

1 **High mid-Holocene accumulation rates over West Antarctica inferred from**
2 **a pervasive ice-penetrating radar reflector**

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20 **Key points**

- 21
- 22 • We estimate mean accumulation rates for the past ~4700 years across the Pine Island,
23 Thwaites, and Institute and Möller ice-stream catchments in West Antarctica using a
24 ubiquitous, ice-core dated internal radar reflection
 - 25 • Accumulation rates were 18% higher during the mid-Holocene compared to modern rates
26 over the Amundsen-Weddell-Ross divide
 - 27 • Spin-up of regional and continental ice-sheet models should include time-varying changes in
28 Holocene accumulation rates from the WAIS Divide Ice Core to generate more realistic
29 grounding-line evolution and past sea level rise contribution across this region

30 **Abstract**

31

32 Understanding the past and future evolution of the Antarctic Ice Sheet is challenged by the availability
33 and quality of observed palaeo-boundary conditions. Numerical ice-sheet models often rely on these
34 palaeo-boundary conditions to guide and evaluate their models' predictions of sea-level rise, with
35 varying levels of confidence due to the sparsity of existing data across the ice sheet. A key data source
36 for large-scale reconstruction of past ice-sheet processes are Internal Reflecting Horizons (IRHs)
37 detected by Radio-Echo Sounding (RES). When isochronal and dated at ice cores, IRHs can be used to
38 determine palaeo-accumulation rates and patterns therein. Using a spatially extensive IRH over Pine
39 Island Glacier, Thwaites Glacier, and Institute and Möller Ice Streams (covering a total of 610 000 km²
40 or 30% of the WAIS), and a local layer approximation model, we infer mid-Holocene accumulation
41 rates over the slow-flowing parts of these catchments for the past ~4700 years. By comparing our results
42 with modern climate reanalysis models (1979 – 2019) and observational syntheses (1651 – 2010), we
43 estimate that accumulation rates over the Amundsen-Weddell-Ross divide were on average 18% higher
44 during the mid-Holocene than modern rates. However, no significant spatial changes in the
45 accumulation pattern were observed. The higher mid-Holocene accumulation-rate estimates match
46 previous palaeo-accumulation estimates from ice-core records and targeted RES surveys over the ice
47 divide, and they also coincide with periods of grounding-line readvance during the Holocene over the
48 Weddell and Ross Sea sectors. We find that our spatially-extensive, mid-Holocene-to-present
49 accumulation estimates are consistent with a sustained late-Holocene period of higher accumulation
50 rates occurring over millennia at the WAIS Divide Ice Core, thus highlighting the spatial
51 representativeness of this ice core to the wider West Antarctic region. We conclude that future regional
52 and continental ice-sheet modelling studies should base their climatic forcings on time-varying
53 accumulation rates from the WAIS Divide Ice Core through the Holocene to generate more realistic
54 predictions of the West Antarctic Ice Sheet's past contribution to sea-level rise.

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56

57 **Key words:** West Antarctica, Internal Reflecting Horizons, Accumulation, Holocene, Ice-Penetrating
58 Radars, Ice-Core, Pine Island Glacier, Thwaites Glacier.

59

60 1. Introduction

61 Improving our knowledge of past climatic changes over the Antarctic Ice Sheet is required if we
62 are to understand its present evolution and model its future under increasingly rapid climatic changes
63 (IPCC, 2021). Most studies of past ice-sheet behaviour over Antarctica have focused on modelling
64 changes in ice volume and grounding-line retreat following the Last Glacial Maximum (LGM, ~20 ka
65 Before Present, BP) (Denton and Hughes, 2002; Golledge et al., 2012; 2013; Hillenbrand et al., 2013;
66 2014; Le Brocq et al., 2011; Kingslake et al., 2018); however, less attention has been paid to ice-sheet
67 evolution during the Holocene (~11.7 ka BP to present). Recent evidence suggests that parts of the
68 grounding line of West Antarctica may have retreated several hundred kilometres inland from its
69 current position at ~10 ka and subsequently readvanced to reach its modern position sometime during
70 the Holocene, due to isostatic rebound and climate-induced changes, particularly over the Weddell
71 Sea and western Ross Sea sectors (Siegert et al., 2013; Bradley et al., 2015; Kingslake et al., 2018;
72 Wearing and Kingslake, 2019; Venturelli et al., 2020; Neuhaus et al., 2021; Johnson et al., 2022).
73 However, the atmospheric and ice-dynamical conditions farther inland, which could also have
74 induced grounding-line migration, remain poorly constrained. An early investigation by Whillans
75 (1976) using radar data near Byrd Ice Core indicated stability during the Late Pleistocene and
76 Holocene epochs. Records of temperature and precipitation from the WAIS Divide Ice Core (hereafter
77 abbreviated as WD14; Fig. 1) in the central West Antarctic Ice Sheet (WAIS) suggest higher
78 accumulation rates during the Holocene than at present (Fudge et al., 2016), a trend that is also
79 observed across small parts of the Amundsen-Weddell-Ross divide (Fig. 1) near the WAIS Divide Ice
80 Core (hereafter referred to as WD14; Fig. 1) where isolated Radio-Echo Sounding (RES) surveys
81 have shown 15-30% higher accumulation rates during the mid-Holocene compared to modern values
82 (Siegert and Payne, 2004; Neumann et al., 2008; Koutnik et al., 2016).

83 Many numerical ice-sheet models that aim to predict Antarctica's long-term (past and future)
84 contribution to sea-level rise use past ice-sheet reconstructions from after the LGM to guide and
85 evaluate their models (Chavaillaz et al., 2013; DeConto and Pollard, 2016; Bracegirdle et al., 2019).
86 However, even well-used ice-sheet reconstructions assume that the ice sheet retreated continuously
87 throughout the Holocene (e.g. RAISED Consortium, 2014), a finding that has been challenged
88 recently for the WAIS (e.g. Kingslake et al., 2018). Further, significant discrepancies between model
89 simulations and the palaeo-proxy record currently impede our ability to predict confidently how the
90 ice sheet will respond to future changes in the climate (e.g. Johnson et al., 2021). While improvements
91 in model parameterisations are needed to close this gap (Bracegirdle et al., 2019; Sutter et al., 2021),
92 considerable improvement in the availability and quality of palaeo-proxy records, particularly during
93 the Holocene, is also needed to provide better constraints for ice-sheet models and ultimately better
94 resolve past ice-sheet changes (Kingslake et al., 2018; Jones et al., 2022). Palaeo-proxy data have
95 traditionally come from point-based measurements, such as ice cores (e.g. Petit et al., 1999; Parrenin
96 et al., 2007; WAIS Divide Project Members, 2013; Buizert et al., 2021), sediment cores (e.g.
97 Hillenbrand et al., 2013; Arnd et al., 2017; Hillenbrand et al., 2017; Kingslake et al., 2018; Venturelli
98 et al., 2020; Neuhaus et al., 2021; Sproson et al., 2022), or surface-exposure dating (e.g. Stone et al.,
99 2003; Suganuma et al., 2014; Johnson et al., 2014; Hein et al., 2016; Nichols et al., 2019; Johnson et
100 al., 2020; Braddock et al., 2022). A complimentary and spatially extensive alternative data source for
101 inferring past climate across an ice sheet is provided by Internal Reflecting Horizons (IRHs) detected
102 by RES. They primarily result from englacial acidity contrasts and are often detected horizontally for
103 hundreds of kilometres on RES data (Harrison, 1973; Bingham and Siegert, 2007). When employed in
104 combination with ice-core stratigraphies, IRHs can be used to extend age-depth relationships away
105 from an ice core by following peaks in electromagnetic power in the radar data (e.g. Beem et al.,
106 2021; Bodart et al., 2021a; Cavitte et al., 2016; Jacobel and Welch, 2005; MacGregor et al., 2015;
107 Whillans, 1976; Winter et al., 2019).

