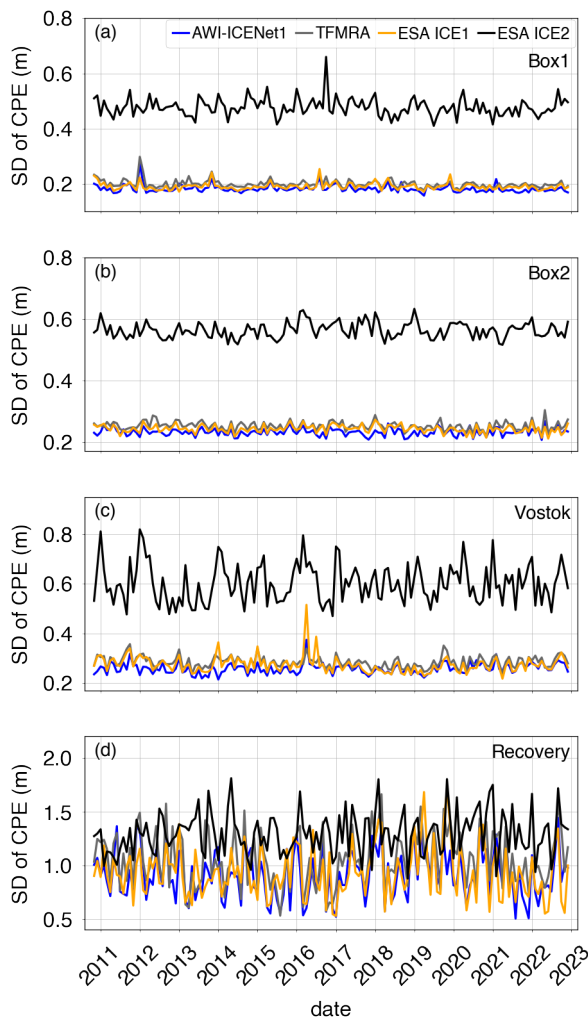
The strong difference in surface characteristics become also evident when comparing the SD of CPE: In the Recovery area  
285 the range of the SD of CPE is about 0.5-1.5 m, roughly a factor of three compared to the Vostok area and the two box areas. In all cases the trend of median and SD of CPE are rather stable over the years with few exceptions. In the area Box2 we find a decrease in median of CPE for ESA ICE2 from 2018 on-wards, and some incremental changes, a few centimeters in size, are observed by ESA ICE1 and TFMRA, while AWI-ICENet1 remains constant over time. The time series of SD from CPE also show generally stable trends, again with few exceptions: in Box1 and Vostok we find a peak in all retracker except ESA ICE2,  
290 but the trend found for the median in Box 2 is not reflected in the SD. The high temporal variability in the SD of CPE in the Recovery area prohibits analysing trends, but all retracker show a variability of about the same order of magnitude.





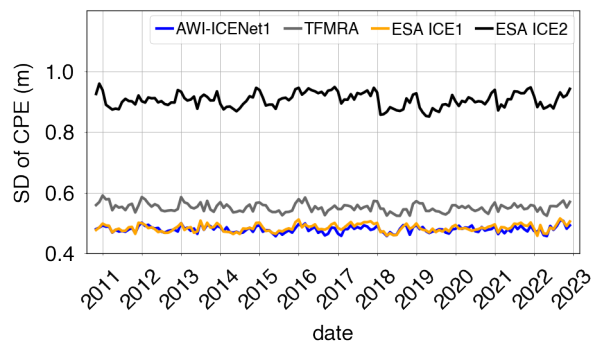
**Figure 8.** Median of cross point error for different areas

In a next step, we broaden the view and discuss the CPE over the entire ice sheet. Figure 10 displays the time series of SD of CPE for the complete LRM zone of Antarctica using the 12 years observational time period of the CryoSat-2 era. Here, AWI-ICENet1 and ESA ICE1 are similar in SD of CPE, whereas TFMRA is considerable higher in SD over the entire ice sheet. ESA ICE2's has particular high SD, almost twice as large as the first two. The temporal variability is rather low and without any trends for all four retracker solutions. The pan-Antarctic gridded median and SD of CPE are shown in Figures 11 and 12. The results reveal, that the crossover difference is reduced by roughly a factor 4 to 5 for ICE1 and TFMRA compared to ICE2 (note the different value range in panel (c) of Fig. 11). However, remaining CPEs with alternating signs in the order of up to  $\pm 0.3$  m are visible all over Antarctic including a prominent pattern in the interior of East Antarctica, close to the polar

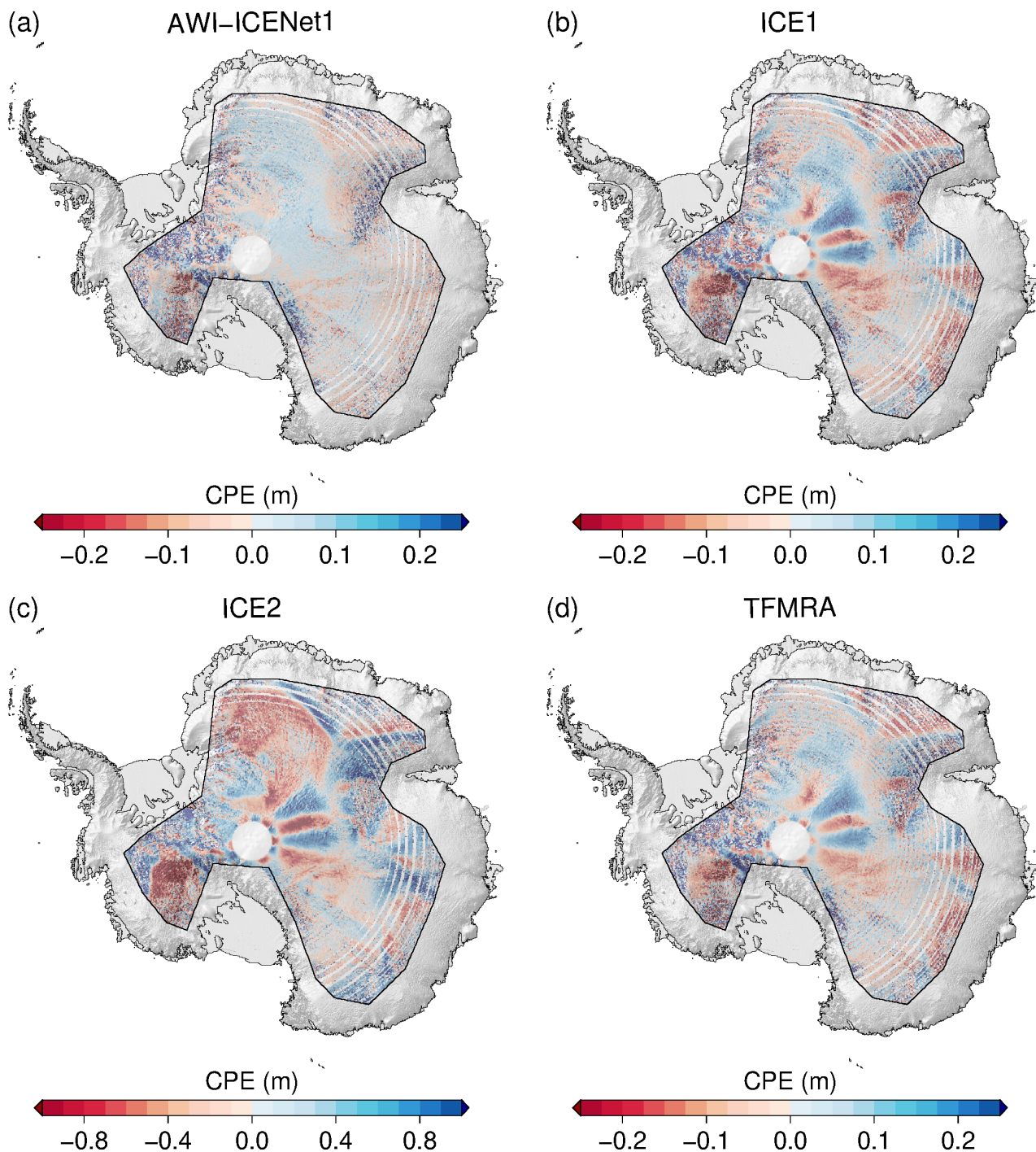


**Figure 9.** SD of cross point error for different areas

300 gap. This pattern, discussed below, is entirely eliminated with AWI-ICENet1, as displayed in panel (a) of Fig. 11. Low median  
of CPE with AWI-ICENet1 can be also seen close to the ice divides north of Dome Fuji (indicated with F in Fig. 1), in southern  
Dronning Maud Land (D) as well as in topographically more complex areas of the Siple Coast (S) and drainage area of Amery  
ice shelf (A). The SD of CPE in Fig. 12 (a), (b) and (d) is very similar for AWI-ICENet1, ESA ICE1 and TFMRA, respectively  
and a factor of four to five smaller than for ESA ICE2 (Fig. 12c). In general lowest values are found in the flat interior and  
305 highest values in the sloped areas with complex topography. This is consistent with our findings for the four test areas, where  
highest SD of CPE are observed in the Recovery area of roughest topography.



**Figure 10.** Time series of SD of CPE over the entire Antarctic LRM zone.



**Figure 11.** Median of CPE pan-Antarctic for all four retracers. Please note, panel c has a different range of the color bar then the other three panels.

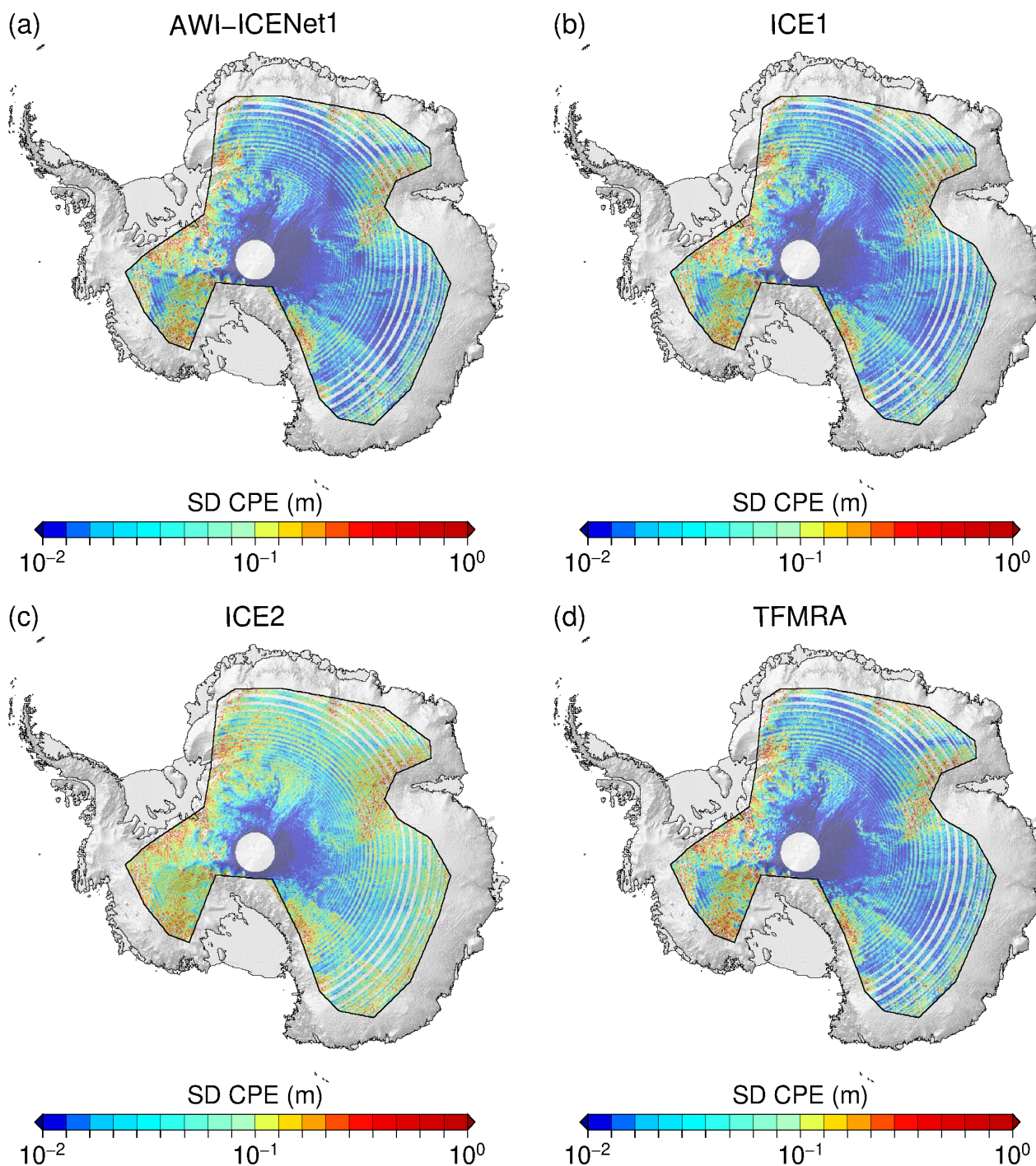


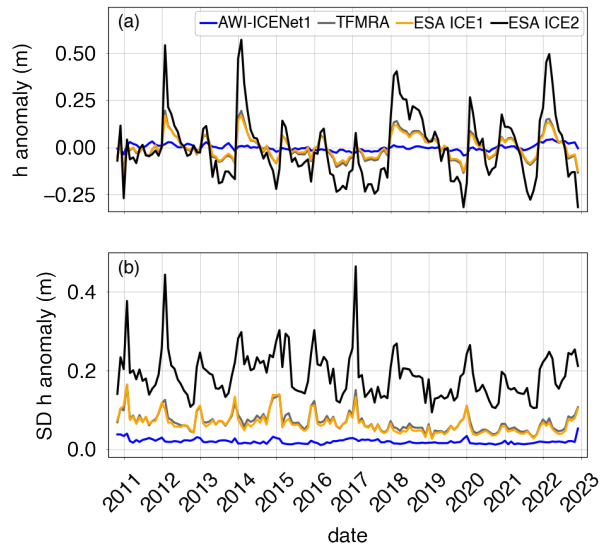
Figure 12. SD of CPE pan-Antarctic for all four retracers.



