

Reply to Referee comments

Manuscript title

Parameterizing anisotropic reflectance of snow surfaces from airborne digital camera observations in Antarctica

Date of reply

5 31 August 2020

Dear Dr. Marie Dumont,

we would like to thank the two anonymous referees and Aleksey Malinka for their comments and suggestions as part of the discussion phase of our submitted manuscript. Please find the point-by-point responses to the comments below followed by a marked-up version of the manuscript.

In addition to the edits made with regard to the referee comments, we implemented the following changes in the manuscript:

- The tabulated data including the retrieved f_{iso} , f_{vol} , and f_{geo} as well as θ_0 , l_{rough} , R_{opt} , and auxiliary information (time and position of the measurements) is available for all research flights in the Publishing Network for Geoscientific & Environmental Data (PANGAEA): <https://doi.pangaea.de/10.1594/PANGAEA.921707>.
The data availability section was updated accordingly.
- The references Ehrlich et al. (2008) and Yang et al. (2014) were updated in the bibliography as they referred to the wrong publications.
- Minor edits in the abstract and conclusions (see marked-up manuscript version).

20 Kind regards on behalf of all authors,
Tim Carlsen

Replies to comments of Anonymous Referee #1

1. My comments mostly pertain to how the terms BRDF and HDRF are used. It still remains a bit confusing in your introduction the difference between the BRDF and the HDRF. Since this is still often used non-correctly in many publications it would make sense to be more careful and not use the terms so interchangeably. My understanding from Schaeppman-Strub et al. 2006 is that the BRDF is not something that can be measured whereas the HDRF is. Thus, it is confusing when you then say on line 26 that comparison of in situ measured BRDF with simulations, as the BRDF is not measured in situ as far as I understand it. The equations for each are shown in the methodology, which is helpful, but it does remain a bit confusing they way it is discussed. Since the HDRF is what is actually being measured from the digital camera, I think that all needs to be stated more clearly upfront and it would be good to add discussion on how Eq.18 relates to Eq. 3 for those less familiar with these topics – i.e. students.

Thank you for this comment. We agree that the used terminology and the differentiation between HDRF and BRDF is done too sloppily in some other publications. We tried to be careful in this matter, however as you correctly pointed out, some inconsistencies slipped our attention. In case the actual measured quantity was the HDRF, we changed the term BRDF to HDRF (although stated differently in the respective publication). We also define an 'effective BRDF' as done in Gatebe and King (2016) in case the FOV is small and the atmospheric influence is considered. We believe this is reasonable for measurements with the CAR instrument and decided to reference the CAR database in the same way as reported by Gatebe and King (2016) as this constitutes an important dataset.

We agree that, so far, the relation between Eq. 18 and Eq. 3 was not obvious. We therefore omit the rather theoretical definition given so far in the manuscript (Eq. 3) and give already the more conceptual definition via the BRDF. This way, Eq. 3 already has the form of Eq. 18 and the relation should become more clear to the reader.

Changes in text:

- Page 3 Line 29: The comparison of in situ measurements of the snow reflectance with simulations is essential in terms of model validation. Observations of the HDRF (or effective BRDF in case the FOV is small and the atmospheric influence is considered) are conducted using a variety of different measurement concepts.
- P4 L10: 'Several studies observed systematically less anisotropy for a typical snow BRDF than estimated from simulations [...]': we replaced 'BRDF' with 'HDRF'
- P4 L19: 'Thus, the latter are more suitable for studying the influence of macroscopic surface roughness on the surface BRDF.': we replaced 'BRDF' with 'HDRF'
- P4 L26: 'and the snow BRDF loses its azimuthal symmetry (Warren et al., 1998)': we replaced 'BRDF' with 'HDRF'
- P5 L8: 'Cox and Munk (1954) analyzed radiance calibrated analog photographs for the parameterization of the ocean BRDF': we replaced 'BRDF' with 'HDRF'

– P5 L10: 'The instantaneous measurement of multiple viewing angles facilitates aerial BRDF measurements with digital cameras.': we replaced 'BRDF' with 'HDRF'

– P6 L5: As the BRDF of an ideal Lambertian surface is $(\pi \text{sr})^{-1}$, the BRF is given by

$$R_{\text{BRF}} = \frac{dI_r(\theta_i, \varphi_i; \theta_r, \varphi_r)}{dF_i(\theta_i, \varphi_i)} \cdot \frac{dF_i(\theta_i, \varphi_i)}{dI_r^{\text{ideal}}(\theta_i, \varphi_i)} = \pi \text{sr} \cdot f_{\text{BRDF}}.$$

5 In Eq. 2 and in the remainder of this section, the spectral dependence is omitted for reasons of simplicity. In atmospheric conditions, both the BRDF and BRF cannot be measured directly as the global irradiance F_0 reaching the surface is composed of a direct (F_i) and diffuse (F_{diff}) component. In this case, the measurable quantity is the HDRF. The definition of the HDRF is analogous to the BRF, but includes irradiance from the entire hemisphere (denoted with 2π):

$$\begin{aligned} R_{\text{HDRF}} &= \pi \text{sr} \cdot \frac{dI_r(\theta_i, \varphi_i, 2\pi; \theta_r, \varphi_r)}{dF_0(\theta_i, \varphi_i, 2\pi)} \\ &= R_{\text{BRF}}(\theta_i, \varphi_i; \theta_r, \varphi_r) \cdot f_{\text{dir}} + R(2\pi; \theta_r, \varphi_r) \cdot (1 - f_{\text{dir}}). \end{aligned}$$

10 – new Eq. 18:

$$R_{\text{HDRF}}(\theta_0, \varphi_0; \theta_r, \varphi_r) = \frac{\pi \text{sr} \cdot I(\theta_r, \varphi_r)}{F^\downarrow(\theta_0, \varphi_0)}.$$

2. I also do not follow why the downward irradiance from SMART could not be used to calculate the HDRF. If it's good enough for the surface albedo, and thus the grain size determination, then I do not follow why it's not good enough for the incoming solar radiation. Also, if as you state you are mostly interested in the shape of the HDRF, not the absolute magnitude of the values, then the calibration shouldn't have mattered? I would like to see a comparison between your modeled and measured incoming solar irradiance.

The grain size determination relies on surface albedo measurements. The calculation of the surface albedo from the raw spectra measured with SMART involves three main steps: (a) the correction for dark and stray light, (b) the correction for the non-ideal angular response of the sensor–spectrometer system, and (c) the cross- calibration of the up- and downward looking optical inlets to account for different sensitivities of the sensors-spectrometer systems. The latter part (c) was performed in the field during the observations and comprised successive measurements of the signals of the up- and downward-looking sensors under the same illumination with an integrating sphere. Thus, for albedo measurements, an absolute calibration converting the digital numbers registered by the spectrometers into units of irradiance is not required. All calibration steps for measurements of the surface albedo could be carried out, which is why the retrieval of the optical-equivalent snow grain size from SMART measurements is not affected by these calibration issues.

This is different for the downward irradiance that requires an absolute calibration converting the raw signal into units of irradiance. The absolute calibration was performed in the laboratory. Normally, the cross-calibration is used to transfer the absolute calibration as measured in the laboratory to the field to account for changes in the transmissivity of the sensor–fiber–spectrometer system. Unfortunately, problems with the power supply of the integrating sphere in the field occurred,

and the integrating sphere got destroyed after the campaign before it could be used later for an additional laboratory calibration of SMART. Due to this failed transfer calibration, an absolute calibration was not possible and the downward irradiance measurements from SMART could not be used for the calculation of the HDRF. For the same reason, a comparison between modeled and measured incoming solar irradiance is not possible.

5 Instead, the global irradiance was simulated along the flight track with libRadtran using DISORT. The simulated irradiance was integrated over the wavelength range of each camera channel and weighted with the RSR function of the camera. The use of simulations limits the validity of absolute values of the measured HDRF to cloudless conditions. It is true that we are mostly interested in the shape of the HDRF. However, the retrieved values for the model parameter f_{iso} would be influenced by the absolute values of the HDRF. We therefore restrict our analysis to cloudless cases when we are confident that the simulated
10 values for the downward irradiance were representative of the actual measurement conditions.

As both reviewers mentioned this, we added a clarification in the manuscript.

Changes in text:

15 – P14 L4: The downward irradiance measurements from SMART could not be used for the calculation of the HDRF due to calibration issues. Note that this pertains to the radiometric calibration only, which converts the digital numbers registered by the spectrometer into units of irradiance. For the albedo measurements with SMART, a relative calibration of the upper and lower sensors is sufficient and an absolute radiometric calibration is not required. Thus, the albedo measurements and the retrieval of the optical-equivalent snow grain size are unaffected by this calibration issues. For the
20 calculation of the HDRF, the global irradiance was simulated [...] instead.

3. What were the cloud vs. clear sky conditions during these flights? Were synoptic cloud observations not also obtained? Could you not see whether or not there were clouds with the SMART measured incoming solar irradiance?

In approximately 75 % of the camera observations, cloudless conditions prevailed. Indeed, cloudless cases were identified from visual synoptic observations during the flights as well as using the downward measured irradiance by SMART. Periods with fast fluctuations in the downward irradiance were flagged as cloudy. Clouds influence the HDRF measurements in changing the direct/diffuse ratio of incoming solar radiation. For a quantitative analysis, the cloud scene during the times of measure-
25 ment (cloud cover, optical thickness, etc.) needs to be well characterized, which is not possible from the observations available during the flights. Thus, we restricted our analysis to cloudless conditions only.

Changes in text:

– P14 L13: The use of simulations limits the validity of absolute values of the measured HDRF to cloudless conditions.
30 Cloudless cases (approximately 75 % of the camera observations) were identified from visual synoptic observations

during the flights as well as using the downward irradiance measured with SMART. Periods with fast fluctuations in the downward irradiance were flagged as cloudy. Although mainly the shape of the HDRF is analyzed within this work (which is independent from the absolute value of F^\downarrow), the analysis is restricted to cloudless conditions only.

4. I don't follow exactly what was done in section 3.6 for the inversion. How was the HDRF used in this context? Also, seems you setting the HDRF equal to the BRDF in Equation 12 but we don't find that out until section 3.7 so be good to mention it earlier. However, in section 3.7 you then mention you don't expect the atmospheric conditions to be large, which is true though as you mention it is wavelength dependent with the blue channel more impacted. Also the discussion then focuses on the BRF and HDRF relationship with the proportion of direct vs. diffuse but again it's the BRDF that you are substituting with the HDRF so I think it would be better to keep the discussion in that context as the interchanging of terms is hard to follow.

Thank you for suggesting this improved structure of the manuscript. We agree that it is helpful for the reader to mention the substitution of the BRDF with the HDRF in Equation 12 earlier. We therefore changed the structure as follows: we switched Sects. 3.6 and 3.7 and added the suggested clarification at the beginning of the new Sect. 3.6. Concerning your second suggestion, we believe it is better to continue discussing the BRF and HDRF relationship as this is what was simulated by Schaepman-Strub et al. (2006). However, we added a reminder about the close relation between the BRF and BRDF by referring to Eq. 2 at this point.

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Changes in text:

- P15 L2: For each averaged HDRF, the weighting functions f_{iso} , f_{vol} , and f_{geo} are retrieved by inverting the RTLSR BRDF model and setting the HDRF equal to the BRDF in Eq. 12. This is necessary as the BRDF is not measurable under atmospheric conditions. However, the atmospheric influence [...]
- 15 - P15 L11: Schaepman-Strub et al. (2006) simulated the difference between the HDRF and BRF ($R_{\text{BRF}} = \pi \cdot f_{\text{BRDF}}$, compare Eq. 2) for snow surfaces
- switched Section 3.6 and 3.7

5. It is well known that the anisotropy increases with increasing solar zenith angle so it would be good to reference some early publications that have already discussed this (i.e. some early work by Warren seems relevant here). It's also well known that surface roughness reduces the overall forward scattering as there is more backscatter, so again referencing earlier work is important here. This is also how data from MISR are currently being used to map surface roughness over ice sheets and sea ice.

We totally agree and added more references in the discussion part of the manuscript.

Changes in text:

- P19 L18: This expected increase in the anisotropy of the snow HDRF for increasing θ_0 is obvious in the changing model parameters f_k and has been discussed in earlier studies (e.g., Dirmhirn and Eaton, 1975; Kuhn, 1985; Warren et al., 1998; Hudson et al., 2006).
- 5 – P20 L1: This is in accordance with Warren et al. (1998) who observed a reduction of the forward reflection peak due to sastrugi during tower measurements at South Pole Station. The effect of surface roughness on the BRDF has been previously investigated in observational (e.g., Grenfell et al., 1994; Hudson and Warren, 2007) and modeling studies (e.g., Leroux and Fily, 1998; Zhuraleva and Kokhanovsky, 2011) and is facilitated by the multi-angular reflectance data from MISR to map surface roughness over ice sheets and sea ice from space (Nolin et al., 2002).

6. I'm surprised that there is no mention of how BRDF uncertainties translate into albedo and absorbed solar energy uncertainties as that is what is really important after all. While the paper is already quite long, this would complete the study. If the MODIS BRDF model is off, how much does it influence the albedo and the energy balance of the ice sheets?

- 10 Thank you for this suggestion. We agree that the influence on the surface energy budget would be an important point. While it would be possible to translate the retrieved model parameters into a spectral albedo corresponding to the wavelength range of the camera channel (490-585 nm for green), it is the broadband albedo that controls the surface energy budget. For the narrow-to-broadband albedo conversion, the MODIS albedo product uses precomputed coefficients and information from all spectral channels of MODIS. The camera observations could ideally only cover a spectral range between 400 to 700 nm (compare
- 15 Figure 3). As the HDRF is strongly wavelength dependent (e.g., Hudson et al., 2006; Marks et al., 2015; Gatebe and King, 2016), it is not possible to imply the same discrepancies between the camera measurements and MODIS at other wavelengths. This holds especially true as the camera observations in the visible wavelength range do not cover longer wavelengths above 1000 nm where ice absorption becomes dominant. All in all, these points would make an attempt at a surface energy budget calculation doubtful in our view.

Replies to comments of Anonymous Referee #2

General Comments

1. Section 1: provides an excellent background literature review of snow reflectance anisotropy, snow BRDF models and verification of the models with existing BRDF measurements.

Thank you! We tried to be more rigorous in distinguishing between the BRDF and HDRF as some inconsistencies in referencing other studies slipped our attention. In case the actual measured quantity was the HDRF, we changed the term BRDF to HDRF (although stated differently in the respective publication).
5 HDRF, we changed the term BRDF to HDRF (although stated differently in the respective publication). We further added a publication from an earlier study using airborne camera data for BRDF observations for forest classification.

Changes in text:

- Koukal, T. and Atzberger, C.: Potential of multi-angular data derived from a digital aerial frame camera for forest classification, IEEE J. Sel. Topics Appl. Earth Observ. Remote Sens., 5, 30–43, <https://doi.org/10.1109/JSTARS.2012.2184527>, 2012.

