



| 1 | High mid-Holocene accumulation rates over West Antarctica inferred from |
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2 a pervasive ice-penetrating radar reflector 3

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Key points

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- We estimate mean accumulation rates for the past ~5000 years across the Pine Island, Thwaites, and Institute and Möller ice-stream catchments in West Antarctica using a ubiquitous radar-sounded and ice core-dated Internal Reflecting Horizon
- Accumulation rates were 18% higher during the mid-Holocene compared to modern rates over the Amundsen-Weddell-Ross Sea divide
 - Spin-up of ice-sheet models must account for time-varying changes in accumulation rates between the Last Glacial Maximum and the present

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Abstract

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Modelling the past and future evolution of the West Antarctic Ice Sheet (WAIS) to atmospheric and ocean forcing is challenged by the availability and quality of observed palaeo-boundary conditions. Key potential data for reconstructing past ice-sheet processes on large spatial scales are Internal Reflecting Horizons (IRHs) detected by Radio-Echo Sounding (RES) techniques. When isochronal and dated at ice cores, IRHs can be used to determine palaeo-accumulation rates and patterns. Using a spatially extensive IRH over Pine Island Glacier, Thwaites Glacier, Institute and Möller Ice Streams, and a local layer approximation model, we infer mid-Holocene accumulation rates over the slow-flowing parts of these catchments for the past ~5000 years. By comparing our results with modern climate reanalysis models and observational syntheses, we estimate that accumulation rates over the Amundsen-Weddell-Ross divide were on average 18% higher than modern rates during the mid-Holocene. However, no significant spatial changes in the accumulation pattern were observed. These higher mid-Holocene accumulation-rate estimates match previous palaeo-accumulation estimates from ice-core and targeted RES surveys over the ice divide, and also coincide with periods of grounding-line re-advance during the Holocene over the Weddell and Ross Sea sectors. Our results highlight the need for ice-sheet models to account for time-varying accumulation rates across the WAIS during the Holocene to provide better estimates of its contribution to past sea-level rise.

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Key words: West Antarctica, Internal Reflecting Horizons, Accumulation, Holocene, Ice-Penetrating Radars, Ice-Core, Pine Island Glacier, Thwaites Glacier.

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1. Introduction

Improving our knowledge of past climatic changes over the West Antarctic Ice Sheet (WAIS) is required if we are to understand its present evolution and model its future under increasingly rapid climatic changes (IPCC, 2021). Most studies of past ice-sheet behaviour over the WAIS have focused on modelling changes in ice volume and grounding-line (GL) retreat following the Last Glacial Maximum (LGM, ~20 ka Before Present, BP) (Denton and Hughes, 2002; Hillenbrand et al., 2013; 2014); however, less attention has been paid to ice-sheet evolution during the Holocene (~11.7 ka BP to present). Recent evidence suggests that the GL may have retreated several hundred kilometres inland from its current position at ~10 ka and subsequently re-advanced to reach its modern position sometime during the Holocene, due to isostatic rebound and climate-induced changes, particularly over the Weddell Sea and western Ross Sea sectors (Siegert et al., 2013; Bradley et al., 2015; Kingslake et al., 2018; Wearing and Kingslake, 2019; Venturelli et al., 2020; Neuhaus et al., 2021; Johnson et al., 2022).

The inland atmospheric and ice-dynamical conditions, which may have partly caused this Holocene GL migration, remain relatively under-studied. An early investigation by Whillans (1976) using radar data near Byrd Ice Core indicated stability during the Holocene and late Pleistocene. Records of temperature and precipitation from the WAIS Divide Ice Core (hereafter abbreviated as WD14; Fig. 1) in the central WAIS suggest higher accumulation rates during the Holocene than the present (Fudge et al., 2016), a trend that is also observed across small parts of the Western divide (hereafter referred to as WD; Fig. 1) near WD14 where isolated Radio-Echo Sounding (RES) surveys have shown 15-30% higher accumulation rates during the mid-Holocene compared to modern values (Siegert and Payne, 2004; Neumann et al., 2008; Koutnik et al., 2016).

At present, modelling studies that aim to predict future changes over the WAIS use past icesheet reconstructions from after the LGM to guide and evaluate their models (Chavaillaz et al., 2013; DeConto and Pollard, 2016; Bracegirdle et al., 2019). However, significant discrepancies between model simulations and the palaeo-proxy record currently impede our ability to predict confidently how the ice sheet will respond to future changes in the climate (e.g. Johnson et al., 2021). While improvements in model physics and parameterisations are needed to close this gap (Bracegirdle et al., 2019; Sutter et al., 2021), a considerable improvement in the availability and quality of palaeo-proxy records, particularly during the Holocene, is also needed to gain a more accurate picture of the past ice-sheet changes (Jones et al., 2022). Palaeo-proxy data over the WAIS have traditionally come from point-based measurements, such as ice cores (WAIS Divide Project Members, 2013; Cuffey et al., 2016; McConnell et al., 2017; Buizert et al., 2021), sediment cores (Hillenbrand et al., 2013; Arnd et al., 2017; Hillenbrand et al., 2017; Kingslake et al, 2018; Venturelli et al., 2020; Neuhaus et al., 2021; Sproson et al., 2022), or from surface-exposure dating (Stone et al., 2003; Johnson et al., 2014; Hein et al., 2016; Nichols et al., 2019; Johnson et al., 2020; Braddock et al., 2022). A spatially extensive alternative data source is provided by Internal Reflecting Horizons (IRHs) sounded by RES (Bingham and Siegert, 2007; Harrison, 1973). When employed in combination with ice-core stratigraphy, IRHs can be used to extrapolate age-depth relationships across large spatial scales by following peaks in electromagnetic power in the radar data (e.g. Beem et al., 2021; Bodart et al., 2021a; Cavitte et al., 2016; Jacobel and Welch, 2005; MacGregor et al., 2015; Whillans, 1976; Winter et al., 2019).

