



New ^{10}Be exposure ages improve Holocene ice sheet thinning history near the grounding line of Pope Glacier, Antarctica

Jonathan R. Adams^{1,2}, Joanne S. Johnson¹, Stephen J. Roberts¹, Philippa J. Mason², Keir A. Nichols²,
Ryan A. Venturelli³, Klaus Wilcken⁴, Greg Balco⁵, Brent Goehring³, Brenda Hall⁶, John Woodward⁷,
5 Dylan H. Rood²

1 British Antarctic Survey, High Cross, Madingley Road, Cambridge CB3 0ET, UK

2 Department of Earth Science & Engineering, Imperial College London, London SW7 2AZ, UK

3 Department of Earth & Environmental Sciences, Tulane University, New Orleans, LA 70118, USA

4 Australia's Nuclear Science and Technology Organisation (ANSTO), New Illawarra Road, Lucas Heights, NSW 2234,
10 Locked Bag 2001, Kirrawee DC 2232, Australia

5 Berkeley Geochronology Center, 2455 Ridge Road, Berkeley, CA 94709, USA

6 School of Earth and Climate Sciences and the Climate Change Institute, University of Maine, Orono, ME 04469 USA

7 Department of Geography and Environmental Sciences, Northumbria University, Newcastle-upon-Tyne, NE1 8ST, UK

Correspondence to: Jonathan R. Adams (j.adams19@imperial.ac.uk)

15 Co-author email list (alphabetical order):

Greg Balco – balcs@bgc.org

Brent Goehring – bgoehrin@tulane.edu

Brenda Hall – brendah@maine.edu

Joanne S. Johnson – jsj@bas.ac.uk

20 Philippa J. Mason – p.j.mason@imperial.ac.uk

Keir A. Nichols – keir.nichols@imperial.ac.uk

Stephen J. Roberts – sjro@bas.ac.uk

Dylan H. Rood – d.rood@imperial.ac.uk

Ryan A. Venturelli – rventurelli@tulane.edu

25 Klaus Wilcken – klausw@ansto.gov.au

John Woodward – john.woodward@northumbria.ac.uk



Abstract. Evidence for the timing and pace of past grounding line retreat of the Thwaites Glacier system in the Amundsen
30 Sea embayment (ASE) of Antarctica provides constraints for models that are used to predict the future trajectory of the West
Antarctic Ice Sheet (WAIS). Existing cosmogenic nuclide surface exposure ages suggest that Pope Glacier, a former
tributary of Thwaites Glacier, experienced rapid thinning in the early to mid-Holocene. There are relatively few exposure
ages from the lower ice-free sections of Mount Murphy (< 300 m asl) that are uncomplicated by either nuclide inheritance or
scattering due to localised topographic complexities; this makes the trajectory for the latter stages of deglaciation uncertain.
35 This paper presents 12 new ^{10}Be exposure ages from erratic cobbles collected from the western flank of Mt Murphy, within
160 m of the modern ice surface and 1 km from the present grounding line. The ages comprise two tightly clustered
populations with mean deglaciation ages of 7.1 ± 0.1 ka and 6.4 ± 0.1 ka (1SE). Linear regression analysis applied to the
age-elevation array of all available exposure ages from Mt Murphy indicates that the median rate of thinning of Pope Glacier
was 0.27 m yr^{-1} between 8.1 – 6.3 ka, occurring 1.5 times faster than previously thought. Furthermore, this analysis better
40 constrains the uncertainty (95 % confidence interval) in the timing of deglaciation at the base of the Mt Murphy vertical
profile (~80 m above the modern ice surface), shifting it to earlier in the Holocene (from 5.2 ± 0.7 ka to 6.3 ± 0.4 ka). Taken
together, the results presented here suggest that early–mid Holocene thinning of Pope Glacier occurred over a shorter
interval than previously assumed and permit a longer duration over which subsequent late Holocene rethickening could have
occurred.

45 1 Introduction

The Amundsen Sea Embayment (ASE), dominated by the Pine Island – Thwaites Glacier system, has recently undergone the
fastest rates of ice mass loss of all sectors of the West Antarctic Ice Sheet (WAIS), which is estimated at 136 Gt yr^{-1} (70 %
from Pine Island – Thwaites) from 2009 - 2017 (Rignot et al., 2019). Ice mass loss from WAIS is driven principally by high
rates of basal melting caused by incursions of warm Circumpolar Deep Water (CDW) onto the continental shelf (Adusumilli
50 et al., 2020; Pritchard et al., 2012). This basal melting has led to increased ice flow velocity (Rignot et al., 2014), faster
grounding line retreat (Konrad et al., 2018; Milillo et al., 2022) and dynamic ice thinning (Pritchard et al., 2009; Shepherd et
al., 2019). However, quantifying how much WAIS will contribute to future global mean sea level rise under different
emissions pathways remains subject to considerable uncertainty (Bamber et al., 2019; SROCC, 2019). Physics based ice
sheet models simulating evolution of WAIS (e.g. (Albrecht et al., 2020; Pollard and DeConto, 2009)) require geologic
55 records over centennial-millennial timescales for validation (Bentley, 2010). However, existing geologic records of ice sheet
change since the Last Glacial Maximum (LGM) in the ASE based on cosmogenic nuclide surface exposure ages (Johnson et
al., 2008, 2014, 2017, 2020; Lindow et al., 2014) are either incomplete or exhibit considerable scatter closest to the modern
ice surface. Here we present new ^{10}Be surface exposure ages from erratic cobbles that improve the thinning history of the
lowest 300 metres of presently exposed rock at Mount Murphy, a volcanic edifice adjacent to Pope Glacier in the central
60 ASE (Fig. 1a). These data fill a critical gap in the Holocene ice sheet thinning history of Pope Glacier.

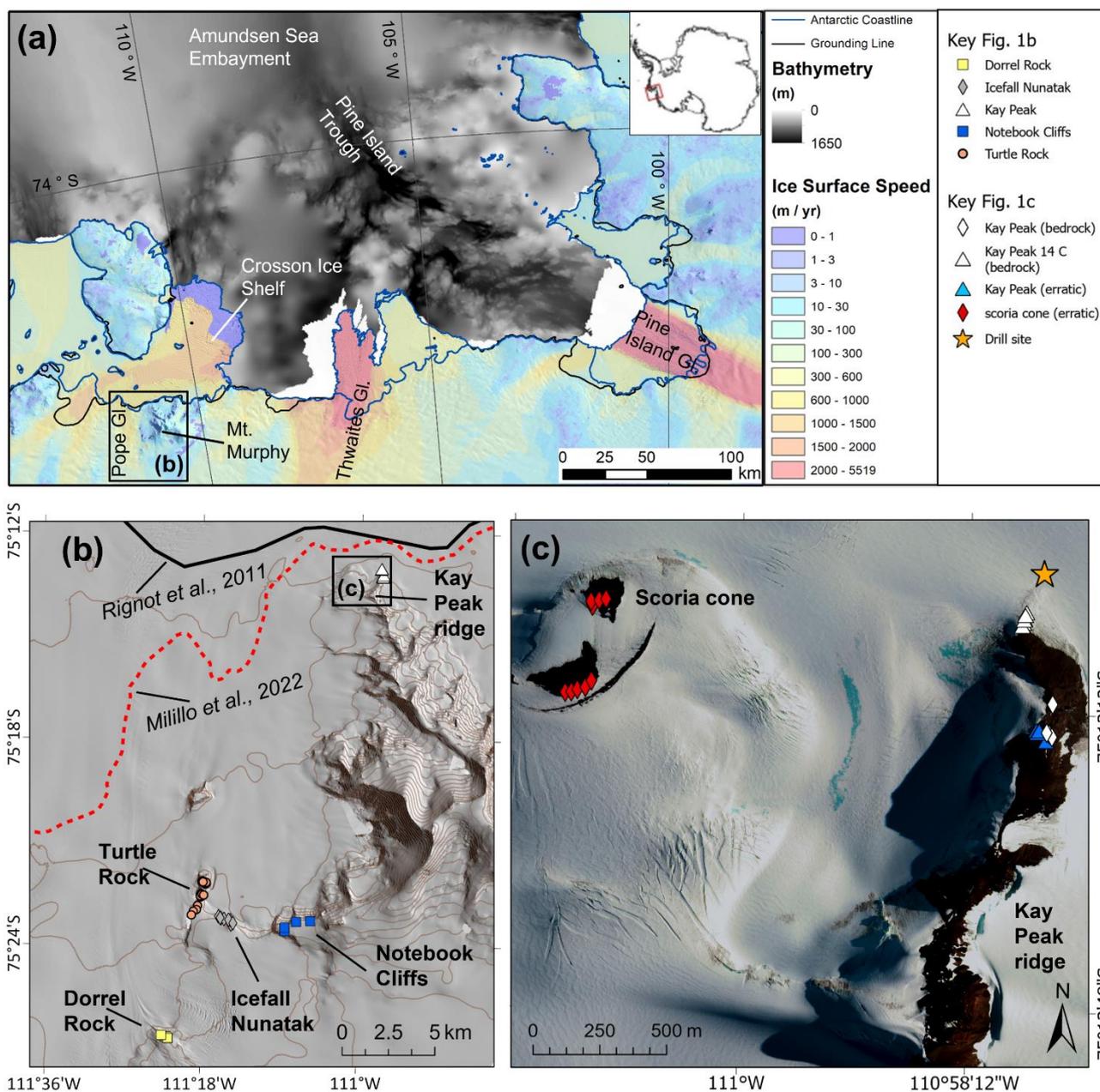


Figure 1: Location of the study area in the context of Mt. Murphy and the wider ASE. (a) ASE area map displaying regional ice streams, ice velocities (Mouginot et al., 2012) overlaid on Reference Elevation Model of Antarctica (REMA) (Howat et al., 2019) regional bathymetry (Arndt et al., 2013), 2011 grounding line (black line; Rignot et al., 2011) and Antarctic coastline (blue; Antarctic Digital database, version 7.2). (b) Mt. Murphy site map generated using REMA derived hill shade, with contours representing 100 m intervals. Locations of previous exposure ages (Johnson et al., 2020, 2008) indicated by coloured symbols shown in Fig. 1b key, also displayed are both the 2011 grounding line (same as in Fig. 1a), and most recent localised grounding line (red dotted; Milillo et al., 2022). (c) High resolution satellite imagery (732 false colour composite), DigitalGlobe Products, WorldView 2 © 2013 DigitalGlobe, Inc., a Maxar company showing scoria cone and Kay Peak Ridge. Drill site indicates subglacial bedrock core recovery by the Geological History Constraints (GHC) subglacial drilling project (Boeckmann et al., 2021).