108 In contrast to East Antarctica and Greenland, IRH extension of WAIS ice cores has so far
109 been challenging due to fewer deep ice cores there and, until recently, the lack of well-dated IRH
110 datasets. However, efforts have intensified in recent years to improve our understanding of ice
111 stratigraphy over this sector. In particular, four recent studies using airborne RES data (Karlsson et al.,
112 2014; Muldoon et al., 2018; Ashmore et al., 2020a; Bodart et al., 2021a) all identified a distinct and
113 bright IRH dated using the Byrd and WD14 ice-core chronologies to 4.72 ± 0.28 ka BP (Muldoon et
114 al., 2018; Bodart et al., 2021a). A comparison of volcanic sulphate deposition within the WD14 and
115 Siple Dome ice cores revealed a large peak in sulphate concentration matching the age and depth of
116 this ubiquitous IRH (Kurbatov et al., 2006; Bodart et al., 2021a; Cole-Dai et al., 2021; Sigl et al.,
117 2022), which we hereafter term the “4.72 ka IRH”. This IRH has now been observed by multiple RES
118 systems and extends throughout much of the slower-flowing ice of the Amundsen and Weddell Sea
119 embayments ($< 400 \text{ m a}^{-1}$), including across the divides demarcating regions draining into the
120 Amundsen, Weddell and Ross Seas.

121 Despite their potential wide-ranging applications, the incorporation of IRHs into ice-sheet
122 models has so far been limited compared to other types of palaeo-proxy data, primarily because the
123 inference of accumulation-rate or ice-flow history from IRHs is an ill-posed inverse problem
124 (Waddington et al., 2007). Previous applications using IRHs to inform regional and continental
125 models include: (a) constraining decadal-scale Surface Mass Balance (SMB) estimates from
126 atmospheric models using annually-resolved IRHs found in the shallow firn (Medley et al., 2013;
127 2014; Van Wessem et al. 2018; Dattler et al., 2019; Kaush et al., 2020; Cavitte et al., 2022); (b)
128 inferring past accumulation rates going back further in time (i.e. 100s to 1000s years) with the aim of
129 comparing past accumulation estimates with modern times (e.g. Leysinger Vieli et al., 2004; Siegert
130 and Payne, 2004; Neumann et al., 2008; MacGregor et al., 2009; 2016; Leysinger Vieli et al., 2011;
131 Cavitte et al., 2018); or (c) integrating both their characteristics (e.g. elevation in the ice) and the
132 information inferred from them (e.g. accumulation or basal-melt rates) to evaluate the output from
133 regional and continental ice-sheet models (Leysinger Vieli et al., 2011; 2018; Holschuh et al., 2017;
134 Sutter et al., 2021). Promisingly, Sutter et al. (2021) recently showed that spatially extensive Antarctic
135 IRHs can provide unique benchmarks for constraining ice-sheet model parameterisations (i.e. climate
136 forcing and simulated ice flow), which are then used to simulate palaeo ice-sheet evolution. Together,
137 these studies indicate multiple avenues for ice-sheet models to assimilate IRHs further improve
138 estimates of past, current and future ice-sheet changes.

139
140 Here, we estimate mid-Holocene accumulation rates across the WAIS from first-order
141 calculations using a one-dimensional (1-D) model, constrained by the spatially extensive 4.72 ka IRH.
142 We first describe the data, the model used and their limitations and uncertainties (Sect. 2). We then
143 present our accumulation-rate estimates and compare to observed and modelled modern accumulation
144 rates to reveal a longer-term perspective on changes between the mid-Holocene and the present (Sect.
145 3). Finally, we place our results in the context of previous studies that consider WAIS evolution
146 during the Holocene (Sect. 4).

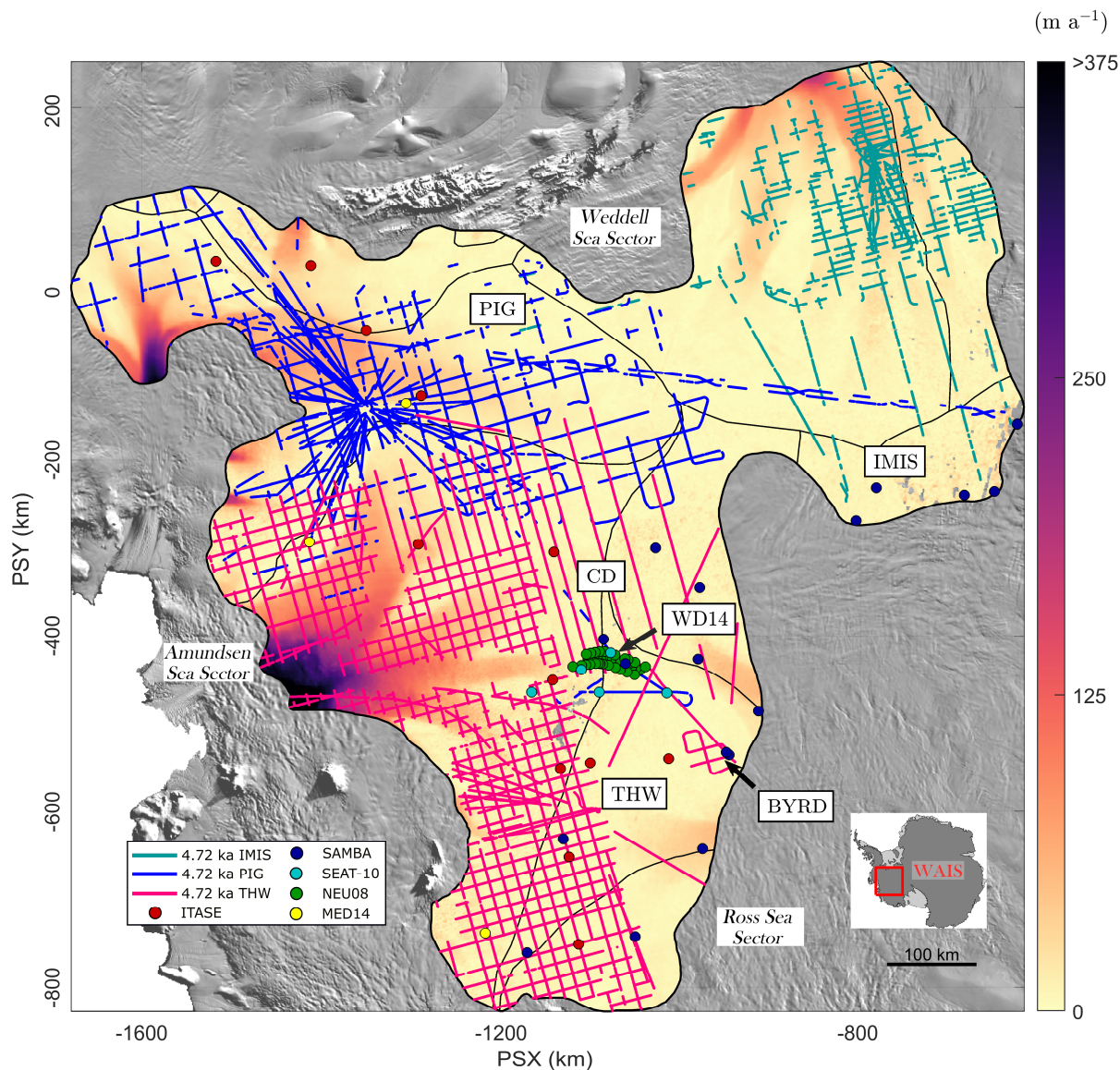
147 **2. Data and methods**

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149 **2.1 Along-track IRH data**

150 We used data from extensive (~91 000 flight-track km) RES surveys acquired across West
151 Antarctica between 2004 and 2018. The main contributing surveys are the University of Texas
152 Institute for Geophysics (UTIG) 2004-2005 AGASEA survey flown over Thwaites Glacier (THW)
153 and Marie Byrd Land which deployed the 60-MHz High Capability Airborne Radar Sounder
154 (HiCARS) radar system (Holt et al., 2006; Peters et al., 2007), and the British Antarctic Survey (BAS)
155 2004-05 BBAS survey over Pine Island Glacier (PIG) and 2010-2011 IMAFI survey over Institute
156 and Möller Ice Streams (IMIS) which deployed the 150-MHz Polarimetric Airborne Survey
157 INstrument (PASIN) radar system (Vaughan et al., 2006; Corr et al., 2007; Ross et al., 2012; Frémand,

158 Bodart et al., 2022) (Fig. 1; Table 1). Additional profiles from NASA’s Operation Ice Bridge (OIB;
 159 MacGregor et al., 2021) 2016 and 2018 surveys, flown with the 190-MHz Multichannel Coherent
 160 Radar Depth Sounder 2 (MCoRDS-2) radar system (CReSIS, 2018), were also used to extract IRH
 161 information near the WD14 Ice Core and upper IMIS catchments (Bodart et al., 2021a; Figure 1 and
 162 Table 1). We refer the reader to the above references for comprehensive details on each system’s
 163 capabilities.