In comparison to East Antarctica and Greenland, extrapolating past ice-sheet records from WAIS ice cores to entire glacier catchments has so far been challenging due to the limited availability of deep ice cores there and, until recently, the lack of suitable datasets. However, efforts have intensified in recent years to improve our understanding of the ice stratigraphy over this sector. In particular, recent studies using airborne RES data over the WAIS (Karlsson et al., 2014; Muldoon et al., 2018; Ashmore et al., 2020a; Bodart et al., 2021a) all identified a particularly bright IRH precisely dated using the Byrd and WD14 ice-core chronologies at 4.72 ± 0.28 ka BP (Muldoon et al., 2018;





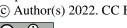
97 Bodart et al., 2021a). Comparison with volcanic sulphate deposition within the WD14 and Siple
98 Dome ice cores revealed a large peak in sulphate concentration matching the age and depth of this
99 ubiquitous layer (Kurbatov et al., 2006; Bodart et al., 2021a; Cole-Dai et al., 2021; Sigl et al., 2022),
100 which is hereafter termed the 4.72 ka IRH. This IRH has been observed by multiple RES systems
101 from different surveys and data providers. It extends through much of the slow-flowing ice of the
102 Amundsen and Weddell Sea embayments, including across the divides demarcating regions draining
103 the Amundsen, Weddell and Ross Seas.

Here, our aim is to estimate mid-Holocene accumulation rates across the WAIS from first-order calculations using a local-layer approximation (LLA) model (Waddington et al., 2007), informed by the spatially extensive 4.72 ka IRH. We first describe the data and the 1-D model, and discuss the methods used to assess the feasibility of the LLA from gradients in horizontal ice-flow, ice thickness and accumulation-rate. We then present our accumulation rate estimates and discuss the effects of spatial and topographic controls on our results in relation to modern accumulation rates using both modelled gridded data from the Regional Atmospheric Climate Model 2 (hereafter RACMO2; Van Wessem et al., 2018) and observational point-based data from snow, firn and ice cores (Neumann et al., 2008; Burgener et al., 2013; Favier et al., 2013; Mayewski and Dixon, 2013; Medley et al., 2014), to generate a long-term perspective on changes between the mid-Holocene and the present. Finally, we place our results in the context of previous studies that consider the evolution of the WAIS during the Holocene.

2. Datasets and methods

2.1 Along-track IRH data

We used data from extensive (~91 000 flight-track km) RES surveys acquired across West Antarctica between 2004 and 2018. The main contributing surveys are the University of Texas Institute for Geophysics (UTIG) 2004-2005 AGASEA survey flown over Thwaites Glacier (THW) and Marie Byrd Land which deployed the 60-MHz High Capability Airborne Radar Sounder (HiCARS) radar system (Holt et al., 2006; Peters et al., 2007), and the British Antarctic Survey (BAS) 2004-05 BBAS survey over Pine Island Glacier (PIG) and 2010-2011 IMAFI survey over Institute and Möller Ice Streams (IMIS) which deployed the 150-MHz Polarimetric Airborne Survey INstument (PASIN) radar system (Vaughan et al., 2006; Corr et al., 2007; Ross et al., 2012; Fremand, Bodart et al., 2022) (Fig. 1; Table 1). Additional profiles from NASA's Operation Ice Bridge (OIB; MacGregor et al., 2021) 2016 and 2018 surveys, flown with the 190-MHz Multichannel Coherent Radar Depth Sounder 2 (MCoRDS-2) radar system (CReSIS, 2018), were also used to extract IRH information near the WD14 Ice Core and upper IMIS catchments (Bodart et al., 2021a; Figure 1 and Table 1). We refer the reader to the above references for comprehensive details on each system's capabilities.



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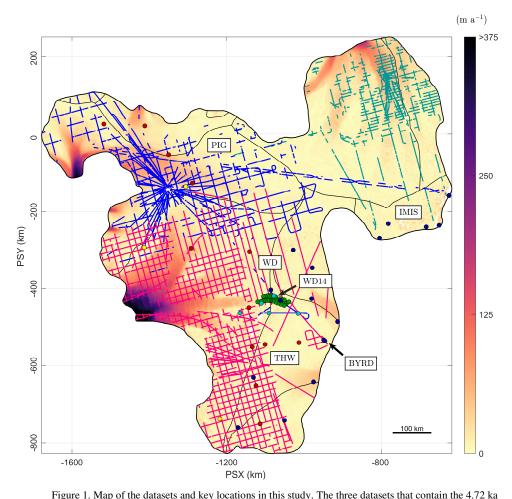
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IRH are color-coded as IMIS (green), PIG (blue), and THW (pink). IRHs falling outside D > 1 (see Section 2.2.1; Figure S1) are excluded. Points on the map represent the snow, firn and ice cores used in this study to compare modern rates of accumulation with our 4.72 ka-to-present estimates (see Section 2.4 for source references). The background colour map shows modern surface speeds from Rignot et al. (2017). Locations mentioned in this paper are abbreviated on the map, as follows: BYRD (Byrd Ice Core), IMIS (Institute and Möller Ice Streams), PIG (Pine Island Glacier), THW (Thwaites Glacier), WAIS (West Antarctic Ice Sheet), WD (Western Divide), WD14 (WAIS Divide Ice Core). Major ice divides are from Mouginot et al. (2017). Projection for all figures in this paper is WGS84 Antarctic Polar Stereographic (PSX/PSY; EPSG: 3031).