This study applies the method of cosmogenic nuclide surface exposure dating to determine the timing and rate of ice surface elevation change since the Last Glacial Maximum (LGM). This ice surface history is achieved by measuring the cosmogenic nuclide content of bedrock and glacial deposits, such as cobbles or boulders, which have undergone englacial transport (Ackert et al., 1999; Stone et al., 2003; Johnson et al., 2014). Several cosmogenic surface exposure studies reconstructing post LGM changes in ice surface thickness have been conducted in the ASE (Lindow et al., 2014; Johnson et al., 2008, 2014, 2020). These previous studies have improved our understanding of the post-LGM deglaciation history of outlet glaciers in the central and eastern embayment. From these previous studies, we have a relatively good understanding of the trajectory of ice sheet thinning after the LGM until the mid-Holocene, but few exposure age arrays extend close to the modern ice sheet surface. At Mt. Murphy, where a high density of samples from higher elevations have been analysed, exposure ages at elevations within < 200 m of the modern ice surface are either absent or scattered (fig. 5, Johnson et al., 2020). This age uncertainty makes it difficult to determine whether the rapid thinning indicated by the higher elevation exposure age data slowed down before reaching its modern elevation, or whether the fast rate of thinning observed between 9-6 ka continued after that time. Furthermore, the time when the ice surface reached its modern elevation cannot be determined from the existing data because no records of thinning in the last 5 ka in the ASE currently exist from above the modern ice surface (Johnson et al., 2014, 2020; Lindow et al., 2014). This existing data is ambiguous and implies either that ice thickness - and by inference grounding line position of Pope Glacier - has been largely stable over this period, or that any evidence for late Holocene thinning/retreat is currently below the present ice surface (Johnson et al., 2021b).

To improve our understanding of Holocene ice sheet history in the Pine Island - Thwaites Glacier region, we focus here on the area within 1 km of the grounding line at Mt Murphy (Milillo et al., 2022) (Fig. 1). The base of Kay Peak ridge was the location for a subglacial bedrock drilling campaign undertaken in 2019-20 by the International Thwaites Glacier Collaboration (Fig. 1c). The aim of that campaign was to collect bedrock cores and measure cosmogenic nuclide concentrations within them to detect whether the ice sheet surface was ever lower than present in the past few millennia. Here we present new ^{10}Be surface exposure ages for 12 erratic cobbles collected from the surface of a scoria cone situated 1.6 km west of the drill site. The current understanding of the thinning history of Pope Glacier and mid - late Holocene ice sheet configuration in the ASE is improved by dating erratic cobbles from this site for the following three reasons. First, the lower section of the Mt Murphy vertical thinning transect is currently poorly constrained, largely due to an absence of exposure ages from 240 – 100 m above the modern ice surface (320 – 180 m asl). There is also spread in the array of existing exposure ages (both ^{10}Be and to a lesser extent ^{14}C) from Kay Peak at the base of the profile 100 – 80 m above the modern ice surface (see fig. 5, Johnson et al., 2020). The scoria cone is situated at the ideal elevation (180 – 240 m asl) to fill this data gap and better constrain the deglacial history. Second, the scoria cone site is situated in close proximity (0.7 km) to the current grounding zone of Pope Glacier (Fig. 2) (Milillo et al., 2022). Thus, this site is expected to be highly sensitive to present and past changes in grounding line position during the mid-late Holocene. Finally, exposure ages from the site will directly inform interpretation of the cosmogenic nuclide concentrations within the subglacial bedrock cores that were



105 recovered from the projection of Kay Peak ridge under Pope Glacier, which are critical to determining if the ice surface was
thinner than present at any time in the last 5 ka.

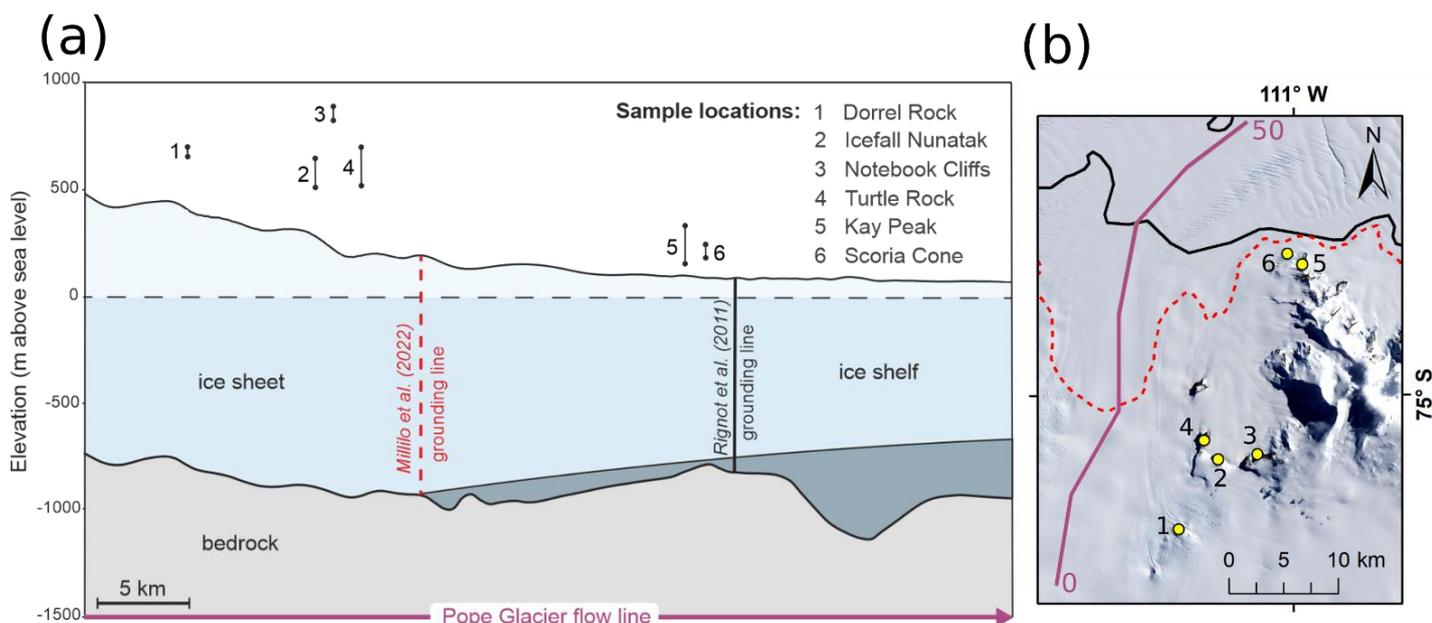


Figure 2: Relationship between nunatak elevation, bedrock and ice sheet topography (a) Cross-section transects showing modern ice surface elevation data (line above sea level; from Reference Elevation Model for Antarctica (REMA); (Howat et al., 2019) and bedrock topography (line below sea level; from BedMachine; (Morlighem et al., 2018)). Note that the glacier flowline profile (solid purple line in panel b) is >10km from the sample sites; elevation range markers on the cross-section represent the highest and lowest elevation sample at each site (m asl) (Johnson et al., 2020) and are displayed at points approximately adjacent to flowline on the right hand panel. (b) Landsat-8 true colour composite (432) image showing location of Pope Glacier flowline with distance downstream from arbitrary starting position. Profile is displayed in the cross section above. Elevation range markers on the cross-section (a) are displayed as yellow dots approximately adjacent to flowline on panel b. Landsat-8 image courtesy of the U.S. Geological Survey; <https://doi.org/10.5066/P9OGBGM6>.