164
 165 Figure 1. Map of the datasets and key locations in this study. The three datasets that contain the 4.72 ka
 166 IRH are colour-coded as IMIS (green), PIG (blue), and THW (pink). IRH data where $D > 1$ are excluded (see
 167 Section 2.2.1; Figure S1). Points represent the snow, firn and ice cores used in this study to compare modern
 168 accumulation rates with those inferred from the 4.72 ka IRH (Sect. 2.4). The background colour map shows
 169 modern surface speeds from Rignot et al. (2017). Locations mentioned in this paper are abbreviated on the map,
 170 as follows: BYRD (Byrd Ice Core), IMIS (Institute and Möller Ice Streams), PIG (Pine Island Glacier), THW
 171 (Thwaites Glacier), WAIS (West Antarctic Ice Sheet), CD (Central Amundsen-Weddell-Ross Divide), WD14
 172 (WAIS Divide Ice Core). Major ice divides are from Mouginito et al. (2017). The background image is the 2014
 173 MODIS mosaic of Antarctica (Haran et al., 2018). For all analysis and figures in this study, the SCAR Antarctic
 174 Polar Stereographic projection is used (EPSG: 3031).

175 These RES surveys were used to track and date six IRHs spanning the Late Pleistocene and
 176 Holocene (25.7 – 2.3 ka BP) that collectively cover much of the WAIS, including IMIS (Ashmore et

177 al., 2020a), PIG (Karlsson et al., 2014; Bodart et al., 2021a) and THW (Muldoon et al., 2018). Here
 178 we only consider the 4.72 ka IRH mapped in all four studies and shown in Figure 1, as it is by far both
 179 the most spatially extensive and the only commonly traced IRH across all studies. We first merged all
 180 data points from the 4.72 ka IRH across the three catchments, resulting in a cumulative distance of
 181 ~40 000 line-km of IRH profiles (44% of the RES surveys' total coverage; Table 1). Although the
 182 along-track RES data were acquired with a trace spacing of between 10 and 35 m, depending on the
 183 dataset used, we re-sampled these points to 500 m in the along-track direction. We then added a
 184 spatially invariant firm correction of 10 m onto the Muldoon et al. (2018) dataset to match the same
 185 firm correction applied by the other studies to correct the IRH depth. Finally, we calculated the median
 186 value of all ice thicknesses and IRH depths falling within each 500 m interval.

187 Table 1. Characteristics of each IRH dataset used in this study that contain the 4.72 ka IRH. 'Reflector
 188 1' in Muldoon et al. (2018) is abbreviated here as 'R1'.

<i>Survey name</i>	<i>Survey provider</i>	<i>RES system</i>	<i>Dataset reference</i>	<i>Cumulative IRH distance (10³ km)</i>
IMAFI	BAS	PASIN 150-MHz	H2 in Ashmore et al. (2020a)	15
BBAS / OIB	BAS / NASA	PASIN 150-MHz / MCoRDS-2 190-MHz	R2 in Bodart et al. (2021a)	6
AGASEA	UTIG	HiCARS 60-MHz	R1 in Muldoon et al. (2018)	19

189 2.2 Inferring accumulation rates

190 To infer accumulation rates from the 4.72 ka IRH, we used the Nye model, a 1-D ice-flow
 191 model widely used for estimating accumulation rates and age-depth relationships over relatively slow-
 192 flowing parts of an ice sheet (Nye, 1957; Fahnestock et al., 2001a). This model invokes the local-layer
 193 approximation (LLA), i.e. it assumes that the time-averaged accumulation rate that the IRH has
 194 experienced since its upstream inception at the surface can be adequately represented by its depth
 195 where it is observed presently. Other 1-D models exist, including the Dansgaard-Johnsen (Dansgaard
 196 and Johnsen, 1969) and the shallow-strain rate model (MacGregor et al., 2016), but were less suitable
 197 for estimating accumulation rates here due to uncertainty in the basal shear layer thickness across our
 198 survey area and because we are limited to only one IRH to constrain the ice-flow model respectively.
 199 The Nye model assumes that ice thickness is constant and therefore that the ice sheet has been in a
 200 steady state since the deposition of the IRH, an acceptable assumption for the period under
 201 investigation here. The Nye model states:

$$202 \quad \dot{b}_a = \ln\left(\frac{z_a}{H} \frac{H}{a}\right), \quad (1)$$

203 where \dot{b}_a is the mean accumulation rate during the Holocene epoch between an IRH of age a and the
 204 present, z_a represents the depth of the IRH dated at the ice core, and H is the ice thickness. The model
 205 assumes that the vertical strain rate, $\dot{\epsilon}_{zz}^a$, is also constant and vertically uniform, so that it exactly
 206 balances the overburden of local ice accumulation:
 207

$$208 \quad \dot{\epsilon}_{zz}^a = \frac{\dot{b}_a}{H}. \quad (2)$$

209 We iterated Eq. (1) over the re-sampled 500-m spaced dataset using the depth of the 4.72 ka
 210 IRH for z_a and used the median radar-derived ice-thickness measurement re-sampled over the 500-m
 211 grid to obtain H , when this information was available. In areas where the radar did not sound the bed,

211 we used the BedMachine Antarctica v2 gridded product to obtain a value for H (Morlighem, 2020).
212 Note that accumulation rate values presented in this study are all reported in m a^{-1} of ice equivalent.

213 **2.2.1 Assessing the suitability of the 1-D model**

214 To quantify the suitability of the LLA from which our accumulation rates are based, we
215 calculated the effects of horizontal gradients in modern ice thickness and accumulation rates along
216 particle paths in their ability to affect IRH depths across our grid, as per Waddington et al. (2007).
217 Where these gradients are large, estimates of accumulation rates from IRHs likely require a more
218 complete treatment of ice flow and its effect upon IRH depths, which multi-dimensional models and
219 more physically complete models can better resolve (e.g. Waddington et al., 2007; Leysinger Vieli et
220 al., 2011; Karlsson et al., 2014; Nielsen et al., 2015; Koutnik et al., 2016;). However, such models are
221 significantly more computationally expensive over such a larger area and depend on well-constrained
222 boundary conditions from along-flow radar profiles which are not often available at an ice-sheet level
223 (MacGregor et al., 2009).

224 We quantified the effect of horizontal gradients on an IRH of age a by first estimating the
225 total horizontal particle path length L_{path} each “particle” of the 4.72 ka IRH has travelled since a , and
226 then the characteristic lengths of variability in ice thickness (L_H) and apparent accumulation rate (L_b)
227 (Supplementary Information). These three components were then combined to generate a non-
228 dimensional parameter D (Fig. S1d), which we used as a confidence metric for our inferred
229 accumulation rates. Both Waddington et al. (2007) and MacGregor et al. (2009) suggested a value of
230 $D \ll 1$ over Antarctica, whereas MacGregor et al. (2016) used a maximum value of $D = 1$ to estimate
231 where the LLA is acceptable over Greenland. Because D cannot be translated simply into an
232 uncertainty in an LLA-inferred accumulation rate, it is not yet clear what exact value is appropriate.
233 Smaller values of D indicate that local horizontal gradients in ice thickness and accumulation rates
234 have a smaller effect on IRH depth of age a , and thus that the LLA may be valid (Waddington et al.,
235 2007; MacGregor et al., 2009; 2016). Where $D \geq 1$, the depth of an IRH is less likely to be the result
236 of accumulation rates at the surface or vertical strain rates further down, and thus a more sophisticated
237 model is likely required (Sect. 2.2.2) (Waddington et al., 2007). However, MacGregor et al. (2009)
238 found that even along a flowband across Lake Vostok where the mean value of D is 0.50 for a 41-ka
239 IRH, the difference in accumulation rate inferred from the LLA and from a more sophisticated
240 flowband model could be relatively small (4-16%). This similarly suggests that accumulation rate can
241 be inferred acceptably using the LLA where D is higher.