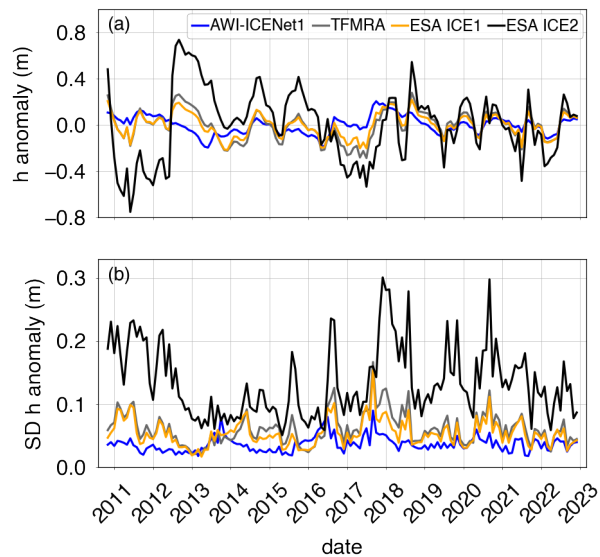
### 3.2.2 Transient penetration

In this section we analyse the time dependent variability of  $\Delta h$ , which has been widely discussed in the literature as transient penetration or penetration bias due to changes in firm properties e.g. (Davis and Zwally, 1993; Michel et al., 2014; Slater et al., 2019) for various retracker. To this end, we present in Fig. 13 the  $\Delta h$  and its standard deviation ( $SD(\Delta h)$ ) for all retracker and over the entire CryoSat-2 time period for the Lake Vostok region, a region with a stable surface height over the last decade (Richter et al., 2014). Except AWI-ICENet1 all retracker are showing a strong variability with ESA ICE2 having the largest range of up to 0.8 m. The temporal variability of ESA ICE1 and 2 and TFMRA are correlated, but ESA ICE1 and TFMRA are lower in magnitude compared to ESA ICE2. The new AWI-ICENet1 exhibits an  $\Delta h$  of only few centimeters and with minor temporal variability, with a slight increase from 2021 on-wards. The SD of the elevation anomaly (Fig. 13b) is largest for ESA ICE2, while ESA ICE1 and TFMRA have a nearly similar SD, significantly lower than ESA ICE2 but the temporal evolution has the same form as for ESA ICE2. AWI-ICENet1 exhibits not only a low  $\Delta h$  but also a low  $SD(\Delta h)$ . Fig. 14 presents the  $\Delta h$  and  $SD(\Delta h)$  for an  $5000 \text{ km}^2$  large high elevated area on the North Greenland plateau ( $79^\circ \text{ N}$ ,  $45^\circ \text{ W}$ ). Large sudden increase of more than 1.0 m for ESA ICE2 are observed in July 2012 and August 2018. Here TFMRA and ESA ICE1 also experience an increase of 0.3 m whereas AWI-ICENet-1 stays at the same level. Both events can be related to unusual heat waves transporting warm air up to high elevations resulting in surface melt conditions as reported by the DMI portal. In all other months the anomalies are correlated and follow a similar course but with strong differences in the observed magnitudes with AWI-ICENet1 showing smallest and ESA ICE2 highest values. For the SD of the elevation anomaly (Fig. 14b) we observe the same as for the Vostok area with lowest values for AWI-ICENet1 and highest for ESA ICE2. Below we will discuss how this relates to findings of other studies, but now continue with inspecting the spatial distribution of the  $\Delta h$ .

Figure 15 presents the spatial distribution of the  $SD(\Delta h)$ . Here the standard deviation is calculated for each single pixel (5 km posting). ESA ICE1 and TFMRA have a close spatial distribution and magnitude with the largest elevation anomaly along Siple Coast (S in Fig. 1) and the drainage basins in the Queen Elisabeth Land feeding the Ronne Ice Shelf (Q in Fig. 1). For ESA ICE2, the SD of elevation anomaly exceeds 30 cm in an extensive area. The lowest SD of elevation anomaly is found for AWI-ICENet1, where the largest values are also found in West Antarctica, but  $SD(\Delta h)$  for the large area of nearly whole East Antarctica is below 5 cm, whereas for ESA ICE1 and TFMRA  $SD(\Delta h)$  is in the range of 8–25 cm.



**Figure 13.** Time dependent elevation anomaly (a) and its SD (b) in the Vostok area.



**Figure 14.** Time dependent elevation anomaly (a) and its SD (b) in an area in Northern Greenland.

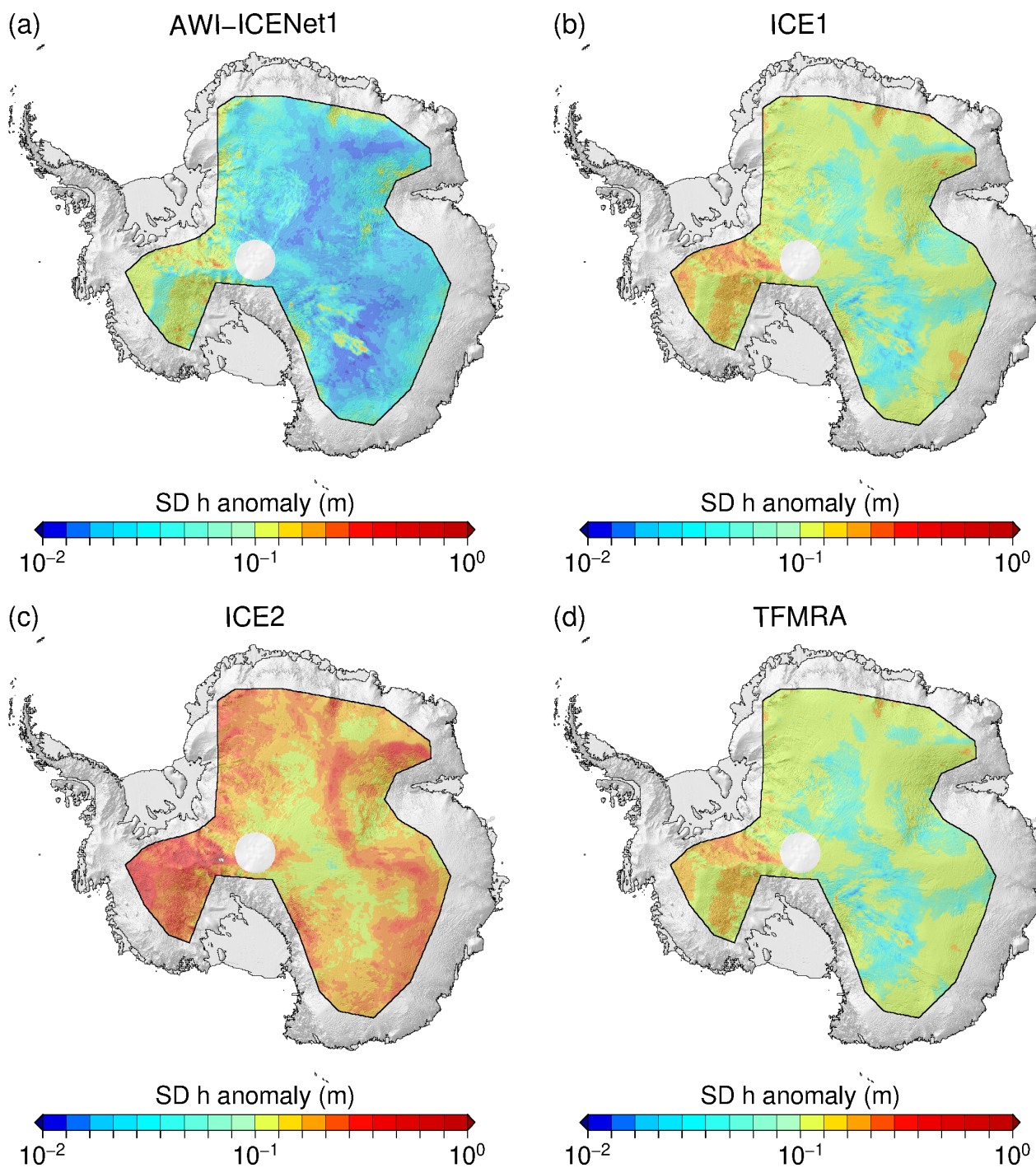


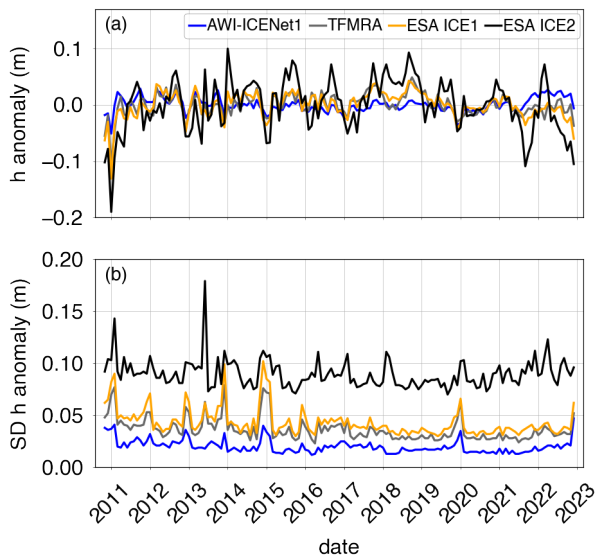
Figure 15. SD of elevation (h) anomaly



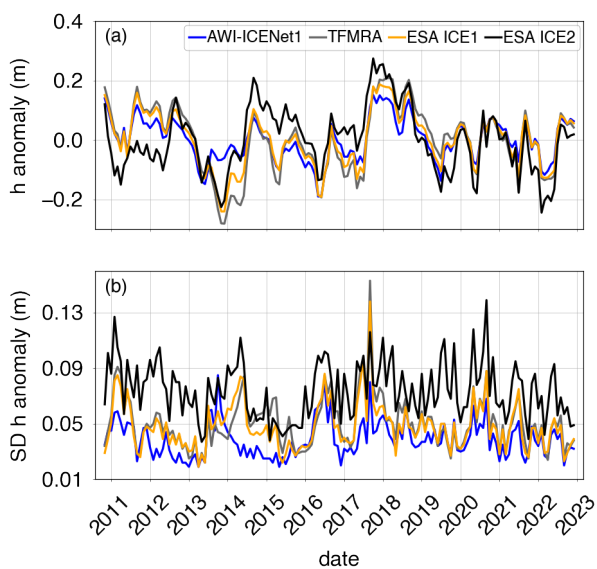


### 3.2.3 Empirical correction of transient penetration compared to AWI-ICENet1

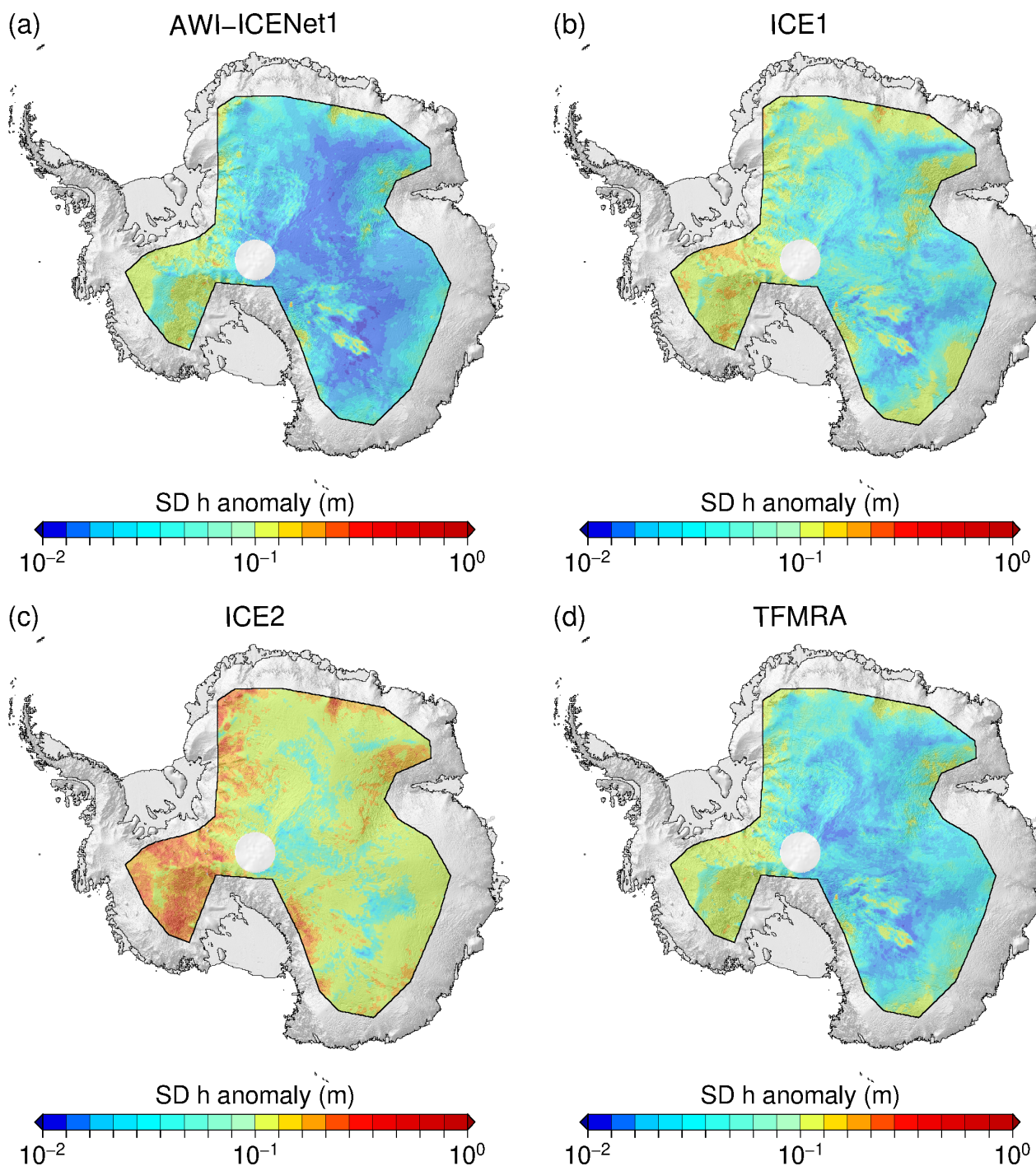
Flament and Rémy (2012); Simonsen and Sørensen (2017); Schröder et al. (2019); Nilsson et al. (2022) used an empirical relation of the elevation anomaly with backscatter, LEW and/or TeS to partly reduce the transient penetration effect. Here, we  
335 make use of the same correction largely following Nilsson et al. (2022). Exemplary we apply the correction using LEW and  
backscatter to all retracker for the Vostok area and LEW only for the North Greenland area, which is line with findings of  
Simonsen and Sørensen (2017). Results of Vostok and North Greenland are shown in Figures 16 and 17, respectively. In both  
cases  $\Delta h$  is strongly reduced for all retracker with exception of AWI-ICENet1. Largest reduction is observed for ESA ICE2  
but still showing largest remaining  $\Delta h$  which exceed the values for AWI-ICENet1 by a factor of two and higher. TFMRA  
340 and ESA ICE1 are similar and closer to AWI-ICENet1 but still show larger magnitudes. The SD of the  $\Delta h$  is reduced for  
all retracker with lowest improvement for AWI-ICENet1. However, our new retracker shows the smallest SD as presented in  
Figure 16 b). In North Greenland the sudden positive elevation increase, due to a change of the dominant scattering regime  
caused by the melt events in summer 2012 and 2018, are strongly suppressed. The different solutions are now correlated over  
the years. Smallest remaining  $\Delta h$  is observed with AWI-ICENet1, followed by TFMRA, ESA ICE1 and ESA ICE2. For North  
345 Greenland, TFMRA and ESA ICE1 show larger negative and positive deviations from AWI-ICENet1 in winter 2013/2014  
and 2017/2018, respectively, as shown in Figure 17 a).  $\Delta h$  in North Greenland is more variable compared to Vostok reaching  
values of 0.1 m for AWI-ICENet-1. Around Vostok only small perturbations of  $<3$  cm are observed with AWI-ICENet-1,  
reflecting the low accumulation regime in East Antarctica. Figure 18 presents the spatial distribution of the SD of elevation  
anomaly corrected with LEW and backscatter. Again, ESA ICE2 show largest anomalies followed by ESA ICE1 and TFMRA.  
350 However, the anomaly correction is shown to be more effective for TFMRA than for ESA ICE1 especially over the East  
Antarctic plateau.  $SD(\Delta h)$  is reduced to 4–10 cm for TFMRA and to 6–12 cm for ESA ICE1. For AWI-ICENet1 only a  
minor reduction of  $SD(\Delta h)$  could be achieved with values to be less than 4 cm for the majority of East Antarctica. It is  
worth mentioning that even the corrected TFMRA in Fig. 18 (d) still show larger values than the uncorrected AWI-ICENet1  
in Fig. 15 (a). To investigate the extent to which the applied correction mitigates the transient penetration for each retracker,  
355 we sort the  $SD(\Delta h)$  of the uncorrected and corrected  $\Delta h$  within the LRM zone in bins of 0.02 m as shown in Fig. 19. The  
cumulative frequency in Fig. 19(a) depicts that approx. 80% of  $SD(\Delta h)$  for the AWI-ICENet1 is less than 6 cm, while this  
is true for less than 10% of TFMRA and ESA ICE1. ESA ICE2 reveals values greater than 15 cm for more than 90% and  
values gretaer than 30 cm for more than 50% of the area. After the correction, as presented in 19(b), 60% of TFMRA and  
40% of ESA ICE1 reach values less than 6 cm. While the effect of the applied correction is significant for all retrackers except  
360 AWI-ICENet1, it is not able to minimize the transient penetration bias to the same order of magnitude as the AWI-ICENet1.  
Since AWI-ICENet1 experiences only a small reduction of the  $SD(\Delta h)$ , we conclude that most of the transient penetration  
bias is already corrected by the retracker. Similar results are obtained for the Greenlandic ice sheet but with in general larger  
magnitudes of  $SD(\Delta h)$ . Corresponding Figures of the spatial distribution of  $SD(\Delta h)$  and the cumulative distribution are  
presented in the Appendix in Figures A9 A10 and A11.