2. Site Description and Methods

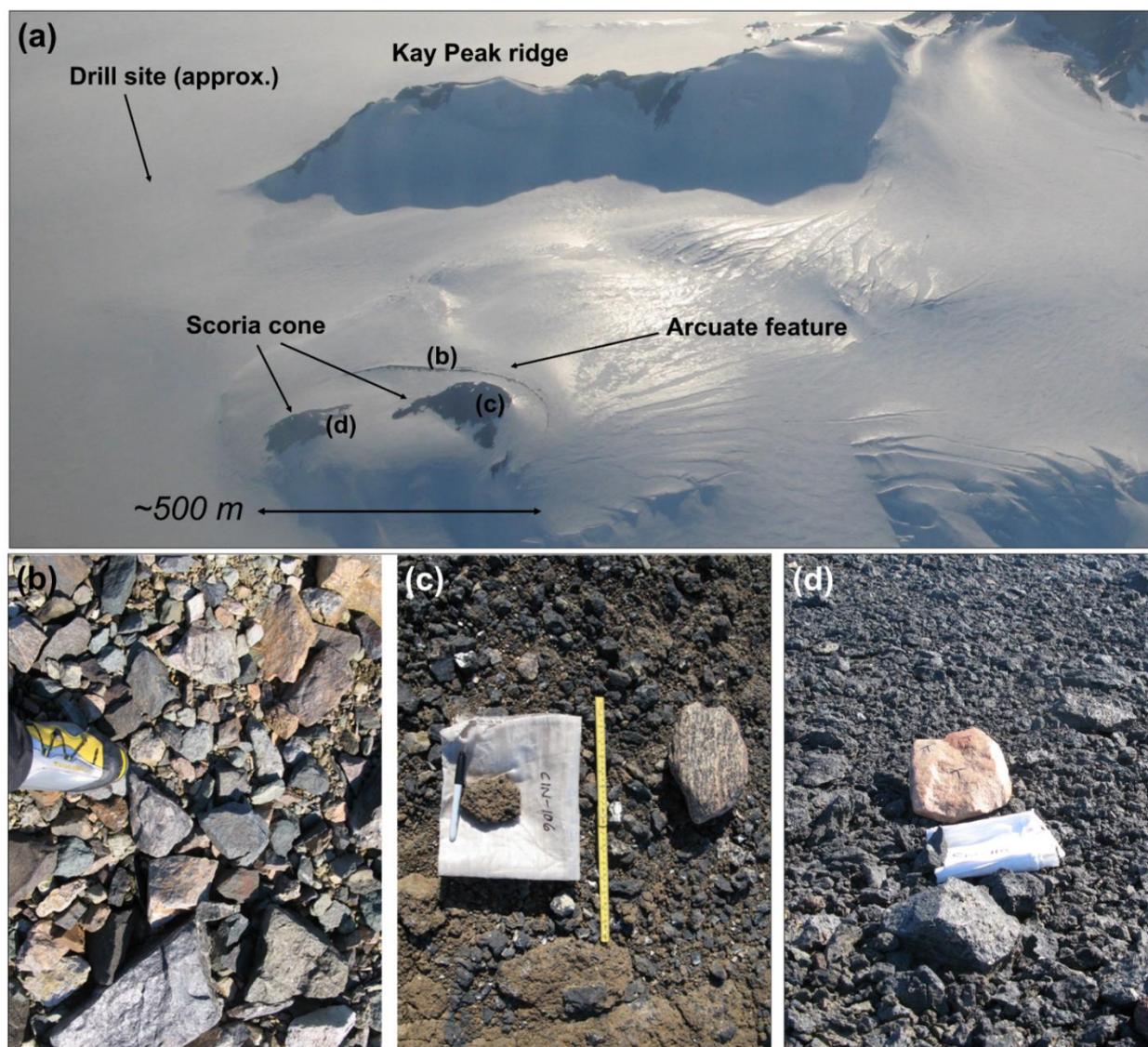
2.1 Site Description

Mt. Murphy, which reaches a maximum elevation of 2703 m asl, is located approximately 50 km from the modern Thwaites
120 Glacier ice stream and is bounded by Crosson Ice shelf to the north and Pope Glacier to the west (Johnson et al., 2020) (Fig
1). Pope Glacier is approximately 14 km wide and flows into the Crosson Ice Shelf at a velocity of 0.8 km yr⁻¹ (Mouginot et
al., 2012). Mt. Murphy hosts an abundance of glacial deposits including numerous erratics and striated bedrock surfaces.
These erratics and striations indicate that glacial ice previously covered and flowed over the lower section of Mt Murphy
including scoria cone and Kay Peak ridge (Johnson et al., 2020).

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We collected 12 glacially deposited erratic cobbles for cosmogenic surface exposure age dating from two outcrops on the scoria cone, which is situated < 1 km from the present grounding line of Pope Glacier (Milillo et al., 2022) (see Fig. 1b). The two bedrock outcrops onto which the erratics were deposited, hereafter referred to as outcrop A (upper) and outcrop B (lower), mostly consist of rubbly oxidised scoria accompanied by smaller exposures of hyaloclastite breccia.



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Figure 3: Geomorphic difference between clasts deposited at the scoria cone. (a) Image showing location of scoria cone site in relation to Kay Peak ridge. Horizontal black line indicates approximate field of view. (b) Clasts forming the arcuate ridge landform. (c and d) Erratic cobbles (samples CIN 106 and CIN 110) perched on the bedrock surface at scoria cone. Approximate locations of clasts displayed in (b – d) shown in top panel. Scale indicated by boot (b), 50 cm ruler (c). Images of other scoria cone erratics collected for exposure dating are provided in the supplementary material (Fig. S3). Photo credit: Joanne S. Johnson and Stephen J. Roberts.



The samples collected from the scoria cone range in altitude from 178-239 m asl., which equates to an elevation above the modern ice surface of ~100-160 m, where the modern ice surface elevation is 80 m asl at the scoria cone site. Erratic cobbles, primarily composed of gneiss and granite rock types, are present on the surface of both outcrops and provide evidence that the site was previously overridden by ice. Of the 12 erratic cobbles that we analysed, 6 are medium cobbles (long axis 15-50 cm) and 6 are small cobbles (long axis < 15 cm). Half the cobbles (n = 6) were sub-rounded and only one of the cobbles collected was angular (CIN -111) (Table S2). The characteristics of erratics deposited on scoria cone, such as clast shape and lithology, differ from those of material deposited at a nearby landform that forms an arcuate ridge (Fig. 1c, Fig. 3). Material from this landform is predominantly angular, indicating it is locally derived (probably from Kay Peak), whilst erratics deposited at scoria cone are more rounded. The roundness of many erratics emplaced at scoria cone is evidence that they were subject to prolonged erosion during englacial transport (Darvill et al., 2015). Moreover, six of the erratics at scoria cone are granitic (granite/aplite), a rock type not common on Mt Murphy or surrounding nunataks. Erratics of similar shape and rock type have been observed at other nunataks surrounding Mt. Murphy [Turtle Rock and Icefall Nunatak (Fig. 1; Johnson et al., 2020)], indicating that the glacial deposits selected for ¹⁰Be analysis from scoria cone originated from a source upstream rather than being locally derived from Kay Peak ridge.

2.2 Analytical Methods

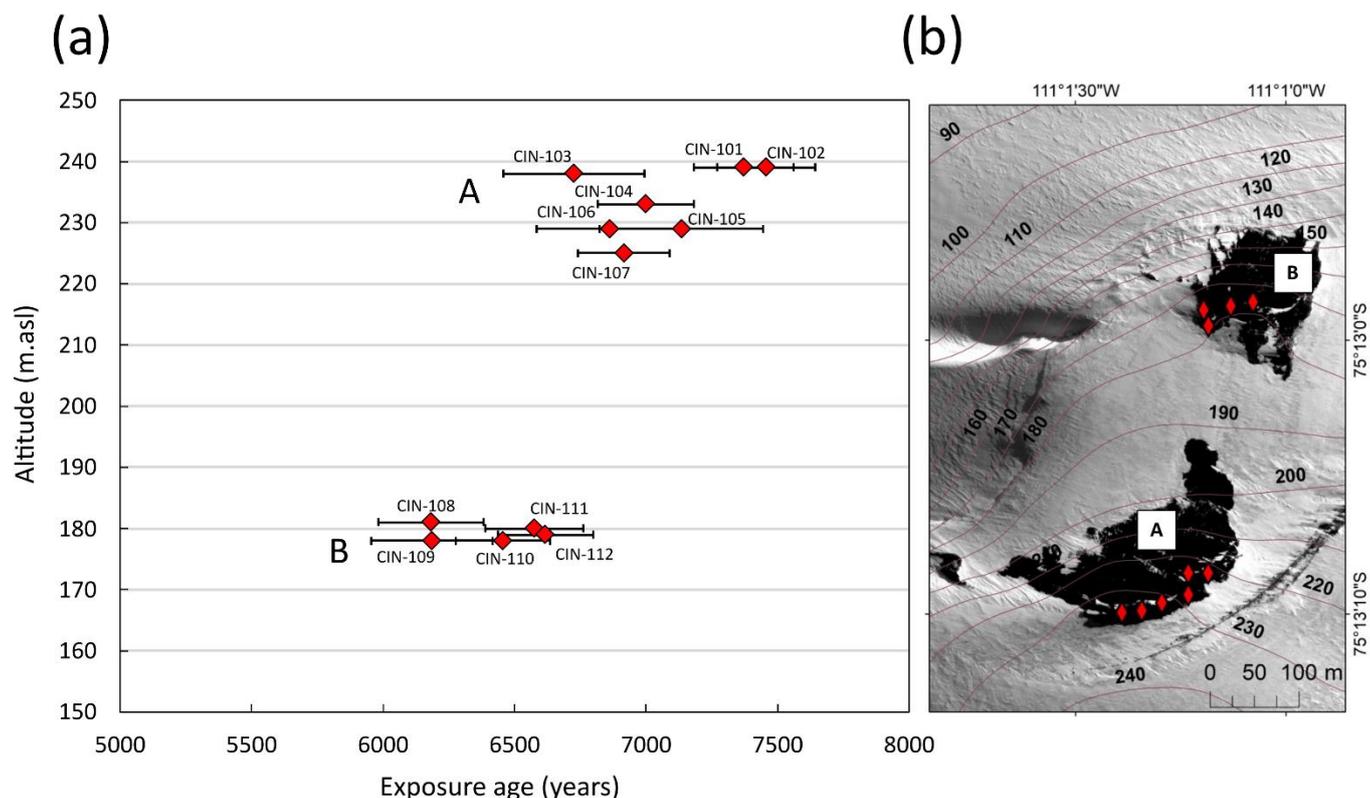
Quartz mineral separation and ¹⁰Be isotope dilution chemistry was conducted in the CosmIC Laboratory at Imperial College London using standard procedures (Kohl and Nishiizumi, 1992; Corbett et al., 2016). After beryllium was extracted and purified from quartz, the samples were loaded into cathodes for ¹⁰Be measurement by accelerator mass spectrometry (AMS). These AMS measurements were performed at the Australian Nuclear Science and Technology Organisation (ANSTO) (Wilcken et al., 2017). Measurements were then normalised to the KN-5-3 standard with an assumed ratio of 6.320×10^{-12} ($t_{1/2} = 1.36$ Ma; (Nishiizumi et al., 2007)). Exposure ages were calculated using version 3 of the online calculators at hess.ess.washington.edu (Balco et al., 2008). The online calculators use the LSDn production rate scaling method for neutrons, protons and muons following Lifton et al., 2014 and summarised in Balco, 2017 and the primary production rate calibration data set of (Borchers et al., 2016). In order to keep the input parameters consistent with those used by Johnson et al., 2020, all exposure ages are reported assuming no erosion or snow cover and a material density of 2.7 g cm^{-3} .