242 For our study area, D values are mostly well below unity (median: 0.19; 25th quartile: 0.09;
243 75th quartile: 0.34), which suggests relatively little effect from ice-dynamical processes upon IRH
244 depths across most of our grid. We used the upper quartile of the D distribution across our model
245 domain (i.e. $D \leq 0.34$) to show areas where we can have confidence that accumulation rate remains
246 the dominant factor influencing the vertical position of our IRHs in the ice column (i.e. where the
247 $D \ll 1$ criterion is likely met; Fig. S1d). While accumulation rates inferred from IRHs situated in the
248 upper quartile (Fig. S1d) may still be valid, we suggest additional caution in interpreting our results
249 there due to the potential impact of larger flow gradients on IRH depths.

250 **2.2.2 Model limitations and uncertainty**

251 One of the main limitations of the Nye model is that it assumes that gradients in sliding
252 velocity are mostly concentrated in a thin layer at the ice-bed interface and that the ice column
253 deforms by pure shear only (Nye, 1957; Fahnestock et al., 2001a). For this reason, the Nye model is
254 generally only appropriate for IRHs found in the upper part of the ice column, as is the case here.
255 Additionally, the use of the model is restricted to areas where ice flow is relatively slow and
256 horizontal strain rates are also relatively low.

257 Here we focus on a shallower IRH situated in the upper part of the ice column (median: 40%;
258 25th quartile: 30%; 75th quartile: 50%; Fig. 2b-c), for which we can reasonably assume that the ice
259 sheet has remained close to steady state and where IRHs are likely shallow enough not to have
260 experienced appreciable flow disturbances that would affect the Nye model assumptions.
261 Additionally, due to the inherent nature of tracking IRHs through RES data, our coverage is limited to
262 areas where ice-flow speeds are relatively low and IRHs are undisturbed. In some portions of our
263 study area, the IRH is found deeper in the ice column or in faster-flowing sections of the ice sheet
264 (e.g. in the downstream sectors of our grid, Fig. 1-2b-c); areas where the assumptions that the 1-D
265 model is based on may be challenged.

266 Estimating uncertainty in accumulation rates from the Nye model is non-trivial. Previous
267 studies have used the misfit between the accumulation rate calculated using multiple proximal IRHs
268 in the ice column (e.g. Fahnestock et al., 2001a; 2001b; Leysinger Vieli et al., 2004; MacGregor et al.,
269 2016). Unfortunately, this method is not suitable here due to the dearth of spatially extensive IRHs
270 younger than 4.72 ka over our model domain.

271 Instead, uncertainty in the Nye-inferred accumulation rates were calculated using: (a) the
272 lowest and highest possible accumulation rates from Eq. (1) using the age uncertainty (± 0.28 ka) of
273 the 4.72 ka IRH and (b) the lowest and highest possible accumulation rates inferred from an additional
274 1-D model (Eq. S5) which accounts for the effect of strain rates on accumulation rates (i.e. the
275 shallow-strain rate model from MacGregor et al. (2016); Supplementary Information; Fig. S2-4).

276 This calculation provides lower and upper bounds for the IRH-inferred accumulation rates
277 (Fig. S4a-b), which were then averaged to generate a relative uncertainty (Fig. S4c). From this
278 assessment, we estimate a median relative uncertainty in the Nye-inferred accumulation rates for the
279 4.72 ka IRH of 14% across our grid. This uncertainty is higher in the downstream edges of our grids,
280 particularly over the PIG, THW and IMIS catchments, and generally low over the Amundsen-
281 Weddell-Ross divide (Fig. S4), reflecting the effect of spatially variable strain rates on the inferred
282 accumulation rates. When combined with the assessment of the suitability of the LLA and exclusion
283 of IRHs where the $D > 1$ (Sect. 2.2.1-2.2.2), we conclude that it supports our application of a 1-D
284 modelling approach here.

285 **2.3 Gridding and filtering**

286 Once IRH depths and accumulation rates for the 4.72 ka IRH were obtained at regular 500-m
287 points along RES flight paths, we filtered the results using a moving-average Gaussian filter of length
288 30 samples (equivalent to ~ 15 km) to reduce along-track noise in the IRH depth, and then gridded the
289 filtered result using a Delaunay-triangulation-based natural neighbour interpolation method onto a 1-
290 km polar stereographic grid. We further smoothed the gridded data using an 18-km square cell mean
291 filter to limit the localised interpolation artefacts in areas of poor survey coverage. Figure S5 shows
292 the maximum distance away from the nearest 500-m along-track point used to produce Figures 2-3,
293 and thus where errors in the interpolated grids are expected to be larger. The median value of this
294 maximum distance is 5 km and its maximum value is 75 km, which is comparable to previous studies
295 that infer SMB from IRHs in the shallow firn (e.g. Medley et al., 2014). We evaluated other possible
296 interpolation methods (e.g. kriging and using different semi-variogram models), but they resulted in
297 similar or poorer quality and were thus discounted.

298 **2.4 Comparison with modern observations**

299 To compare our inferred accumulation estimates for the past 4.72 ka with modern values
300 (defined here as 1651-2019), we derived information on modern accumulation rates from two sources,
301 one modelled (gridded) and one from a series of observational (point-based) datasets.

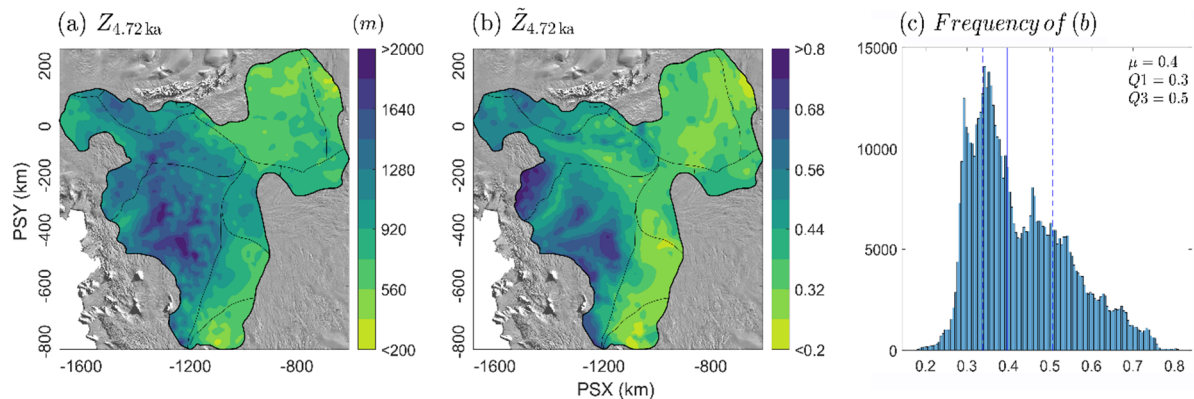
302 We used modelled gridded accumulation rates from the Regional Atmospheric Climate Model
 303 2.3p2 (hereafter RACMO2) 1979-2019 SMB product forced at its margin with the ERA-Interim
 304 product (native resolution: 27 km) as an estimate for modern accumulation rates (Van Wessem et al.,
 305 2018). Although SMB is not technically equivalent to the accumulation rate, runoff and sublimation
 306 are negligible in our survey area (Medley et al., 2013) so we assume SMB is equal to accumulation
 307 rate in this region. We converted modelled values from $\text{kg m}^{-2} \text{a}^{-1}$ to m a^{-1} of ice equivalent using an
 308 ice density value of 917 kg m^{-3} , calculated the 40-year mean, and then bi-linearly interpolated the
 309 gridded RACMO2 product to the same 1-km grid resolution as our 4.72 ka-to-present accumulation
 310 grid (Sect. 2.3) to ensure consistency when comparing both datasets.

311 Observational point-based measurements were obtained from a series of snow, firn and ice
 312 cores from the ITASE (Mayewski and Dixon, 2013), MED14 (Medley et al., 2014), SAMBA (Favier
 313 et al., 2013), and SEAT-10 (Burgener et al., 2013) datasets, as well as from a network of centennially-
 314 averaged modern accumulation rates derived from shallow IRHs traced on ground-based RES data
 315 over the central divide and dated using a shallow ITASE Ice Core (Neumann et al., 2008) (Fig. 1).
 316 This resulted in 79 point-based accumulation measurements from cores covering the period 1651-
 317 2010 CE (Common Era) and spread across our model domain (see Figure 1). Further detail on these
 318 datasets can be found in the above references.