2. Section 2: Methodology – the definitions of reflectance quantities are articulated well, however, the use of symbols is inconsistent. For example, in subsection 2.1, the symbol for the solar zenith angle is given as $\theta(i)$ (see Equation 1), then the subscript changes to something else, $\theta(0)$ (see Equation 3) without any explanation. In subsection 2.2, the symbol for the relative azimuth angle are different both in Equation 4 and Equation 7.

Thanks a lot for spotting these typos. We introduce the reflectance quantities and the modeling (Sects. 2.1 and 2.2) with the general illumination zenith and azimuth angles (θ_i and φ_i) as, in principle, the illumination source does not need to be the Sun. Only when referring to specific measurements (from Sect. 3
15 on), we use θ_0 and φ_0 as the solar zenith and azimuth angle, respectively. We changed the subscripts in Eqs. 3-11 accordingly and use the consistent symbol $\Delta\varphi$ for the relative azimuth angle in Eqs. 4-11.

Changes in text:

- Changed subscripts in Eqs. 3-11.

3. Subsection 3.2 - optical-equivalent grain size retrieval from spectral surface albedo measurements. It was reported elsewhere in the manuscript (subsection 3.4) that “the downward irradiance measurements from SMART could not be used for the calculation of the HDRF due to calibration issues.” So does this imply that the optical-equivalent snow grain size retrieved from SMART measurements may be affected by those calibration issues? It’s also odd that the “optical-equivalent snow grain size from SMART and analogous ground-based measurements were validated against grain size observations utilizing reflectance measurements with MODIS.” Which retrievals are assumed to be the "truth"?

All calibration steps for measurements of the surface albedo could be carried out, which is why the retrieval of the optical-equivalent snow grain size from SMART measurements is not affected by these calibration issues. This is different for the downward irradiance that requires an absolute radiometric calibration converting the raw signal into units of irradiance. The calibration issues were related to the radiometric calibration only. Thus, the downward irradiance measurements from SMART could not be used for the calculation of the HDRF. The other reviewer asked the same question, for more details please see our answer to Comment #2 of Reviewer #1.

As both reviewers mentioned this, we added a clarification in the manuscript.

Thank you also for the second part of your comment. It was not our intention to assume one retrieval to be the truth. We rephrased the sentence accordingly.

Changes in text:

- Page 14 Line 4: The downward irradiance measurements from SMART could not be used for the calculation of the HDRF due to calibration issues. Note that this pertains to the radiometric calibration only, which converts the digital numbers registered by the spectrometer into units of irradiance. For the albedo measurements with SMART, a relative calibration of the upper and lower sensors is sufficient and an absolute radiometric calibration is not required. Thus, the albedo measurements and the retrieval of the optical-equivalent snow grain size are unaffected by this calibration issues. For the calculation of the HDRF, the global irradiance was simulated [...] instead.

- P10 L10: The retrieved R_{opt} from SMART and analogous ground-based measurements were compared to grain size observations utilizing reflectance measurements with MODIS (Carlsen et al., 2017).

4. Subsection 3.3.2 – Radiance calibration and image post-processing. This subsection provides important details about the post processing, but there are no details about the pre-deployment calibration. Given that the flights were performed between 24 December 2013 and 5 January 2014, it may be shown that each flight need a different calibration, which would depend on the calibration stability of the digital camera Canon EOS-1D Mark III.

Thank you for your comment. Due to the problems with the integrating sphere during the campaign
5 (compare Comment #2 from Reviewer #1), no transfer calibration in the field could be analyzed. However, unlike for SMART, there is no reason to expect a change of the camera calibration as no mechanical parts are unplugged during the deployment (laboratory vs. field). A comparison of laboratory calibrations before and after the campaign and for different temperature regimes did not show significant changes for the camera calibration.

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Specific Comments

1. Pg. 1. Line 17: The study finds that “MODIS observations generally underestimated the anisotropy of the surface reflection”, but it would help to provide the degree to which MODIS underestimates the anisotropy of surface reflection.

Suggestion taken, thank you. We added the average factor of underestimation for the volumetric model weight (which is the dominant contribution to the anisotropy in the analyzed cases). The factor 10 was calculated when neglecting the possible artifacts such as the latitudinal band around 76° S with MODIS
15 values for f_{vol} of exactly zero. We added this information in the abstract, results and discussion, and conclusions.

Changes in text:

- P1 L16: For the analyzed cases, MODIS observations (545-565 nm wavelength band) generally underestimated the anisotropy of the surface reflection. The largest deviations were found for the volumetric model weight f_{vol} (average underestimation by factor of 10).
- P24 L32: For the analyzed cases, MODIS underestimates the volumetric model weight f_{vol} on average by at least a factor of 10 compared to the airborne measurements (neglecting MODIS values of exactly zero).
- P26 L10: The airborne values for f_{vol} are larger than the corresponding MODIS retrievals (by at least a factor of 10).

2. Pg. 2, Line 4: The statement, “Satellites monitor the reflectance (i.e., reflected radiance in units of $\text{Wm}^{-2}\text{sr}^{-1}$) at the top of the atmosphere (TOA). However, they are restricted in terms of the number of available observation angles and spectral bands as well as their temporal resolution.” Is this in reference to polar orbiting satellites or geostationary satellites?

Thank you for this question. The statement is in reference to polar-orbiting satellites such as relevant for the MODIS instrument onboard Aqua/Terra satellites. The temporal resolution for geostationary satellites is of course much better, however the use of geostationary satellites for global surface albedo measurements comes with the drawback of less global coverage (especially missing polar regions) and the need to generate consistent estimates from different sensors. Eventually, this leads to different algorithms for surface albedo estimates for polar-orbiting and geostationary satellites. We clarified the relevance for polar-orbiting satellites in the manuscript.

Changes in text:

- P2 L4: Polar-orbiting satellites monitor the reflectance [...]
- P2 L6: The processing of measurements from polar-orbiting satellites for monitoring the broadband surface albedo [...]

3. Pg. 2, Line 8, the statement, “During the first step, the TOA reflectance is converted into a surface reflectance by means of an atmospheric radiative transfer parameterization,” need to be clarified. What is “atmospheric radiative transfer parameterization?”

With ‘atmospheric radiative transfer parameterization’ we mean that the atmospheric correction algorithm has to correct for gaseous and aerosol scattering and absorption and surface-atmosphere-coupling effects (e.g. due to the anisotropic BRDF of the surface) using radiative transfer modeling. To avoid confusion in that regard, we clarified the statement in the manuscript. Thanks for this comment.

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Changes in text:

- P2 L8: During the first step, the TOA reflectance is converted into a surface reflectance by correcting for gaseous and aerosol scattering and absorption applying radiative transfer modeling (e.g., Vermote and Kotchenova, 2008).

4. Pg. 2, line 10, the recommended terminology: bidirectional reflectance-distribution function (BRDF). Note the “reflectance-distribution” is a compound adjective and should be hyphenated.

- 10 Thank you for pointing this out. As this might be correct, we prefer omitting the additional hyphen for reasons of consistency applying the terminology as defined in Schaepman-Strub et al. (2006).

5. Pg. 3, line 19, change the statement, “Comparing this asymptotic model to in situ observations of the BRDF...” to “Evaluating this asymptotic model with in situ observations of the BRDF...”

Changed as suggested. To be more consistent with regard to the impossibility to measure the BRDF in situ, we also rephrased it to ‘in situ measurements of snow reflectance’.

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Changes in text:

- P3 L21: Evaluating this asymptotic model with in situ measurements of the snow reflectance, Kokhanovsky et al. (2005) found a generally good agreement but a reduced model accuracy in the solar principal plane at large observation angles.

6. Pg. 3, in the paragraph, marked by lines 8-25, various snow reflectance terms are mentioned: snow BRDF (line 10), a polarized BRDF (line 12), reflectance of snow (line18), inherent optical properties (line 22), and bidirectional reflectance (line 24). This may be confusing to readers and it would help to clarify or use consistent terms.

Thanks for pointing this out. We rephrased, where possible, some of the terms in order to make it easier for the reader. However, the terms mentioned in the paragraph are not all interchangeable and need to remain as stated in order to stay rigorous. We added some details for 'inherent optical properties' for clarity. The bidirectional reflectance (former line 24) cannot be replaced, as Malinka (2016) calculate the BRDF which is not yet defined in the manuscript and we prefer to leave it like that to avoid confusion.

10 **Changes in text:**

- P3 L14: changed 'polarized BRDF model' to 'snow BRDF model including polarization'
- P3 L20: changed 'to model the reflectance of snow' to 'to model the BRDF of snow'
- P3 L24: added details for 'inherent optical properties': 'inherent optical properties (extinction coefficient, single scattering albedo, scattering phase function, and the polarization properties)'

7. Pg. 3, line 29, the use of "solar azimuth" and "azimuthal directions" in this sentence is confusing, "Kuhn (1985) observed a peak in reflectance in the azimuthal directions up to 60° to both sides of the solar azimuth that becomes more prominent with increasing solar zenith angle and snow grain size." What is the zero azimuth in this case?

15 To avoid a possible confusion, we rephrased this sentence. The zero azimuth is the solar azimuth angle, the $\pm 60^\circ$ refer to the azimuth angle range of reflected radiation relative to the solar azimuth angle.

Changes in text:

- P3 L32: Kuhn (1985) observed a peak in reflectance that is contained within $\pm 60^\circ$ to both sides of the solar azimuth angle. This peak becomes more prominent with increasing solar zenith angle and snow grain size.

8. Pg. 4, line 28, change “480” to “340”. The smallest CAR wavelength is 340 nm.

Thank you for spotting this. We updated the CAR wavelength limits to 340 and 2324 nm and reference
5 the earlier Gatebe et al. (2003) publication (instead of Gatebe et al., 2005).

Changes in text:

- P4 L30: The effective BRDFs were acquired by the Cloud Absorption Radiometer (CAR, Gatebe et al. 2003) over a 30-year period between 1984 and 2014. The CAR is a scanning radiometer covering
10 14 spectral channels between 340 and 2324 nm
- Gatebe, C., King, M., Platnick, S., Arnold, G., Vermote, E., and Schmid, B.: Airborne spectral measurements of surface-atmosphere anisotropy for several surfaces and ecosystems over southern Africa, *J. Geophys. Res.*, 108, 8489, <https://doi.org/doi:10.1029/2002JD002397>, 2003.

9. Pg. 5, line 29, the BRDF symbol used here is not consistent with Schaepman-Strub et al. (2006) as stated/line 27.

Thank you for this comment. It is true that Schaepman-Strub et al. (2006) denote the BRDF with
15 $f_r(\theta_i, \varphi_i; \theta_r, \varphi_r)$, the BRF with $R(\theta_i, \varphi_i; \theta_r, \varphi_r)$, and the HDRF with $R(\theta_i, \varphi_i, 2\pi; \theta_r, \varphi_r)$. To better distinguish between the reflectance functions $f()$ (in units sr^{-1}) and the dimensionless reflectance factors R , we changed the notation for the BRF and HDRF to R_{BRF} and R_{HDRF} . This is more consistent with Schaepman-Strub et al. (2006), however we decided to keep the acronyms in the indices of the reflectance factors as the difference in notation between the BRF ($R(\theta_i, \varphi_i; \theta_r, \varphi_r)$) and HDRF ($R(\theta_i, \varphi_i, 2\pi; \theta_r, \varphi_r)$)
20 is only marginal and hard to identify, especially for someone not familiar with the details.

Changes in text:

- We changed the notation throughout the manuscript from f_{BRDF} and f_{HDRF} to R_{BRDF} and R_{HDRF} , respectively.

10. Pg. 5, line 30, the statement “the reflected radiance for all reflection angles...” is unclear. The entire definition of the spectral BRDF need to be clarified.

Thanks, we clarified the definition of the spectral BRDF by including the definition of the direction of reflection.

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Changes in text:

- P5 L33: The spectral BRDF (f_{BRDF} , unit of sr^{-1}) provides for each solar zenith (θ) and azimuth angle (φ) of incident direct irradiance $F_i(\theta_i, \varphi_i; \lambda)$ the reflected radiance $I_r(\theta_i, \varphi_i; \theta_r, \varphi_r; \lambda)$ for all directions of reflection (defined by the reflection zenith and azimuth angles θ_r and φ_r) by [...]

11. Pg. 13, line 14, the location of each pixel is given by x and y , but “ x ” and “ y ” are not defined elsewhere in the manuscript.

10 x and y define the location of each pixel in terms of rows and columns on the sensor. We added this detail when first defining x and y in Sect. 3.3.2 right before Equation (16).

Changes in text:

- P13 L6: The calibration factor k_c is defined at each pixel location on the sensor (row x , column y) as

15

[...]

12. Pg. 6, the derivation of Equation (3) is not clear.

Thanks for pointing this out. However, the derivation of Equation (3) is lengthy and beyond the scope of this manuscript. We refer to Schaepman-Strub et al. (2006) for details.

Changes in text:

- P6 L13: The second step in Eq. 3 assumes that the incident diffuse radiation is isotropic (for details, see Schaepman-Strub et al., 2006).

13. Pg. 13, line 19, the author need to show how “...the viewing angles are corrected” using the “the roll, pitch and yaw angles of the aircraft at the time of measurement.”

So far, the correction of the viewing angles was explained earlier in the text. However, to avoid this potential confusion, we decided to provide all information about the aircraft attitude correction in one
5 paragraph.

Changes in text:

- P13 L25: The camera viewing angles are calculated from the geometric calibration for each pixel (x , y). As the camera is fixed to the aircraft frame, a correction for the aircraft attitude was implemented
10 to obtain the reflection angles θ_r and φ_r . Thus, beside the geometric calibration, the observed reflection angles are determined by the attitude angles of the aircraft. Utilizing the data from the internal navigation system (INS) and the global positioning system (GPS) on Polar 6, the viewing angles are corrected depending on the roll, pitch and yaw angles of the aircraft at the time of measurement. In this regard, Euler rotations are applied as described in Ehrlich et al. (2012).

14. Pg. 13, Equation 18, the viewing and illumination geometry variables need to be defined taking into account the aircraft orientation. The downward flux F should also be described here immediately after the equation is listed, unless it's previously defined.

15 Thank you for this comment. The reflection geometry is obtained from the viewing geometry by means of the aircraft attitude correction. We believe in moving the aircraft attitude correction paragraph to directly above Eq. 18 (see comment #13 above), the illumination and reflection geometry are now sufficiently defined. We added the illumination angles θ_0 and φ_0 to Eq. 18. The downward flux F^\downarrow is already defined in Sect. 3.2, but we added a description for clarity.

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Changes in text:

– P13 L30: Finally, the observed HDRF is calculated by:

$$R_{\text{HDRF}}(\theta_0, \varphi_0; \theta_r, \varphi_r) = \frac{\pi_{\text{SF}} \cdot I(\theta_r, \varphi_r)}{F^\downarrow(\theta_0, \varphi_0)}.$$

$F^\downarrow(\theta_0, \varphi_0)$ denotes the downward solar irradiance at the time of measurement (with solar zenith and azimuth angles θ_0 and φ_0).

Reply to short comment by Aleksey Malinka

Dear Authors, I have a little note. Kokhanovsky and Zege (2004) do not represent the snow grains as fractal particles. They consider any particles with no regard to their shape. Truly yours, Aleksey Malinka.