These RES surveys were used to track and date six IRHs spanning much of the Holocene and Late Pleistocene across IMIS (Ashmore et al., 2020a), PIG (Bodart et al., 2021a) and THW (Muldoon et al., 2018), collectively spanning much of the WAIS. Here we only consider the 4.72 ka IRH mapped in all three studies and shown in Figure 1, as it is by far both the most spatially extensive and the only commonly traced IRH across all studies. We first merged all data points from the 4.72 ka IRH across the three catchments, resulting in a cumulative distance of ~40 000 line-km of IRH profiles (44% of the RES surveys' total coverage; Table 1). Although the along-track RES data were acquired with a trace spacing of between 10 and 35 m, depending on the dataset used, we re-sampled these points to 500 m in the along-track direction. We then added a spatially invariant firn correction





of 10 m onto the Muldoon et al. (2018) dataset to match the same firn correction applied by the other studies to correct the IRH depth, and finally calculated the median value of all ice thicknesses and IRH depths falling within the nearest 500 m interval.

Table 1. Characteristics of each IRH dataset used in this study that contain the 4.72 ka IRH. 'Reflector 1' in Muldoon et al. (2018) is abbreviated here as 'R1'.

| Survey name | Survey provider | RES system | Dataset reference | Cumulative IRH distance (km) |
|----------------|--------------------|-------------------------------------|--------------------------------|------------------------------------|
| IMAFI | BAS | PASIN 150-MHz | H2 in Ashmore et al. (2020a) | 1.5x10 ⁴ |
| BBAS / OIB | BAS / NASA | PASIN 150-MHz / MCoRDS-2 190-MHz | R2 in Bodart et al. (2021a) | $0.6x10^4$ |
| AGASEA | UTIG | HiCARS 60-MHz | R1 in Muldoon et al. (2018) | 1.9 x10 ⁴ |

2.2 Inferring accumulation rates

To infer accumulation rates from the 4.72 ka IRH, we used the Nye model, a 1-D ice-flow model widely used for estimating accumulation rates and age-depth relationships over relatively slow-flowing parts of an ice sheet (Nye, 1957; Fahnestock et al., 2001). Other 1-D models exist, including the Dansgaard-Johnsen (Dansgaard and Johnsen, 1969) and the shallow-strain rate model (MacGregor et al., 2016), but neither were suitable for this study due to uncertainties in the basal shear layer thickness across our survey area and because we are limited to only one IRH to constrain the ice-flow model. The Nye model assumes that ice thickness is constant and therefore that the ice sheet has been in a steady state since the deposition of the IRH, a realistic assumption for the period under investigation here. The Nye model states:

$$\dot{b}_a = \ln\left(\frac{z_a}{H}\frac{H}{a}\right),\tag{1}$$

where \dot{b}_a is the mean surface-accumulation rate during the Holocene epoch between an IRH of age a and the present, z_a represents the depth of the IRH dated at the ice core, and H is the ice thickness. The model assumes that the vertical strain rate, $\dot{\varepsilon}_a$, is also constant and vertically uniform, so that it exactly balances the overburden of local ice accumulation:

$$\dot{\varepsilon}_a = \frac{\dot{b_a}}{H}.\tag{2}$$

We iterated Eq. (1) over the re-sampled 500-m spaced dataset using the depth of the 4.72 ka IRH for z_a and used the median radar-derived ice-thickness measurement re-sampled over the 500-m grid to obtain H, when this information was available. In areas where the radar did not sound the bed, we used the BedMachine v2 gridded product to obtain a value for H (Morlighem, 2020). Note that the accumulation values presented in this study are all reported in m a^{-1} of ice equivalent.

Uncertainties in accumulation rates are calculated by iterating Eq. (1) using the lowest and highest possible age of the 4.72 IRH based on the maximum and minimum age uncertainty (\pm 0.28 ka) calculated from RES and ice-core depth uncertainties (see details in Muldoon et al. (2018) and Bodart et al. (2021a)). This results in lower and upper bounds in accumulation rates over our model domain (Fig. S2a-b), which are then combined to provide a relative uncertainty in accumulation rates for the 4.72 ka IRH (Fig. S2c). The maximum relative uncertainty in accumulation rates for the 4.72 ka IRH is 3.3%, with a median value of <0.5% across our grid. Note that these values do not include





uncertainties due to the model approximation itself. As discussed in the next two sections, we believe model uncertainties to be small for the domain considered.

2.2.1 Assessing the suitability of the 1-D model

To quantify the suitability of the LLA onto which the 1-D model is based, we calculated the effects of horizontal gradients in modern ice thickness and accumulation rates along particle paths in their ability to affect IRH depths, as per Waddington et al. (2007). In places where these gradients are too large, estimates of accumulation rates from IRHs likely require a more complete treatment of ice flow and strain rates to account for disturbances in IRH depths, which only multi-dimensional models are able to resolve (Waddington et al., 2007; Koutnik et al., 2016; MacGregor et al., 2016).