319 To compare the Holocene gridded product with the point-based measurements, we first
 320 calculated the average value of the accumulation rate at the point measurement for the entire period.
 321 We converted these values to ice-equivalent accumulation rates and then extracted two paired values,
 322 i.e. the value for the point-measurement for modern accumulation rates and the value for the nearest
 323 grid cell in the gridded 4.72 ka-to-present accumulation estimates to this measurement.

324 Results

325 The final grids for depth and accumulation rates for the 4.72 ka IRH are shown in Figures 2 and 3.
 326 In total, these grids are made of $\sim 89\,000$, 500-m spaced points, which cover an area of $\sim 610\,000 \text{ km}^2$,
 327 or $\sim 30\%$ of the total surface area of the WAIS. The grids span most of the PIG and THW glacier
 328 catchments, as well as the Ronne (upper Rutford, Institute, and Möller) and upper western Ross
 329 (Bindschadler, Kamb, MacAyeal, and Whillans) catchments (IPY Antarctic boundaries G-H, J-Jpp,
 330 and Ep-F; Mouginito et al. (2017); Fig. 1-2). Overall, the 4.72 ka IRH is shallower within the IMIS
 331 and upper PIG and THW catchments, as well as on the Ross side of the central divide where ice
 332 thickness is particularly deep (Fig. 2b). Conversely, the 4.72 ka IRH is deeper in the ice near a 400-m
 333 high bedrock plateau that separates the northern and southern basins of PIG (Vaughan et al., 2006)
 334 and at two locations in the upstream parts of the main trunk of THW where ice flows over highs in
 335 subglacial topography (Fig. 2b).



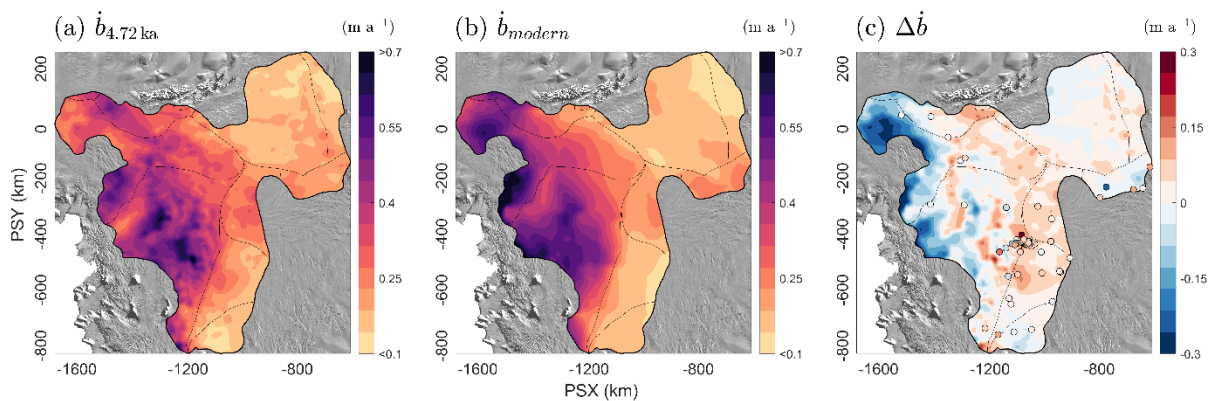
336

337 Figure 2. Gridded depths for the 4.72 ka IRH across the model domain covering the PIG, THW, and
 338 Institute and Möller ice-stream catchments. (a) Gridded depth of the 4.72 ka IRH. (b) Normalised depth of the

339 4.72 ka IRH relative to ice thickness. (c) Histogram showing the distribution of values in (b) with the median ($\bar{\mu}$)
 340 and interquartile range (i.e. 25th (Q1) and 75th (Q3) quartiles) shown as solid and dashed blue lines respectively.
 341 The background image is the 2014 MODIS mosaic of Antarctica (Haran et al., 2018).

342 3.1. Catchment-scale accumulation estimates

343 Figure 3 shows a comparison of the ice-equivalent accumulation rates we inferred for the 4.72
 344 ka IRH (Fig. 3a) and modern SMB estimates from RACMO2 (Fig. 3b). We observe that the IRH
 345 accumulation rate pattern for the last 4.72 ka is similar to the modern pattern of accumulation rates for
 346 the Amundsen Sea sector of the WAIS, which is dominated by higher coastal accumulation rates that
 347 progressively decrease inland to reach their lowest rates over the Ross side of the divide (Fig. 3a-b).
 348 Differences in accumulation rates between the 4.72 ka-to-present estimates and modern values are
 349 mainly observed directly upstream of the main trunks of PIG and THW, where modern rates are much
 350 higher (up to 0.2 m a^{-1} ice equivalent) than for the 4.72 ka-to-present estimates (Fig. 3c). In
 351 comparison, higher accumulation rates for the last 4.72 ka relative to modern rates are observed for
 352 the entire stretch of the Amundsen-Weddell-Ross divide (Fig. 3c; Table 2). Over the IMIS catchment,
 353 little change is observed between the two periods. Over the entire model domain, we observe a
 354 median percentage change value of 6% higher accumulation since 4.72 ka compared with modern
 355 rates (Fig. 4); however, when considering only the values that fall within 100 km of either side of the
 356 Amundsen-Weddell-Ross divide (i.e. in the accumulation zone of the Amundsen, Weddell, and Ross
 357 Sea sectors and where mean surface speeds average $\sim 7 \text{ m a}^{-1}$), we obtain a median percentage change
 358 value of 18% higher accumulation compared with modern accumulation rates (Fig. 4).



359
 360 Figure 3. Gridded estimates of ice-equivalent accumulation rates for the last 4.72 ka and modern times.
 361 (a) Gridded accumulation rates inferred from the 4.72 ka IRH. (b) Modern (1979 – 2019) modelled SMB rates
 362 from RACMO2. (c) Difference between 4.72 ka-to-present and modern accumulation rates (red = 4.72 ka-to-
 363 present accumulation higher than modern times, blue = 4.72 ka-to-present accumulation lower than modern
 364 times). The dots represent the difference between the value for the nearest grid cell in (a) and time-averaged
 365 accumulation rates at each of the 79 core locations (see Section 2.4; Fig. S6). The background image is the 2014
 366 MODIS mosaic of Antarctica (Haran et al., 2018).