Dear Aleksey Malinka,

Thank you for your interest in our manuscript and your efforts to improve it. We checked Kokhanovsky and Zege (2004) again following your comment: they indeed checked different particle models for the phase function of ice crystals, namely the spherical model, fractal model, and random particle model (which we believe you are referring to in your comment). However, in the last paragraph on page 5, they state: "We choose the fractal model here because it has no free parameters for nonabsorbing ice particles and corresponds closely to the random-particle model at extreme values of its randomness parameters." Kokhanovsky et al. (2005) compare the developed approximate asymptotic theory with in situ measurements. On page 2 they state: "Kokhanovsky and Zege (2004) derived the following relationship [...] in the framework of the approximation of fractal snow grains." Therefore, we are confident that the Kokhanovsky and Zege (2004) work is referenced correctly in our Introduction. But thanks again for looking at this and encouraging us to double-check.

Kind regards from the authors.

Changes in text:

- Page 3 Line 21: Their approach represents the snow grains for example as fractal particles and, thus, accounts for their non-sphericity.

Parameterizing anisotropic reflectance of snow surfaces from airborne digital camera observations in Antarctica

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Abstract. The angular surface reflection of solar radiation ~~by snow surfaces~~ comprises an important boundary condition for solar radiative transfer simulations. In polar regions above snow surfaces, the surface reflection is particularly anisotropic due to low sun elevations and the highly anisotropic scattering phase function of ~~ice crystals. This anisotropy needs to be considered in the angular modeling of the surface reflection, which the snow crystals.~~ The characterization of this surface reflection anisotropy is essential for satellite remote sensing ~~techniques over both the Arctic and Antarctica.~~ To quantify the angular snow reflection properties, the hemispherical-directional reflectance ~~function factor~~ (HDRF) of snow surfaces was derived from airborne measurements ~~using in Antarctica during austral summer in 2013/14. For this purpose,~~ a digital 180° fish-eye camera (green channel, 490-585 nm wavelength band) ~~in Antarctica during austral summer in 2013/14. This function was was used.~~ The HDRF was measured for different surface ~~roughnesses~~ roughness conditions, optical-equivalent snow grain sizes, and solar zenith angles. The airborne observations covered an area of around 1000×1000 km² in the vicinity of Kohonen station (75°0′ S, 0°4′ E) at the outer part of the East Antarctic Plateau. The observations ~~over Dronning Maud Land~~ include regions with higher (coastal areas) and lower (inner Antarctica) precipitation amounts and frequencies. The digital camera ~~was calibrated in terms of spectral radiances and installed in a downward-looking configuration on a research aircraft. It provides~~ provided upward, angular-dependent radiance measurements from the lower hemisphere. The comparison of ~~HDRF data the measured~~ HDRF derived for smooth and rough snow surfaces (sastrugi) showed significant differences, which are superimposed on the diurnal cycle. By inverting a semi-empirical, kernel-driven bidirectional reflectance distribution function (BRDF) model, the measured HDRF of snow surfaces was parameterized ~~with respect to as a function of~~ solar zenith angle, surface roughness, and optical-equivalent snow grain size. This allows a direct comparison of the HDRF measurements with the BRDF derived from the Moderate Resolution Imaging Spectroradiometer (MODIS) satellite product MCD43. For the analyzed cases, MODIS observations (~~545-565 nm wavelength band~~) generally underestimated the anisotropy of the surface reflection. ~~Some of these deviations between airborne and MODIS satellite retrievals. The largest deviations were found for the volumetric model weight f_{vol} (average underestimation by factor of 10). These deviations~~ are likely linked to short-term changes in snow properties.

1 Introduction

Surface reflection in the polar regions plays an important role in the Earth's climate system. The snow surface albedo is highly variable both on a temporal and spatial scale. This causes uncertainties in determining the surface solar radiative energy budget in these areas (e.g., Kuipers Munneke et al., 2008). As such it is essential to monitor the reflective properties of snow surfaces for accurately predicting future climate change in the polar regions. However, their remote and harsh environment makes them difficult to access and requires remote sensing techniques to observe snow surface reflection and its influencing parameters.

The high spatial and temporal variability of the snow surface albedo necessitates continuous observations with global coverage, which are provided by satellite instruments. ~~Satellites~~ Polar-orbiting satellites monitor the reflectance (i.e., reflected radiance in units of $\text{W m}^{-2} \text{sr}^{-1}$) at the top of the atmosphere (TOA). However, they are restricted in terms of the number of available observation angles and spectral bands as well as their temporal resolution. The processing of ~~satellite measurements~~ measurements from polar-orbiting satellites for monitoring the broadband surface albedo typically requires three steps (e.g., Qu et al., 2015): Atmospheric correction, modeling of the angular reflectance, and narrow-to-broadband conversion. During the first step, the TOA reflectance is converted into a surface reflectance by ~~means of an atmospheric radiative transfer parameterization~~ correcting for gaseous and aerosol scattering and absorption applying radiative transfer modeling (e.g., Vermote and Kotchenova, 2008). For the calculation of the narrowband surface albedo, an accurate knowledge of the bidirectional reflectance distribution function (BRDF) of the surface is required due to the limited number of available observation angles. The BRDF describes the scattering of electromagnetic (EM) radiation from one incident direction into another direction of the hemisphere. By using a linear combination of the discrete spectral band measurements with specific weighting, the broadband surface albedo is calculated (e.g., Brest and Goward, 1987; Klein and Stroeve, 2002; Liang et al., 2003; Pohl et al., 2020). The largest uncertainty in this three-step process is introduced by the angular dependence of the surface BRDF, especially when the reflection of the underlying surface is highly anisotropic. This is the case for the reflection of solar radiation at snow surfaces in polar regions due to low sun elevations and the anisotropic scattering phase function of snow crystals. As the snow surface of the Antarctic ice sheet is also used for the validation and cross-calibration of ~~polar-orbiting~~ polar-orbiting satellites (Jaross and Warner, 2008), a thorough understanding of the anisotropic reflection is needed. However, as an infinitesimal quantity, the BRDF of a surface cannot directly be measured under natural illumination conditions. From a rigorous physical point of view, most satellite, airborne, and ground-based instruments measure the hemispherical-directional reflectance factor (HDRF, Schaepman-Strub et al., 2006) if the reflectance is constant over the full cone angle of the instrument's field of view (FOV). In contrast to the BRDF, the HDRF includes incident irradiance from the entire hemisphere. An effective BRDF can only be derived from HDRF observations if the atmospheric influence is considered and the FOV is sufficiently small (e.g., Gatebe et al., 2003; Lyapustin et al., 2010).

For example, the retrieval of the aerosol optical depth (AOD) over highly reflecting surfaces is complicated, as it is hard to separate the surface and aerosol contributions to the measured reflectance at the TOA (Tomasi et al., 2015), especially for the comparably low values of AOD in polar regions (Mei et al., 2013). Due to its high surface albedo, most of the solar radiation incident on a snow surface is reflected, which depends strongly on the BRDF of the snow surface and, thus, on the illumination

and viewing geometry. In general, most of the photons are scattered by the snow grains into the forward scattering direction. [Yang et al. \(2014\)](#) [Yang et al. \(2014\)](#) compared Moderate Resolution Imaging Spectroradiometer (MODIS) measurements of AOD at 550 nm over Eastern China with ground-based data from the Aerosol Robotic Network (AERONET). Using a Lambertian model, 54 % of the regression points fell within the expected error envelope around the 1:1 line. This result is improved to 69 % if a more sophisticated, non-Lambertian model accounting for anisotropic reflection at the land surface is applied.

Yet, the BRDF varies with snow grain morphology, the solar zenith and azimuth angles, the liquid water content of the snowpack, the snow impurity type and concentration, the dimension and orientation of surface roughness structures, and the wavelength. To account for these effects, an accurate model of the snow BRDF is needed. A BRDF model can be either physical, empirical, or semi-empirical. Physical BRDF models (e.g., Cook and Torrance, 1982) accurately simulate the scattering of an EM wave at a surface by applying physical laws. The high accuracy comes at the cost of very complex computations. Empirical BRDF models (e.g., Phong, 1975; Walthall et al., 1985) mimic the surface reflection by means of a simple, non-physical formulation. However, the drawback of the rather simple computations is their restricted accuracy. Semi-empirical models (e.g., Martonchik et al., 1998; Lucht et al., 2000) use simple, direct parameterizations of a more complex physical BRDF with a limited number of independent parameters.

The anisotropic reflection by snow surfaces was investigated by Li (1982) with simulations of the snow BRDF using Mie theory and the doubling and adding method. Applying Mie theory and the discrete ordinate method, Han (1996) retrieved the surface albedo from satellite measurements in the Arctic. Leroux and Fily (1998) developed a snow BRDF model including the effect of sastrugi by means of regularly spaced, identical, and rectangular protrusions. Leroux et al. (1998) and Leroux et al. (1999) employed the doubling and adding method, Mie theory, and ray tracing to develop a [polarized BRDF model snow BRDF model including polarization](#). Comparing the simulated values with observations in the principal plane, they found that the BRDF in the near-infrared wavelength range is strongly affected by the snow grain shape, whereby simulations assuming hexagonal particle shapes yield an improved agreement with the observations compared to assuming spheres. Accordingly, Aoki et al. (2000) stressed the importance of the particle shape assumption for the scattering phase function used in their BRDF model (applying both Mie theory and the Henyey-Greenstein phase function) when comparing to observations between 0.52 μm and 2.21 μm wavelength. Kokhanovsky and Zege (2004) demonstrated the use of an asymptotic analytical equation to model the [reflectance BRDF](#) of snow. Their approach represents the snow grains [for example](#) as fractal particles and, thus, accounts for their non-sphericity. [Comparing-Evaluating this asymptotic model to in situ observations of the BRDF with in situ measurements of the snow reflectance](#), Kokhanovsky et al. (2005) found a generally good agreement but a reduced model accuracy in the solar principal plane at large observation angles. In contrast to considering snow grains as independent scatterers of fractal shape, Malinka (2014) provided a framework to calculate the inherent optical properties ([extinction coefficient, single scattering albedo, scattering phase function, and the polarization properties](#)) applying the concept of stereology that considers snow as a random mixture of ice and air. Malinka et al. (2016) combined the stereological approach with analytical, asymptotic formulas to calculate the bidirectional reflectance. Their model showed a high accuracy when compared to albedo observations of snow-covered sea ice.

The comparison of in situ ~~measured BRDF~~ measurements of the snow reflectance with simulations is essential in terms of model validation. Observations of the ~~BRDF or HDRF~~ HDRF (or effective BRDF in case the FOV is small and the atmospheric influence is considered) are conducted using a variety of different measurement concepts. For ground-based applications, manual or automated gonio-spectrometer systems are employed (e.g., Painter et al., 2003; Pegrum et al., 2006; Bourgeois et al., 2006a). Kuhn (1985) observed a peak in reflectance ~~in the azimuthal directions up to 60°~~ to that is contained within $\pm 60^\circ$ to both sides of the solar azimuth ~~that angle~~. This peak becomes more prominent with increasing solar zenith angle and snow grain size. Marks et al. (2015) measured the HDRF of snow surfaces between 400 and 1600 nm wavelength with the Gonio Radiometric Spectrometer System (GRASS, Pegrum et al., 2006) at Dome C, Antarctica. Their observations also showed enhanced forward scattering. In addition, they observed a larger anisotropy of the surface reflection at longer wavelengths. Employing a gonio-spectrometer, Bourgeois et al. (2006b) measured strong variations of the HDRF between 0.6 and 13 (in the wavelength range 350-1050 nm) depending on the solar zenith angle and the surface roughness at Summit, Greenland. Measurements with the Automated Spectro-Goniometer (ASG, Painter et al., 2003) showed a decrease of snow HDRF at all wavelengths between 400 to 2500 nm when the snow grain size increased from 80 to 280 μm (Painter and Dozier, 2004b). Further comparisons with the results of a forward discrete ordinates radiative transfer model (Stamnes et al., 1988) revealed larger deviations between the simulations and observations for more complex crystals. The importance of the crystal habit for the anisotropy of reflection at snow surfaces was ~~further~~ also emphasized by Dumont et al. (2010) and Stanton et al. (2016). The latter measured increasing anisotropy of the surface reflection during the growth of surface hoar in the laboratory. Several studies observed systematically less anisotropy for a typical snow ~~BRDF~~ HDRF than estimated from simulations (Warren et al., 1998; Painter and Dozier, 2004a; Hudson et al., 2006; Hudson and Warren, 2007). In the solar principal plane, the models mainly overestimate the forward scattering and underestimate the backward scattering. Implementing non-spherical grains in the BRDF models (e.g., Mishchenko et al., 1999; Xie et al., 2006; Jin et al., 2008) improves the comparison with observations. The nonspherical model of Jin et al. (2008) agrees within $\pm 10\%$ for viewing zenith angles less than 60° with observations in Antarctica performed by Hudson et al. (2006). However, the asymmetry between forward and backward scattering still remains. This highlights the need to further incorporate macroscopic effects such as the roughness of the snow surfaces into the models (Leroux and Fily, 1998; Hudson and Warren, 2007).

Ground-based instruments observe the directional reflectance of a characteristic, homogenous surface, whereas airborne and satellite observations average over a larger measurement area. Thus, the latter are more suitable for studying the influence of macroscopic surface roughness on the surface ~~BRDF~~ HDRF. Nolin et al. (2002) employed multiangular measurements with the multi-angle imaging spectroradiometer (MISR) for the characterization of surface roughness over Greenland and Antarctica. Measurements with the clouds and the Earth's radiant energy system (CERES) showed monthly regional biases between -12 and 7.5 W m^{-2} in the cloudless TOA solar irradiance. These biases were attributed to the effect of sastrugi, which introduce a significant solar azimuth dependence: Kuchiki et al. (2011) observed a diurnal cycle in MODIS reflectances over the South Pole. In general, sastrugi decrease the forward scattering by casting shadows, increase the backward scattering due to a lower effective incident angle caused by the sastrugi slope, and the snow ~~BRDF~~ HDRF loses its azimuthal symmetry (Warren et al., 1998). Zhuravleva and Kokhanovsky (2011) ~~observed~~ simulated a larger effect for a higher density of the sastrugi field.