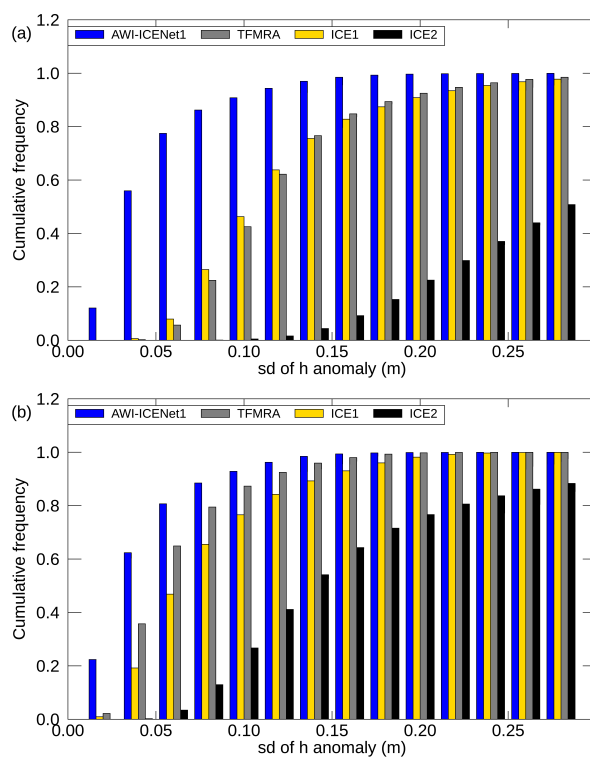
**Figure 16.** Time dependent elevation anomaly (a) and its SD (b) in the Vostok area corrected for transient penetration using LEW and backscatter.



**Figure 17.** Time dependent elevation anomaly (a) and its SD (b) in an area in Northern Greenland corrected for transient penetration using LEW only.



**Figure 18.** SD of elevation (h) anomaly corrected using a correlation with leading edge width and backscatter



**Figure 19.** SD of elevation (h) anomaly accumulated over the LRM zone in Antarctica for all retrackerers. (a) uncorrected and (b) corrected h anomalies using a correlation with LEW and backscatter.



## 365 4 Discussion

### 4.1 Assessment of AWI-ICENet1 for simulated waveforms

Several waveform models have been developed to study the effects of physical parameters on waveform shape and their effect on retracked surface height, such as Ridley and Partington (1988); Femenias et al. (1993); Legrésy and Rémy (1997); Adams and Brown (1998); Arthern et al. (2001). Martin et al. (1983) showed that the shape of the waveform echo are affected by large  
370 scale surface undulations. The surface slope has a large effect on the width of the leading edge (Femenias et al., 1993). Femenias et al. (1993) also showed that the error on range retrieval due to volume backscatter can be up to 50 cm and mainly affects the trailing edge of the echo. Small scale roughness can account for  $\pm 2$  cm of the height bias. Our waveform model captures large scale undulations and surface slope because we use a subset of the REMA elevation model as input for each individual simulated waveform. We do not consider macro-roughness (meter scale sastrugis) and small scale roughness (centimeter scale  
375 ripples) as these are not the main source of error in the height estimates. The subsurface signal can be attenuated by various physical processes and parameters, e.g. absorption losses, scattering from ice grains, multiple reflections within the snow pack caused by the stratification of the subsurface due to density contrasts (Ridley and Partington, 1988; Femenias et al., 1993; Legrésy and Rémy, 1997). In our model, we do not resolve the various parameters, but consider their combined attenuation effect as an overall or bulk attenuation rate that changes the waveform. In section 3.1, simulated sample waveforms for high  
380 and low  $L_A$  were shown in Figure 6. In addition to the very different TeS, the broadening of the LEW for low  $L_A$  is striking and reflects the observations of Femenias et al. (1993). This subsequently leads to incorrect estimates of the TFMRA-retracked surface elevation, as the retracked range is estimated to be 25% of the leading edge (Helm et al., 2014). As a result, the retracked range is estimated within the first meter of the snow pack, depending on the width of the leading edge, rather than the true surface which is shown in Fig. 7. In contrast AWI-ICENet1 strongly suppresses this penetration effect and positions  
385 the retracked range much closer to the true surface and is therefore less sensitive to errors caused by changes in the volume scattering contribution. Because AWI-ICENet1 has only been trained over the LRM zone, it may not be optimally trained for the very complex regions near the ice sheet margins. This could lead to larger errors when applied to PLRM data included in the ESA Baseline E data product. If AWI-ICENet1 is to be used for other missions such as EnviSat, SARAL/Altika, or Sentinel-3 (PLRM), new simulation and training would be required to best represent mission-specific sensor characteristics such as centre  
390 frequency, bandwidth, or antenna gain pattern.

### 4.2 Assessment of AWI-ICENet1 for satellite altimetry data

#### 4.2.1 Cross point error analysis

In section 3.2.1, CPE results were presented as pan-Antarctic gridded median and SD of CPE, Figures 11 and 12. A prominent feature in the central part (around the pole gap) in East Antarctica was identified as a static crossover pattern. This pattern  
395 was explained by Armitage et al. (2014) as the result of an an-isotropic dependence of the extinction coefficient on the angle between the radar polarization and wind-induced properties of the firm. As a consequence of the Armitage et al. (2014) results,

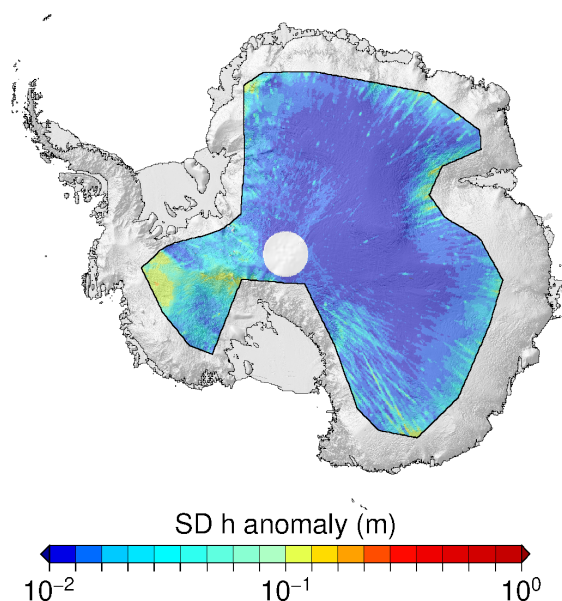


Box1 and Box2 were selected to investigate whether the crossover pattern is static over time and can be reduced with AWI-ICENet1. We have shown, that the results in this area reveal a reduction in CPE of about a factor 4 for ICE1 and TFMRA compared to ICE2 (note the different range of values in panel (c) of Fig. 11). A similar observation has already been made  
400 by Helm et al. (2014), where TFMRA showed slightly better results than ESA ICE1. This can be confirmed here as well. Moreover, the pattern is completely eliminated with AWI-ICENet1, as displayed in panel (a) of Fig. 11. We conclude that the influence of anisotropic dependence of the extinction coefficient on wind-driven directional anisotropy of the ice sheet surface and firn on surface height measurements as described by Legresy et al. (1999) is strongly suppressed by the new AWI-ICENet1 retracker. These results are consistent with those of Arthern et al. (2001), who found that the extinction coefficient, or as we  
405 refer to it, the bulk attenuation rate, decorrelates between ascending and descending tracks. Because AWI-ICENet1 is able to minimize the effect of bulk attenuation rate on range measurements it is also able to remove the directional effect from ice sheet height measurements.

#### 4.2.2 Transient penetration

Time dependent elevation anomalies as measured by radar satellite altimetry is a composite signal of ice dynamical processes,  
410 changes in the surface mass balance (SMB), changes in the firn compaction rate and time dependent radar penetration. The latter one can exceed SMB and firn compaction anomalies by one order of magnitude in areas of low accumulation rate such as the majority of the vast East Antarctic ice sheet. To reliably estimate volume and mass change it is very important to correct for the penetration bias as the true elevation change is only driven by changes in SMB, firn compaction and ice dynamics. In Section 3.2.2 the time dependent variability of the elevation anomaly over the Lake Vostok area was presented. All retracker  
415 usually used over ice sheets show undulations of the surface height of a couple of decimeters in an area which is known to be stable over the last decades (Richter et al., 2014). Similar findings for this area were already discussed in the literature for CryoSat-2 as well as for Envisat and ways to partly suppress the penetration bias presented. As this undulation is driven by time dependent changes in the firn properties the widely accepted method is to use the correlation of the  $\Delta h$  to backscatter and/or waveform parameters LEW and TeS as they are also affected by firn properties and thus also change in time.

420 We applied this correction to all retracker in 3.2.2 and showed how effectively this correction method works. However, it is still unclear how much penetration bias explains the measured anomaly and how effectively the applied corrections can mitigate the bias. To address this issue we compare our results with ICESat-2 observations and focus on the time 2019/01–2021/12 period when both missions were in operation. Since ICESat-2 is a laser altimeter, a negligible penetration bias is assumed. This is illustrated in Fig. 20, which shows the  $\Delta h$  observed by ICESat-2 in the given time period. In more than 90%  
425 of the area the SD of the  $\Delta h$  is less than 0.04 m, indicating that for most of the area there is little variation in the SMB and/or firn compaction. Only in the Amundsen Sea Embayment (ASE) and at the Siple coast in West Antarctica are h anomalies of more than 0.1 m found. The latter could be a dynamic thickening signal of the stagnant Camp Ice Stream (Nield et al., 2016). The higher variability in the ASE can be explained by the extreme precipitation events in this area reported by Davison et al. (2023) during the winters of 2019 and 2020. In Fig. 21 we compare the cumulative frequency of the  $SD(\Delta h)$  with ICESat-2  
430 for the uncorrected and corrected data in (a) and (b), respectively. AWI-ICENet-1 is already in very close agreement with

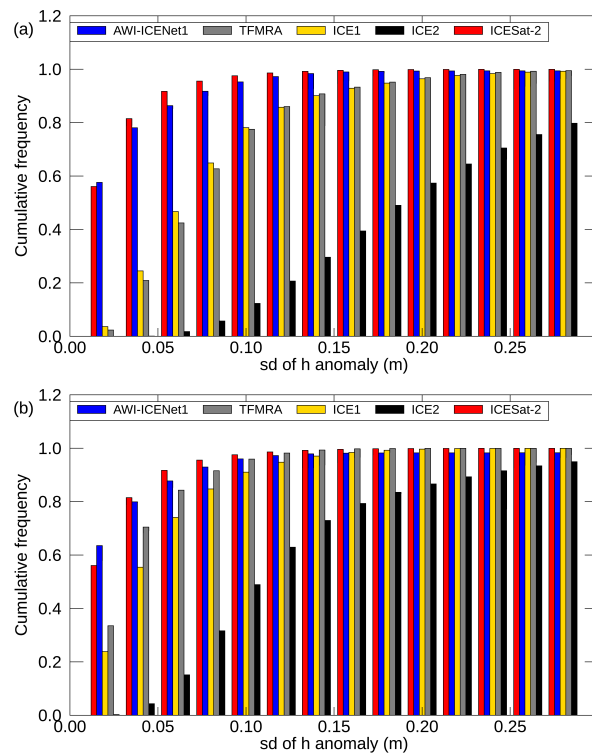


**Figure 20.** ICESat-2 derived SD of elevation ( $h$ ) anomaly over Antarctica in time span from 2019/01–2021/12.

ICESat-2 without any correction and only slightly improved with the correction. However, for TFMRA and ESA ICE1 the applied correction is very effective but still remaining signals are present. The correction is not able to suppress  $SD(\Delta h)$  to less than 4 cm as only 20% of the area for TFMRA and ESA ICE1 but 60% for ICESat-2 and AWI-ICENet1 are observed. A different interpretation of the results is shown in Fig. 22. Here, the proportion of the penetration bias ( $P_{Pbias}$ ) in the measured elevation anomaly signal summarized for the entire LRM zone is presented.  $P_{Pbias}$  is estimated using the following equation:

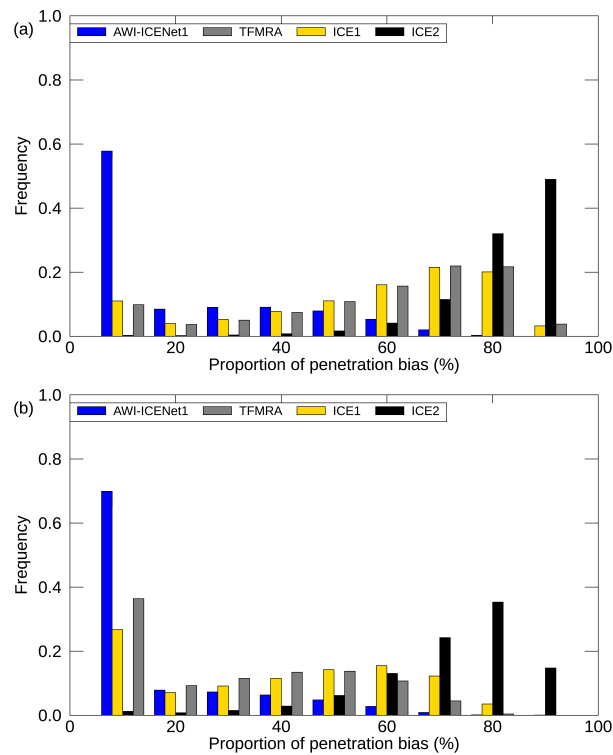
$$P_{Pbias} = 1 - \frac{SD(\Delta h_{ICESat-2})}{SD(\Delta h_{rt})}$$

with the SD of the anomaly of each retracker ( $SD(\Delta h_{rt})$ ) and of ICESat-2 ( $SD(\Delta h_{ICESat-2})$ ), respectively. In the uncorrected time series in Fig. 22 (a) 60% of the AWI-ICENet1 observed anomaly is due to >90% of true elevation change and less than 10% due to penetration artifacts. In contrast, for ESA ICE2 90 % of the observed signal is due to penetration artifacts. TFMRA and ESA ICE1 are again very similar. For only 10% of data, the signal can be explained by 90% of the actual elevation change, but 25% is dominated by 70% radar penetration distortions. After correction (Fig. 22 (b)) the signal is explained by 90% true elevation change in about 35% and 25% of the TFMRA and ESA ICE1 data, respectively. With AWI-ICENet-1 this value exceeds 70%. Our results illustrate, that backscatter and LEW corrections even if they are applied on a relatively short time span of 3 years, significantly improve the data, but are not able to fully correct erroneous signal contribution due to radar penetration. For longer time series, as in Fig. 19, the effect of the correction for ice-sheet-wide observations is even smaller.