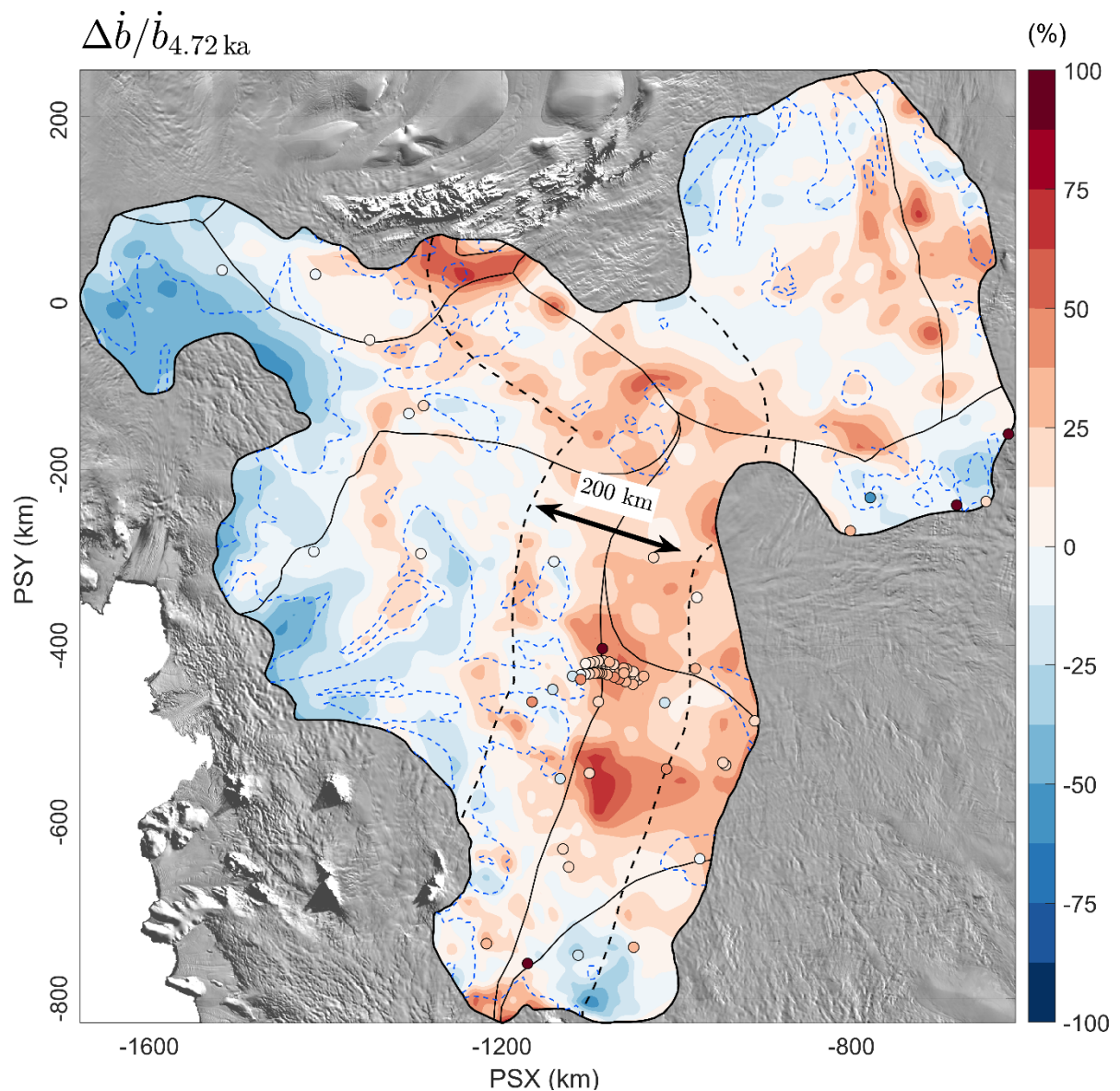
367 Comparison between our 4.72 ka-to-present accumulation-rate estimates and 79 core-derived
 368 point-based accumulation measurements for modern times (1651-2010 CE) are shown in Figures 3-4
 369 and S6. This evaluation shows that the 4.72 ka-to-present accumulation-rate estimates for the nearest
 370 grid cell to each point measurement are, on average, 22% higher for cores situated across the entire
 371 grid ($p < .0015, n=79$) and 23% higher for cores found within 100 km of the divide compared with
 372 modern accumulation rates ($p < .0001, n=59$; Figs. 4 and S6). In comparison, a similar analysis
 373 between grid cells from the 4.72 ka-to-present accumulation-rate estimates and RACMO2 at these 79
 374 core locations shows mid-Holocene accumulation rate estimates are, on average, 32% ($P < .00002,$
 375 $n=79$) higher for cores situated across the entire grid and 36% higher for cores found within 100 km
 376 of the divide ($p < .00001, n=59$; Fig. S6). This result confirms that the relative change for gridded

377 accumulation rates between the 4.72 ka-to-present and modern modelled accumulation rates is
 378 consistent with modern rates from point-based measurements.

379 Table 2. Summary statistics for the modern (modelled and observational) and 4.72 ka-to-present ice-
 380 equivalent accumulation rates at the catchment-scale and over the Amundsen-Weddell-Ross divide (abbreviated
 381 CD here). Values for the Amundsen-Weddell-Ross divide are for all points that fall within 100 km of either side
 382 of the divide (see dashed line in Figure 4). $\tilde{\mu}$ refers to the median and IQR represents the Interquartile Range
 383 calculated by computing the difference between the 75th and 25th percentiles. Note that the values provided in
 384 the text represent the median relative change from the cell-by-cell change between each grid (Fig. 4), rather than
 385 the relative change of the median values provided here.

Accumulation rate ($m a^{-1}$)	Catchment-wide		CD only	
	$\tilde{\mu}$	IQR	$\tilde{\mu}$	IQR
Modern (model)	0.23	0.23	0.22	0.10
Modern (cores)	0.24	0.12	0.24	0.09
4.72 ka-to-present	0.27	0.18	0.27	0.11

386

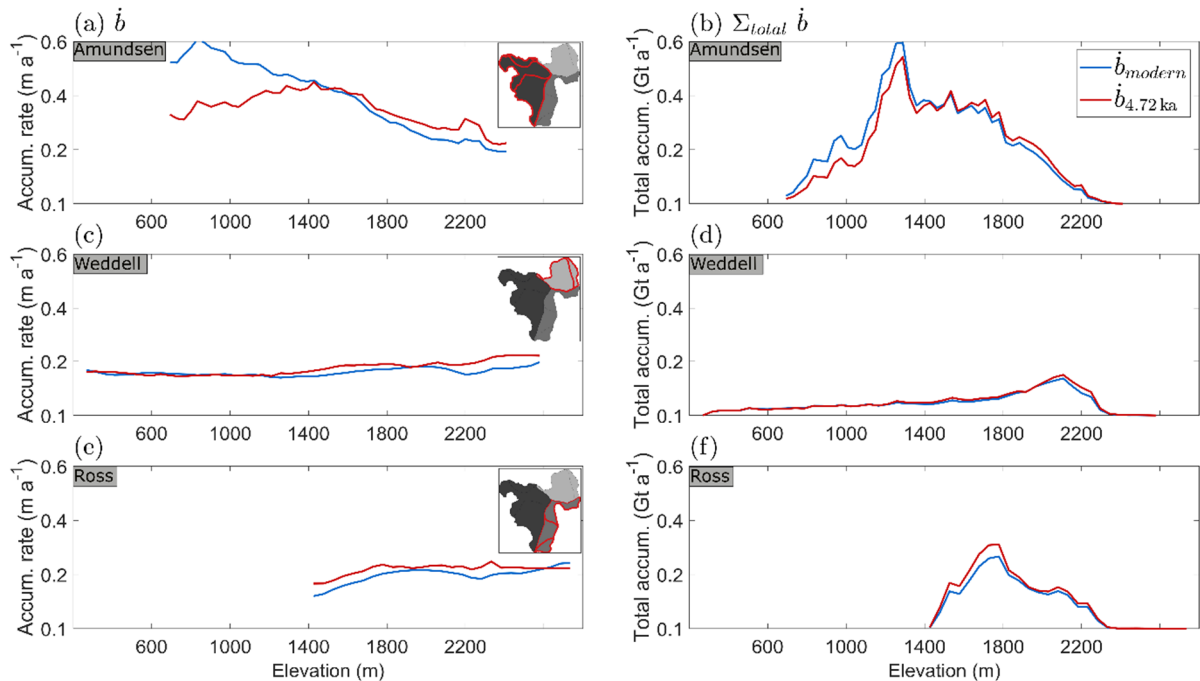


387

388 Figure 4. Relative change in accumulation rates between the 4.72 ka-to-present estimates and modern
 389 rates. The points on the map represent the relative change in ice-equivalent accumulation rate between the
 390 nearest grid cell in the 4.72 ka-to-present grid and the 79 modern observations from cores (Figs. 1 and S6; Sect.
 391 2.4). The dashed black outline line represents the 100-km boundary on either side of the Amundsen-Weddell-
 392 Ross divide used to provide the summary statistics in Section 3.1 and Table 2. The dashed blue line shows the
 393 contours of the upper limit of the interquartile range for the D parameter ($D \leq 0.34$) (Sect. 2.2.1-2.2.2). The
 394 background image is the 2014 MODIS mosaic of Antarctica (Haran et al., 2018).