Airborne measurements can be employed for the validation of BRDF models and the comparison with the large pixel size of satellite observations. Gatebe and King (2016) provided an extensive database of airborne spectral [\(effective\) BRDFs](#) for various surface types, e.g., ocean, vegetation, snow, desert, and clouds. The [effective BRDFs](#) were acquired by the Cloud Absorption Radiometer ~~(CAR, Gatebe et al., 2005)~~ [\(CAR, Gatebe et al., 2003\)](#) over a 30-year period between 1984 and 2014. The CAR is a scanning radiometer covering 14 spectral channels between [480-340](#) and 2324 nm. The effect of surface roughness on the [effective BRDF](#) was studied by Lyapustin et al. (2010) with CAR measurements during the Arctic Research of the Composition of the Troposphere from Aircraft and Satellites (ARCTAS) spring campaign in April 2008 (Jacob et al., 2010; Matsui et al., 2011). Their results showed an agreement between the kernel-based Ross-Thick Li-Sparse-Reciprocal (RTLSR) BRDF model (Lucht et al., 2000) used in the operational MODIS BRDF/albedo product (Schaaf et al., 2002) and the CAR measurements to within ± 0.05 . The RTLSR model performed better in the forward scattering direction, whereas the kernel-based Modified Rahman-Pinty-Verstraete (MRPV) BRDF model (Martonchik et al., 1998) used for the processing of data from MISR performed better in the backscattering direction. Jiao et al. (2019) proposed an additional snow kernel within the kernel-driven BRDF model framework to better account for the anisotropic reflection of pure snow surfaces. Incorporating the additional snow kernel yielded a correlation coefficient of 0.9 (compared to 0.65 for the original 3-kernel model) and only small biases between the model and different BRDF validation data from ground-based and satellite observations.

Cox and Munk (1954) analyzed radiance calibrated analog photographs for the parameterization of the ocean [BRDF HDRF](#). Nowadays, digital cameras are increasingly applied in vegetation and soil monitoring ~~(Lebourgeois et al., 2008)~~ [\(e.g., Lebourgeois et al., 2008\)](#). The instantaneous measurement of multiple viewing angles facilitates aerial [BRDF HDRF](#) measurements with digital cameras. The high angular resolution allows the detection of features in the anisotropy of the reflection, which could be missed with HDRF measurements allowing only a limited number of observation angles. This was demonstrated by Goyens et al. (2018) for ground-based measurements of the snow HDRF comparing radiance measurements with a spectroradiometer at discrete viewing angles with hyperangular observations using a fish-eye radiance camera. Ehrlich et al. (2012) used a commercial digital camera equipped with a wide-angle lens with a FOV of 100° for measurements of the HDRF of snow-covered sea ice, ocean, and clouds (viewing zenith angle up to around 60°).

This study presents a method to derive the snow HDRF at visible wavelength (spectral band 490-585 nm) from airborne digital camera measurements in Antarctica during austral summer in 2013/14. Concurrent measurements of the optical-equivalent snow grain size retrieved from spectral surface albedo measurements, and surface roughness determined by means of a laser scanner, allow for investigating their effect on the snow HDRF in separate case studies. Subsequently, the measurements are used to parameterize the snow HDRF applying the RTLSR model (Lucht et al., 2000). The presented methodological approach allows for a direct comparison of the airborne HDRF measurements with satellite observations from MODIS.

The definition of reflectance quantities used within this work, the modeling of the BRDF, and the inversion of a semi-empirical, kernel-driven BRDF model are presented in Sect. 2. The field work and instrumentation are presented in Sect. 3 together with the detailed calibration involved in the measurement of the snow HDRF. The dependence of the snow HDRF with respect to the solar zenith angle, the surface roughness, and the optical-equivalent snow grain size is studied in Sect. 4

based on three cases. Subsequently, the airborne measurements are compared with satellite-derived BRDF from MODIS. In the concluding Sect. 5, the findings of this work are summarized and perspectives for future studies are given.

2 Methodology

2.1 Definition of reflectance quantities

- 5 The definition of the reflectance quantities applied within this work follows Schaepman-Strub et al. (2006). The intrinsic reflectance properties of a surface are given by its BRDF. It quantifies the reflection of the incident radiation at the surface and its scattering from one into another direction of the hemisphere. The spectral BRDF (f_{BRDF} , unit of sr^{-1}) provides for each solar zenith ($\theta_{\text{zenith}}(\theta_i)$) and azimuth angle (φ_i) of incident direct irradiance $F_i(\theta_i, \varphi_i; \lambda)$ the reflected radiance for all reflection angles $I_r(\theta_r, \varphi_r; \theta_i, \varphi_i; \lambda)$ by $I_r(\theta_i, \varphi_i; \theta_r, \varphi_r; \lambda)$ for all directions of reflection (defined by the reflection zenith and azimuth angles θ_r and φ_r) by:

$$f_{\text{BRDF}} = \frac{dI_r(\theta_i, \varphi_i; \theta_r, \varphi_r; \lambda)}{dF_i(\theta_i, \varphi_i; \lambda)}. \quad (1)$$

The bidirectional reflectance factor (BRF, dimensionless) is obtained when the BRDF of a sample surface is divided by the BRDF of a diffuse (Lambertian) standard surface illuminated under the same conditions (identical beam geometry). The BRDF of an ideal Lambertian surface is $(\pi \text{sr})^{-1}$, the BRF is given by

$$f_{\text{BRF}} = \frac{dI_r(\theta_i, \varphi_i; \theta_r, \varphi_r)}{dF_i(\theta_i, \varphi_i)} \cdot \frac{dF_i(\theta_i, \varphi_i)}{dI_r^{\text{ideal}}(\theta_i, \varphi_i)} = \pi \text{sr} \cdot f_{\text{BRDF}}. \quad (2)$$

In Eq. 2 and in the remainder of this section, the spectral dependence is omitted for reasons of simplicity. In atmospheric conditions, both the BRDF and BRF cannot be measured directly as the global irradiance F_0 reaching the surface is composed of a direct (F_i) and diffuse (F_{diff}) component. In this case, the measurable quantity is the HDRF. The definition of the HDRF is analogous to the BRF, but includes irradiance from the entire hemisphere (denoted with 2π):

$$R_{\text{HDRF}} = \pi \text{sr} \cdot \frac{dI_r(\theta_i, \varphi_i, 2\pi; \theta_r, \varphi_r)}{dF_0(\theta_i, \varphi_i, 2\pi)} = R_{\text{BRF}}(\theta_i, \varphi_i; \theta_r, \varphi_r) \cdot f_{\text{dir}} + R(2\pi; \theta_r, \varphi_r) \cdot (1 - f_{\text{dir}}). \quad (3)$$

f_{dir} denotes the fraction of direct incident radiation (i.e. $f_{\text{dir}} \in [0, 1]$). The second step in Eq. 3 assumes that the incident diffuse radiation is isotropic (for details, see Schaepman-Strub et al., 2006). The additional integration over all reflection angles leads to the bihemispherical reflectance, generally called surface albedo α .

2.2 Modeling of the bidirectional reflectance

- 25 The shape of the BRDF is described by a weighted sum of trigonometric functions, generally referred to as kernels for volumetric scattering (K_{vol}), geometric scattering (K_{geo}), and isotropic scattering (K_{iso}).

The kernel-driven semi-empirical Ross-Li model (Lucht et al., 2000), which forms the basis of the MODIS 16-day BRDF/albedo product (Schaaf et al., 2002), is applied within this study. The BRDF is given as a linear combination of the kernels with corresponding non-negative weighting functions f_{iso} , f_{vol} , and f_{geo} :

$$f_{\text{BRDF}}\left(\theta_{\underline{r}_i}, \theta_{\underline{0}_r}, \Delta\varphi, \lambda\right) = f_{\text{iso}}(\lambda) \cdot K_{\text{iso}} + f_{\text{vol}}(\lambda) \cdot K_{\text{vol}}\left(\theta_{\underline{r}_i}, \theta_{\underline{0}_r}, \Delta\varphi\right) + f_{\text{geo}}(\lambda) \cdot K_{\text{geo}}\left(\theta_{\underline{r}_i}, \theta_{\underline{0}_r}, \Delta\varphi\right), \quad (4)$$

5 with the viewing zenith angle θ_r , solar zenith angle θ_0 , relative viewing azimuth angle $\Delta\varphi$, and wavelength λ . The kernels,

$$K_{\text{iso}} = 1, \quad (5)$$

$$K_{\text{vol}} = \frac{\left(\frac{\pi}{2} - \theta\right) \cos \theta + \sin \theta}{\cos \theta_0 + \cos \theta_r} - \frac{\pi}{4}, \quad \text{and} \quad (6)$$

$$K_{\text{geo}} = \mathcal{O}\left(\theta_{\underline{r}_i}, \theta_{\underline{0}_r}, \Delta\Phi\varphi\right) - \sec \theta_0 - \sec \theta_r + \frac{1}{2}(1 + \cos \theta) \cdot \sec \theta_0 \cdot \sec \theta_r, \quad (7)$$

depend on the scattering angle θ and the function \mathcal{O} :

$$10 \quad \cos \theta = \cos \theta_0 \cdot \cos \theta_r + \sin \theta_0 \cdot \sin \theta_r \cdot \cos \Delta\Phi \cos \Delta\varphi, \quad (8)$$

$$\mathcal{O}\left(\theta_{\underline{r}_i}, \theta_{\underline{0}_r}, \Delta\Phi\varphi\right) = \frac{1}{\pi} (\mathcal{C} - \sin \mathcal{C} \cdot \cos \mathcal{C}) \cdot (\sec \theta_0 + \sec \theta_r). \quad (9)$$

The functions \mathcal{C} and \mathcal{D} depend solely on the viewing and illumination geometry:

$$\cos \mathcal{C} = \frac{2 \cdot \sqrt{\mathcal{D}^2 + (\tan \theta_r \cdot \tan \theta_0 \cdot \sin \Delta\Phi)^2}}{\sec \theta_0 + \sec \theta_r} \frac{2 \cdot \sqrt{\mathcal{D}^2 + (\tan \theta_r \cdot \tan \theta_0 \cdot \sin \Delta\varphi)^2}}{\sec \theta_0 + \sec \theta_r}, \quad (10)$$

$$\mathcal{D} = \frac{\sqrt{\tan^2 \theta_r + \tan^2 \theta_0 - 2 \tan \theta_r \cdot \tan \theta_0 \cdot \cos \Delta\Phi}}{\sqrt{\tan^2 \theta_r + \tan^2 \theta_0 - 2 \tan \theta_r \cdot \tan \theta_0 \cdot \cos \Delta\varphi}}. \quad (11)$$

15 Originally, the concept of the model was developed studying typical features in observations of the bidirectional reflectance of various surface types (Roujean et al., 1992). First, bare soil surfaces exhibit strong backscattering characteristics and show the effect of geometrical structures on the surface. Secondly, dense leaf canopies typically feature a minimum reflectance close to the nadir-direction that increases off-nadir for all azimuthal directions. In their approach, Roujean et al. (1992) represented the BRDF model as a linear combination of these two observational characteristics; the former being accounted for by K_{geo} ,
 20 the latter by K_{vol} . The volumetric kernel K_{vol} stems from volume scattering radiative transfer models (Ross, 1981). Thereby, randomly located facets that absorb and scatter incident radiation are modeled under the single-scattering approximation. The formula given in Eq. 6 corresponds to the Ross-Thick kernel for a dense leaf canopy. The geometric kernel K_{geo} is derived from surface scattering and geometric shadow casting theory (Li and Strahler, 1992) and expresses effects caused by inter-crown gaps within vegetation. It represents randomly oriented vertical protrusions on a flat and horizontal surface that isotropically
 25 reflect radiation. Eq. (7) gives the Li-Sparse geometric kernel in reciprocal form describing the casting of shadows by a sparse ensemble of surface objects.

Hu et al. (1997) validated kernel-based BRDF models with 27 multi-angular data sets from various land cover types. The accuracy of the model was high. The correlation coefficient between model and observations was above 0.7 for all and above 0.9 for more than half of the data sets. Although originally calculated for surfaces covered with vegetation, the Ross-Thick and Li-Sparse-Reciprocal kernels are also applied for snow surfaces for the MODIS BRDF/albedo product. Stroeve et al. (2005) and Stroeve et al. (2013) assessed the accuracy of the 16-day albedo product of eleven years of measurements at 17 automatic weather stations on the Greenland ice sheet. They retrieved physically realistic ice sheet albedo values with an overall mean bias between MODIS and the in situ measurements of 0.022.

2.3 Inversion of semi-empirical, kernel-driven model

The main benefit of retrieving the weighting functions f_{iso} , f_{vol} , and f_{geo} from the HDRF measurements is that they encompass all of the angular radiance information. The retrieval is done by inverting the RTLSR BRDF model. Thus, information about the complete two-dimensional shape of the HDRF is used.

The modeled $f_{\text{BRDF}}(\theta_r, \theta_0, \Delta\Phi, \lambda)$ from Eq. 4 can be written in the form of a sum as:

$$f_{\text{BRDF},l} = \sum_{k=1}^3 f_k \cdot K_{kl}. \quad (12)$$

The spectral dependence is omitted here and the index l denotes a particular viewing and illumination geometry $(\theta_r, \theta_0, \Delta\Phi)_l$. Considering an observation with N directional measurements ρ_l ($l = 1, \dots, N$), the error function \mathcal{E}^2 is defined as the difference between the observed and the modeled reflectances such that,

$$\mathcal{E}^2 = \frac{1}{d} \sum_{l=1}^N \frac{(\rho_l - f_{\text{BRDF},l})^2}{w_l}. \quad (13)$$

The degree of freedom d is $(N - 3)$ and w_l denotes weights which are assigned to the respective observations. In general, w_l could take the values 1, ρ_l , or ρ_l^2 . The goal of the inversion is the determination of the model weighting functions f_k such that \mathcal{E}^2 is minimized. Strahler et al. (1996), amongst others, presented the analytical solutions for the three f_k following this minimization and Lewis (1995) showed that \mathcal{E}^2 has a global minimum in f_k .

The inversion depends on the choice of the error function \mathcal{E}^2 and, thus, on the choice of the weights w_l . Minimizing the absolute error (using weights equal to unity) leads to smaller errors in angular domains with a large reflectance. In contrast, minimizing the relative error (e.g., $w_l = \rho_l, \rho_l^2$) performs better in angular regions with lower reflectance. Subsequently, different choices for w_l were tested for the retrievals performed in this work and $w_l = \rho_l^2$ was chosen as it produced the lowest retrieval errors across the entire angular domain.