**Figure 21.** SD of elevation (h) anomaly accumulated over the LRM zone for all retrackerers and ICESat-2 for the time span 2019/01–2021/12. (a) uncorrected and (b) corrected h anomalies using a correlation with LEW and backscatter.





**Figure 22.** Proportion of the penetration bias in the measured elevation anomaly signal for the entire LRM zone for all retrackers. (a) uncorrected and (b) corrected h anomalies using a correlation with LEW and backscatter.



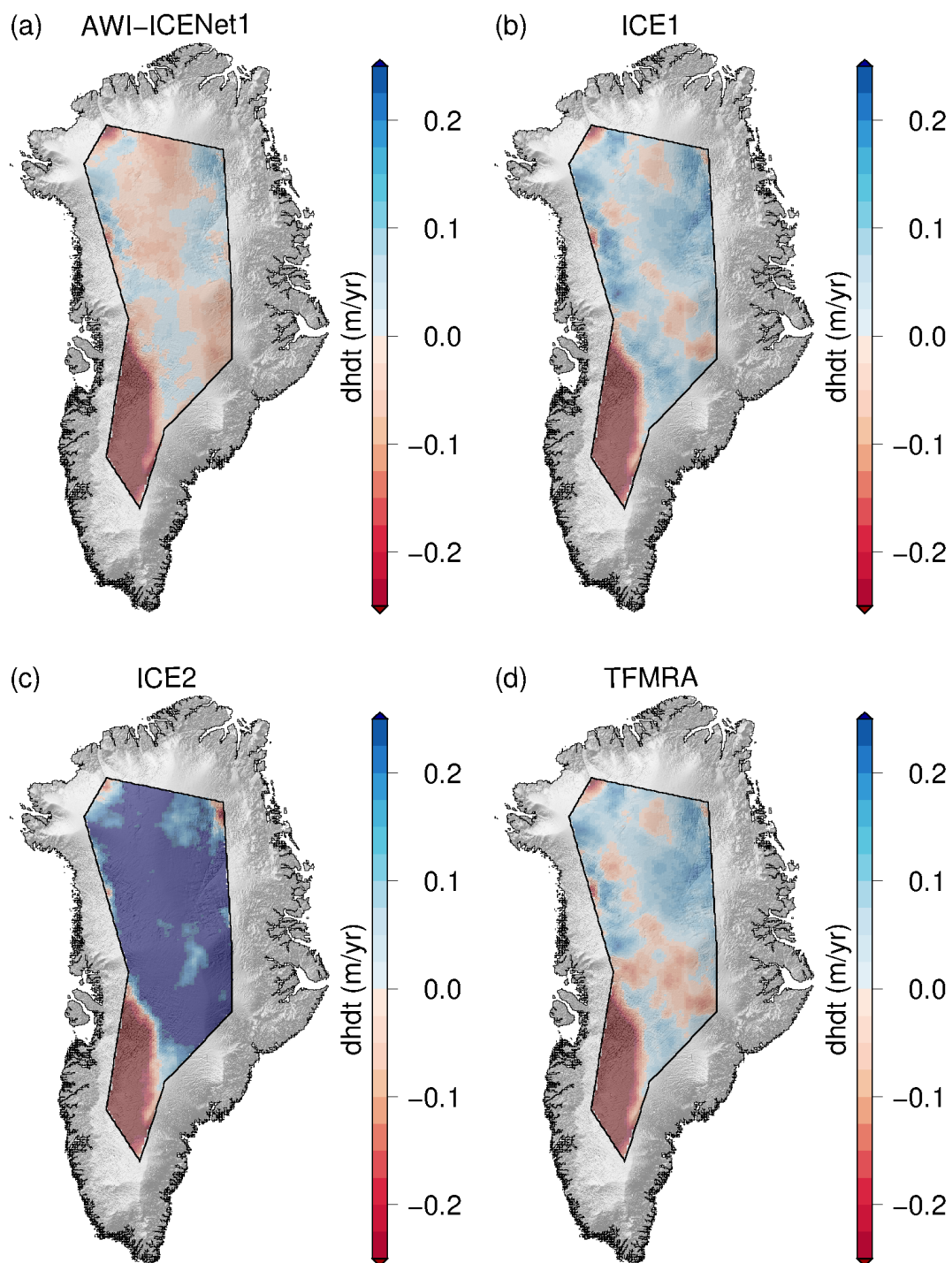
### 4.2.3 Sudden changes in scattering properties and its effect on elevation change

Nilsson et al. (2015, 2016); McMillan et al. (2016); Simonsen and Sørensen (2017) presented and discussed the impact of the unusual melt event on Cryosat-2 observations that occurred over the Greenland Ice Sheet in July 2012. Nilsson et al. (2015) reported an improved performance by using a 20% threshold retracker for the LRM data with less sensitivity to changes in  
450 near-surface scattering properties, which is in close agreement with findings of Helm et al. (2014). On the other hand McMillan et al. (2016) introduced a step function to mitigate the observed elevation step. This procedure however, is not applicable as a general approach as multiple melt events, which can be expected in future, cannot be taken into account. Simonsen and Sørensen (2017) studied in detail, by using the ESA LRM\_L2 data, the implication of different waveform parameters (LEW and backscatter) to correct for the scattering properties in Cryosat-2 observations for the time span November 2010 until  
455 November 2014. Their findings suggest for the LRM zone as best approach the LEW correction. They also found that the bias of temporal changes is not entirely removed which agrees with our results for ESA ICE1 and ICE2. In order to evaluate if the AWI-ICENet1 is capable to handle sudden scattering changes we applied the dhdt processing for all retracker to the time span January 2011 until January 2015 covering nearly the same time span as investigated in Simonsen and Sørensen (2017). In Fig. 23 the uncorrected dhdt and in Fig. A15 the corrected dhdt using LEW and backscatter are shown. The same unusual elevation  
460 increase with mean rates of  $0.1\text{--}0.2\text{ m/yr}$  is observed with ICE1 and TFMRA for large parts of the high elevated area. ICE2 shows rates of  $> 0.25\text{ m/yr}$  for almost the whole LRM zone. Only the South Eastern part experiences a surface lowering. In contrast, the AWI-ICENet1 derived dhdt ranges between  $-0.05$  to  $0.05\text{ m/yr}$  and does not seem affected by the change of scattering characteristics due to the surface melt event. When compared to the corrected dhdt estimates we find the best agreement with TFMRA. In agreement with Simonsen and Sørensen (2017) we find that in the corrected ESA ICE1 and ICE2  
465 result still some remaining signal is left. The application of a low threshold retracker as suggested by Nilsson et al. (2015) strongly reduces the sensitivity to scattering changes but still needs to be combined with waveform parameter corrections to provide reliable results. Only the AWI-ICENet1 incorporates most of the correction directly within the retracking. The differences between uncorrected and corrected  $dV/dt$  estimates are within the uncertainty of  $8\text{ km}^3/\text{yr}$  as listed in Table 2. AWI-ICENet1 is in close agreement with the ESA ICE1-LEW estimates. These are found to match best to surface elevation  
470 changes measured by LiDAR during the Operation Icebridge campaigns, as presented in Simonsen and Sørensen (2017), and agree well with our corrected TFMRA estimates.



Correction	AWI-ICENet1	ICE1	ICE2	TFMRA
none	$-62 \pm 8$	$-28 \pm 13$	$166 \pm 32$	$-36 \pm 14$
LEW	$-58 \pm 8$	$-59 \pm 11$	$20 \pm 19$	$-77 \pm 11$
Backscatter	$-63 \pm 8$	$-7 \pm 12$	$208 \pm 29$	$-22 \pm 12$
LEW and backscatter	$-55 \pm 8$	$-33 \pm 11$	$39 \pm 19$	$-52 \pm 10$

**Table 2.** Volume change  $dV/dt$  estimates for the time span 2011/01–2013/12 for Greenland derived by different retrackers and corrections. All values are given in ( $km^3/yr$ ).



**Figure 23.** Greenland dhdt for the time span 2011/01–2013/12 covering the melt event July 2012.



#### 4.2.4 Comparison of elevation change estimates to ICESat-2

As an additional way to assess the performance of the retrackerers, we conduct a comparison of elevation change (dhdt) to ICESat-2 for the time span 2019/01–2021/12. The difference in elevation change (dhdt-CryoSat-2 – dhdt-ICESat-2) (dhdt difference) is displayed in Fig. 24 for each retrackerer individually. Again the pattern of ESA ICE1 and TFMRA are similar spatially as well as magnitude-wise. ESA ICE2 is farthest off the ICESat-2 elevation change. The new AWI-ICENet1 deviates by less than  $\pm 0.1$  m/yr from ICESat-2 over approx. 90% of the entire area. In Table 3 estimated volume changes ( $dV/dt$ ) observed in the LRM zone for the time span 2019/01–2021/12 are listed for ICESat-2 and all retrackerers including the applied corrections for both Ice Sheets. In Antarctica, the difference between switching the correction on/off is tremendously high for ICE2 ranging from 293 to  $-60$  km<sup>3</sup>/yr, followed by ICE1 with a 150 km<sup>3</sup>/yr and TFMRA with a 92 km<sup>3</sup>/yr spread. AWI-ICENet1 shows the smallest range with only 11 km<sup>3</sup>/yr indicating that most of the corrections are already covered by the retracker itself. The results also show the absolute need to apply the corrections for all retrackerers except AWI-ICENet1 as otherwise  $dV/dt$  and in consequence mass balance estimates are unreliable, especially for East Antarctica. The uncorrected estimates are prone to errors which are strongly related to the elevation anomalies originating from changes in the electromagnetic properties of the upper layers of the snow/ice pack. In Greenland the same conclusion can be drawn. Here, the spread of  $dV/dt$  is smaller. The spread is 50 km<sup>3</sup>/yr for ICE2, 14 km<sup>3</sup>/yr for both ICE1 and TFMRA, and only 8 km<sup>3</sup>/yr for AWI-ICENet1. In all cases the uncertainty of the estimates could be reduced using the combined correction of LEW and backscatter. The smallest spread between the different retracker solutions is also found for LEW and backscatter corrections applied for Antarctica and Greenland, respectively. Interestingly, as shown in Figures 15, 18, A9 and A10 none of the corrections applied to ICE2, ICE1 and TFMRA are capable to reduce the  $\Delta h$  as strong as the AWI-ICENet1. This is reflected especially for Greenland in Table 3 where none of the  $dV/dt$  estimates are reaching values around 11 km<sup>3</sup>/yr as observed by ICESat-2. For Greenland and Antarctica AWI-ICENet1 also reveal smallest offsets to ICESat2, in both cases the difference of  $dV/dt$  is roughly 30 km<sup>3</sup>/yr, where ICESat-2 tend to show larger values (compare to panel (a) of Figures 24, 25, A13 and A14). This small  $< 1$  cm/yr difference in dhdt to ICESat-2 is most likely a residual signal due to radar penetration and not related to any kind of ICESat-2 mission bias as observed for the predecessor mission ICESat-1 (NSIDC, 2021). Sea-level trends computed from ICESat-2 observations agree with independent measurements from radar altimetry and tide gauges as shown by Buzzanga et al. (2021) and do not indicate any trend bias, at least over oceans. In Antarctica the best match with AWI-ICENet1 is found for TFMRA with LEW and backscatter corrections applied. Especially the correction with backscatter tend to introduce errors for ICE1 and ICE2, which most likely originate from an error in the power scaling factors in the L1B LRM data as reported by ESA (Exprivia, 2021). The issue has been solved in Baseline E. This caused the backscatter computed at Level 1 to differ between Baseline D and Baseline E resulting in a jump in backscatter in August 2019 when Baseline E was operational (Exprivia, 2021). In our processing of the TFMRA data we applied a fix by using the transmit power parameter provided in the Level 1B product to correct for this issue, which seems to solve most of the problems, especially for Antarctica as indicated by a generally higher correlation between  $\Delta h$  and backscatter anomaly for TFMRA (see Figures A6 and A8). The fact that the correlations



505 for AWI-ICENet1 are lower is in the nature of things, since the seasonal h anomalies are already strongly suppressed leading to a much lower correlation with backscatter or LEW and thus reducing the correction.

Area	Correction	ICESat2	AWI-ICENet1	ESA ICE1	ESA ICE2	TFMRA
ANT	none		$64 \pm 10$	$149 \pm 22$	$293 \pm 54$	$153 \pm 22$
	LEW		$55 \pm 9$	$111 \pm 18$	$142 \pm 38$	$69 \pm 14$
	Backscatter	$86 \pm 8$	$56 \pm 9$	$-1 \pm 17$	$-60 \pm 43$	$88 \pm 13$
	LEW and backscatter		$53 \pm 9$	$17 \pm 14$	$-19 \pm 33$	$61 \pm 11$
GRE	none		$-22 \pm 7$	$-51 \pm 10$	$-87 \pm 19$	$-61 \pm 10$
	LEW		$-25 \pm 6$	$-43 \pm 9$	$-53 \pm 14$	$-48 \pm 9$
	Backscatter	$11 \pm 7$	$-17 \pm 6$	$-55 \pm 8$	$-103 \pm 14$	$-54 \pm 8$
	LEW and backscatter		$-20 \pm 5$	$-57 \pm 7$	$-76 \pm 12$	$-53 \pm 7$

**Table 3.** Volume change  $dV/dt$  estimates for the time span 2019/01–2021/12 for Greenland (GRE) and Antarctica (ANT) derived by different retrackers compared to ICESat2. All values are given in ( $km^3/yr$ ).