3. Results

165 The ^{10}Be exposure ages obtained from the scoria cone range from 7.5 ± 0.5 ka to 6.2 ± 0.4 ka (1σ external errors on
individual ages throughout, unless otherwise noted), and the average exposure age is 6.8 ka ($n=12$) (Table 1). The exposure
ages are clustered in two separate groups (Fig. 4a), which correspond to outcrops A and B (Fig. 4b). Outcrop A contains a
cluster of ages at an elevation range 225–239 m asl, whereas outcrop B ages are all clustered at a narrower elevation range of
 ~ 180 m asl. We interpret the ages as representing the timing of the most recent deglaciation of the two scoria cone outcrops.

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175 **Figure 4: Exposure ages from the two scoria cone outcrops.** (a) ^{10}Be exposure ages versus sample altitude. Error bars displayed represent internal uncertainty of AMS measurement on samples (1σ). (b) Very high-resolution panchromatic satellite image of scoria cone site, credit; DigitalGlobe Products, WorldView 2 © 2013 DigitalGlobe, Inc., a Maxar company. Purple elevation contour lines in units of metres. Red diamonds represent the location of each sample. The separate outcrops are labelled as A and B and correspond to the upper and lower population clusters displayed in panel a, respectively. Note only 10 sample locations are displayed on panel b as some of the samples were so close to each other they were recorded with the same GPS location. See Table S1 for analytical data.



180 The two clusters of ages each correspond to a distinct outcrop; the higher elevation cluster ($n = 7$) were all obtained from
 outcrop A and the lower elevation cluster ($n = 5$) from outcrop B (Fig 4b). The reduced χ^2 (χ^2v) and probability (p) values for
 ages from each outcrop (Outcrop A: $\chi^2v = 1.67$, p value ≥ 0.01 ; Outcrop B: $\chi^2v = 1.11$, p value ≥ 0.01) indicate that the
 scatter in exposure ages at each outcrop can be attributed to analytical uncertainty alone (Balco, 2011). However, when the
 same statistical test is performed across both outcrops the reduced χ^2v value is 4.28 and p value is < 0.01 . Taken together,
 185 these statistical analyses are consistent with the interpretation that the ages from Outcrop A and B are two statistically
 different populations and represent distinct times of deglaciation. Therefore, the error weighted mean standard error provides
 the best estimate for the outcrop deglaciation ages: 7.1 ± 0.1 ka for outcrop A and 6.4 ± 0.1 ka for outcrop B.

Sample Group Name	No. of samples (n)	Sample range elevation (m asl)	Sample range elevation above modern ice (m)	Error weighted mean (ka)	Std. error (yrs.)	Ext. error (yrs.)	Reduced χ^2 (χ^2v)	p value
Scoria cone outcrop A	7	239 - 225	159 - 145	7.1	80	426	1.67	0.1240
Scoria cone outcrop B	5	181 - 178	101 - 98	6.4	86	390	1.11	0.3521
Scoria cone (comb.)	12	239 - 178	159 - 98	6.8*			4.28	0.0000
Lower Kay Peak (^{14}C)	6	170 - 150	100 - 80	5.8*			15.37	0.0000
Rev. lower Kay Peak (^{14}C)	4	167 - 154	97 - 84	6.5	252	303	3.71	0.0110

190 **Table 1: Summary statistics of new ^{10}Be exposure ages from scoria cone and previously published ^{14}C Kay Peak ridge bedrock
 ages.** Table includes error weighted mean exposure age, reduced χ^2 (χ^2v) value, and p value from each sample group. When p value > 0.01 ,
 the error weighted mean value and standard error is reported for that statistically significant population. When the reduced χ^2v p value is $<$
 0.01, the arithmetic mean value indicated by * is provided without an associated uncertainty because the reduced χ^2v and p value are not
 195 consistent with a single statistically significant population (see 3. Results section for details). The sample elevation range above the
 modern ice surface is relative to the ice sheet elevation adjacent to Kay Peak ridge reported in Johnson et al., (2020) and using 80 m asl as
 the modern ice surface elevation at scoria cone site.

We now compare the new scoria cone exposure ages to previously published in situ ^{14}C bedrock exposure ages from lower
 200 Kay Peak ridge (Fig. 1c). The existing six Kay Peak ridge samples measured for in situ ^{14}C are from a similar elevation to
 the scoria cone samples. They range from 100 – 80 m above the modern ice surface (170 – 150 m asl) (Johnson et al., 2020),
 with exposure ages spread between 8.0 ± 0.6 ka and 3.7 ± 0.3 ka. The range in elevation of samples measured for ^{10}Be from
 scoria cone is similar, 160 – 100 m above the modern ice surface (240 – 180 m asl), but the age range is smaller than at
 lower Kay Peak ridge (7.5 ± 0.4 ka to 6.2 ± 0.5 ka, $n=12$). The average in situ ^{14}C age for the lower Kay Peak ridge samples
 205 is 5.8 ± 1.4 ka (1SD, $n=6$; Johnson et al., 2020), and they have $\chi^2v = 15.37$ and p value < 0.01 (Table 1). This indicates that
 the six in situ ^{14}C ages have scatter above that expected from analytical uncertainties alone and, therefore, are not from a
 single statistically significant age population (cf. section 2.3, Jones et al., 2019).



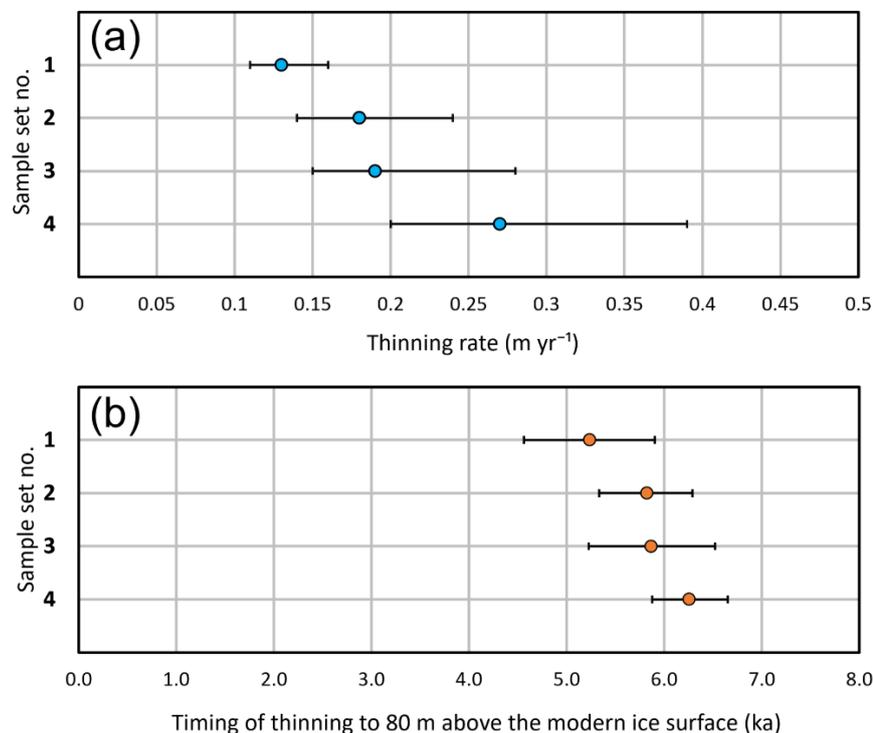
210 The six Kay Peak ridge samples were further evaluated by performing a kernel density estimation (Lowell, 1995). In the
kernel density estimate plot (Fig. A1), the two youngest in-situ ^{14}C ages from Kay Peak (from samples KAY-105 and KAY-
109), constitute a second distinctive but smaller peak. Removing these two in-situ ^{14}C ages results in a more normal
distribution (Fig. A2) consistent with a single age population, implying that in-situ ^{14}C exposure ages measured from KAY-
105 and KAY-109 are outliers. The removal of KAY-105 and KAY-109 provides a revised ^{14}C based mean deglaciation age
215 for Kay Peak ridge of 6.5 ± 0.3 ka (1SE, $n=4$), with a reduced $\chi^2_\nu = 3.29$ and p value ≥ 0.01 (Table 1). This revised mean
deglaciation age falls between the mean ^{10}Be deglaciation ages for scoria cone outcrops A and B (7.1 ± 0.1 and 6.4 ± 0.1 ka,
respectively). The two in-situ ^{14}C ages for KAY-105 and KAY-109 were included in the previously published linear
regression analysis of Johnson et al. (2020). Further linear regression analysis was conducted to understand the impact of
removing these two exposure ages on our interpretations of the thinning history of Pope Glacier. The results and implications
of this sensitivity test are described in section 3.1.

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3.1 Results of exposure age linear regression analysis

In the following section, we discuss the ice surface thinning rates and best fit constraints for end of thinning, which we
define as the time that the ice surface lowered to 80 m above the modern ice surface, i.e., the elevation of the lowermost
sample included in the linear regression transect. The thinning rates and best fit were calculated using the iceTEA Monte
225 Carlo linear regression model (Jones et al., 2019). This model randomly applies a least squares regression to exposure ages
which are normally distributed through a Monte Carlo simulation (Jones et al., 2019). All scoria cone samples were first
evaluated for their inclusion in the revised input dataset following the principles outlined in Johnson et al., 2020, whereby
samples would be removed if they 1) exhibited ^{10}Be inheritance or 2) were anomalously young (> 2 standard deviations from
the mean). Average ice surface lowering rates for Pope Glacier were then calculated using 5000 iterations of linear
230 regression through the ^{10}Be and ^{14}C ages and their internal uncertainties.