3 Measurements and instrumentation

The airborne measurements were performed with the Polar 6 research aircraft (Wesche et al., 2016) operated by the Alfred Wegener Institute, Helmholtz Centre for Polar and Marine Research (AWI), Bremerhaven, Germany. Polar 6 was based at

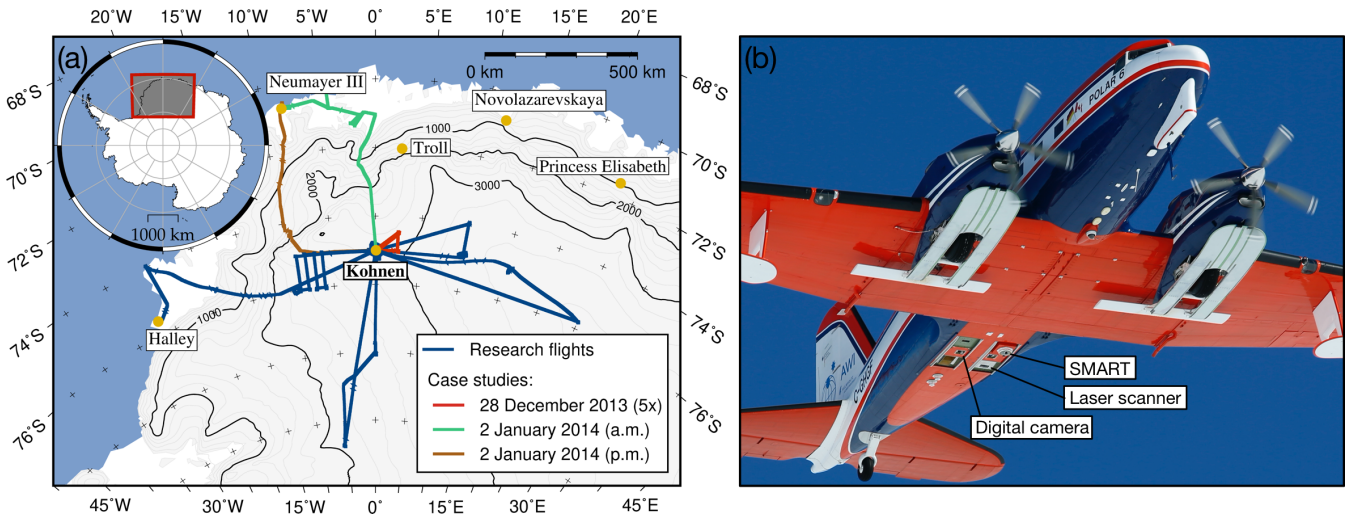


Figure 1. (a) Illustration of research flights with Polar 6 in Dronning Maud Land between 24 December 2013 and 5 January 2014. The five flights on 28 December 2013 (marked in red) and the two flights on 2 January 2014 (green and brown) are used for the case studies. The parameter ranges covered by the case studies are marked with the same colors within the histograms in Fig. 2. (b) Photograph of Polar 6 illustrating the mounting of the digital camera, SMART, and the laser scanner.

the Kohnen research station ($75^{\circ}0' S$, $0^{\circ}4' E$). Figure 1a illustrates the tracks of the research flights with Polar 6 (60 flight hours) covering an area of $1000 \times 1000 \text{ km}^2$ around Kohnen station in Dronning Maud Land. The flights were performed between 24 December 2013 and 5 January 2014. The observations comprised regions with higher (coastal areas) and lower precipitation amounts and frequencies (inner Antarctica), as well as a variety of surface roughness structures. An overview of the airborne instrumentation is given in Carlsen et al. (2017); the mounting of the digital camera, laser scanner, and the Spectral Modular Airborne Radiation measurement sysTEM (SMART) for surface albedo measurements on the underside of the aircraft is illustrated in Fig. 1b. The variability of the solar zenith angle is illustrated in Fig. 2a in the form of a histogram. During the flights, the solar zenith angle varied between 49° and 73° .

3.1 Surface roughness measurements using a laser scanner

10 The snow surface topography was measured using the airborne laser scanner RIEGL VQ580. In time-of-flight laser ranging, a near-infrared laser beam (1064 nm wavelength) is emitted downward and subsequently reflected upward by the snow surface before the echo is acquired by the sensor. From the time lag between emission and detection, the distance to the ground is calculated with a precision of about 25 mm depending on flight altitude. A fast-rotating polygonal mirror with a FOV of 60° and 10 to 150 scans per second allows for fully linear, unidirectional, and parallel scan lines.

15 After a correction for the aircraft attitude, a $1 \times 1 \text{ km}^2$ digital elevation model (DEM) is generated from the geotagged laser point cloud with a resolution of 1 m. Subtracting the large scale topography (smoothed DEM), the residual field contains the roughness information. The standard deviation of the residual field is interpreted as the surface roughness at the central

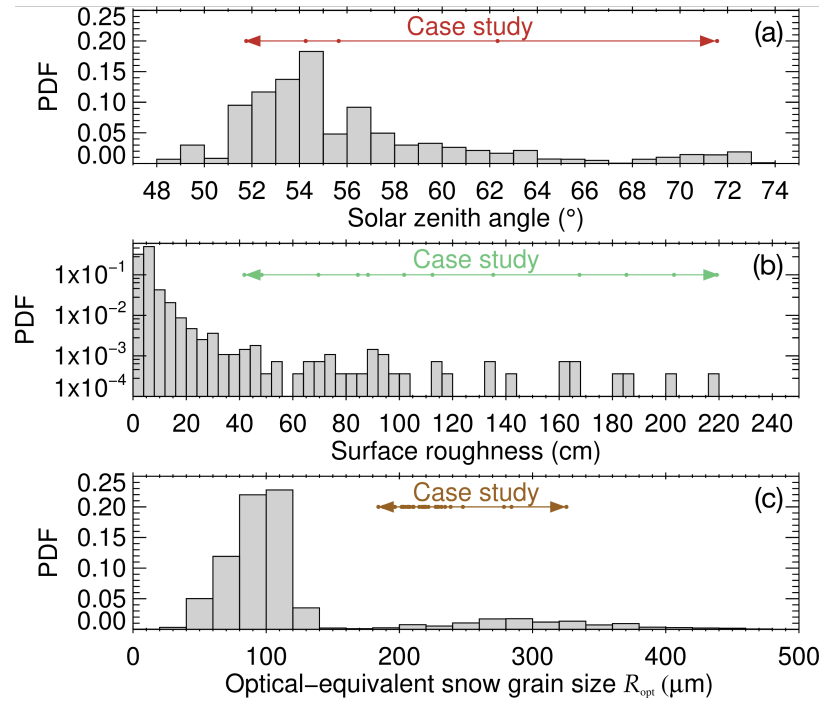


Figure 2. Variability of different parameters during research flights shown as histograms of the probability density function (PDF). The parameter ranges covered by the case studies are marked with the same colors as the corresponding flight tracks in Fig. 1. **(a)** Variability of solar zenith angle (binsize: 1°). **(b)** Variability of surface roughness (binsize: 4 cm). Note the logarithmic scale. **(c)** Variability of optical-equivalent snow grain size (binsize: 20 μm).

coordinate of the DEM. Thus, roughness data is given for one data point per 1 km along the flight track of the aircraft. The uncertainty of the absolute height measurements (used for DEM generation) is less than 0.1 m. The relative analysis applied to the measurements yields an even higher accuracy.

The range in surface roughness is illustrated in Fig. 2b in the form of a histogram. The snow surface was generally smooth with roughness structures mostly below 5-10 cm (note the logarithmic scale, maximum: 2.2 m).

3.2 Optical-equivalent snow grain size as retrieved from spectral surface albedo measurements

Solar spectral radiation measurements were conducted using SMART (Wendisch et al., 2001; Ehrlich et al., 2008b) (Wendisch et al., 2001; . The upward and downward spectral irradiance [$F^\uparrow(\lambda)$, $F^\downarrow(\lambda)$] is measured within 0.3 to 2.2 μm wavelength applying an active horizontal adjustment system of the optical inlets to compensate for aircraft movement. From the irradiance measurements, the spectral snow surface albedo was obtained (Wendisch et al., 2004). The spectral resolution is 2 to 3 nm between 0.3 and 1.0 μm and 15 nm up to 2.2 μm wavelength; the temporal resolution is in the order of 1 s. Combining the errors associated with the signal-to-noise ratio (1.3-3.0 %), the accuracy of the dark correction (0.1 %), the wavelength calibration (1.0 %), the accuracy of the cross-calibration (1.0-4.5 %), the non-ideal cosine response of the optical inlets (3.5 %), and the horizontal

stabilization (1.0 %), the surface albedo measurements with SMART have an estimated uncertainty between 4.1 % and 8.1 % depending on wavelength (Carlsen et al., 2017).

The optical-equivalent snow grain size is defined as the radius R_{opt} of a collection of spheres with the same volume-to-surface ratio compared to the actual non-spherical snow grains. The retrieval of R_{opt} from spectral surface albedo measurements is described in Carlsen et al. (2017). In principle, the snow grain size and pollution amount (SGSP) algorithm (Zege et al., 2011) was extended to spectral ratios of surface albedo at 1280 and 1100 nm wavelength. Being independent of systematic measurement uncertainties (e.g., cross-calibration of the optical inlets), this approach decreases the uncertainty of the retrieved R_{opt} compared to the single-wavelength approach. The retrieved R_{opt} from SMART and analogous ground-based measurements were ~~validated against~~ compared to grain size observations utilizing reflectance measurements with MODIS (Carlsen et al., 2017). The grain size as retrieved from SMART measurements (mean value: 105 μm) is slightly higher than the MODIS retrievals (89 μm). The larger dataset of ground-based observations (72 μm) agrees well with the MODIS retrievals within the ranges given by the measurement uncertainties (linear correlation coefficient: 0.78, root-mean-square error: 24 μm).

The variability of R_{opt} during the research flights is illustrated in Fig. 2c in the form of a histogram. The optical-equivalent snow grain size varied between 16 and 480 μm , but was mostly below 120 μm depending on precipitation and snow metamorphism processes.

3.3 Directional radiance measurements using a digital camera

3.3.1 Camera specifications

The digital camera Canon EOS-1D Mark III was used for the airborne HDRF measurements of the snow surface. It is a digital single-lens reflex camera with a complementary metal oxide semiconductor (CMOS) image sensor. The CMOS sensor covers 3908 \times 2600 picture elements (pixels) on a sensor area of 28.1 \times 18.7 mm (Advanced Photo System APS-H format). During the observations, the camera was configured with the 8 mm F3.5 EX DG Circular Fish-eye lens by Sigma. Color information is obtained by means of a color filter array (CFA) in front of the CMOS sensor. The CFA consists of red (R), green (G), and blue (B) color filters, which determine the spectral response of the underlying pixel.

In principle, the measured signal s (in digital numbers, DN) should be saved essentially verbatim in the proprietary Raw format CR2 (Canon Raw version 2). This would ensure highest flexibility in the data analysis as no interpolation or white balance color correction is applied to the raw data. Full control over the post-processing steps is a prerequisite for radiometric measurements. During the airborne measurements, the camera was set to produce sRaw output format (small raw) with a resolution of 1944 \times 1296 pixels. At some point, internal chrominance subsampling is applied by the camera (Kerr, 2015).

The sRaw photos are decoded utilizing the open source tool *dcraw* (Coffin, 2017). The darkness level was set to 0 DN (see below) and the saturation level to 16383 DN as the camera captures images with color depth of 14 bit. In between, the raw data is linearly interpolated. The multipliers for all channels are set to one meaning no white balance color correction is applied. The dynamic range of the output file is 16 bit (saturation at 65536 DN).

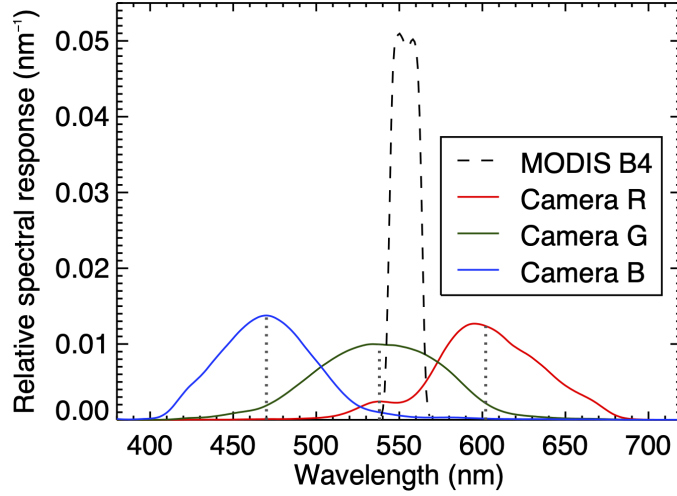


Figure 3. Relative spectral response functions for the three camera channels (colored lines: R, G, and B). The vertical dotted lines denote the respective central wavelengths. Black dashed line: RSR function for MODIS spectral band 4.

To characterize the linearity of the sensor, photos from the radiation emitted by an integrating sphere with varying output intensities were taken with the camera in the laboratory (Carlsen, 2018). The camera sensor showed a linear sensitivity over a large part of the dynamic range up to 53000 DN, the coefficient of correlation was exceptionally high with 0.99998. Photographs that measured higher signals (0.4 % of total number) were excluded from the analysis.

- 5 Compared to the dynamic range, the dark current (5 DN) and the read-out noise (7 DN) can be neglected.

Information on the spectral response of the camera is needed for the radiometric calibration. It determines the extent to which radiation of a certain wavelength passes the fish-eye lens as well as the CFA and is registered by the photodiodes. In this regard, the relative spectral response (RSR) function in units nm^{-1} of the three camera channels is defined as

$$\text{RSR}(\lambda) = \frac{T_c(\lambda)}{\int_0^\infty T_c(\lambda) d\lambda}. \quad (14)$$

- 10 $T_c(\lambda)$ with $c = R, G, B$ denotes the dimensionless spectral transmission coefficients. The RSR function is normalized such that

$$\int_0^\infty \text{RSR}(\lambda) d\lambda = 1. \quad (15)$$

The RSR function of the camera was measured in the laboratory using an integrating sphere as a radiation source and varying the outgoing radiation between 300 and 750 nm in increments of 5 nm wavelength by means of a grating monochromator.

- 15 Figure 3 shows the measured RSR functions for the three camera channels. The central wavelengths of the non-Gaussian RSR functions are 602 nm (R), 538 nm (G), and 470 nm (B). Their full width at half maximum (FWHM) varies between 68 nm for the blue channel and 95 nm for the green channel.

The FOV of the camera is 180° due to the optics of the fish-eye lens. However, the camera is installed slightly above the lower aircraft body so that the camera is protected especially during take-off and landing. Therefore, parts of the aircraft frame are constantly in the FOV of the camera, which is why the effective FOV is reduced to approximately 160°.

3.3.2 Radiance calibration and image post-processing

5 The post-processing of each raw image involves (a) the radiometric calibration, (b) the geometric calibration, (c) the aircraft attitude correction, and (d) the calculation of the HDRF. A detailed description of the different steps can be found in Carlsen (2018).

10 The measured signal was converted into the physical quantity of radiance (unit of $\text{W m}^{-2} \text{nm}^{-1} \text{sr}^{-1}$) by means of a radiometric calibration in the laboratory. The pixel respond differently to an uniform illumination due to manufacturing tolerances (e.g., irregularities in the used silicon), contamination with dust particles, and optical effects at the edges of the lenses. These deviations lead to the photo response non-uniformity (PRNU) and need to be corrected.

15 The camera was positioned in front of an integrating sphere that served as an uniform radiation source. However, a distinct decrease in brightness is visible from the center to the edges of the sensor. This vignetting effect is typical for digital cameras (e.g., Lebourgeois et al., 2008). From the data sheet, the spectral radiance $I_{\text{sphere}}(\lambda)$ that the integrating sphere emits at the specific optometer current is known. The calibration factor k_c is defined at each pixel location ~~(x, y)~~ on the sensor (row x , column y) as

$$k_c(x, y) = \frac{I_{\text{sphere}}(\lambda)}{s(x, y)} \cdot t_{\text{exp}}, \quad (16)$$

20 and carries the unit of $\text{W m}^{-2} \text{nm}^{-1} \text{sr}^{-1} (\text{DN/s})^{-1}$. During the calibration, the exposure time t_{exp} was set to 1/1000 s. Not only does k_c correct for the PRNU, it simultaneously performs the absolute calibration transforming the measured digital signal into the physical quantity of radiance with units $\text{W m}^{-2} \text{nm}^{-1} \text{sr}^{-1}$.