To quantify the effect of horizontal gradients on IRH of age a, we computed the total horizontal particle path length L_{path} and characteristic lengths L_H and L_b representing the gradients in ice thickness and accumulation rates respectively for age a (Supplementary Information). These three components were then combined to generate a non-dimensional parameter D (Fig. S1d). Waddington et al. (2007) and MacGregor et al. (2009; 2016) used a maximum D value of unity to estimate where the LLA is acceptable, but because the value of D cannot yet be translated quantitatively into an uncertainty value for an LLA-inferred accumulation rate, it is not clear what value is appropriate. Smaller values of D indicate that local horizontal gradients in ice thickness and accumulation rates have a smaller effect on IRH depth of age a, and thus we assume that the LLA is valid (Waddington et al., 2007; MacGregor et al., 2009; 2016). However, if D > 1, we assume that the depth of the IRH is unlikely to be the sole result of accumulation rates at the surface and that a more sophisticated model is required (Sect. 2.2.2). In our case, most of the survey areas has D values that are well below the unity, except for a limited number of IRH profiles near the onset of PIG's tributaries and over THW's central trunk, which we excluded from our analysis.

2.2.2 Model limitations

One of the limitations of the Nye model is that it assumes that gradients in sliding velocity are mostly concentrated in a thin layer at the ice-bed interface and that the ice column deforms by pure shear only (Nye, 1957; Fahnestock et al., 2001). For this reason, the Nye model is generally only appropriate for IRHs found in the upper part of the ice column, as is the case here. Additionally, the use of the model is restricted to areas where ice flow is slow and horizontal strain rates are low.

Multi-dimensional models would likely improve the accumulation estimates for IRHs found in the lower half of the ice column or in more disrupted or faster flowing areas (e.g. Waddington et al., 2007; Neumann et al., 2008; MacGregor et al., 2009; Leysinger Vieli et al., 2011; Karlsson et al., 2014; Koutnik et al., 2016; Sutter et al, 2021). However, we focus on a shallower IRH situated within the upper 40% of the ice column (Fig. 2b-c), where we can be reasonably confident that the ice sheet has remained close to steady-state and where IRHs are likely shallow enough not to have sustained appreciable disturbances that would affect the Nye model assumptions (Sect. 2.2).

Additionally, due to the inherent nature of tracking IRHs through RES data, our coverage is limited to areas where ice-flow speeds are low and IRHs are relatively undisturbed. An assessment of strain rates over our model domain suggests limited disturbance over the WD and most of our grid, apart from near the onset of faster flow at the boundaries of our grid with the trunks of PIG and THW where higher strain rates are observed (Fig. S3). This pattern, combined with the assessment of the suitability of the LLA (Sect. 2.2.1) and exclusion of IRHs outside of the D > 1 boundary, supports our application of a 1-D modelling approach here.

2.3 Gridding and filtering





Once IRH depths and accumulation rates for the 4.72 ka IRH were obtained at regular 500-m points along RES flight paths, we filtered the results using a moving-average Gaussian filter of length 30 samples (equivalent to ~15 km) to reduce along-track noise in the IRH depth, and then gridded the filtered result using a Delaunay-triangulation-based natural neighbour interpolation method onto a 1-km polar stereographic grid. We further smoothed the gridded data using an 18-km square cell mean filter to limit the localised artefacts arising from the interpolation, which can be problematic in areas with poor data coverage. Figure S4 shows the maximum distance away from the nearest 500-m along-track point used to produce Figures 2-3, and thus where errors in the interpolated grids are expected to be larger. The median value of this maximum distance is 5 km and its maximum value is 75 km, which is comparable to previous studies (e.g. Medley et al., 2014). We evaluated other possible interpolation methods (e.g., kriging and using different semi-variogram models), but they resulted in similar or poorer quality, and were thus discounted.

2.4 Comparison with modern observations

To compare our inferred accumulation estimates for the past 4.72 ka with modern times (defined here as 1651-2019), we derived information on modern accumulation rates from two sources, one modelled (gridded) and one from a series of observational (point-based) datasets.

We used modelled gridded accumulation rates from the RACMO 2.3p2 1979-2019 Surface Mass Balance (SMB) product forced at its margin with the ERA-Interim product (native resolution: 27 km) as an estimate for modern accumulation rates (Van Wessem et al., 2018). Although SMB is not technically equivalent to the accumulation rate, runoff and sublimation are negligible in our survey area (Medley et al., 2013) so we consider SMB equal to accumulation rate in this region. We converted modelled values from kg m⁻² a⁻¹ to m a⁻¹ of ice equivalent using an ice density value of 917 kg m⁻³, calculated the 40-year mean, and then bi-linearly interpolated the gridded RACMO2 product to the same 1-km grid resolution as our 4.72 ka-to-present accumulation grid (Sect. 2.3) to ensure consistency when comparing both datasets.

Observational point-based measurements were obtained from a series of snow, firn and ice cores from the ITASE (Mayewski and Dixon, 2013), MED14 (Medley et al., 2014), SAMBA (Favier et al., 2013), and SEAT-10 (Burgener et al., 2013) datasets, as well as from a network of centennially-averaged modern accumulation rates derived from shallow IRHs traced on ground-based RES data over the WD and dated using a shallow ITASE Ice Core (Neumann et al., 2008) (Fig. 1). This results in 79 point-based accumulation measurements covering the period 1651-2010 CE (Common Era) and spread across our model domain (see Figure 1). Further detail on these datasets can be found in the above references. To compare the Holocene gridded product with the point-based measurements, we first calculated the average value of the accumulation rate at the point measurement for the entire period. We converted this value to ice-equivalent accumulation rates (as above) and extracted the two paired values, i.e., the value for the point-measurement for modern accumulation rates and the value for the nearest grid cell in the gridded 4.72 ka-to-present accumulation estimates to this measurement.