395 3.2 Elevation-dependent accumulation estimates

396 While Figures 3 and 4 help to assess potential differences in patterns and rates across spatial
 397 scales, considering accumulation-rate differences in terms of elevation can inform how topography
 398 influences accumulation and whether this has changed over time. We binned the ice-equivalent
 399 accumulation values by 50-m elevation bands across the three main catchments covering our grid
 400 (Amundsen, Weddell and Ross) for both the 4.72 ka-to-present estimates and modern model rates and
 401 calculated the mean accumulation rate and the total accumulation rate for each bin over the entire
 402 elevation gradient (Fig. 5). We again find that the accumulation-rate estimates for the period since
 403 4.72 ka are lower at low elevations ($\sim 700 - 1400$ m) over the Amundsen sector compared with
 404 RACMO2, but begin to exceed RACMO2 near the 1400-m elevation band where the 4.72 ka-to-
 405 present accumulation rate is higher than modern times across the divide (Fig. 5a-b). We also note that
 406 whilst an elevation-dependent gradient in accumulation rates, dominated by high accumulation at the
 407 coast decreasing inland, exists over this sector for the mid-Holocene, it is much less marked than for
 408 present rates. This is not surprising, as this sector is where we observe the largest relative
 409 uncertainties in inferred accumulation rates across our grid (Fig. S4), indicating that the 1-D model is
 410 less able to produce realistic accumulation rates in the downstream end of our grid where ice flow is
 411 faster and strain rates are likely higher. In comparison to the Amundsen sector, accumulation rates
 412 since 4.72 ka are generally higher at all elevations for the Weddell and Ross sectors compared with
 413 the present, although this difference is less than over the Amundsen sector (Fig. 5 c-f).



414
 415 Figure 5. Comparison of ice-equivalent accumulation rates between the 4.72 ka-to-present estimates and
 416 modern rates (RACMO2) binned by 50-m elevation bands across the three main catchments considered here
 417 (Amundsen, Weddell, and Ross). (a, c, e) Mean accumulation rate averaged per 50-m elevation band across the
 418 specific catchment area in m a^{-1} . (b, d, f) Total accumulation rate per 50-m elevation band across the specific
 419 catchment area in Gigatonnes per annum (Gt a^{-1}).

420 4. Discussion

421 4.1. Comparison with other Holocene accumulation estimates

422 Previous studies of past accumulation rates over the WAIS have shown that accumulation varied
423 temporally during the Holocene. Using a single airborne RES profile over the Amundsen Sea sector,
424 Siegert and Payne (2004) showed that accumulation rates were approximately the same at 3.1 ka
425 compared with modern rates, but $\sim 0.3 \text{ m a}^{-1}$ greater ($\sim 15 \%$) than current rates between 3.1-6.4 ka,
426 before which accumulation was $\sim 50\%$ of modern rates between 6.4 and 16.0 ka. Similarly, Neumann
427 et al. (2008) found that accumulation rates at the Amundsen-Weddell-Ross divide were $\sim 30\%$ higher
428 between 3-5 ka than modern values based on a dense network of IRHs traced on ground-based RES
429 data, while Karlsson et al. (2014) found that accumulation patterns had likely changed twice during
430 the early to mid-Holocene over PIG from the lack of a model fit between the depths and ages of two
431 prominent IRHs. Using the updated WD14 record, Fudge et al. (2016) showed that accumulation rates
432 were higher there in the mid to late-Holocene (19% between 4.72 ka BP and the present), a trend that
433 was also observed by Koutnik et al. (2016), who found a 20% increase in accumulation rates between
434 2-4 ka compared with modern rates from a ground-based RES profile across the ice divide.

435 These studies together point to a period of increasing accumulation observed at the WD14 Ice
436 Core from ~ 7 ka onwards (Fudge et al., 2016; their Figure 2), with its peak matching the age of the
437 4.72 ka IRH used here. Thus, our accumulation-rate estimates likely form part of a wider pattern of a
438 sustained increase in accumulation across the Amundsen-Weddell-Ross divide over several millennia.
439 In showing that mean accumulation rates since 4.72 ka were 18% greater than modern rates modelled
440 from RACMO2 across the Amundsen-Weddell-Ross divide, our results provide much wider regional
441 support for the hypothesis that accumulation rates during the mid-Holocene exceeded modern rates
442 across central West Antarctica. A possible explanation for the higher accumulation rates during the
443 mid-Holocene compared with modern values is that they represent a continued climatic transition
444 from the LGM (Steig et al., 2001). Alternatively, it has been suggested that seasonal or interannual
445 variability, such as a weaker circumpolar vortex (van Den Broeke and van Lipzig, 2004; Neumann et
446 al., 2008), or teleconnections to tropical Pacific Ocean warming (Sproson et al., 2022), may also lead
447 to such difference. We did not find evidence for significant changes in accumulation patterns between
448 the mid-Holocene and modern times, suggesting that the current spatial pattern of high accumulation
449 on the Amundsen side of the divide transitioning to low accumulation on the Ross side of the divide
450 was stable throughout the mid-Holocene, as previously suggested by others (Siegert and Payne, 2004;
451 Neumann et al., 2008; Koutnik et al., 2016).

452 We also find that accumulation estimates for the 4.72 ka-to-present are smaller than modern rates
453 in the lowest elevation bands ($< 1400 \text{ m}$), particularly over the Amundsen Sector (Fig. 5 a-d). This
454 pattern was also found by Medley et al. (2014), who compared modern observational and modelled
455 data over this sector and hypothesised that this discrepancy at low elevations resulted primarily from a
456 lack of sufficient accumulation measurements in the lower sections of their survey area. In our case,
457 these low-elevation values are close to the boundary where we consider the LLA acceptable for the
458 4.72 ka IRH, albeit where D values are higher than for the rest of the catchment (Figure S1d), so it is
459 more likely that accumulation rates calculated there are affected by ice-flow gradients and their
460 influence upon IRH depths leading to lower accumulation rates there. Despite this caveat, Figures 5b
461 and 5d show that values at low elevations contribute relatively little to the total accumulation (by
462 mass) over our survey area.

463 We suggest that future ice-sheet modelling studies investigate the difference in accumulation rates
464 inferred from our 1-D model using multi-dimensional flowband models to assess effects of divergent
465 and convergent flow on IRH depth and ultimately accumulation rates, as previously considered
466 elsewhere in Antarctica (MacGregor et al., 2009). This could be conducted along a flowline
467 transitioning from the slow-flowing regions directly downstream of the Amundsen-Weddell-Ross

468 divide to the coastal margins of our grid, particularly over THW where we observe the largest
469 uncertainties in accumulation rates. In addition, we suggest that future modelling studies use the
470 accumulation-rate variability from the WD14 Ice Core as a climate forcing in their ice-sheet models.
471 Koutnik et al. (2016) previously showed that the WD14 record is unique in that it provides a reliable
472 record of accumulation-rate variability during the Holocene, which other East Antarctic ice-core
473 records often used to reconstruct the evolution of the WAIS do not possess. We found that these
474 higher accumulation rates are spatially extensive across nearly one third of the WAIS, further
475 suggesting that the WD14 Ice Core is indeed representative of the wider WAIS and can be used in
476 regional or continental ice-sheet models as a reliable climate forcing for the region. Future regional
477 and continental ice-sheet models should make use of this record to adjust their climatic boundary
478 conditions to provide improved estimates of ice-elevation change and grounding-line evolution over
479 Antarctica.

480 **4.2 Impact for ice-sheet elevation change during the Holocene**

481 Model results from Steig et al. (2001) suggest that the maximum elevation of the WAIS was most
482 likely reached during the early to mid-Holocene (around ~7 ka) following higher accumulation rates
483 at the late glacial–interglacial transition, after which the WAIS slowly declined to present conditions
484 as the sea-level-rise-induced kinematic wave reached the ice-sheet interior and outpaced the increase
485 in accumulation rates. However, higher accumulation rates in the mid-Holocene relative to the
486 present, which our results suggest occurred spatially across the WAIS, would likely delay the timing
487 of this thinning by several thousand years (Steig et al., 2011).

488 Using a flowband model, Koutnik et al. (2016) suggested that an increase of up to 40% in
489 accumulation rates for the period 9 – 2 ka would likely have led to an increase in ice thickness of tens
490 of metres during the mid-Holocene. Although this finding was warranted by physical assumptions
491 around the response time of the ice-sheet interior to adjust to an increase in accumulation in the
492 model, it points to the potential for the divide to have thickened by several metres over a relatively
493 short period of time from increased accumulation rates alone. Because the WAIS is also sensitive to
494 ice-dynamical changes at the ice-sheet margins (e.g. grounding-line retreat or calving), an increase in
495 accumulation rates in the upper part of the ice sheet may not necessarily result in enough thickening to
496 counteract potential dynamical losses farther downstream (Jones et al., 2022). Conway and
497 Rasmussen (2008) reported that the Amundsen-Ross Divide is currently thinning and migrating
498 towards the Ross Sea at a speed of 10 m a⁻¹, but they were unable to determine whether this was in
499 response to long-term (last two millennia) accumulation-rate changes there or short-term (last few
500 centuries) ice-dynamical forcing from the coastal margins of the Amundsen and Ross sectors. More
501 recently, Balco et al. (2023) showed that Thwaites and Pope glaciers experienced 35 m of thickening
502 in the mid-to-late Holocene, when accumulation rates were higher than present. While this thickening
503 relative to present was attributed to glacio-isostatic rebound, it is also possible that higher
504 accumulation rates in the upstream sections of the WAIS contributed to this thickening, if sustained
505 over millennia.