25 The geometric calibration of the camera-lens system relates each sensor pixel to a viewing zenith and azimuth angle (θ_v , φ_v). Often, the process of geometric calibration of a camera involves calibration equipment or the use of planar targets such as checkerboard patterns (e.g., Tsai, 1987; Urquhart et al., 2016). Within this work, a stellar calibration method is applied (e.g., Schmid, 1974; Klaus et al., 2004; Mori et al., 2013; Urquhart et al., 2016) utilizing the high precision to which the positions of stellar objects are known. In this regard, 684 different star positions identified in subsequent pictures of the night sky were utilized to calculate θ_v and φ_v for each sensor pixel.

30 ~~As the camera is fixed to the aircraft frame, a correction for the aircraft attitude was implemented. Thus, beside the geometric calibration, the viewing angles θ_v and φ_v of the camera are determined by the attitude angles of the aircraft. The movement of the aircraft needs to be corrected to obtain the reflection angles θ_r and φ_r . In this regard, Euler rotations were applied as described in Ehrlich et al. (2012).~~

3.4 Calculation and uncertainties of measured surface HDRF

The camera took pictures with a temporal resolution of about 8 s. Exemplarily, Fig. 4 demonstrates the derivation of the snow HDRF from a raw image taken by the camera on 2 January 2014 (8:16 UTC, see Fig. 4a). First, the pixels receiving radiation from the direction of the aircraft frame are excluded. For each pixel location (x, y) , the radiance $I(x, y)$ (in units $W m^{-2} nm^{-1} sr^{-1}$) is calculated from the measured signal s using the absolute calibration factor k_c and the exposure time t_{exp} :

$$I(x, y) = \frac{s(x, y)}{t_{exp}} \cdot k_c(x, y) \quad (17)$$

The camera viewing angles are calculated from the geometric calibration for each pixel (x, y) . As the camera is fixed to the aircraft frame, a correction for the aircraft attitude was implemented to obtain the reflection angles θ_r and φ_r . Thus, beside the geometric calibration, the observed reflection angles are determined by the attitude angles of the aircraft. Utilizing the data from the internal navigation system (INS) and the global positioning system (GPS) on Polar 6, the viewing angles are corrected depending on the roll, pitch and yaw angles of the aircraft at the time of measurement. In this regard, Euler rotations are applied as described in Ehrlich et al. (2012). Finally, the observed HDRF is calculated by:

$$fR_{HDRF}(\theta_0, \varphi_0; \theta_r, \varphi_r) = \frac{\pi \cdot I(\theta_r, \varphi_r)}{F^\downarrow} \frac{\pi sr \cdot I(\theta_r, \varphi_r)}{F^\downarrow(\theta_0, \varphi_0)} \quad (18)$$

$F^\downarrow(\theta_0, \varphi_0)$ denotes the downward solar irradiance at the time of measurement (with solar zenith and azimuth angles θ_0 and φ_0). Based on the resulting reflection angles, a polar plot of the measured fR_{HDRF} is created (camera channel G, see Fig. 4b). To achieve comparability, each image is rotated into the azimuthal direction of the Sun.

~~Note that the~~ The downward irradiance measurements from SMART could not be used for the calculation of the HDRF due to calibration issues. ~~Instead, the~~ Note that this pertains to the radiometric calibration only, which converts the digital numbers registered by the spectrometer into units of irradiance. For the albedo measurements with SMART, a relative calibration of the upper and lower sensors is sufficient and an absolute radiometric calibration is not required. Thus, the albedo measurements and the retrieval of the optical-equivalent snow grain size are unaffected by this calibration issues. For the calculation of the HDRF, the global irradiance was simulated along the flight track with the library for radiative transfer libRadtran (Mayer and Kylling, 2005) using the discrete ordinate radiative transfer solver DISORT by Stamnes et al. (1988) instead. Vertical profile data of air temperature and relative humidity measured by radiosondes at Kohnen, Neumayer III (König-Langlo, 2014), and Halley (Durre et al., 2016) research stations were used as input for the simulations. The simulated irradiance was integrated over the wavelength range of each camera channel and weighted with the RSR function of the camera. The use of simulations limits the validity of absolute values of the measured HDRF to cloudless conditions. ~~However, within this work~~ Cloudless cases (approximately 75 % of the camera observations) were identified from visual synoptic observations during the flights as well as using the downward irradiance measured with SMART. Periods with fast fluctuations in the downward irradiance were flagged as cloudy. Although mainly the shape of the HDRF is analyzed within this work (which is independent from the absolute value of F^\downarrow), the analysis is restricted to cloudless conditions only.

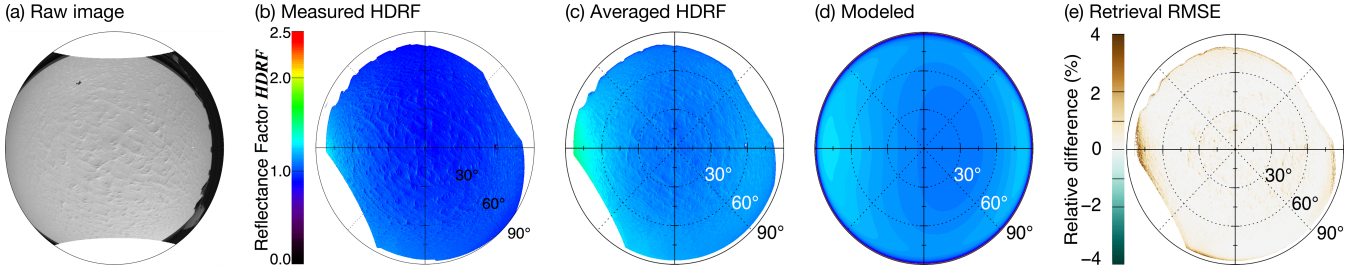


Figure 4. Exemplary post-processing and retrieval of model parameters for HDRF measurement. **(a)** Raw image taken on 2 January 2014 at 8:16 UTC ($\theta_0 = 58.9^\circ$). **(b)** Polar plot of observed snow HDRF ρ_l calculated from raw image for camera channel G. The image is rotated into the azimuthal direction of the Sun such that the Sun is on the left. **(c)** Averaged HDRF measurement using 3 images, the footprint is 1670 m. **(d)** Modeled snow HDRF $f_{\text{HDRF},l} R_{\text{HDRF},l}$ after retrieval of model parameters $f_{\text{iso}} = 1.12$, $f_{\text{vol}} = 0.17$, and $f_{\text{geo}} = 0.01$ from (c). **(e)** Relative difference $(\rho_l - f_{\text{HDRF},l}) / f_{\text{HDRF},l} (\rho_l - R_{\text{HDRF},l}) / R_{\text{HDRF},l}$. The accumulated RMSE is 0.04.

The overall relative uncertainty of the HDRF measurements with the digital camera range in the order of 4.5%. The error in the absolute value of $f_{\text{HDRF},l} R_{\text{HDRF},l}$ might be higher depending on the atmospheric conditions due to the usage of simulated values for the global irradiance F^\downarrow in Eq. (18). The uncertainties in the HDRF measurements stem from sensor characteristics (estimated with 0.5% due to signal-to-noise ratio, dark current, linearity, read-out noise, chrominance subsampling), the radiometric calibration (4%), the geometric calibration and the correction for the aircraft attitude (combined estimate of 1.0%).

3.5 Averaging

From trigonometric considerations, the footprint of the camera (twice the radius r of the disc on the ground pictured by the camera) depends on the flight altitude z and the FOV. At an altitude of 100 m, the radius is 570 m and the footprint approximately 1 km. The footprint grows with increasing altitude to 5.7 km ($z = 500$ m) and 11.3 km ($z = 1000$ m). To get independent of local roughness features, averaging is performed over time intervals of around 30 s (including 3 to 4 pictures). This interval was chosen trading off the gained independence from local features against a growing footprint and thus reduced comparability with the MODIS BRDF measurements that are provided on a 500×500 m grid. Generally, averaging over 30 s yields footprints between 1 and 2 km. However, individual footprints depend on flight altitude, exact averaging time, and aircraft velocity. The averaged HDRF for the exemplary measurement on 2 January 2014 for a footprint of 1670 m is shown in Fig. 4c.

3.6 Approximation of surface BRDF with HDRF measurements

For each averaged HDRF, the weighting functions f_{iso} , f_{vol} , and f_{geo} are retrieved by inverting the RTLSR BRDF model and setting the HDRF equal to the BRDF in Eq. 12. This is necessary as the BRDF is not measurable under atmospheric conditions. However, the atmospheric influence is assumed to be small due to the high surface elevation and the dry, aerosol-free atmosphere on the Antarctic plateau: the mean AOD at Kohnen station was 0.015 between 2001 and 2006 (at 500 nm, Tomasi et al., 2007).

and the mean integrated atmospheric water vapor was 1.1 kg m^{-2} between December 2013 and January 2014 (Carlsen et al., 2017). Simulations with the library for radiative transfer libRadtran (Mayer and Kylling, 2005) of the mean direct fraction of the global irradiance f_{dir} at Kohnen station between 10 December 2013 and 31 January 2014 indicate rather weak atmospheric scattering effects. For the central wavelengths of the three camera channels, the mean direct fraction is 87 % (for 602 nm wavelength, channel R), 81 % (538 nm, G), and 69 % (470 nm, B). The values for f_{dir} are even higher when contributions from a small scattering cone around the incident direction are considered. Schaepman-Strub et al. (2006) simulated the difference between the HDRF and BRDF ($R_{\text{BRDF}} = \pi \cdot f_{\text{BRDF}}$, compare Eq. 2) for snow surfaces for different fractions f_{dir} at 550 nm wavelength. They found that with an increasing diffuse fraction of the incident irradiance, the shape of the HDRF is smoothed in comparison to the BRDF ($f_{\text{dir}} = 1$, see Eq. 3). For $f_{\text{dir}} = 0.8$ at $\theta_0 = 30^\circ$, the shape of the HDRF is still close to that of the BRDF. Hence, the measured HDRF shapes can serve as approximations for the intrinsic BRDF of the underlying snow surface for the measurements analyzed within this study ($f_{\text{dir}} > 0.8$).

3.7 Quality of the inversion

An exemplary inversion is performed on the snow HDRF measurement from 2 January 2014 at 8:16 UTC ($\theta_0 = 58.9^\circ$). Figure 4c shows the averaged observations that are used as input for the algorithm. The inversion leads to the retrieved model parameters of 1.12 (f_{iso}), 0.17 (f_{vol}), and 0.01 (f_{geo}). With the calculated kernels K_{kl} , the modeled HDRF is obtained (see Fig. 4d). The shapes of modeled and measured HDRF are similar; their relative difference is mostly within the range of $\pm 5\%$ (see Fig. 4e). The largest deviations occur in the forward scattering direction (up to 3 %) because the location of maximum is not mimicked perfectly by the modeled HDRF. However, up to viewing zenith angles of around 75° , the relative difference is below 1 %.

The root-mean-square error (RMSE),

$$\sigma_{\text{RMSE}} = \mathcal{E} = \sqrt{\frac{1}{N-3} \cdot \sum_{l=1}^N \frac{(\rho_l - f_{\text{HDRF},l})^2}{w_l}} \sqrt{\frac{1}{N-3} \cdot \sum_{l=1}^N \frac{(\rho_l - R_{\text{HDRF},l})^2}{w_l}}, \quad (19)$$

serves as a criterion for the quality of the inversion. Only RMSE below 0.1 are considered full inversions within the MODIS retrieval (e.g., Stroeve et al., 2005). For the example shown in Fig. 4, it is 4.0 %. σ_{RMSE} strongly depends on the anisotropy of the observed HDRF. Figure 5 shows the RMSE for all airborne retrievals as a histogram, with 96 % of the retrievals showing an RMSE below 0.1. Hence, the upper bound for the uncertainty of the inversion is in the range of 10 %.

The influence of the angular sampling on the retrieval of the model parameters can be investigated using the weights of determination (WoD, Lucht and Lewis, 2000), which depend both on the angular sampling and the number of directional measurements N . WoDs (also: noise amplification factors) below one indicate a retrieval uncertainty that is smaller than the RMSE of the inversions. Due to the large N and consistent sampling of the viewing hemisphere, the WoD for the model parameters f_k within this study are mostly below 0.0002. For comparison, the sparse angular sampling of the reflectance measurements from MODIS observations can lead to WoD above one and only model fits with WoD below 2.5 are considered full inversions.

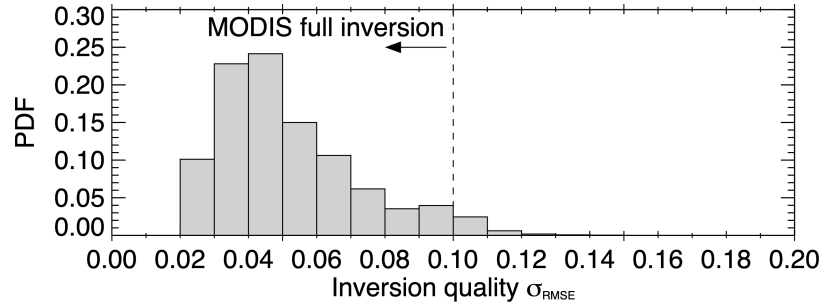


Figure 5. Statistics of the quality of the inversion. RMSE of all airborne retrievals (PDF, binsize: 0.01). Vertical dashed line shows RMSE threshold of 0.1 for MODIS full inversions.

To study the stability of the inversion as a result of the uncertainties in the reflectance measurements, the stability of the retrieval of f_k was studied for a 5 min flight leg over a homogeneous snow surface on 28 December 2013: the surface roughness varied between 2 and 4 cm, R_{opt} varied between 80 and 90 μm , and the solar zenith angle was around 55.3° . The inversion was performed for each individual camera picture (no averaging). The mean and standard deviation of the retrieved model parameters along the flight leg were 1.104 ± 0.030 (f_{iso}), 0.296 ± 0.027 (f_{vol}), and 0.031 ± 0.003 (f_{geo}). The same area was sampled five times at different solar zenith angles on 28 December 2013, the influence of θ_0 on the standard deviations of the f_k is negligible. The mean standard deviations of 0.031 (f_{iso}), 0.020 (f_{vol}), and 0.003 (f_{geo}) are used as error bars for the f_k within this study.