3. Results

The final grids for depth and accumulation rates for the 4.72 ka IRH are shown in Figures 2-3. In total, these grids are made of ~89 000, 500-m spaced points, which cover an area of ~610 000 km², or ~30% of the total surface area of the WAIS. The grids span most of PIG and THW glacier catchments, as well as the Ronne (upper Rutford, Institute, and Möller) and upper western Ross (Bindschadler, Kamb, MacAyeal, and Whillans) catchments (IPY Antarctic boundaries G-H, J-Jpp, and Ep-F; Mouginot et al. (2017); Fig. 1-2). Overall, the 4.72 ka IRH is shallower within the IMIS and upper PIG and THW catchments, as well as on the Ross side of the WD where ice thickness is particularly deep (Fig. 2b). Conversely, the 4.72 ka IRH is deeper in the ice near a 400-m high bedrock plateau that separates the northern and southern basins of PIG (Vaughan et al., 2006) and at



two locations in the upstream parts of the main trunk of THW where ice flows over highs in subglacial topography (Fig. 2b).

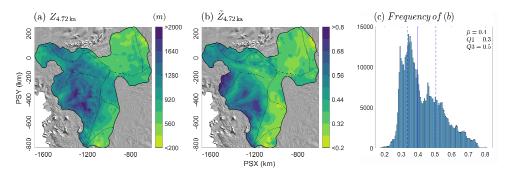


Figure 2. Gridded depths for the 4.72 ka IRH across the model domain covering the PIG, THW, and Institute and Möller ice-stream catchments. (a) Gridded depth of the 4.72 ka IRH. (b) Normalised depth of the 4.72 ka IRH relative to ice thickness. (c) Histogram showing the distribution of values in (b) with the median $(\bar{\mu})$ and the interquartile range (i.e. 25th (Q1) and 75th (Q3) quartiles) shown as solid and dashed blue lines respectively. The background image is the 2014 MODIS mosaic of Antarctica (Haran et al., 2018).

3.1. Catchment-scale accumulation estimates

Figure 3 shows a comparison of the ice-equivalent accumulation rates we inferred for the 4.72 ka IRH (Fig. 3a) and modern SMB estimates from RACMO2 (Fig. 3b). We observe that the IRH accumulation rate pattern for the last 4.72 ka is similar to the modern pattern of accumulation rates for the Amundsen Sea sector of the WAIS, which is dominated by higher coastal accumulation rates that progressively decrease inland towards the ice divides (Fig. 3a-b). Differences in accumulation between the 4.72 ka-to-present estimates and modern values are mainly observed directly upstream of the main trunks of PIG and THW, where modern rates are much higher (up to 0.2 m a⁻¹ ice equivalent) than for the 4.72 ka-to-present estimates (Fig. 3c). In comparison, higher accumulation rates for the last 4.72 ka compared to modern rates are observed for the entire stretch of the WD (Fig. 3c; Table 2). Noticeably over the IMIS catchment, little change is observed between the two periods. Over the entire model domain, we observe a median relative increase of 13% in accumulation rates since 4.72 ka compared with modern rates (Fig. 4; Table 2); however, when considering only the values that fall within 100 km of either side of the WD (where mean surface speeds equal ~7 m a⁻¹), we obtain a median value of 18% higher accumulation compared with modern accumulation rates (Fig. 4).

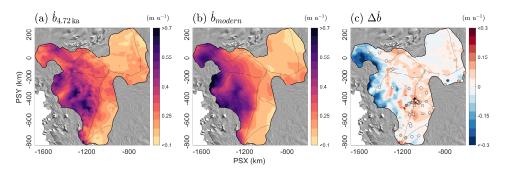


Figure 3. Gridded estimates of ice-equivalent accumulation rates for the last 4.72 ka and modern times. (a) Gridded accumulation rates inferred from the 4.72 ka IRH. (b) Modern (1979 – 2019) modelled SMB rates from RACMO2. (c) Difference between 4.72 ka-to-present and modern accumulation rates (red = 4.72 ka-to-





present accumulation higher than modern times, blue = 4.72 ka-to-present accumulation lower than modern times). The dots represent the difference between the value for the nearest grid cell in (a) and averaged accumulation rates at 79 core locations (see Section 2.4). The background image is the 2014 MODIS mosaic of Antarctica (Haran et al., 2018).

Comparison between our 4.72 ka-to-present accumulation-rate estimates and 79 point-based accumulation measurements for modern times (1651-2010 CE) are shown in Figure 3 and 4. This evaluation shows that the 4.72 ka-to-present accumulation-rate estimates for the nearest grid cell to each point measurement are, on average, 18% (p < 0.0015, n=79) higher over the entire grid and 19% (p < 0.0001, n=59) higher within 100 km of the divide than modern accumulation rates (Fig. 4). This result confirms that the relative difference in gridded accumulation rates between the 4.72 ka-to-present and modern modelled accumulation rates is consistent with modern rates from point-based measurements.