We used four different age datasets for the linear regression analysis. These different input datasets permit quantification of
the improvement of our new ages on the thinning history of Pope Glacier and enabled us to conduct a sensitivity analysis on
the inclusion or exclusion of different age data. The four sample sets are: Sample set 1 – the original dataset used in Johnson
235 et al., 2020, which serves as a baseline comparison for sample sets 2 – 4; sample set 2 – the Johnson et al., 2020 dataset and
our 12 new ^{10}Be exposure ages from scoria cone; sample set 3 – the Johnson et al., 2020 dataset with KAY-105 and KAY-
109 ^{14}C exposure ages removed; and sample set 4 – the Johnson et al., 2020 dataset with KAY-105 and KAY-109 ^{14}C
exposure ages removed and the 12 new ^{10}Be exposure ages from scoria cone included.



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Figure 5: Thinning rates and age constraints from linear regression analysis. (a) range in thinning rates (m yr^{-1}) compiled from linear regression histograms (Fig. A3) and (b) uncertainty range in best fit timing of thinning to 80 m above the modern ice surface (ka) calculated for each of the different input data to the linear regression Monte Carlo simulation (Fig. A4). Blue circles in (a) represent the median thinning rate and orange circles (b) represent the endpoint of the best fit thinning line. Error bars represent 68 % confidence interval uncertainty estimates in plot (a) and 95 % confidence interval uncertainty estimates in plot (b). The sample set no. indicates the four different age datasets used for linear regression analysis, which are described in the text above (section 3.1). Note plot b, sample sets 1 and 4 linear regression best fit end ages correspond to the linear regression transects displayed in Figure 6.

Our preferred input dataset, sample set 4, which includes the addition of our new exposure age data from scoria cone and omission of KAY-105 and KAY-109, changes the average thinning rate from $0.13 +0.03/-0.02 \text{ m yr}^{-1}$ between $\sim 9 - 5 \text{ ka}$ (Fig. 5a) (Johnson et al., 2020) to $0.27 +0.12/-0.07 \text{ m yr}^{-1}$ between $8.1 - 6.3 \text{ ka}$ (Fig. 5b) The difference in the median thinning rate is 0.14 m yr^{-1} , which is an increase of 52% on the previously published thinning rate. The ranges of thinning rates (Fig. 5a) derived using Sample Set 1: $0.11 - 0.15 \text{ m yr}^{-1}$ and Sample Set 4: $0.2 - 0.39 \text{ m yr}^{-1}$ do not overlap within 68 % confidence intervals. The range in thinning rates derived from Sample Set 2 of $0.14 - 0.24 \text{ m yr}^{-1}$, however, does overlap within the 68 % confidence interval uncertainty range with the previously published rate from Johnson et al., 2020, with an increase in the median thinning rate of 28 %. Considered in isolation, the removal of KAY-105 and KAY-109 (Sample Set 3) increases the median rate of thinning (m yr^{-1}) by 32 % compared to the data published in Johnson et al., 2020.

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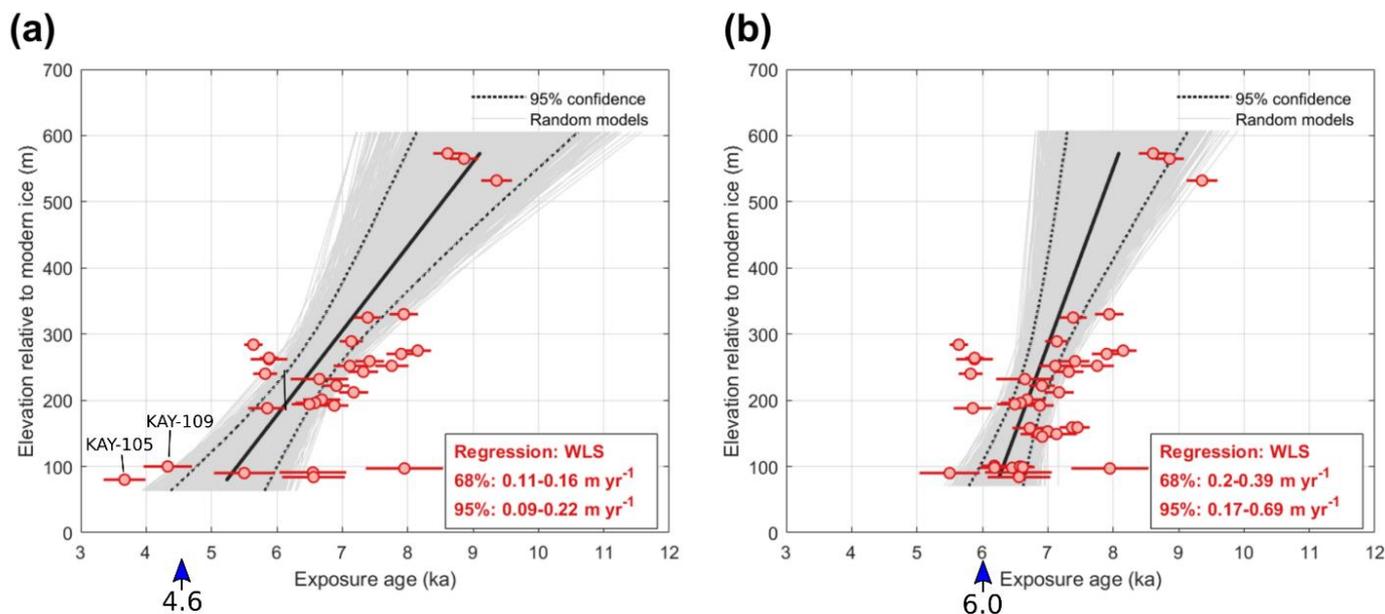


Figure 6: Linear regression transects of ice surface lowering over time (ka) relative to the modern ice surface (m). (a) Results of linear regression using the same exposure ages as Johnson et al. (2020). (b) Results of linear regression using new scoria cone ^{10}Be exposure ages with KAY-105 and KAY-109 outliers removed. In both panels, exposure ages are plotted relative to elevation above the modern ice surface and display (1σ) external errors. Modern ice surface elevations are as reported in Johnson et al., 2020, and using 80 m asl as the modern ice surface elevation adjacent to the scoria cone site. The dotted black lines delimit the 95% confidence intervals of the thinning profile (also shown in Fig. 5b). The straight black line displays the model best fit line and grey lines represent all model fits to the data. Surface exposure age elevations were input normalised to the height above the modern ice surface at Kay Peak (80 m). The blue arrows indicate timing ice surface reached its current elevation based on extrapolation of the best fit line for each transect. See Fig. A4 for the linear regression transects generated from Sample Sets 2 and 3.

The new ^{10}Be exposure ages from scoria cone better constrain the timing that Pope Glacier lowered to ~ 80 m above its present elevation, where the present or modern ice surface is defined as 80 m asl adjacent to the scoria cone site. The best fit end of thinning is changed from 5.2 ± 0.7 ka (95 % confidence interval) to 5.8 ± 0.5 ka (95 % confidence interval) using the same input sample set as Johnson et al., 2020 plus the 12 new ^{10}Be ages. The omission of in situ ^{14}C ages from samples KAY-105 and KAY-109 further changes the best fit end of thinning and shifts the timing of ice surface lowering to ~ 80 m above the modern ice elevation to 6.3 ± 0.4 ka (95 % confidence interval). The best fit onset of thinning is also shifted from 9.1 ka ± 1.1 to 8.1 ± 0.9 ka. Removing these two outliers is justified because they were shown to exhibit a non-normal distribution (Table 1, Fig. A1). When only the scoria cone ages are used in the linear regression analysis, with no re-evaluation of the Kay Peak data, the modern ice surface is reached at 5.4 ka. Extrapolation of the Johnson et al., 2020 best fit line suggests Pope Glacier lowered to its present elevation by 4.6 ka (Fig. 6a), whereas the revised best fit line indicates Pope Glacier lowered to its present thickness by 6.0 ka (Fig. 6b).