3.8 Approximation of surface BRDF with HDRF measurements

Note that this inversion yields the parameters of a BRDF model derived from a HDRF measurement as the BRDF is not measurable under atmospheric conditions. However, the atmospheric influence is assumed to be small due to the high surface elevation and the dry, aerosol-free atmosphere on the Antarctic plateau: the mean AOD at Kohnen station was 0.015 between 2001 and 2006 (at 500 nm, Tomasi et al., 2007) and the mean integrated atmospheric water vapor was 1.1 between December 2013 and January 2014 (Carlsen et al., 2017). Simulations with the library for radiative transfer libRadtran (Mayer and Kylling, 2005) of the mean direct fraction of the global irradiance f_{dir} at Kohnen station between 10 December 2013 and 31 January 2014 indicate rather weak atmospheric scattering effects. For the central wavelengths of the three camera channels, the mean direct fraction is 87% (for 602 wavelength, channel R), 81% (538, G), and 69% (470, B). The values for f_{dir} are even higher when contributions from a small scattering cone around the incident direction are considered. Schaepman-Strub et al. (2006) simulated the difference between HDRF and BRDF for snow surfaces for different fractions f_{dir} at 550 wavelength. They found that with an increasing diffuse fraction of the incident irradiance, the shape of the HDRF is smoothed in comparison to the BRDF ($f_{\text{dir}} = 1$, see Eq. 3). For $f_{\text{dir}} = 0.8$ at $\theta_0 = 30^\circ$, the shape of the HDRF is still close to that of the BRDF. Hence, the measured HDRF shapes can serve as approximations for the intrinsic BRDF of the underlying snow surface for the measurements analyzed within this study ($f_{\text{dir}} > 0.8$).

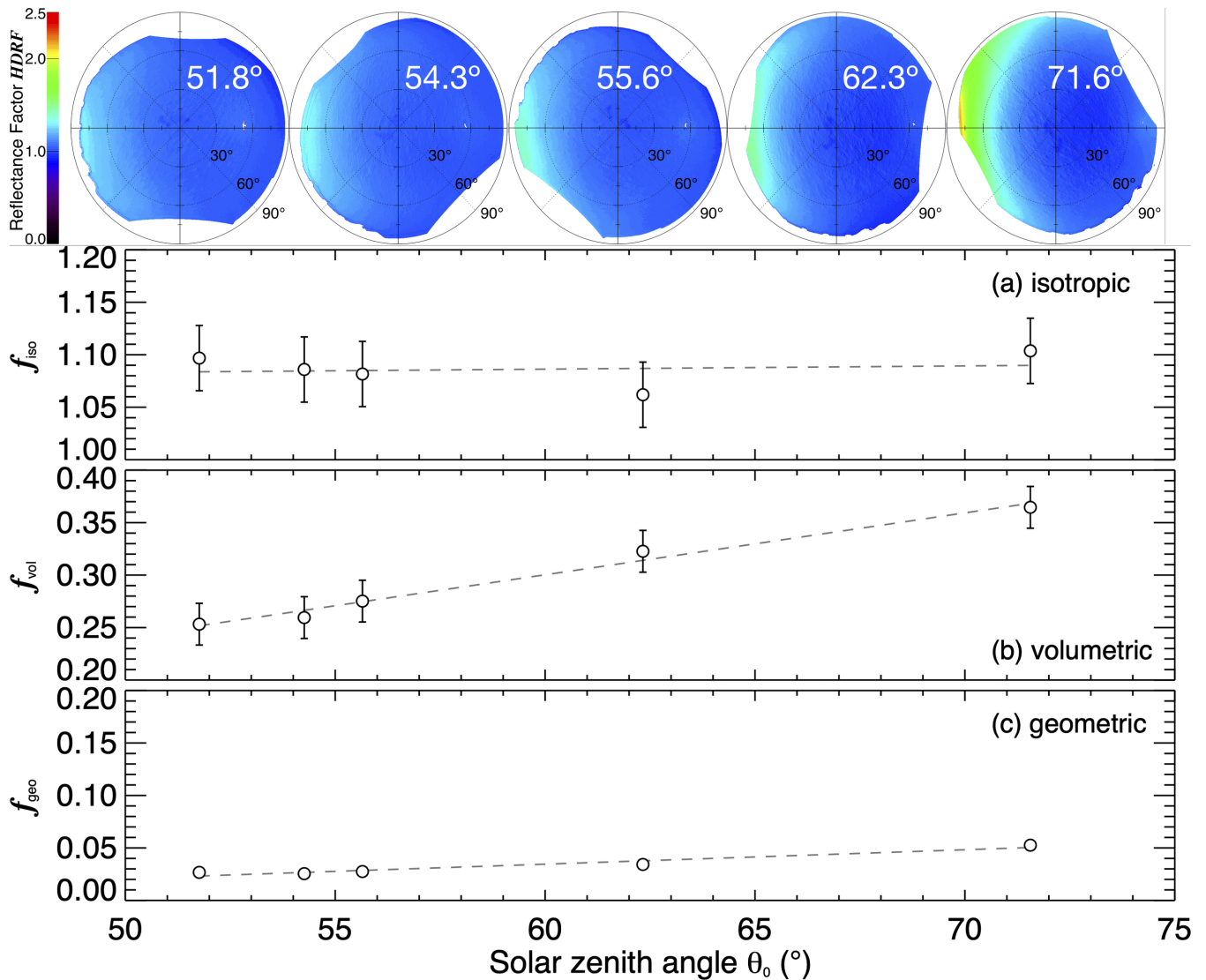


Figure 6. Top: Averaged snow HDRF (490-585 nm wavelength) over 30 s-segment during easternmost flight leg for the five consecutive research flights on 28 December 2013. **Below:** Dependence of the retrieved f_k from the solar zenith angle θ_0 for the five flights. The dashed lines represent linear regression fits of the parameters with respect to θ_0 . **(a)** The isotropic weight f_{iso} . **(b)** The volumetric weight f_{vol} . **(c)** The geometric weight f_{geo} .

3.8 Comparison with BRDF derived from MODIS satellite observations

The use of the RTLSR model in the retrieval of the model parameters f_k from airborne observations provides the framework to compare with satellite observations. The MODIS BRDF/Albedo Model Parameters product (MCD43A1) provides the f_k in 500 m resolution utilizing multi-date, atmospherically corrected, and cloud-cleared input data over a period of 16 days (Schaaf

et al., 2002). Version 6 of the MCD43A1 product uses both Terra and Aqua data and is temporally weighted to the 9th day of a 16-day retrieval window, thereby putting higher emphasis on the actual day of interest than previous versions. The combination of MODIS on Terra and Aqua provides a higher number of high quality reflectance measurements, resulting in better temporal and angular sampling. As the retrieval uncertainty increases with larger solar zenith angle, a thorough quality assessment of the satellite data is needed especially in polar regions. Stroeve et al. (2005) compared the MODIS albedo product with in situ observations on the Greenland ice sheet and recommended to only use full inversion data (RMSE < 0.1, WoD < 2.5). Hence, this study restricts the satellite data to full inversions based on the quality flag information provided in the MCD43A2 product (quality flag values 0 or 1). Lower quality magnitude inversions would facilitate archetypal BRDF parameters to improve the retrieval in case of less than 7 high quality reflectance observations or poor angular sampling. The BRDF data from MODIS spectral band 4 (0.545-0.565 μm) is used as it coincides best with the green camera channel (see Fig. 3).

The MODIS data was resampled on the flight track using nearest neighbour with respect to the great circle distance. For the comparison, the aircraft measurements are filtered analogously: only clear-sky observations with RMSE < 0.1 (96 % of total measurements), WoD < 2.5 (valid for all observations), and a solar zenith angle lower than 70° were used. Stroeve et al. (2005) restricted the analysis to $\theta_0 < 75^\circ$, however, the number of comparable observations between MODIS and the aircraft is not reduced by the stricter condition. In fact, only 434 (21 %) MODIS observations can be compared to the 2078 airborne observations that in principle fulfill the quality requirements. For the remaining observations, no full inversion retrieval could be performed from the MODIS observations.

4 Results and discussion

To study the influence of the solar zenith angle, the surface roughness, and the optical-equivalent snow grain size on the snow HDRF, the effects of the individual parameters need to be separated. For this purpose, several case studies were selected: (a) measurements at different θ_0 (marked with red circles in Fig. 2a, five flights on 28 December 2013), (b) measurements at different surface roughness (marked with green circles in Fig. 2b, morning flight on 2 January 2014), and (c) measurements at different R_{opt} (marked with brown circles in Fig. 2c, afternoon flight on 2 January 2014). The remaining two parameters were kept constant throughout the individual cases. After discussing the individual cases, the airborne observations are compared with the BRDF derived from MODIS.

4.1 Influence of solar zenith angle

The research flights with Polar 6 on 28 December 2013 were exploited to study the influence of θ_0 on the snow HDRF. Five consecutive flights were conducted in an area northeast of Kohnen (see red flight track in Fig. 1a).

For each flight, the snow HDRF was averaged over a 30 s-segment during the easternmost flight leg; the results are shown in the top panel of Fig. 6. The solar zenith angle varied between 51.8° and 71.6° . As Polar 6 took identical routes throughout the flight, the respective flight leg always covered the same area. The central points of the five consecutive segments are separated by less than 1 km. Therefore, the surface roughness (retrieved from the laser scanner) of approximately 6 to 7 cm remained

constant. Within the retrieval uncertainties, the same holds true for the optical-equivalent snow grain size. R_{opt} varied between 70 μm and 85 μm . This minimizes a possible influence of R_{opt} or l_{rough} on the snow HDRF that would be superimposed on the effect of θ_0 . Cloudless conditions prevailed during the flights.

With increasing θ_0 , the maximum of the HDRF in the forward scattering direction becomes more pronounced (see Fig. 6). Correspondingly, the anisotropy gets larger. Figure 6 shows the dependence of the f_k from θ_0 for the green camera channel (490-585 nm wavelength). f_{iso} shows no clear dependence on θ_0 and varies between 1.06 and 1.10. f_{geo} weakly increases with θ_0 from 0.03 at $\theta_0 = 51.8^\circ$ to 0.05 at 71.6° . During the first three research flights, θ_0 did vary only slightly. Therefore, the strongest trends are visible between flights 3 and 5. In particular, f_{vol} increases from 0.25 to 0.36.

The probability for photons entering the snowpack to leave it after just a few scattering events increases for lower Sun elevation. With increasing θ_0 , the reflection properties of the snow layer converge to the single-scattering properties of ice crystals. In addition, longer shadows are cast by surface roughness structures at larger θ_0 . This expected increase in the anisotropy of the snow HDRF for increasing θ_0 is obvious in the changing model parameters f_k [and has been observed in earlier studies \(e.g., Dirmhirn and Eaton, 1975; Kuhn, 1985; Warren et al., 1998; Hudson et al., 2006\)](#).

4.2 Influence of surface roughness

On 2 January 2014, a research flight was conducted from Kohnen in the direction of the coastline to Neumayer III research station. The flight track is shown in Fig. 1a (green). Between 8:50 and 9:00 UTC, Polar 6 crossed a sastrugi field at an elevation between 100 and 250 m above sea level. This fostered large variations in the surface roughness between 42 cm and 2.2 m which became distinguishable by visual observation from the aircraft. At the same time, the solar zenith angle was constant at around 55.5° . For these values, only little variation of the snow HDRF with θ_0 was found (compare Fig. 6). The optical-equivalent snow grain size varied between 240 and 320 μm , the flight altitude remained at around 600 m above the surface leading to a footprint of approximately 3.5 to 4.0 km. During this flight leg, 14 measurements were used to study the influence of the surface roughness on the snow HDRF. Two examples of the snow HDRF are shown in the top panel of Fig. 7.

For the smooth surfaces, the maximum in forward scattering direction is more pronounced. With increasing surface roughness, the anisotropy gets lower. ~~This is in accordance with Warren et al. (1998) who observed a reduction of the forward peak due to sastrugi during tower measurements at South Pole Station.~~ Roughness structures enhance the backscatter by changing the effective angle of incidence. In addition, they cast shadows that reduce the forward scatter. The evolution of the retrieved model parameters f_k as illustrated in Fig. 7 demonstrates the decreasing anisotropy: f_{iso} and f_{geo} show a slight decreasing trend whereas, most prominently, f_{vol} decreases from 0.33 to 0.24 with increasing surface roughness. [This is in accordance with Warren et al. \(1998\) who observed a reduction of the forward reflection peak due to sastrugi during tower measurements at South Pole Station. The effect of surface roughness on the BRDF has been previously investigated in observational \(e.g., Grenfell et al., 1994; Hudson and Warren, 2007\) and modeling studies \(e.g., Leroux and Fily, 1998; Zhuravleva and K](#) [and is facilitated by the multi-angular reflectance data from MISR to map surface roughness over ice sheets and sea ice from space \(Nolin et al., 2002\).](#)

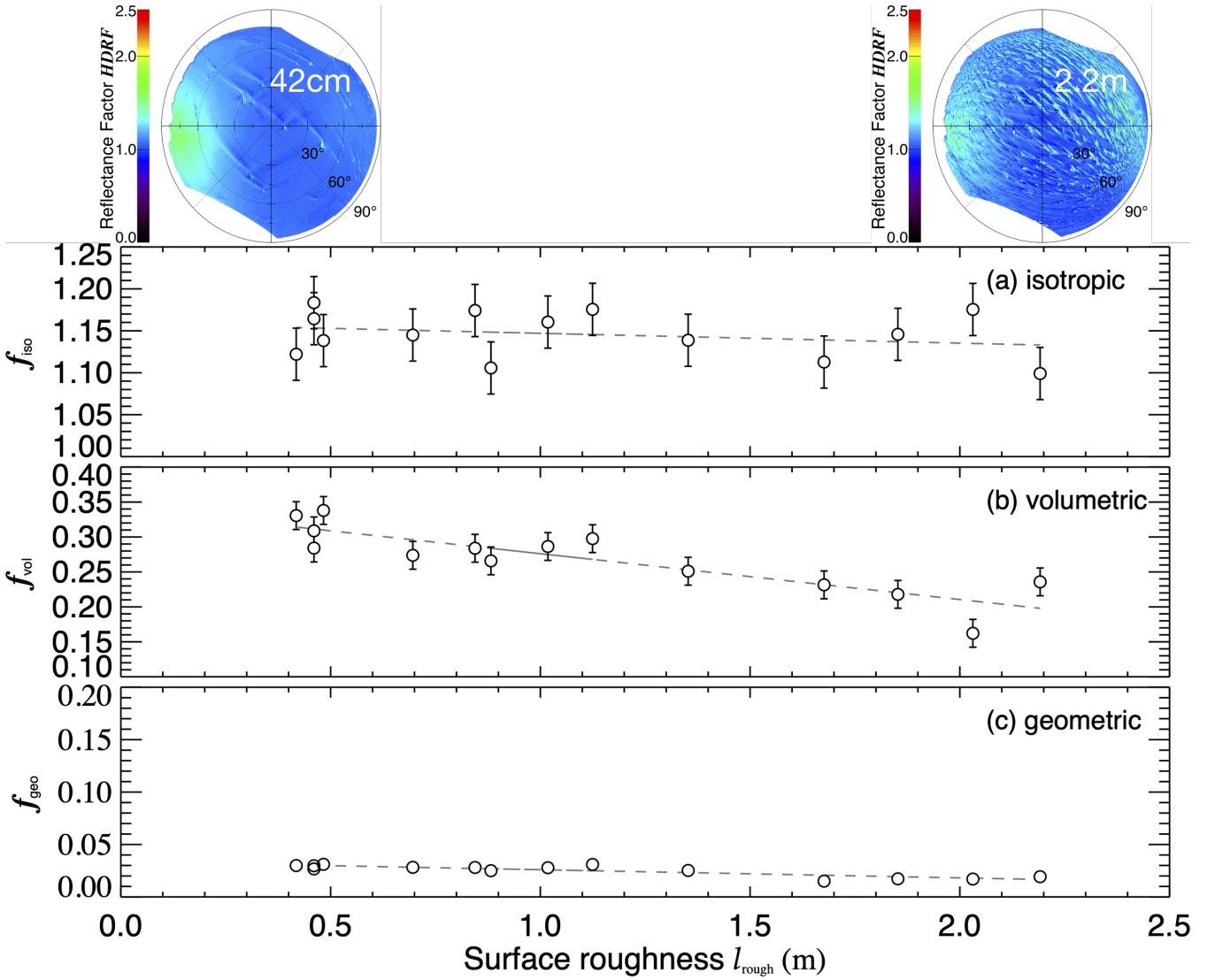


Figure 7. Top: Two examples of snow HDRF (490-585 nm wavelength) at different times during an overflight over a sastrugi field on 2 January 2014 (morning flight). **Below:** Dependence of the retrieved f_k from the surface roughness. The dashed lines represent linear regression fits of the parameters with respect to l_{rough} . **(a)** The isotropic weight f_{iso} . **(b)** The volumetric weight f_{vol} . **(c)** The geometric weight f_{geo} .