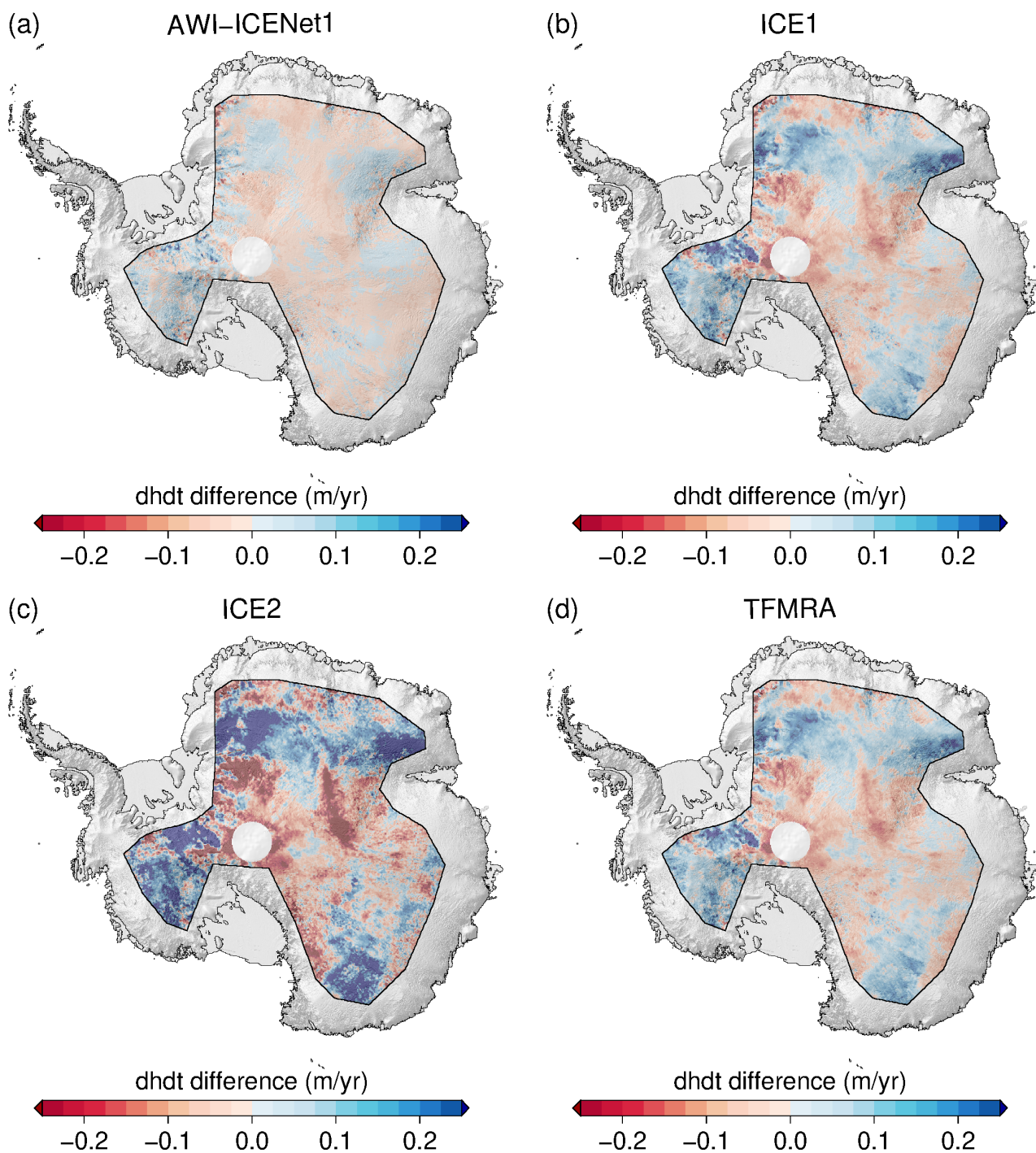
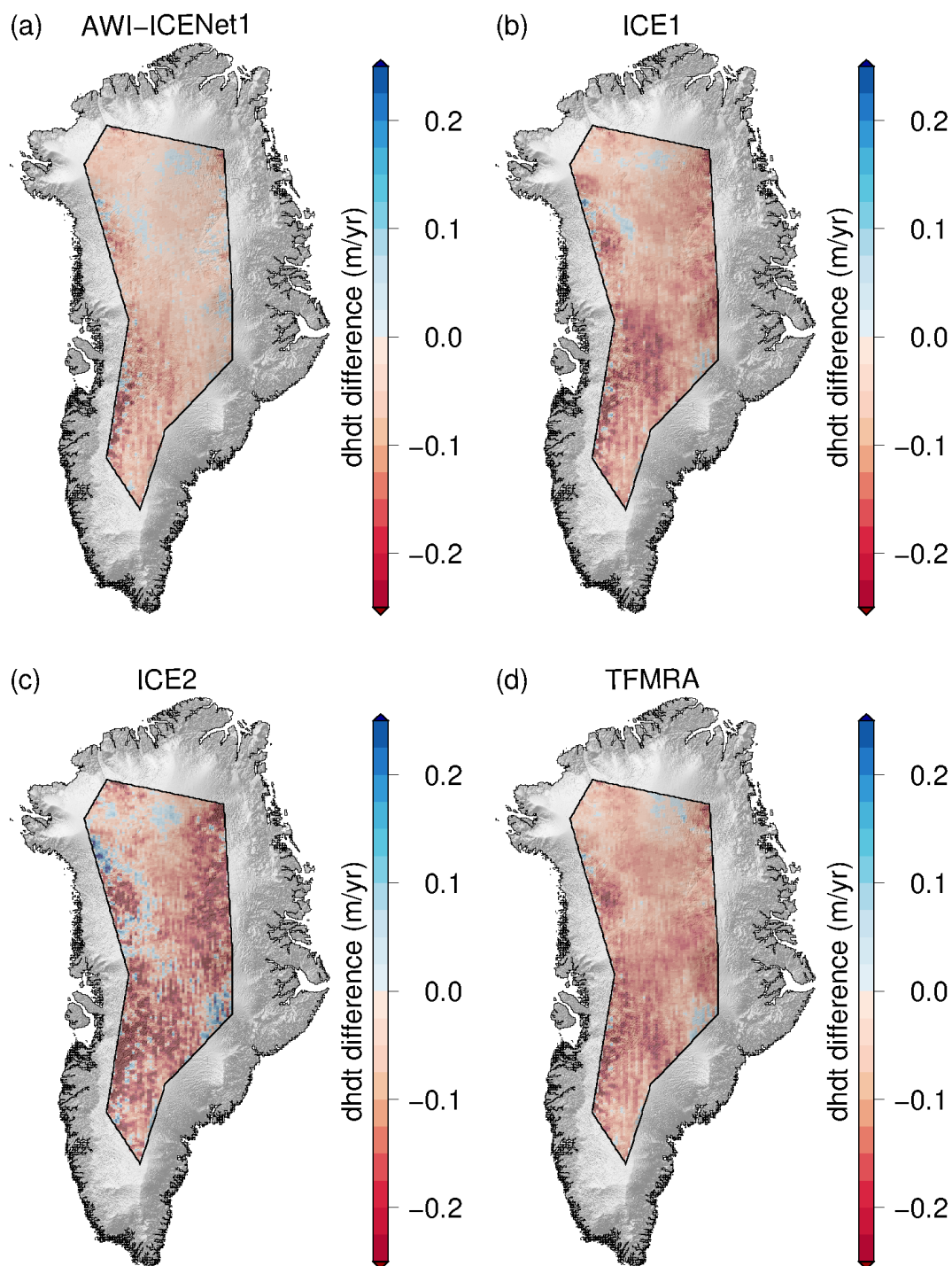


Figure 24. dhdt difference to ICESat2 for the span 2019/01–2021/12



**Figure 25.** Greenland wide dhdt difference to ICESat2 for the span 2019/01–2021/12.

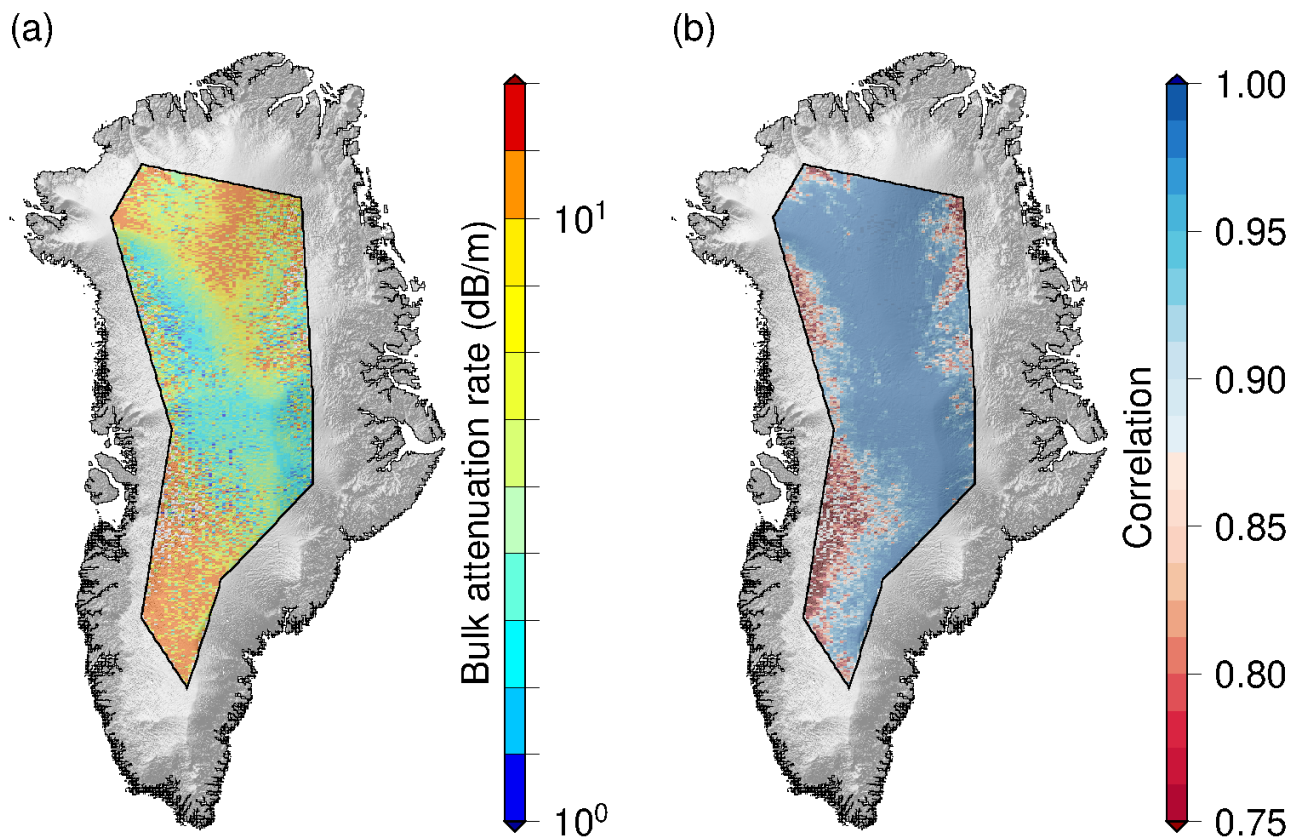




## 5 Conclusions

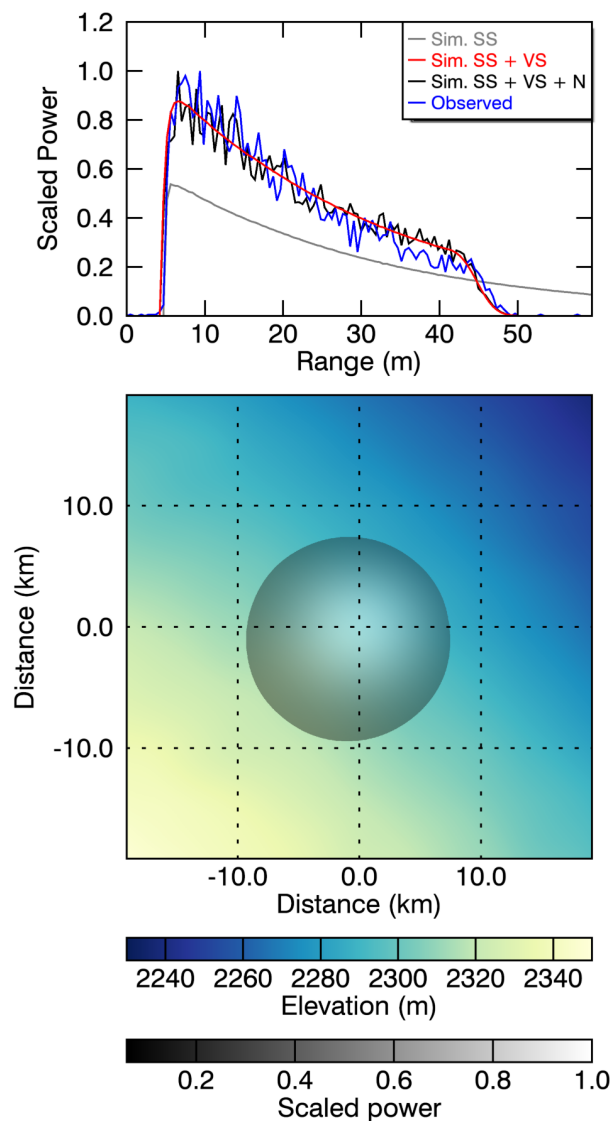
We conclude that the use of the AWI-ICENet1 a new retracker for ice altimetry based on a convolutional network provide better performance in deriving ice sheet wide elevation products than currently used retracker, such as the TFMRA or the ESA Level2  
510 ICE1 and ICE2 products. In all terrain types, the performance of AWI-ICENet1 in terms of CPE is similar to that of ESA Level2  
ICE1, the highest precision retracker. The main improvement obtained with our new approach is the large reduction in sensitivity to transient signal penetration due to temporal changes in near-surface scattering properties of the firn compared to other approaches. The transient penetration bias can be in the order of several decimeters and is reduced to a few centimeters by AWI-ICENet1. AWI-ICENet1 can handle abrupt changes in the prevailing scattering mechanism due to surface melt, as  
515 occurred over Greenland in July 2012, as well as temperature-induced changes of the bulk attenuation rate of the upper firn layers, which are mainly observed as a seasonal signal in East Antarctica. The contribution of the transient penetration to the measured elevation anomaly is greatly reduced compared to equivalent products, even when compared to empirically corrected data using a correlation between elevation changes and changes in backscatter or waveform shape parameters such as LEW.  
520 This results in a much more accurate sampling of the true elevation change due to changes in SMB or firn compaction, and estimates agree well with ICESat-2 derived estimates that are not susceptible to transient penetration. Using the new retracking method for CryoSat-2 LRM data, we determine the volume change in the LRM zone of the Antarctic and Greenland Ice Sheets for the period January 2019 to December 2021 to be  $64 \pm 10 \text{ km}^3 \text{ a}^{-1}$  and  $-22 \pm 7 \text{ km}^3 \text{ a}^{-1}$ , respectively. This agrees well with ICESat-2 estimates obtained over both ice sheets during the same period ( $86 \pm 8 \text{ km}^3 \text{ a}^{-1}$  and  $11 \pm 7 \text{ km}^3 \text{ a}^{-1}$ ). The improvements in our new retracking approach reduce the need of correlation corrections in post-processing, which can introduce  
525 additional bias as a result of anomalous backscatter variations due to different processing baselines. In addition, empirical corrections are sensitive to the length of the time series used to apply the correction. The new retracking method presented, provides a higher intrinsic accuracy in the measured surface elevation and thus lower uncertainty in derived products such as elevation change, and volume change estimates. This will lead to an improved understanding of the response of ice sheets to climate change, especially in areas of low elevation change such as the East Antarctic Plateau, where high measurement  
530 accuracy is required to separate true elevation change from noise originating from transient radar penetration. In the future, AWI-ICENet1 should also be applied to pulse-limited altimetry products (LRM and PLRM) of past (Envisat, ERS), present (Sentinel-3 A/B, SARAL/Altika), and future (CRISTAL) altimetry missions to extend the time series and to realize a consistent elevation change product with the least possible penetration biased estimates. Furthermore an extension of the waveform simulator to SAR waveforms would be of high interest to be able to apply a similar approach to CRISTAL and  
535 Sentinel-3 SAR data.