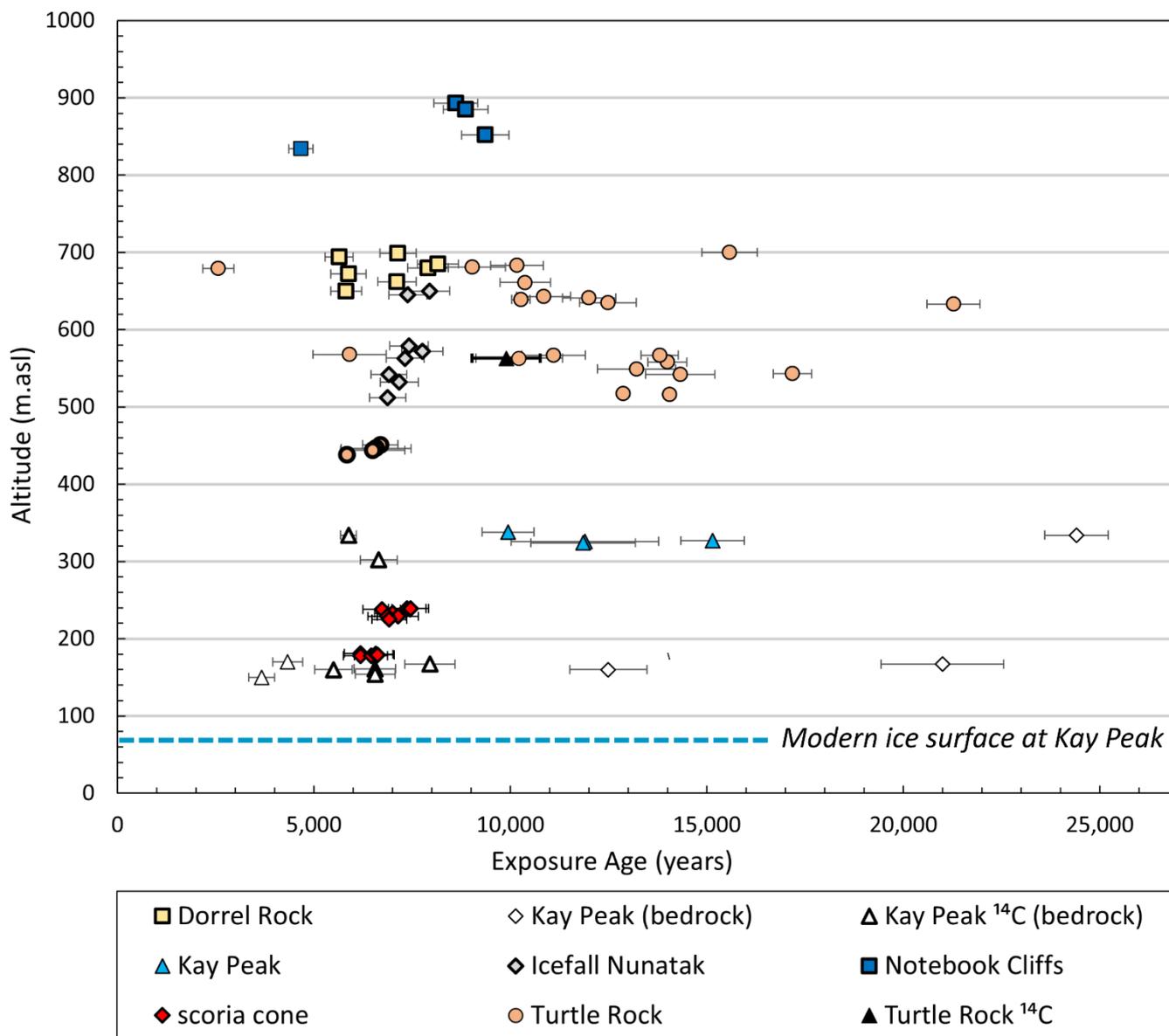
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4. Discussion

4.1 Wider context of the scoria cone ^{10}Be exposure ages

Here we discuss the wider context for the new thinning constraints presented above. The scoria cone site is situated 180–240 m asl, an elevation range not covered by existing surface exposure ages from Mt Murphy (Fig. 7). We interpret these new
285 ages as reflecting the timing of deglaciation of the site. Specifically, the error weighted mean deglaciation ages of the upper outcrop A (240 m asl) of 7.1 ± 0.1 ka and the lower outcrop B (180 m asl) of 6.4 ± 0.1 ka suggest that the surface of Pope Glacier lowered by at least 60 m in less than 1000 years (see Fig. 6b), which is equivalent to a rate of 0.06 m yr^{-1} . The scoria cone ages are, in addition, tightly clustered with no outliers (Fig. 5a), i.e., no individual clasts appear to have been subject to
290 prior exposure or post-depositional disturbance that would make their exposure ages more scattered or skew older or younger. This tight clustering of ages shows that the cobbles collected from each outcrop most likely experienced the same history of exposure, and hence the average age from each outcrop is thus likely to reflect its true deglaciation age (Balco, 2011). The lack of geologic scatter permits greater confidence in the error weighted mean exposure age calculated for each outcrop, and by extension, the thinning trajectory of Pope Glacier over the elevation range of 240 – 180 m asl. Evidence for minimal geologic scatter in the exposure ages is further strengthened by reduced χ^2 values close to 1 and p values > 0.01 (see
295 Results section, Table 1). In the context of the whole Mt Murphy age versus elevation profile (Fig. 7), the scoria cone exposure ages suggest that the ice surface over the lower nunataks thinned at a similar rate to that detected at the higher elevation sites (e.g., Icefall Nunatak), and did not slow significantly as it neared its present ice thickness.



300 **Figure 7: New exposure ages obtained from scoria cone in comparison with exposure age data reported by Johnson et al. (2008; 2020) for nunataks surrounding Mt. Murphy.** Exposure ages are plotted as ¹⁰Be and in-situ ¹⁴C exposure ages (years) versus altitude (m asl). Filled symbols represent erratics and open symbols represent bedrock samples, all symbols represent ¹⁰Be exposure ages unless specified as in-situ ¹⁴C in the key. Bold outline indicates samples used for Sample Set 4 linear regression analysis. Dashed blue lines show the modern ice surface elevation near Kay Peak ridge (which is < 1 km from the present grounding line; Fig. 1c). Note symbols for previous and new exposure ages are the same as in Fig. 1b and Fig. 1c.

305



The deglacial history of the lower exposed ridges of Mt. Murphy (below 300 m asl) was, until now, inferred from ^{10}Be and in-situ ^{14}C cosmogenic nuclide measurements of samples from Kay Peak ridge (Fig. 1c; Johnson et al., 2020). The Kay Peak exposure ages differ from the scoria cone exposure ages in several ways. Firstly, many ^{10}Be Kay Peak bedrock exposure ages are considerably older than ^{10}Be exposure ages from erratics on the scoria cone, where the maximum deglaciation age is 7.4 ± 0.5 ka. In contrast, only one Kay Peak ridge bedrock sample analysed for ^{10}Be yielded an exposure age younger than the LGM ($< \sim 20$ ka). ^{10}Be exposure ages of Kay Peak erratics (~ 330 m asl) are also younger than most Kay Peak ^{10}Be bedrock exposure ages and all ^{10}Be exposure ages from erratics post-date the LGM. Measuring ^{10}Be concentrations in erratics is often preferred to bedrock because ^{10}Be inheritance is less likely in erratics due to removal of previously accumulated nuclides by glacial erosion and transport. However, the complete removal of the previous nuclide inventory is not guaranteed (Heyman et al., 2011). Kay Peak erratic exposure ages, while younger than the LGM (9.9 – 15.1 ka), are still 2.5 – 7.5 ka older than the maximum exposure age from the scoria cone. These exposure ages would suggest Kay Peak erratics deglaciated much earlier than samples from higher elevation sites, for example Icefall Nunatak, where ^{10}Be exposure ages are 7.9 – 6.9 ka at 650 – 560 m asl (Johnson et al., 2020).

In addition, Kay Peak bedrock exposure ages are more scattered than the scoria cone erratic exposure ages. Some of this spread is accounted for by the greater elevation range of Kay Peak samples between 330 – 150 m asl. The scoria cone elevation range of 240 – 180 m asl is much smaller. Greater variability in Kay Peak exposure ages would therefore be expected, yet there is significantly more scatter in Kay Peak bedrock exposure ages, even over very small elevation ranges (< 20 m). Inheritance in Kay Peak ^{10}Be bedrock exposure ages can partly explain this scatter, but the scatter also extends to in-situ ^{14}C exposure ages, which limits how well we can constrain thinning closest to the modern ice surface. At lower Kay Peak ridge 150 – 170 m asl, individual in-situ ^{14}C bedrock exposure ages range from 3.7 ± 0.3 – 8.0 ± 0.6 ka. The > 5 ka scatter in the ^{14}C bedrock exposure ages was speculated to be due to the complexity of snow/ice cover related to the curvature of the Kay Peak ridge crest (Johnson et al., 2020). Furthermore, there is no other location in Antarctica to date where so many samples have been measured for in-situ ^{14}C , thus the apparent scatter in ^{14}C ages above that expected by analytical uncertainties alone could be in part due to underestimation of the measurement uncertainty for ^{14}C concentrations.

In summary, these observations imply: i) that ^{10}Be ages of most of the Kay Peak ridge bedrock samples reflect inheritance of ^{10}Be from an earlier period of exposure, and the Kay Peak erratics were likely similarly affected (the younger in-situ ^{14}C exposure ages from the same bedrock samples, including KAY-101 (6.0 ± 0.6 ka), KAY-107 (5.5 ± 0.6 ka), and KAY-108 (8.2 ± 0.9 ka) lend support to this hypothesis); and ii) that there is more scatter within the in-situ ^{14}C bedrock exposure age data than in the scoria cone ^{10}Be erratic ages. Scatter beyond analytical uncertainty in the Kay Peak ridge ^{14}C exposure ages is likely primarily due to a complex localised deglaciation history caused by non-contiguous deglaciation of fringing ice along the ridge axis.



340 4.2 Implications of scoria cone exposure ages for the thinning history of Pope Glacier

In this section, we discuss the implications of our new data for the thinning history of Pope Glacier. The median thinning rate determined from the revised exposure age dataset (sample set 4; $0.27 \pm 0.12 / -0.07$ m yr⁻¹) is the most different from the previously published median rate for Pope Glacier ($0.13 \pm 0.03 / -0.02$ m yr⁻¹; Johnson et al., 2020) of all the rates we calculated. Using this rate implies that the ice surface of Pope Glacier lowered up to 52 % faster than previously estimated.

345 This falls in the middle of the range of thinning rates from elsewhere in Antarctica calculated using the same method (Small et al., 2019, Table 3).

Even though the mid-Holocene thinning of Pope Glacier occurred over only a few thousand years, it appears to have been much slower than contemporary changes in the region. The fastest thinning rate we calculated using the upper limit of the
350 95% confidence interval on our median rate ($0.27 \pm 0.12 / -0.07$ m yr⁻¹) is over an order of magnitude slower than contemporary thinning rates of 4-7 m yr⁻¹ detected above the 2020 grounding line of Pope Glacier (Milillo et al., 2022). However, it is important to consider the relative resolutions of the datasets. Paleo thinning rates for Pope Glacier averaged over a millennial timescale might be perceived as an oversimplification because only a single average is calculated over the period of thinning being examined. However, linear thinning rates averaged over longer timescales are thought to be more
355 indicative of the basin average rather than localised changes in the glacier trunk (Small et al., 2019). Therefore, linear thinning rates of Pope Glacier although less sensitive to short-term fluctuations are extremely relevant for validating model simulations which are generally regional or larger in scale e.g. (Pollard et al., 2016; Johnson et al., 2021a).