Table 2. Summary statistics for the modern (modelled and observational) and 4.72 ka-to-present ice-equivalent accumulation rates at the catchment-scale and over the WD. Values for the WD are for all points that fall within 100 km of either side of the divide (see dashed line in Figure 4). $\bar{\mu}$ refers to the median and IQR represents the Interquartile Range calculated by computing the difference between the 75th and 25th percentiles.

| | Catchment-wide | | WI | WD only | |
|--|----------------|------|------------------|---------|--|
| Accumulation rate (m a ⁻¹) | μ | IQR | $\overline{\mu}$ | IQR | |
| Modern (model) | 0.23 | 0.23 | 0.22 | 0.10 | |
| Modern (cores) | 0.24 | 0.12 | 0.24 | 0.09 | |
| 4.72 ka-to-present | 0.27 | 0.18 | 0.27 | 0.11 | |

For WD14 specifically, the nearest grid node to the ice-core site shows a 22% higher accumulation rate from 4.72 ka to present compared with modern accumulation rates (Fig. 4). There, our IRH-inferred accumulation rate is 0.28 ± 0.01 m a^{-1} , with the nearest IRH point situated 1.2 km away from WD14 showing an accumulation rate of 0.27 ± 0.01 m a^{-1} . This is only slightly higher than the direct ice-core reconstruction of accumulation rates at WD14, which show up to 19% (0.25 m a^{-1}) higher accumulation than modern rates (0.21 m a^{-1}) since 4700 years BP (Fudge et al., 2017).





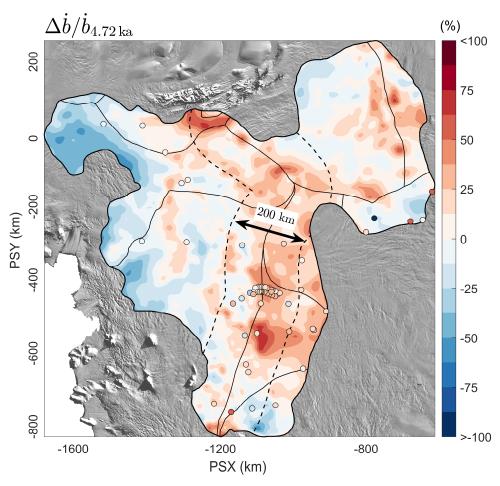


Figure 4. Relative difference in accumulation rates between the 4.72 ka-to-present estimates and modern rates. The points on the map represent the relative difference in ice-equivalent accumulation rate between the nearest grid cell in the 4.72 ka-to-present grid and the 79 observational accumulation measurements from snow, firn, and ice cores (Fig. 1; see Sect. 2.4). The dashed outline line represents the 100-km boundary on either side of the WD used to provide the summary statistics in Section 3.1 and Table 2. The background image is the 2014 MODIS mosaic of Antarctica (Haran et al., 2018).

3.2 Elevation-dependent accumulation estimates

While Figures 3 and 4 help to assess potential differences in patterns and rates across spatial scales, considering accumulation-rate differences in terms of elevation can inform how topography influences accumulation and whether this has changed over time. As a result, we binned the ice-equivalent accumulation values for each 50-m elevation bands across our model domain for both the 4.72 ka-to-present estimates and modern (RACMO2), and calculated the mean accumulation rate, the total accumulation rate, and the cumulative sum of total accumulation rate for each bin over the entire elevation gradient of the two grids (Fig. 5). As above, we observe that the accumulation-rate estimates for the period since 4.72 ka are lower at lower elevations (~700 – 1400 m) compared with RACMO2, and overtake RACMO2 near the 1400-m elevation band where 4.72 ka-to-present accumulation is higher than modern times across the divide up until ~2600 m in elevation (Fig. 5a). The lack of a large difference between the two datasets at elevations of ~250 – 700 m is primarily dominated by



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accumulation rates over IMIS, which, in contrast to the PIG and THW regions, is less exposed to the large accumulation gradients characteristic of the Amundsen Sea Embayment, i.e. higher coastal accumulation decreasing inland.

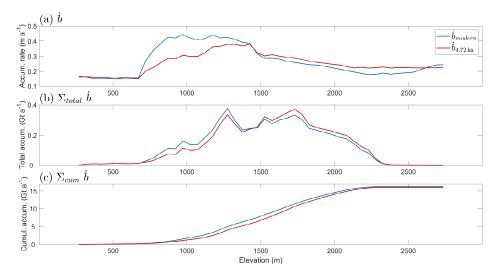


Figure 5. Comparison of ice-equivalent accumulation rates between the 4.72 ka-to-present estimates and modern rates (RACMO2) binned for each 50-m elevation bands across our study. (a) Mean accumulation rate averaged per 50-m elevation band across the survey area in m a^{-1Fb} . (b) Total accumulation rate per 50-m elevation band in Gigatons per annum (Gt a^{-1}). (c) Cumulative sum of total accumulation rate per 50-m elevation band in Gt a^{-1} .