*Data availability.* The complete simulated reference data set, elevation change grids as well as monthly elevation anomalies will be uploaded to the World Data Center PANGAEA.

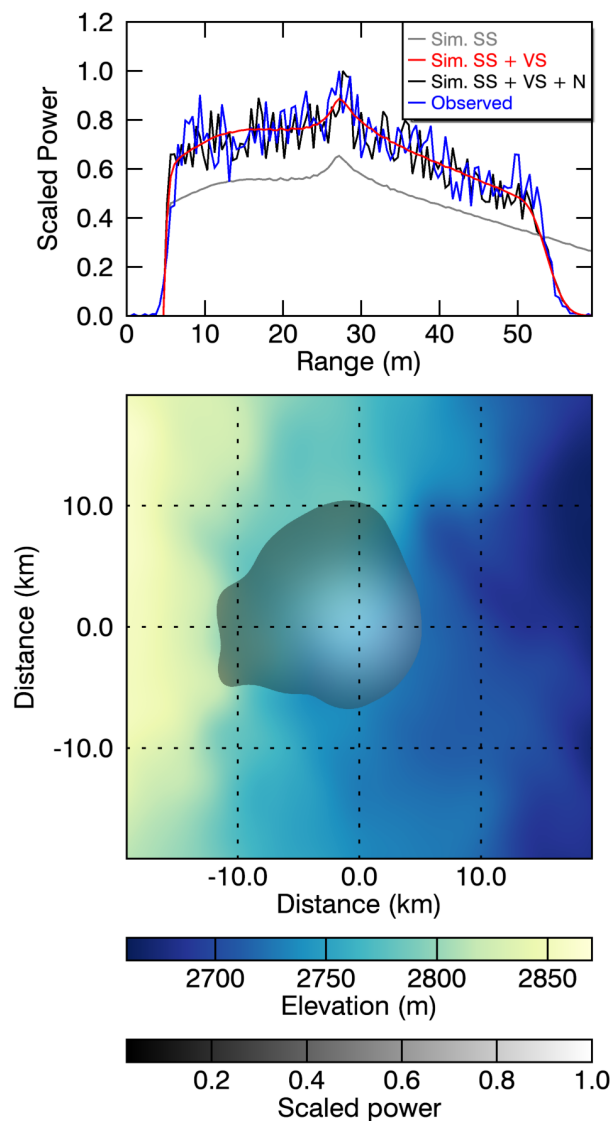


**Figure A1.** Results of the test where a flat surface waveform model is fitted to real CryoSat-2 waveforms by adjusting the attenuation  $L_A$  as the only parameter. (a) gridded median of the attenuation rate estimated by a MLE fit and (b) median of the correlation between observed and fitted waveforms.

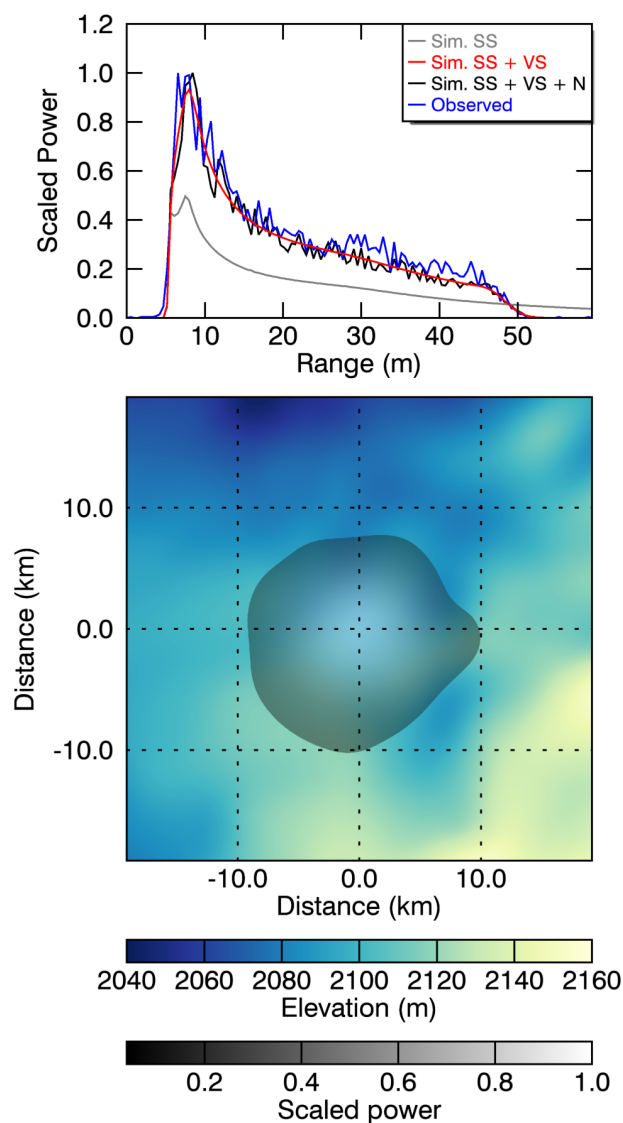
## Appendix A



**Figure A2.** Example of the different simulation steps shown for a typical waveform around the NEEM ice core site in Greenland. Blue line depicts the observed CryoSat-2 LRM waveform and the gray line the simulated surface waveform based on Eq. 1. The Red line shows the simulated waveform including volume scattering according to Eq. 3, while the black line displays the final simulated waveform including surface and volume scattering as well as noise as given in Eq.5. The lower panel represents the 2D elevation model and scaled  $P_{r_s}$ , which is mainly controlled by the Gaussian antenna pattern.



**Figure A3.** Example of the different simulation steps shown for a typical waveform in the Amery drainage basin (region A in 1). Blue line depicts the observed CryoSat-2 LRM waveform and the gray line the simulated surface waveform based on Eq. 1. The Red line shows the simulated waveform including volume scattering according to Eq. 3, while the black line displays the final simulated waveform including surface and volume scattering as well as noise as given in Eq.5. The lower panel represents the 2D elevation model and scaled  $P_{r_s}$ , which is mainly controlled by the Gaussian antenna pattern.



**Figure A4.** Example of the different simulation steps shown for a typical waveform in the Recovery area (see 1).. Blue line depicts the observed CryoSat-2 LRM waveform and the gray line the simulated surface waveform based on Eq. 1. The Red line depicts the simulated waveform including volume scattering according to Eq. 3, while the black line shows the final simulated waveform including surface and volume scattering as well as noise as given in Eq.5. The lower panel represents the 2D elevation model and scaled  $P_{r_s}$ , which is mainly controlled by the Gaussian antenna pattern.

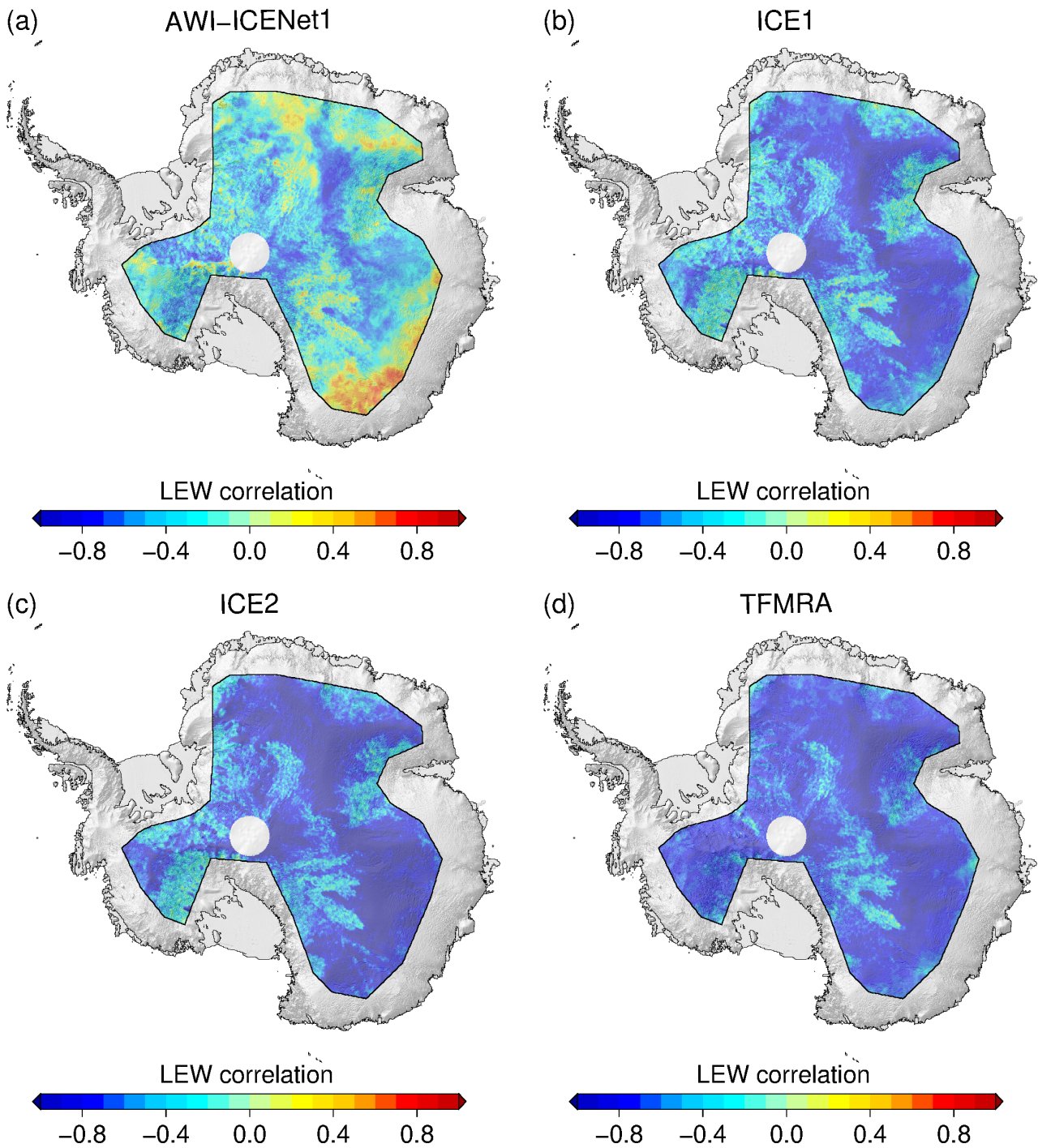
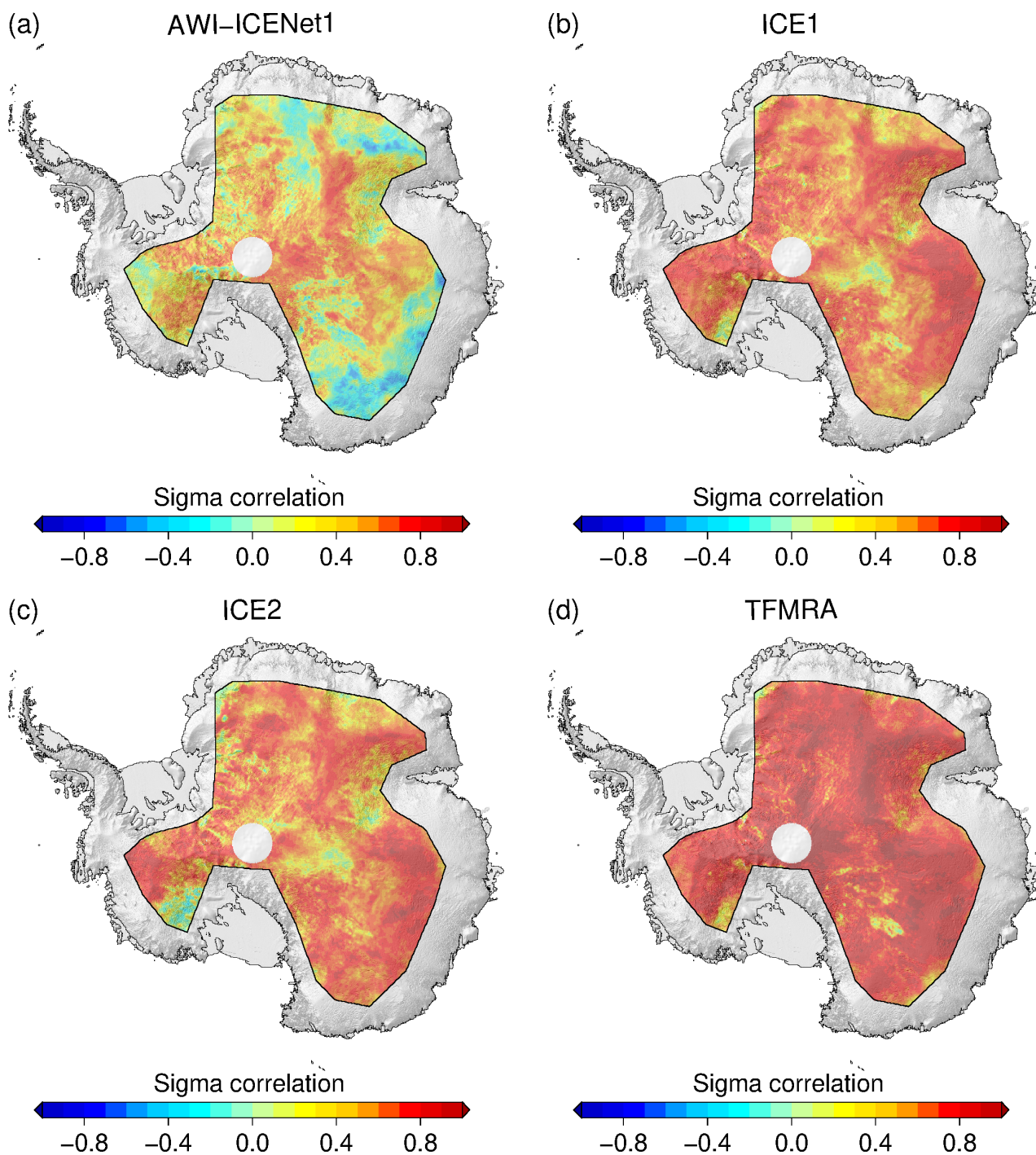
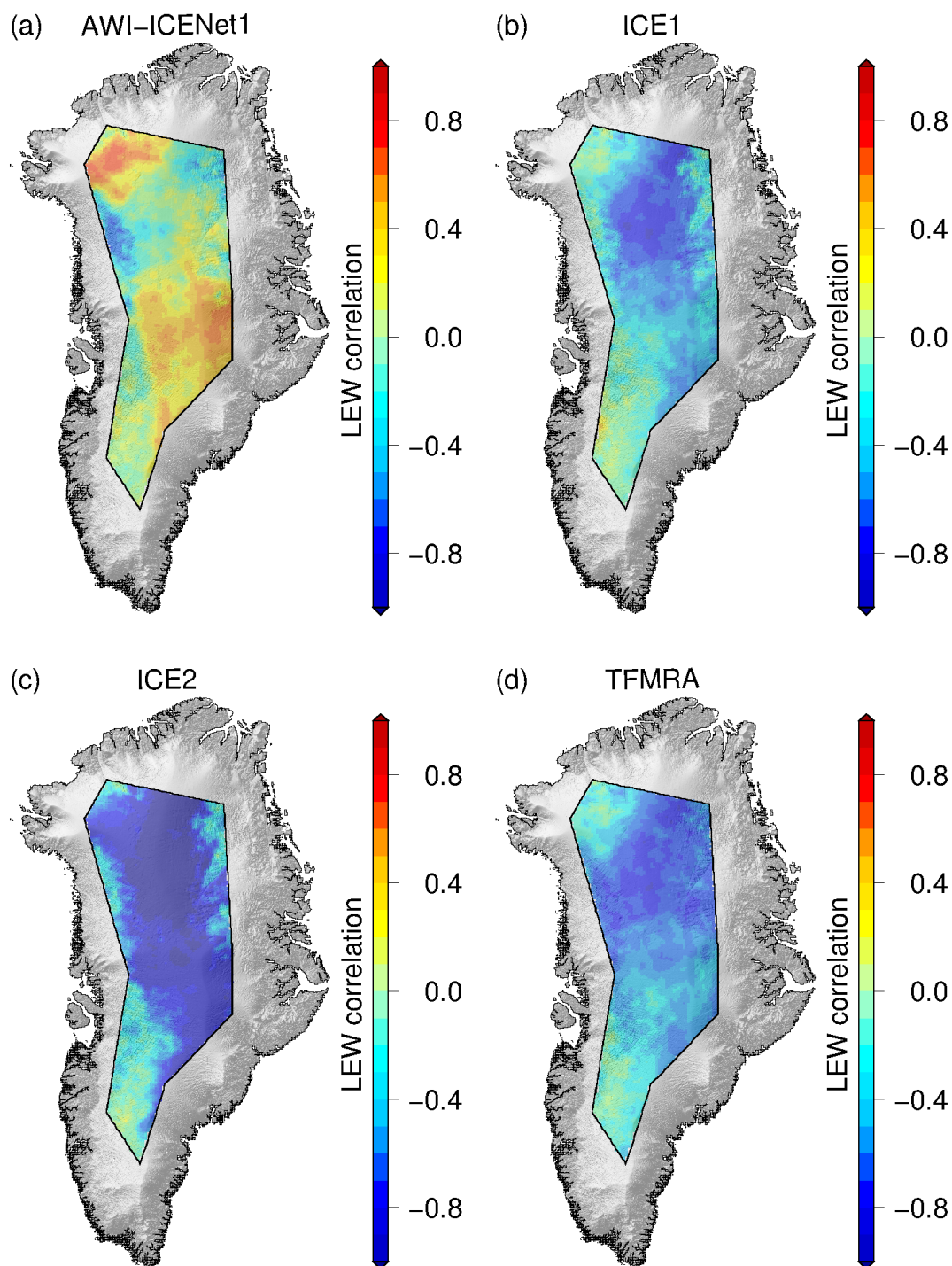