506 The lack of an ice-dynamical component in the model used here precludes us from evaluating
507 any ice-surface-elevation change associated with changing accumulation rates. However, 18% higher
508 accumulation rates during the mid-Holocene relative to the present across 30% of the WAIS could be
509 consistent with an elevation increase of several tens of metres in ice thickness, according to Koutnik
510 et al. (2016). Even if tens of metres of ice-surface-elevation change occurred, it is still unlikely to
511 significantly affect the steady-state assumption of the 1-D model used here (constant ice thickness
512 over time), because such changes are small (a few per cent of the ice thickness) and that ice thickness
513 exceeds 3500 m in places over our survey area.

514 **4.3 Impact for grounding-line evolution during the Holocene over the WAIS**

515 Finally, we consider the possibility for Holocene ice thickening at the divide from increased
516 accumulation rates to affect downstream grounding-line evolution. Recent evidence from ice-sheet
517 modelling and field measurements suggest that grounding-line retreat during the Holocene was not
518 monotonic, particularly at the Ross and Weddell Sea sides of the WAIS (Bradley et al., 2015;
519 Kingslake et al., 2018; Neuhaus et al., 2021). Rather, Kingslake et al. (2018) showed that the
520 grounding-line position in the Ross and Weddell Sea sectors initially retreated from the LGM inland
521 until $\sim 10.2 - 9.7$ ka, and then readvanced to its modern position sometime during the Holocene.
522 Although they attributed this change in grounding-line position to the solid Earth viscoelastic
523 response due to ice-sheet mass change and the subsequent re-grounding around pinning points, it has
524 also been suggested that an increase in accumulation rates upstream of the grounding line could lead
525 to a readvance via ice thickening there and a subsequent increase in ice flow (Steig et al., 2001;
526 Koutnik et al., 2016; Jones et al., 2022). Across parts of the Weddell Sea Embayment, several studies
527 have produced evidence for stability of the LGM ice thickness there until the early to mid-Holocene
528 (Ross et al., 2011; Hein et al., 2016; Ashmore et al., 2020a), contrary to most of the WAIS, after
529 which abrupt thinning of ~ 400 m contributed $\sim 1.4 - 2$ m of sea level rise (Hein et al., 2016). A
530 possible explanation for this delayed thinning in the Weddell Sea Embayment is that increased
531 snowfall in the upper WAIS might have counteracted ice-dynamical processes at the coast until the
532 mid-to-late Holocene (Hein et al., 2016; Spector et al., 2019). Similarly, over part of the Ross Sea
533 sector, Neuhaus et al. (2021) showed that the grounding line over Whillans, Kamb, and Bindschadler
534 ice streams retreated to its minimum Holocene position in the mid to late-Holocene, and then
535 readvanced between $2 - 1$ ka, coinciding with periods of warmer and colder climates, respectively.
536 They concluded that the reported grounding-line migration was likely dominated by modest climate-
537 induced changes upstream rather than ice dynamics further downstream, as suggested for the Weddell
538 Sea sector (Hein et al., 2016).

539 Our results, which provide strong and widespread evidence for higher accumulation along the
540 Amundsen-Weddell-Ross divide during the mid-Holocene compared with the present, support these
541 hypotheses further, as higher accumulation rates at the divide would likely result in upstream
542 thickening (Sect. 4.2). In the absence of ice-dynamical processes counter-balancing this increase in
543 accumulation rates, the grounding-line should advance in these regions. However, we note that the
544 pattern of grounding-line retreat and readvance has not been observed over the Amundsen Sea sector
545 (Kingslake et al., 2018; Johnson et al., 2020; 2021; Braddock et al., 2022) despite the accumulation-
546 rate increase we also observed along the Amundsen-Weddell-Ross divide and the recent results from
547 Balco et al. (2023). This complication may indicate that the Amundsen sector is more strongly
548 influenced by coastal changes in ice dynamics, for which even moderate changes in accumulation rate
549 cannot compensate.

550 **5. Conclusion**

551 Using a ubiquitous internal reflecting horizon found across most of the Pine Island, Thwaites, and
552 Institute and Möller ice-stream catchments, we have estimated mid-Holocene accumulation rates in
553 the relatively slow-flowing parts of West Antarctica, representing 30% of total surface area of the
554 WAIS.

555 By comparing our Holocene accumulation-rate estimates with a modern climate reanalysis model
556 and observational syntheses, we estimated that accumulation rates across the Amundsen-Weddell-
557 Ross Sea divide since 4.72 ka were, on average, 18% higher than modern values. Our results suggest
558 that spatial patterns of accumulation across the WAIS have remained stable during this period, i.e.
559 higher accumulation rates on the Amundsen side of the divide transitioning to lower accumulation
560 rates on the Ross side of the divide. The higher accumulation rates reported here for the mid-Holocene
561 compared to the present agree well with earlier, spatially-focused studies of accumulation rates, all of
562 which indicate higher accumulation rates (+15 - 30%) over the past ~ 5 ka. This change in magnitude

563 occurred at a time of asynchronous grounding-line migration over the WAIS, including readvances of
564 the grounding line in the Weddell and Ross sectors and evidence for delayed deglaciation in the
565 Weddell Sea side of the WAIS.

566 The higher mid-Holocene accumulation estimates inferred here over large sectors of the WAIS
567 occurred at a time of sustained, millennial-scale increase in accumulation rates found at the WAIS
568 Divide Ice Core. This pattern indicates that the ice core is suitably representative of the climatic
569 conditions of the wider region over time. We suggest that future regional or continental ice-sheet
570 modelling studies base their palaeoclimate forcing on modern spatial SMB products that are
571 modulated over time using the WAIS Divide Ice Core record. This will enable those models to obtain
572 a more realistic climatic forcing representative of the past conditions of the wider WAIS, and
573 ultimately, constrain ice-sheet volume change and grounding-line evolution during the Holocene.

574 **Code availability**

575 All the codes used to produce the results presented in this paper will be made available on the
576 GitHub page of Julien A. Bodart (<https://github.com/julbod>, last accessed: 15 October 2022) and on
577 Zenodo (Bodart et al., 2023) upon acceptance of this manuscript.

578 **Data availability**

579 The IRH information for each of the three surveys used in this paper are archived in open-
580 access repositories (Ashmore et al., 2020b; Bodart et al., 2021b; UTIG R1 layer to be made available
581 via USAP-DC in due course) with references and links provided in the reference list. The BAS
582 airborne radar data which were used to extract the IRHs used in this paper are fully available at the
583 UK Polar Data Centre via the Polar Airborne Geophysics Data Portal (see Fremand, Bodart et al.,
584 2022). The RACMO2 product is available on request from j.m.vanwessem@uu.nl or
585 m.r.vandenbroeke@uu.nl. Links to access the observational point-based datasets used here are
586 available from the respective references mentioned in the text (Section 2.4). The gridded depth and
587 accumulation output from this study will be archived on Zenodo upon acceptance of this manuscript
588 (Bodart et al., 2023).

589 **Author contribution**

590 J.A.B. designed the study with supervision from R.G.B., D.A.Y., and D.D.B. J.A.B
591 performed the data processing, gridding, and 1-D modelling, with contributions from J.A.M. for the
592 modelling approach. J.A.B. interpreted the results with input from R.G.B., D.A.Y., D.D.B., and
593 J.A.M. J.A.B. wrote the paper, with edits from R.G.B., D.A.Y., J.A.M., D.W.A., E.Q., A.S.H.,
594 D.G.V., D.D.B.

595 **Competing interests**

596 The authors declare that they have no conflict of interest.

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