4. Discussion

Previous studies of past accumulation rates over the WAIS have shown that accumulation varied temporally during the Holocene. Using a single airborne RES profile over the Amundsen Sea sector, Siegert and Payne (2004) showed that accumulation rates were approximately the same at 3.1 ka compared with modern rates, but ~0.3 m a⁻¹ greater (~15 %) than current rates between 3.1-6.4 ka, before which accumulation was ~50% of modern rates between 6.4 and 16.0 ka. Similarly, Neumann et al. (2008) found that accumulation rates at the WD were ~30% higher between 3-5 ka than modern values based on a dense network of IRHs traced on ground-based RES data, while Karlsson et al. (2014) found that accumulation patterns had likely changed twice during the early to mid-Holocene over PIG from the lack of a model fit between the depths and ages of two prominent IRHs. Using the updated WD14 record, Fudge et al. (2016) showed that accumulation was higher there in the mid to late-Holocene (19% between 4.72 ka BP and the present), a trend that was also observed by Koutnik et al. (2016), who found a 20% increase in accumulation rates between 2-4 ka compared with modern rates from a ground-based RES profile across the ice divide. In showing that mean accumulation rates since 4.72 ka were 18% greater than modern rates across the WD, our results provide a much wider regional picture (across ~30 % of the WAIS) that accumulation during the mid-Holocene exceeded modern rates over large parts of West Antarctica.

We gain confidence in the ability of the Nye model to estimate past accumulation rates from the 4.72 ka IRH, given that its outputs match relatively well with the direct reconstruction of mean accumulation rates at the WD14 Ice Core, with values at the ice core of 0.27 ± 0.01 m a^{-1} (this study) compared with 0.25 m a^{-1} for Fudge et al. (2016). This also suggests that the WD14 Ice Core suitably represents atmospheric conditions across the wider WD. A possible explanation for the higher accumulation rates during the mid-Holocene compared with modern values is that they represent a



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379 continued climatic transition from the LGM (Steig et al., 2001). Alternatively, it has been suggested 380 that seasonal or interannual variability, such as a weaker circumpolar vortex (Van Den Broeke and 381 Van Lipzig, 2004; Neumann et al., 2008) or teleconnections to tropical Pacific Ocean warming 382 (Sproson et al., 2022), may also lead to such difference. We did not find evidence for significant 383 changes in accumulation patterns between the mid-Holocene and modern times, suggesting that the 384 current spatial pattern of high accumulation at the coast, decreasing inland towards the divide has 385 been stable throughout the mid-Holocene over PIG and THW, as previously suggested by others 386 (Siegert and Payne, 2004; Neumann et al., 2008).

Accumulation estimates for the 4.72 ka-to-present are smaller than modern rates in the lowest elevation bands (i.e. <1400 m) (Figs. 4, 5). This pattern was also reported by Medley et al. (2014), who compared modern observational and modelled data within the ASE and hypothesised that this discrepancy at low elevations resulted primarily from a lack of sufficient accumulation measurements in the lower sections of their survey area. In our case, these low-elevation values are close to the boundary where we consider the LLA acceptable for the 4.72 ka IRH, albeit where D values are higher than for the rest of the catchment (Figure S1d), so it is more likely that accumulation rates calculated there are affected by ice-flow gradients and their influence upon IRH depths. Despite this caveat, Figure 5b shows that values at low elevations (250 – 1200 m) contribute relatively little to the total accumulation (by mass) over our survey area.

Our results reinforce the evidence that accumulation rates have varied temporally across West Antarctica during the Holocene, a finding that must be considered by future modelling studies that simulate past sea-level rise from Antarctica since the LGM. Model results from Steig et al. (2001) suggest that the maximum elevation of the WAIS was most likely reached during the early to mid-Holocene (around ~7 ka) following higher accumulation rates at the late glacial-interglacial transition, after which elevations slowly declined to present conditions as the sea-level-rise-induced kinematic wave reached the ice-sheet interior and outpaced the increase in accumulation rates. However, a moderate mid-Holocene increase in accumulation rates, which our results suggest occurred widely across the WAIS, would, if sustained, likely delay the timing of the decline in elevation by several thousand years (Steig et al., 2011). Using a flowband model, Koutnik et al. (2016) suggested that the increase of up to 40% in accumulation rates for the period 9-2 ka would likely lead to an increase in ice thickness of tens of meters during the mid-Holocene. However, because the WAIS is also sensitive to ice-dynamical changes at the ice-sheet margins (e.g. GL retreat and/or calving), an increase in accumulation rates in the upper part of the ice sheet may not necessarily result in enough thickening to counteract potential losses from ice dynamics further downstream (Jones et al., 2022). The lack of an ice-dynamical component in the model used here precludes us from reaching such a conclusion; however, higher accumulation rates of up to 18% during the mid-Holocene across 30% of the WAIS would likely be consistent with an increase of several tens of meters in ice thickness (e.g. Figure 10 of Koutnik et al., 2016). This potential increase in surface elevation is unlikely to affect the steady-state assumption of the 1-D model used here, considering that these changes are small (a few per cent of the ice thickness) and that ice thickness exceeds 3500 m in places over our survey area. We encourage future ice-sheet models to test a range of scenarios that would account for variable accumulation rates between the LGM and the present over the WAIS by using the WD14 reconstructed accumulation rates from Fudge et al. (2016) as a guide, which, as we show here, suitably represents the pattern of accumulation variability over both time and space in West Antarctica.

We may also consider the possibility for Holocene ice thickening at the divide from increased accumulation to affect GL evolution over the WAIS. Recent evidence from ice-sheet modelling and field measurements suggests that GL retreat during the Holocene was not monotonic, particularly at the Ross and Weddell Sea sides of the WAIS (Bradley et al., 2015; Kingslake et al., 2018; Neuhaus et al., 2021). Rather, Kingslake et al. (2018) showed that the GL position in the Ross and Weddell Sea



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sectors initially retreated from the LGM inland until ~9.7 – 10.2 ka, and then re-advanced to its modern position sometime during the Holocene. Although they attributed this change in GL position to the solid Earth viscoelastic response due to ice-sheet mass change and the subsequent re-grounding around pinning points, it has also been suggested that an increase in accumulation rates upstream of the GL could lead to a re-advance via ice thickening there and a subsequent increase in ice flow (Steig et al., 2001; Koutnik et al., 2016; Jones et al., 2022).