Figure A5. Correlation between elevation anomalies and LEW for different retracker over Antarctica.

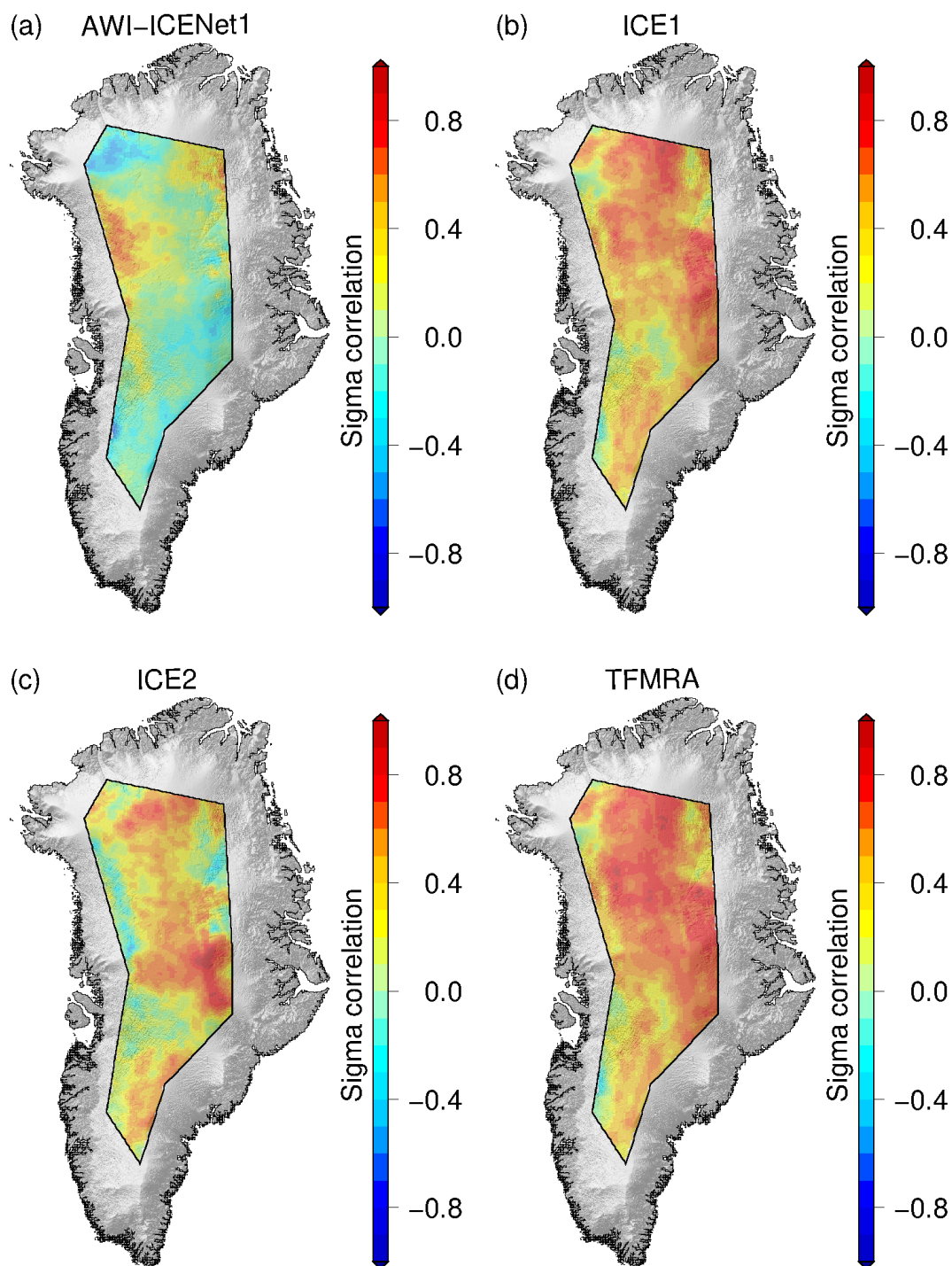


**Figure A6.** Correlation between elevation anomalies and backscatter for different retracker over Antarctica.



**Figure A7.** Correlation between elevation anomalies and LEW for different retracker over Greenland.





**Figure A8.** Correlation between elevation anomalies and backscatter for different retracker over Greenland.

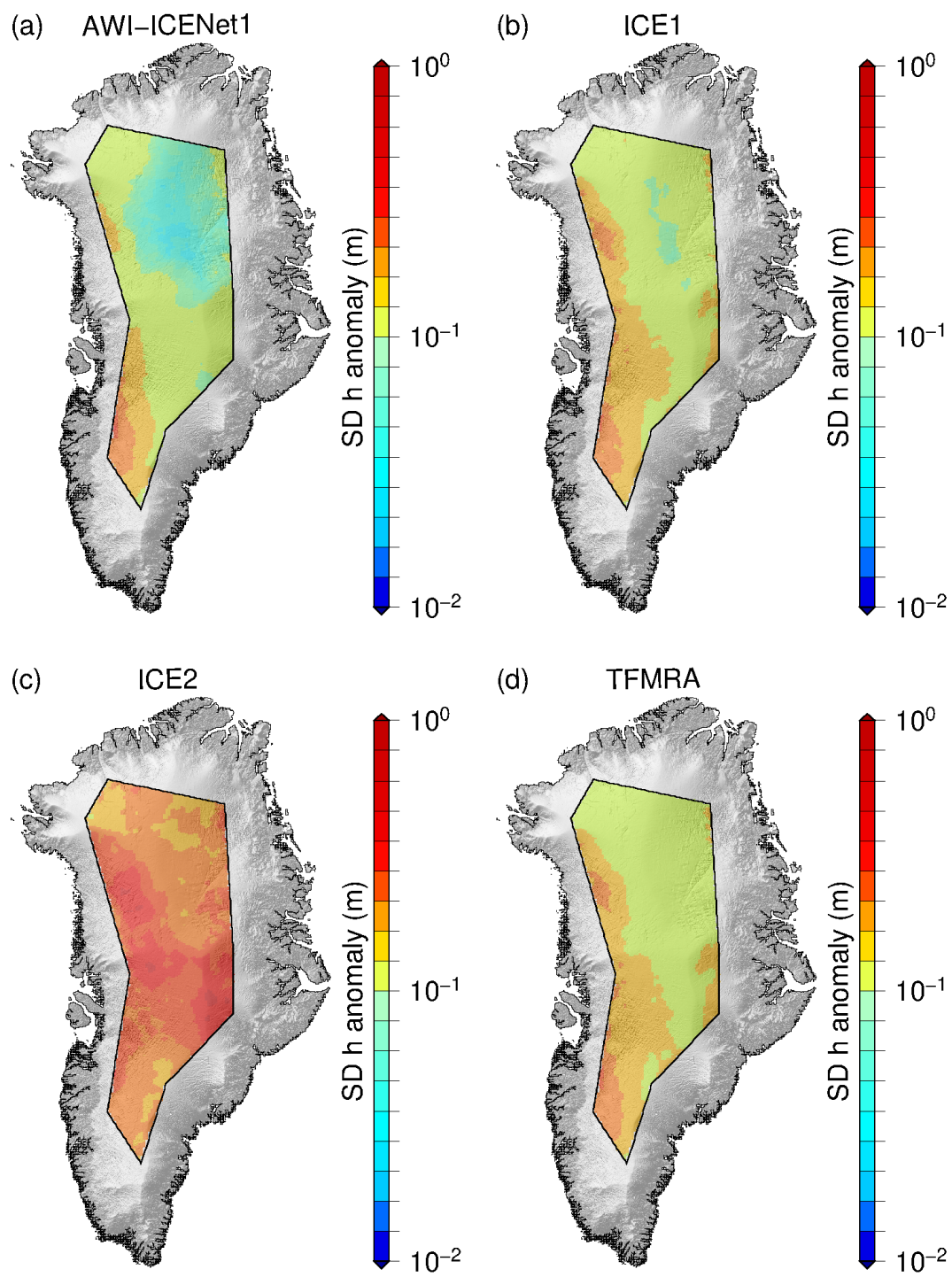
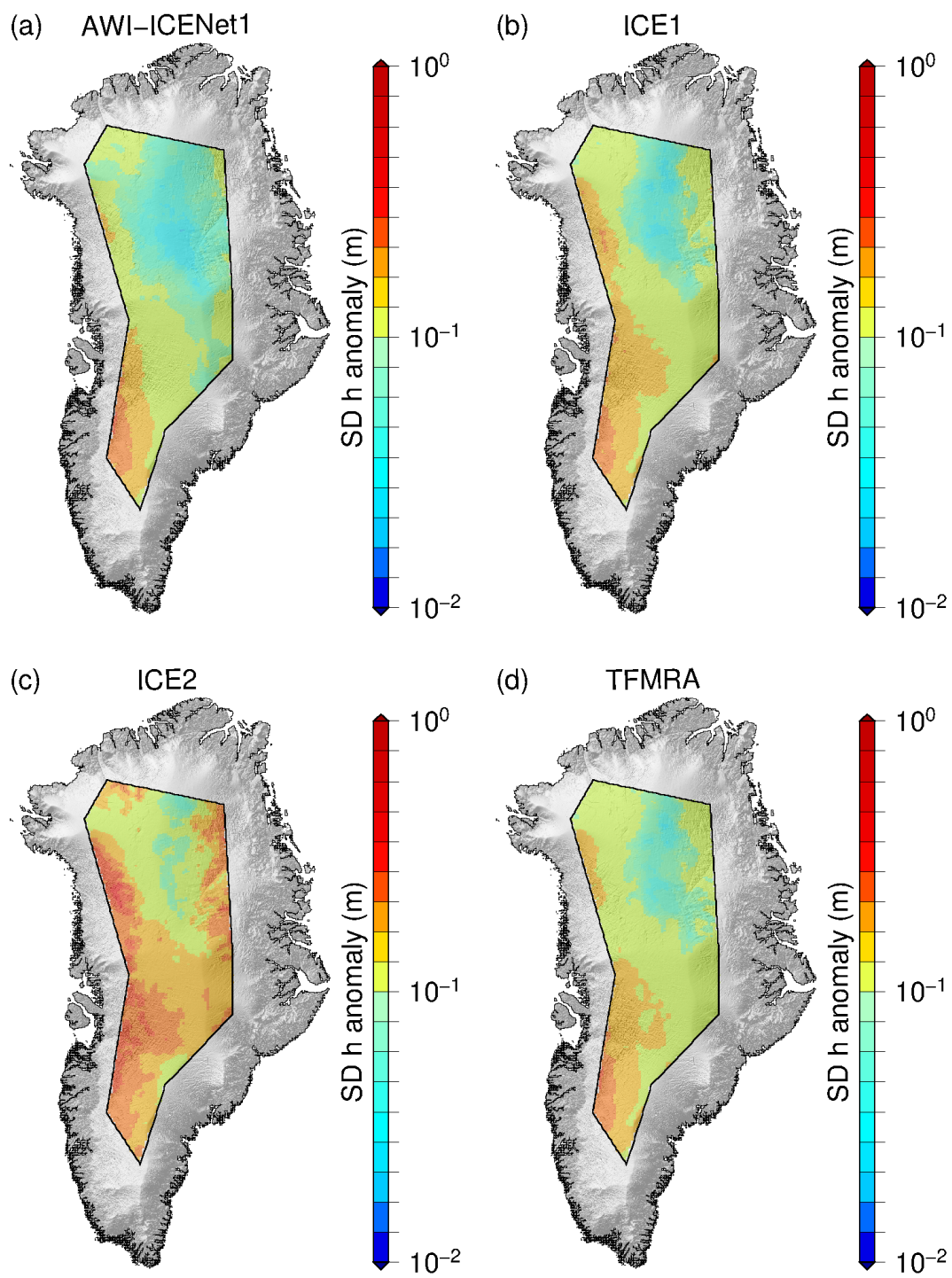
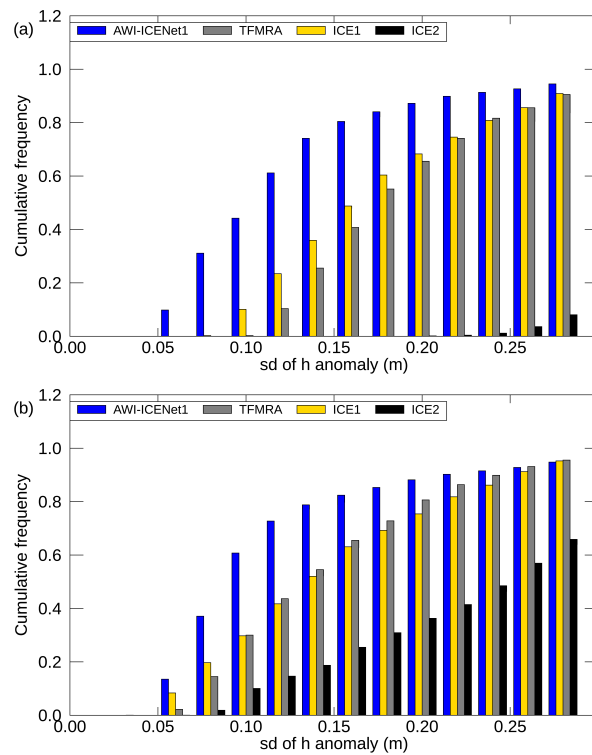