Our new exposure ages also have implications for the timing of the later stages of thinning, closest to the modern ice surface.
360 The revised best fit onset and end of thinning (Fig. 8, red dotted line) from 8.1 – 6.3 ka indicates more abrupt thinning compared to the previously published estimate of 9.1 – 5.2 ka (Fig. 8, black dashed line) (Johnson et al., 2020). The trajectory of thinning indicated by the revised best fit line equates to ≥ 560 m lowering of the ice surface at Pope Glacier in approximately half the time: 1800 years compared to a previous duration of 3900 years (Johnson et al., 2020). In comparison with other parts of the ASE, this revised time span is similar to the duration of early Holocene thinning detected at Mt
365 Moses, in the eastern ASE (adjacent to Pine Island Glacier). The lower 142 m of presently-exposed outcrop at Mt. Moses deglaciated between 7.4 ± 0.7 – 5.4 ± 0.7 ka, over a period of approximately 2,000 years (Johnson et al., 2014). However, at Mt Moses, the ice sheet thinned in two distinct phases at different rates. This appears to contrast with the thinning pattern observed at Mt Murphy, although the relative density of data from the two sites is not similar so fluctuations in thinning rates at Mt Moses may not have been detected.

370

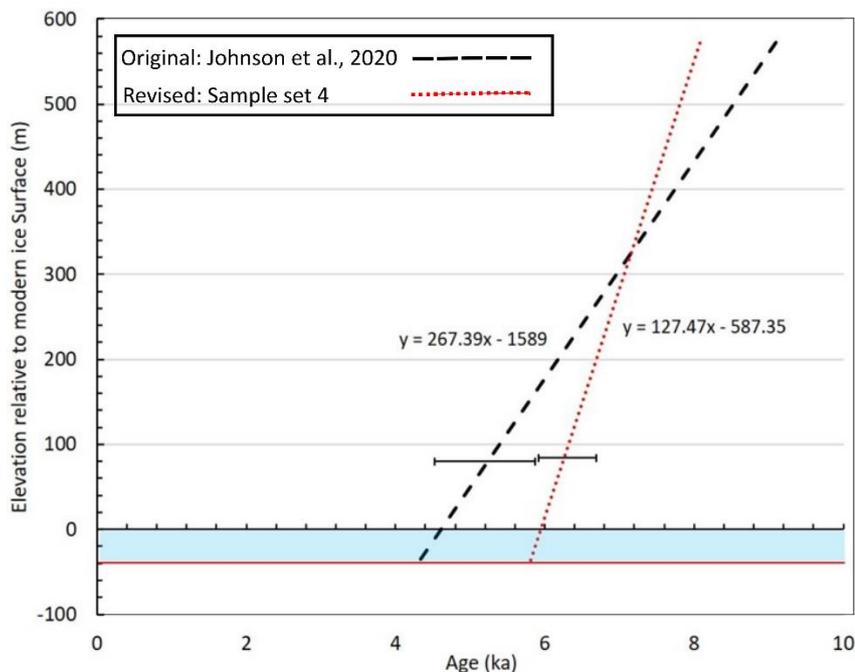


Figure 8: Predicted timing of exposure at or near the surface of subglacial bedrock cores drilled at Kay Peak ridge Predictions of timing for bedrock exposure extrapolated from the two endmember best fit line thinning rate scenarios calculated from different input combinations of exposure ages. The best fit ice surface lowering rate is determined from the published Mt. Murphy exposure age dataset of Johnson et al., 2020 is contrasted with the best fit from the revised dataset (sample set 4) which includes exposure ages from scoria cone and omits KAY-105 and KAY-109. Error bars represent 95 % confidence intervals of best fit timing at which the samples closest to the modern ice surface were deglaciated. Below the 95 % confidence bars the best fit line has been extrapolated based on the linear equations shown in the figure. Blue shaded area represents the ice thickness that was drilled (~40 m) to reach the subglacial bedrock. Red horizontal line is a one-dimensional representation of the bedrock surface at the point of core recovery. Note the 95 % confidence interval of best fit revised end thinning is at a slightly higher elevation (84 m) as KAY-105 is the sample at Mt. Murphy is closest to the modern ice surface (80 m).

Better constraints on the timing and pace of thinning of Pope Glacier during the Holocene allows us to make a prediction for the hypothetical timing of exposure of subglacial bedrock cores drilled at Kay Peak ridge (-75.215, -110.960). Previous work suggested that the ice surface lowered to the elevation of the modern ice (80 m asl) by 4.6 ka (Johnson et al., 2020). Performing an extrapolation on our revised best fit linear thinning history implies that Pope Glacier reached its present thickness considerably earlier in the Holocene, at 6 ka (Fig. 8). Further extrapolating our best fit ice thinning trajectory to below the surface of the modern ice, allows us to predict when subglacial bedrock cores from below Kay Peak ridge could have been exposed at or near to the surface. Based on the previous thinning rate estimates (Johnson et al., 2020), the subglacial bedrock from ~40 m below ice surface would have been exposed at 4.2 ka. With the new exposure ages and revised Kay Peak ages, our linear regression analysis suggests that the likely onset of exposure of the subglacial bedrock occurred ~1,500 years earlier, at 5.7 ka. However, we acknowledge the assumption of linear constant thinning is only valid



over the range of elevations specified by our linear regression analysis (Fig. 5); as such, our chronology is robustly constrained only to 80 m above the modern ice surface. Nevertheless, our results suggest that early- to mid-Holocene ice thinning at Mt. Murphy occurred over a shorter interval than previously assumed and implies a longer duration over which any subsequent rethickening of ice could have occurred.

5. Conclusions

We present 12 new cosmogenic ^{10}Be exposure ages which provide constraints on the timing of the last deglaciation of the western flank of Mt Murphy, in the Amundsen Sea Embayment. The ages were derived from erratic cobbles collected from two outcrops on a scoria cone situated within ~160 m of the modern ice surface and ~1 km from the present grounding line of Pope Glacier. Outlier detection was applied to both the new exposure ages and to existing exposure ages below 300 m asl to better constrain the rate and timing of thinning of Pope Glacier during the Holocene. The new ^{10}Be exposure ages represent two statistically distinct populations, which correspond to two rock outcrops within an elevation range of 240 – 180 m asl and have an error weighted mean age and standard error of 7.1 ± 0.1 ka and 6.4 ± 0.1 ka, respectively.

Linear regression analysis undertaken for this study implies that Pope Glacier thinned by ≥ 560 m at a median rate of 0.27 m yr^{-1} over a period of 1,800 years during the early-mid Holocene. This is 1.5 times faster than previously assumed. Furthermore, the tighter constraints placed by the new data on the timing of deglaciation of the lowest currently exposed section of Mt. Murphy suggest that the ice surface of Pope Glacier had thinned to within 80 m of its present elevation by 6.3 ± 0.4 ka. This is 1,100 years earlier than the previous estimate of Johnson et al. (2020).

These results have implications for bedrock cores collected for a parallel study from below the ice sheet near the lowermost outcrop of Kay Peak ridge, close to our study site: the revised thinning trajectory suggests that the top of those cores could have been exposed at or near the ice sheet surface as early as 5.7 ka. In summary, the results suggest that early- to mid-Holocene thinning of Pope Glacier occurred more rapidly, and earlier, than previously thought. They therefore permit either a longer period of ice sheet stability in the mid- to late-Holocene, or alternatively a longer duration over which late Holocene rethickening could have occurred had the ice sheet not remained stable.

420



425 Appendices

Appendix A

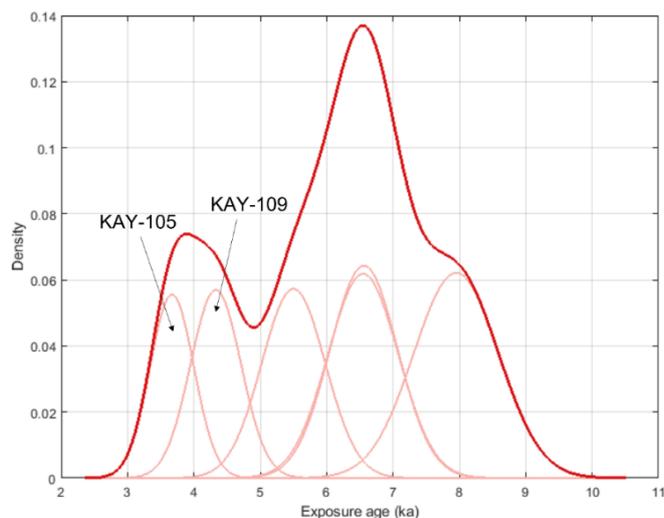
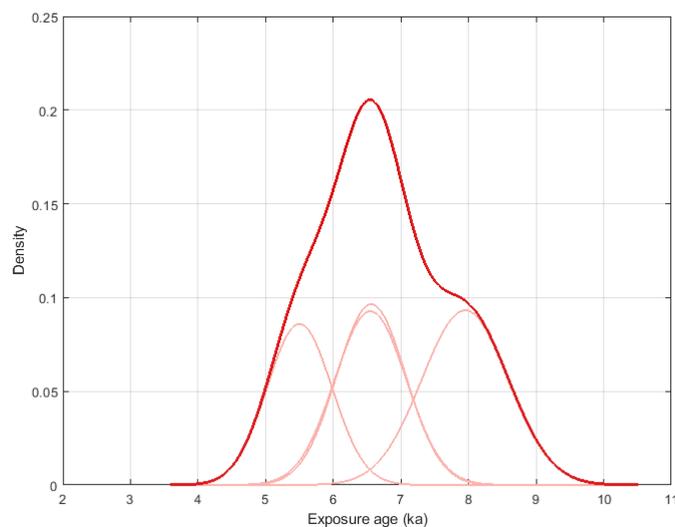