Across parts of the Weddell Sea Embayment, several studies (Ross et al., 2011; Hein et al., 2016; Ashmore et al., 2020a) have produced evidence for stability of the LGM ice thickness there until the early to mid-Holocene, contrary to most of the WAIS, after which abrupt thinning of ~400 m contributed ~1.4 – 2 m of sea level rise (Hein et al., 2016). A possible explanation for this delayed thinning in the Weddell Sea Embayment is that increased snowfall in the upper WAIS might have counteracted ice-dynamical processes at the coast until the mid to late Holocene (Hein et al., 2016; Spector et al., 2019). Similarly, over part of the Ross Sea sector, Neuhaus et al. (2021) showed that the GL over Whillans, Kamb, and Bindschadler ice streams retreated to its minimum Holocene position in the mid to late-Holocene, and then re-advanced between 1-2 ka, coinciding with periods of warmer and colder climates, respectively. They concluded that the reported GL migration was likely dominated by modest climate-induced changes upstream rather than ice dynamics further downstream, as suggested for the Weddell Sea sector (Hein et al., 2016). Our results, which provide strong and widespread evidence for higher accumulation rates at the WD during the mid-Holocene, supports further this hypothesis, as more accumulation at the divide would likely result in upstream thickening, which, in the absence of ice-dynamical processes counter-balancing this increase in accumulation, would result in GL advance in the Ross and Weddell Sea regions. We note that the increase in accumulation at ~4 – 5 ka represents the peak of a period of higher accumulation initiated from ~7 ka onwards at the WD14 Ice Core (Fudge et al., 2016), thus it is likely that our accumulation estimates form part of a wider pattern of sustained accumulation across the WD over several millennia. This noted, the pattern of GL retreat and re-advance has not been observed over the Amundsen Sea sector (Kingslake et al., 2018; Johnson et al., 2020; 2021; Braddock et al., 2022) despite the accumulation increase we observed along the WD, potentially indicating that this sector is more controlled by changes in ice dynamics for which even moderate changes in accumulation rate cannot compensate.

5. Conclusion

Using a ubiquitous Internal Reflecting Horizon found across most of the Pine Island, Thwaites, and Institute and Möller ice-stream catchments, we have estimated mid-Holocene accumulation rates in the slow-flowing parts of West Antarctica (representing 30% of total surface area of the WAIS) using a 1-D ice-flow model. By comparing our Holocene accumulation estimates with modern climate reanalysis models and observational syntheses, we estimated that accumulation rates across the Amundsen-Weddell-Ross Sea divide since 4.72 ka were, on average, 18% higher than modern accumulation rates. While the accumulation rates have therefore varied temporally, our results suggest that spatial patterns of accumulation between regions across the WAIS have remained similar during this period, i.e., higher accumulation rates at the coast and lower accumulation rates at the ice divides. The higher accumulation estimates reported here for the mid-Holocene agree well with previous, more spatially-focused studies of accumulation rates across the WAIS, which all indicate higher accumulation rates of between 15 and 30% over the last ~5 ka. Finally, our results further support the use of the WAIS Divide ice-core record to assess past atmospheric conditions across the wider Western Divide and even a substantial portion of the WAIS, making it a powerful dataset for ice-sheet models. The higher accumulation rates reported here occurred at a time of asynchronous groundingline migration over the WAIS, including re-advances of the grounding-line in the Weddell and Ross sectors and evidence for delayed deglaciation in the Weddell Sea side of the WAIS. We suggest that





ice-sheet models account for the evolution of accumulation rates over time when predicting past and
 future sea level coming from West Antarctica instead of using a fixed Last Glacial Maximum value.

Code availability

All the codes used to produce the results presented in this paper will be made available on the GitHub page of Julien A. Bodart (https://github.com/julbod, last accessed: 15 October 2022) and on Zenodo (Bodart, 2023a) upon acceptance of this manuscript.

Data availability

The IRH information for each of the three surveys used in this paper are archived in open-access repositories (Ashmore et al., 2020b; Bodart et al., 2021b; UTIG R1 layer to be made available via USAP-DC in due course) with references and links provided in the reference list. The BAS airborne radar data which were used to extract the IRHs used in this paper are fully available at the UK Polar Data Centre via the Polar Airborne Geophysics Data Portal (see Fremand, Bodart et al., 2022). The full RACMO 2.3p2 product is available on request from j.m.vanwessem@uu.nl or m.r.vandenbroeke@uu.nl. Links to access the observational point-based datasets used here are available from the respective references mentioned in the text (Section 2.4). The gridded depth and accumulation output from this study will be archived at the UK Polar Data Centre upon acceptance of this manuscript (Bodart et al., 2023b).

493 Author contribution

J.A.B. designed the study with supervision from R.G.B., D.A.Y., and D.D.B. J.A.B performed the data processing, gridding, and 1-D modelling, with contributions from J.A.M. for the modelling approach. J.A.B. interpreted the results with input from R.G.B., D.A.Y., D.D.B., and J.A.M. J.A.B. wrote the paper, with edits from all co-authors.

Competing interests

The authors declare that they have no conflict of interest.

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