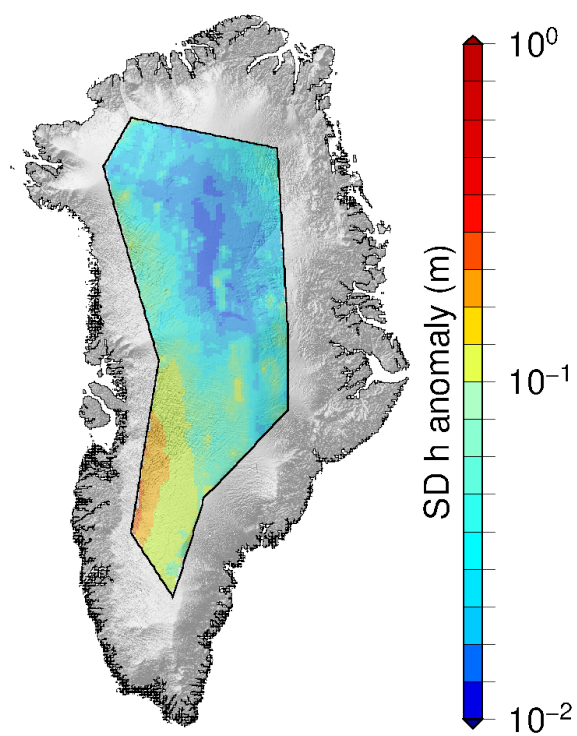
Figure A9. SD of elevation (h) anomaly over Greenland



**Figure A10.** SD of elevation (h) anomaly corrected using a correlation with leading edge width and backscatter over Greenland



**Figure A11.** SD of elevation (h) anomaly accumulated over the LRM zone in Antarctica for all retrackers. (a) uncorrected and (b) corrected h anomalies using a correlation with LEW and backscatter.



**Figure A12.** ICESat-2 derived SD of elevation (h) anomaly over Greenland in the time span from 2019/01–2021/12.

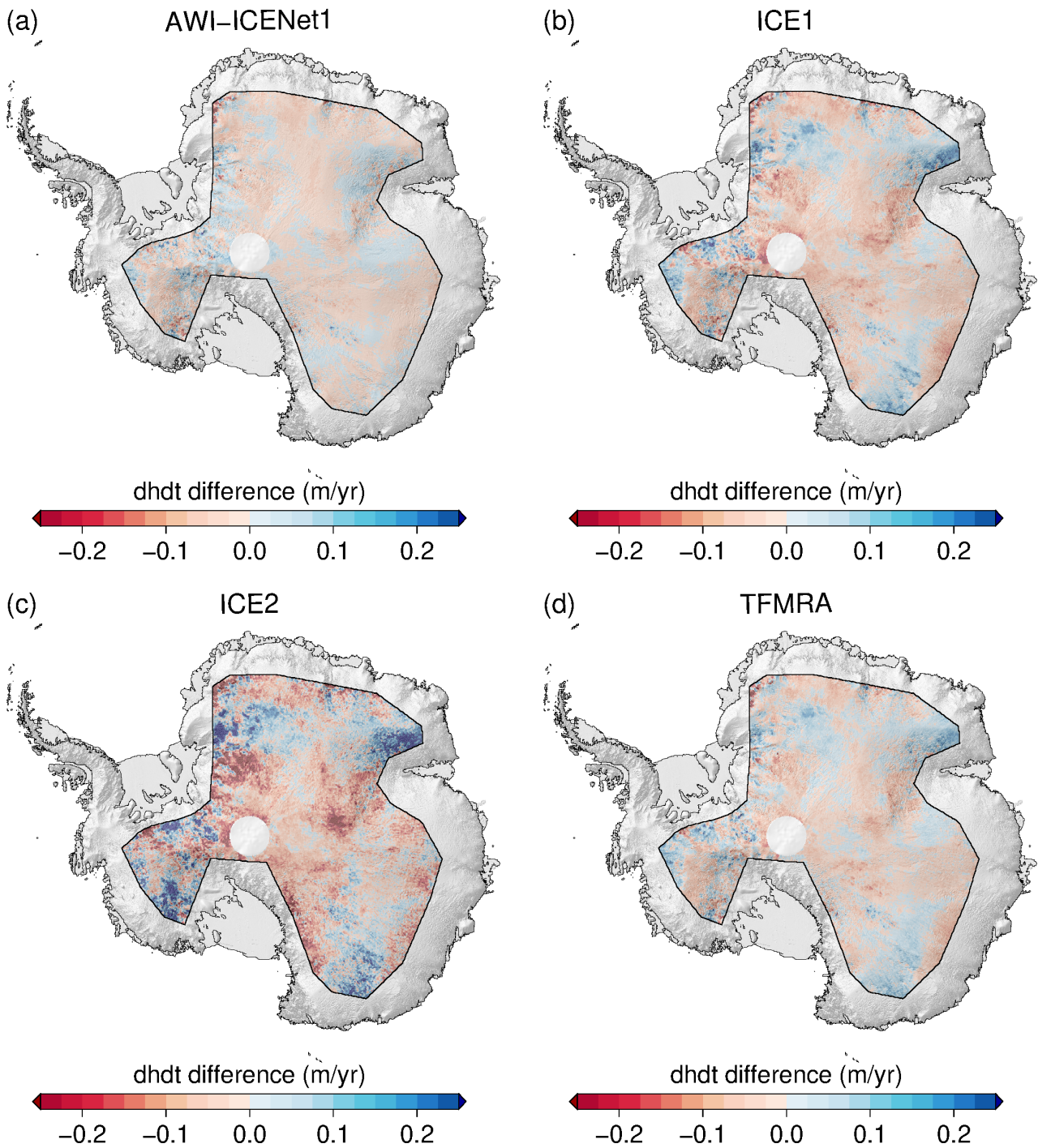
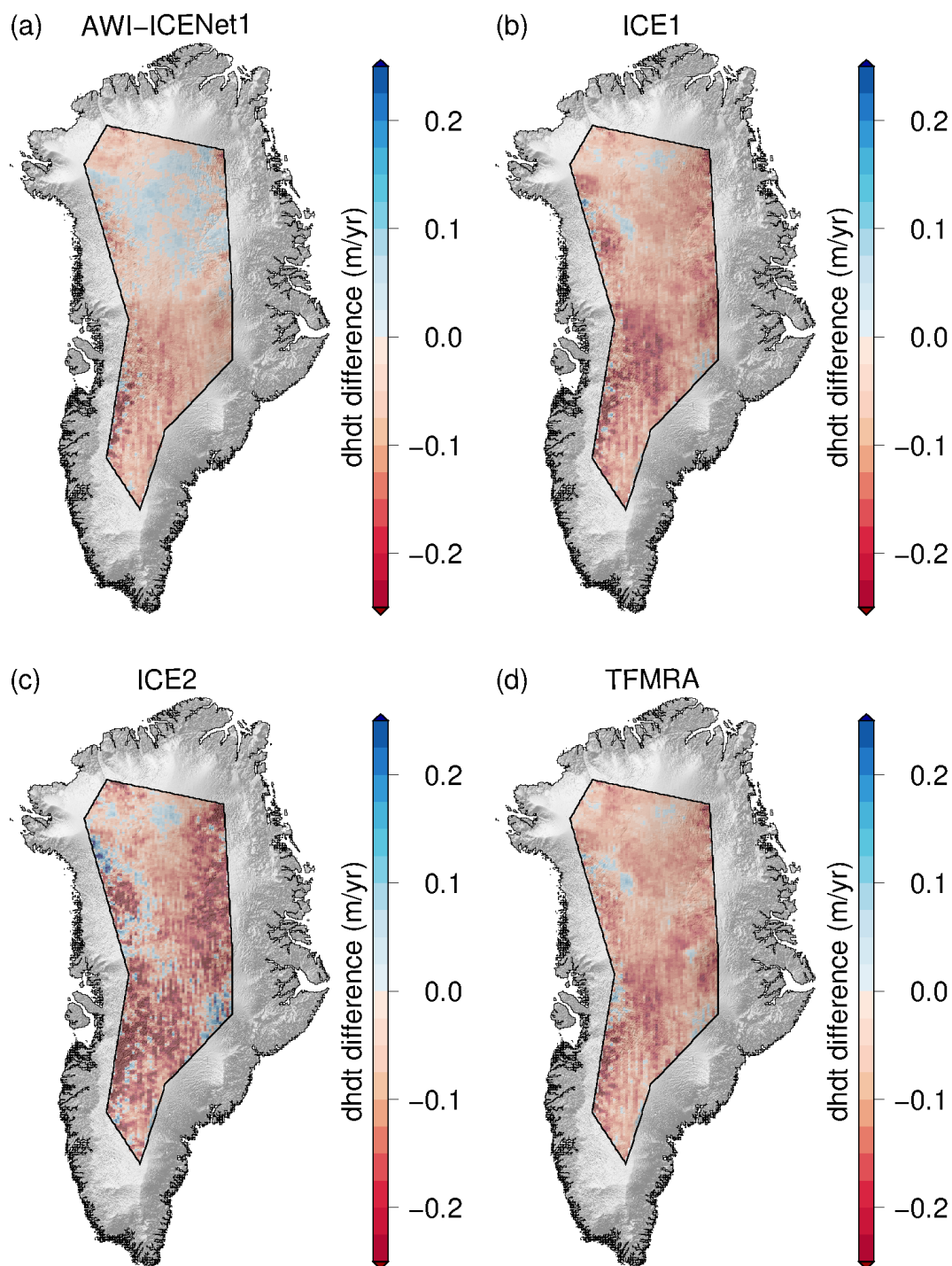
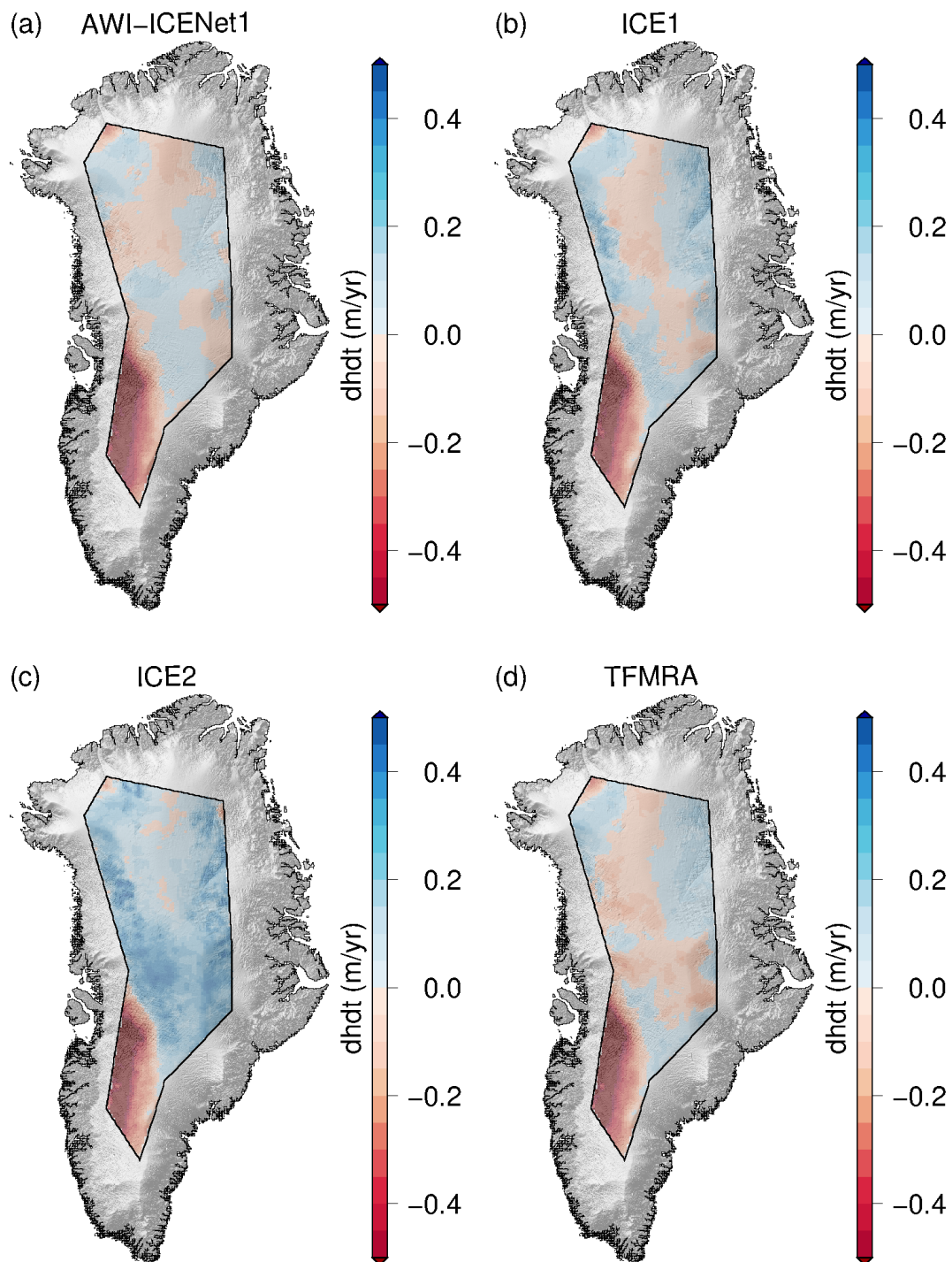


Figure A13. Antarctic wide dhdt difference to ICESat2 for the span 2019/01–2021/12 with LEW and backscatter correction



**Figure A14.** Greenland wide dhdt difference to ICESat2 for the span 2019/01–2021/12 with LEW and backscatter correction



**Figure A15.** Greenland dhdt including corrections for LEW and backscatter for the time span 2011/01–2013/12 covering the melt event July 2012.





540 *Author contributions.* V.H. led the study, simulated the waveforms, applied AWI-ICENet1 to satellite data, analysed the results and contributed figures. A.D. and E.L. developed the ConvNet. A.D. analysed results and contributed figures. V.H. and A.H. wrote large parts of the text. R.H. wrote machine learning specific sections of the text. M.H. and V.H. acquired funding for this research. All authors discussed the results and contributed to editing the manuscript.

*Competing interests.* The authors declare no competing interests.

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