Figure A1: Kernel density estimate plot of ^{14}C ages ($n = 6$) distribution for lower Kay Peak ridge. Error weighted mean: 5173; reduced chi-squared: 15.37; chi-squared p-value: 0.0000

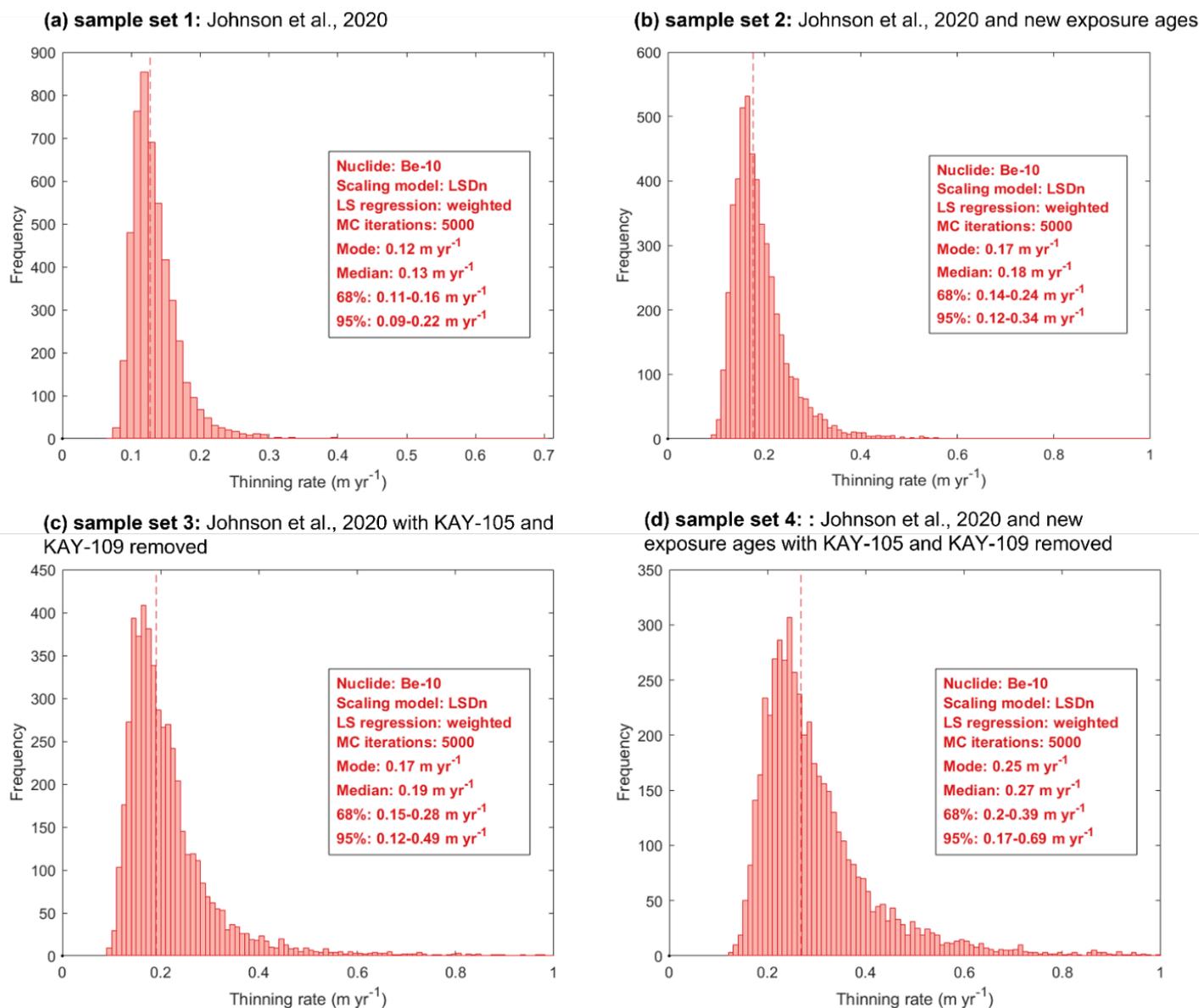


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Figure A2: Kernel density estimate plot of ^{14}C ages distribution for lower Kay Peak Ridge ($n=4$), with the two youngest ages (KAY 105, KAY 109) removed. Error weighted mean: 6472; Standard error: 266; External error: 205; reduced chi-squared: 3.29; chi-squared p-value: 0.0196



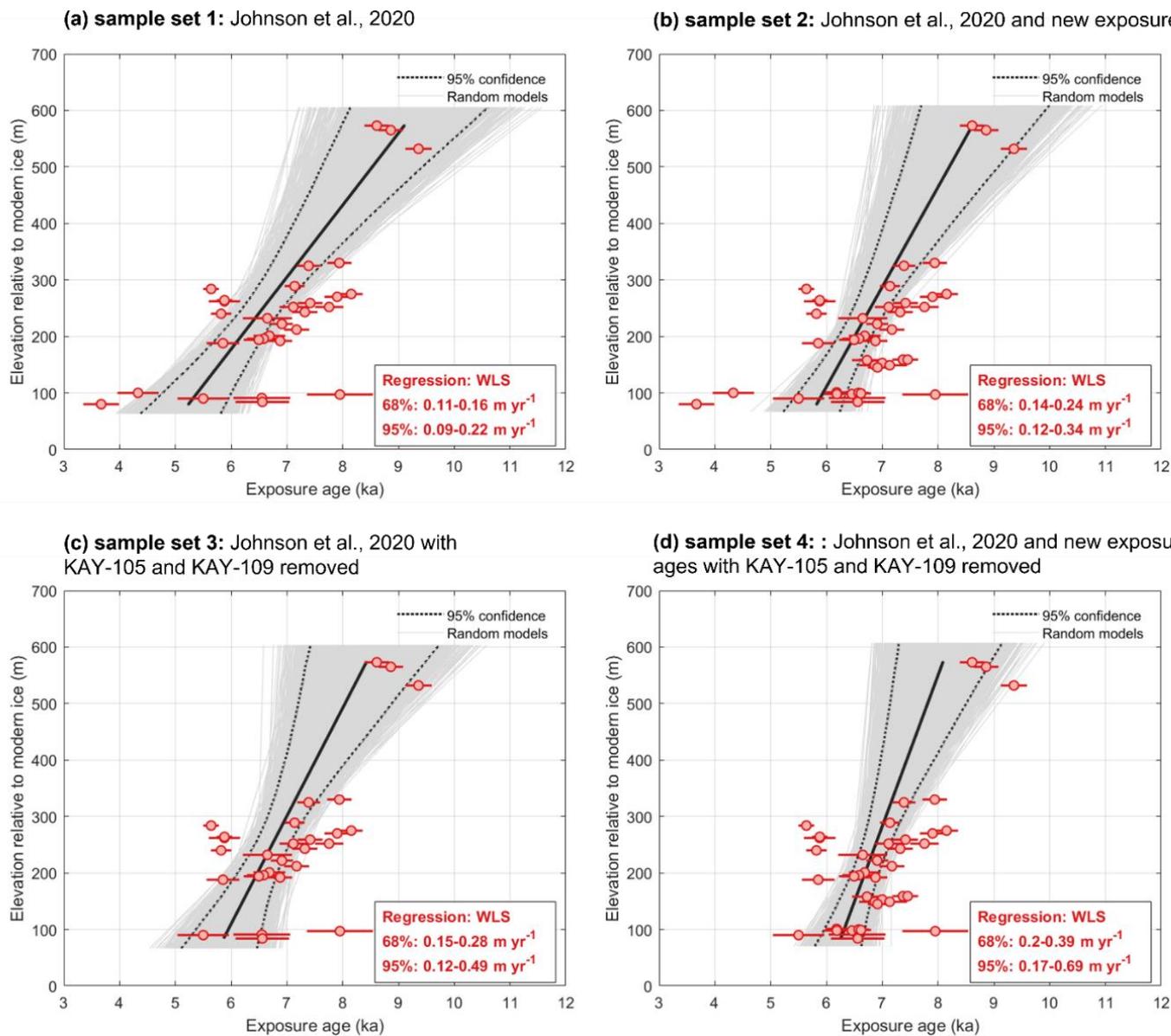
Appendix B



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Figure B1: Histogram outputs of ice surface thinning rates generated by iceTEA (Jones et al., 2019). Results of linear regression: (a) the original sample set of Johnson et al. (2020), (b) incorporating new data from Scoria Cone, (c) removing young ages identified as outliers and (d) incorporating new data from Scoria Cone as well as removing young ages identified as outliers. Panel boxes inset in each subfigure display the median and modal values for thinning rates as well as the range of 68 % and 95 % confidence intervals.

440



445 **Figure B2: Linear Regression Transect of ice surface thinning generated with iceTEA:** Linear regression transects showing (a) the original sample set of Johnson et al. (2020), (b) incorporating new data from Scoria Cone, (c) removing young ages identified as outliers and (d) incorporating new data from Scoria Cone as well as removing young ages identified as outliers. Black solid line is ‘best fit’ averaged from 5000 Monte Carlo simulations. Dotted black lines represent 95 % confidence intervals. Exposure ages displayed as red circles with internal uncertainties.



450 **Data availability**

Exposure age data shown in Figure 4 are publicly available in the NERC Polar Data Centre (doi: in progress).

Author Contributions

JRA led the study with supervision from JSJ, DHR, SJR, and PJM. JRA prepared the samples for AMS analysis and wrote
455 the first manuscript draft with support from KAN and DHR. JSJ and SJR collected the samples. JRA and RAV prepared the
figures with input from all co-authors. DHR supervised analytical work and KW performed AMS analyses. All authors
contributed to writing the manuscript.

Competing Interests

460 The authors declare that they have no conflict of interest.

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Contribution No. XX.

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