4.3 Influence of optical-equivalent snow grain size

The separation of the two parameters θ_0 and l_{rough} showed that they influence the snow HDRF in opposite ways. Ice absorption becomes dominant in the near-infrared part of the solar EM spectrum. Hence, even though a wide variety of optical-equivalent snow grain sizes was covered during the flights (16 to 480 μm), no large effect on the snow HDRF was expected as the camera

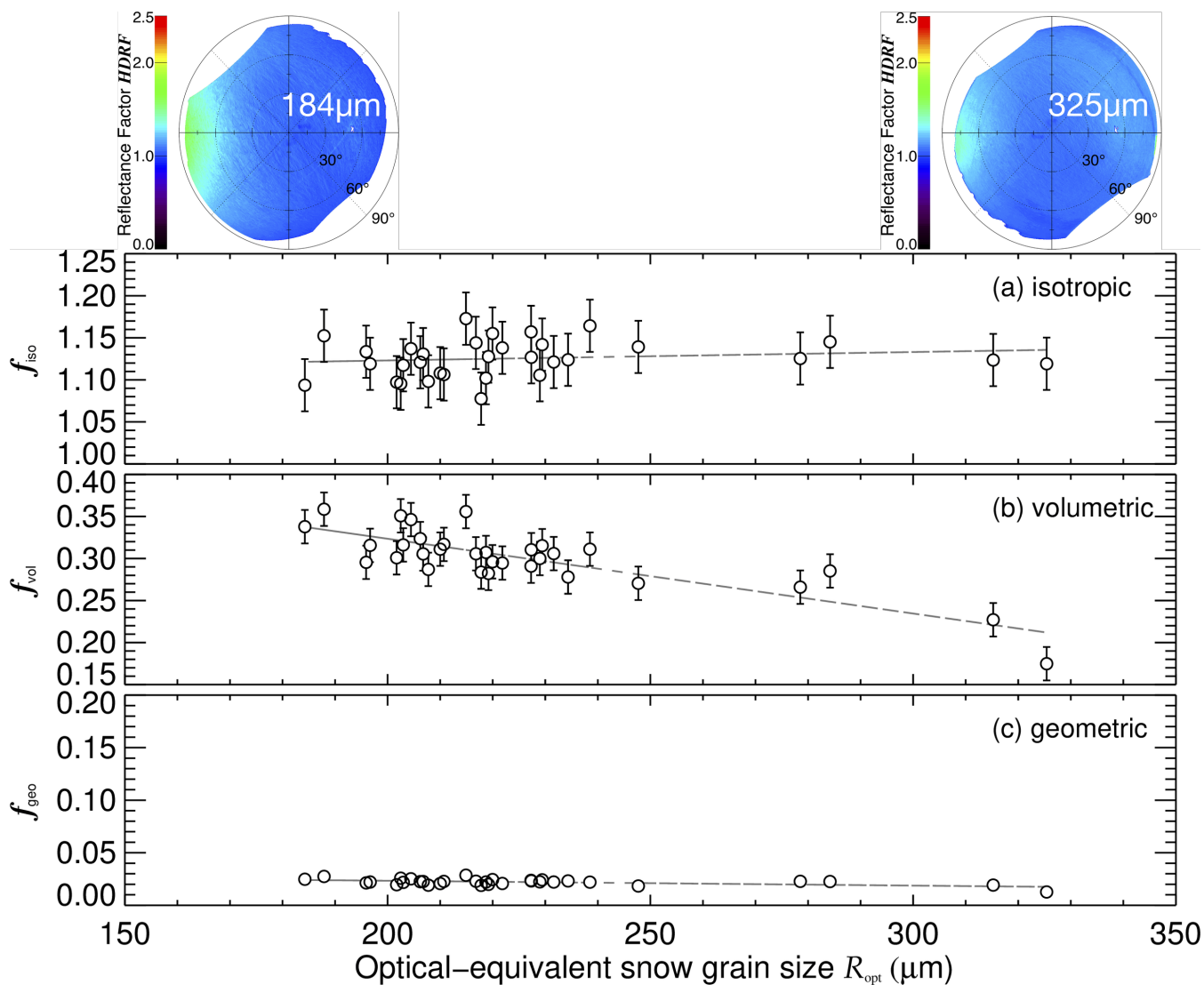


Figure 8. Top: Two examples of snow HDRF (490-585 nm wavelength) at different times on 2 January 2014 (afternoon flight). **Below:** Dependence of the retrieved f_k from the optical-equivalent snow grain size. The dashed lines represent linear regression fits of the parameters with respect to R_{opt} . **(a)** The isotropic weight f_{iso} . **(b)** The volumetric weight f_{vol} . **(c)** The geometric weight f_{geo} .

is only sensitive in the visible wavelength range (Warren et al., 1998, compare Fig. 3). However, 32 observations between 14:45 and 15:15 UTC on 2 January 2014 (see brown flight track in Fig. 1a) and depicted in Fig. 8 show a decrease of f_{vol} from 0.34 at 184 µm to 0.17 at 325 µm snow grain size. This is consistent with the decrease in anisotropy as visible in the averaged HDRF measurements (top panel in Fig. 8).

- 5 However, this stands in contrast to findings by Steffen (1987), who measured the bidirectional reflectance on a glacier in Northwestern China for powder snow (grain size 0.15 mm), new snow (0.5-1.0 mm), and old snow (1-3 mm) and found higher

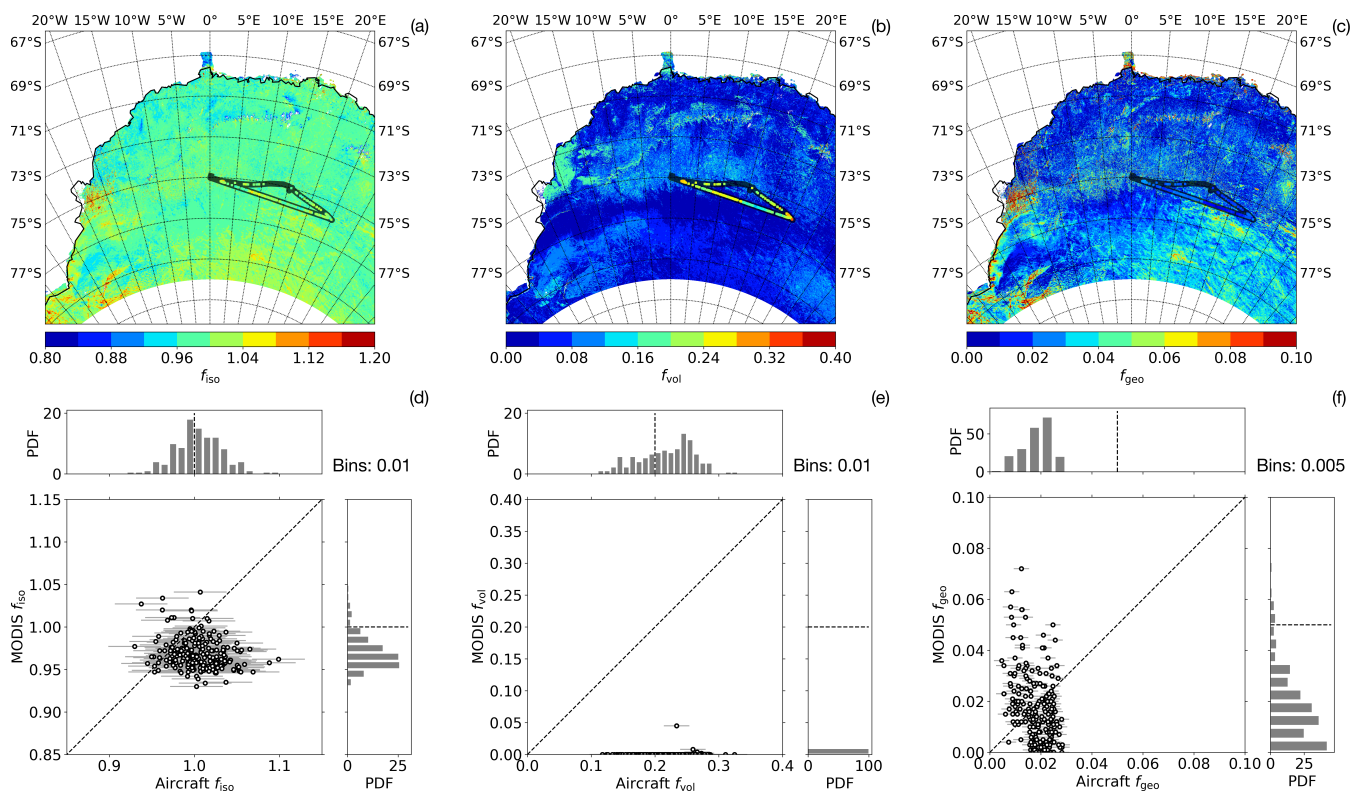


Figure 9. Comparison between airborne **and** MODIS retrievals of f_k (490-585 nm wavelength) **and** MODIS (545-565 nm) retrievals of f_k . **Top:** (a) Map of the area covered by the research flights, the colored contours correspond to the MODIS retrieval of f_{iso} . Black: flight track on 29 December 2013. Colored circles on black flight track: airborne retrieval of f_{iso} . (b) As in (a), but for f_{vol} . (c) As in (a), but for f_{geo} . **Below:** (d) Corresponding combined scatter plot and histograms between airborne and satellite (resampled on flight track) retrieval for f_{iso} . Error bars correspond to the mean standard deviation of the airborne retrieval. Dashed in scatter plot: 1:1 line. Dashed in histograms: center of axis. (e) As in (d), but for f_{vol} . (f) As in (d), but for f_{geo} .

anisotropy for larger snow grains. It should be noted that the range of snow grain sizes covered within this study is significantly smaller and the footprint of the measurements is larger compared to the local study by Steffen (1987). Even though other parameters were kept constant during this case study (θ_0 : 51.3-53.4°, l_{rough} : 4-7 cm, footprint: 1.3-1.7 km), effects of other influencing parameters cannot be ruled out completely. For example, Steffen (1987) reported changing ice crystal shapes
 5 between the measurements and suggested a possible influence of dust particles within the snow. This could partly explain the opposite conclusions drawn within this study.

4.4 Parameterization of snow HDRF

The most significant effects of θ_0 , l_{rough} , and R_{opt} were observed on the model parameter f_{vol} , for which the parameterizations in terms of linear regression fits (black dashed lines in Figs. 6b-8b) are given by:

$$f_{\text{vol}}(\theta_0) = 0.0059 \cdot \theta_0 - 0.0537 \quad (\text{for } l_{\text{rough}} = 6 - 7 \text{ cm and } R_{\text{opt}} = 70 - 85 \mu\text{m}), \quad (20)$$

$$f_{\text{vol}}(l_{\text{rough}}) = -0.0656 \cdot l_{\text{rough}}(\text{m}) + 0.3416 \quad (\text{for } \theta_0 = 55.5^\circ \text{ and } R_{\text{opt}} = 240 - 320 \mu\text{m}), \quad (21)$$

$$f_{\text{vol}}(R_{\text{opt}}) = -0.0009 \cdot R_{\text{opt}}(\mu\text{m}) + 0.5013 \quad (\text{for } \theta_0 = 51.3 - 53.4^\circ \text{ and } l_{\text{rough}} = 4 - 7 \text{ cm}). \quad (22)$$

The linear correlation coefficients (R^2) are 0.98 (for θ_0), 0.75 (l_{rough}), and 0.68 (R_{opt}). Note that these parameterizations are only valid for the wavelength range of about 490-585 nm (green camera channel) and the ranges of all three parameters as given within the descriptions of the case studies.

4.5 Comparison with satellite observations

Figure 9 shows the comparison between retrieved f_k from airborne and satellite measurements for the research flight on 29 December 2013, which showed the largest amount of full inversion retrievals for MODIS (235). The flight covered a triangular pattern Southeast of Kohnen station. The top panel shows contour plots with color-coding for the three parameters as retrieved from MODIS on that day: (a) f_{iso} , (b) f_{vol} , and (c) f_{geo} . The colored circles along the flight track (black) correspond to the airborne retrieved f_k . The corresponding scatter plots between the MODIS and aircraft retrievals are shown in Figs. 9d-f. The isotropic and geometric model weights f_{iso} and f_{geo} from airborne and MODIS retrievals are in the same order of magnitude and partly agree within the measurement uncertainties. f_{iso} values from aircraft are slightly higher than the corresponding MODIS retrievals. This is evident from the difference in the PDFs between aircraft and MODIS retrievals as illustrated in Figs. 9d-f. The comparison of f_{geo} values shows no robust difference between airborne and MODIS retrievals. However, the lowest correlation is found for the volumetric weights f_{vol} . f_{vol} values retrieved from aircraft range mostly between 0.1 and 0.3 and are much larger than the MODIS retrievals. The map in Fig. 9b shows that the flight covered an area where MODIS retrieved f_{vol} values of exactly or close to zero. In particular, this holds true for a latitudinal band around 76° S. Even though full inversions were reported for this area, this might still be a measurement artifact. But even for the cases with non-zero values, the aircraft retrievals lead to much higher f_{vol} values and, thus, larger anisotropy of the surface reflection than shown in the MODIS retrieval.

The combined scatter plots and histograms for the three model weights including all 434 observations from 7 different research flights are shown in Fig. 10 and indicate that the conclusions drawn from the flight on 29 December 2013 are valid for all research flights: f_{vol} shows the strongest discrepancies between airborne and MODIS retrievals, whereas f_{iso} values are slightly higher when retrieved from the aircraft measurements. The correlation between airborne and satellite measurements is low for all three model parameters, R^2 is 0.18 (for f_{iso}), 0.004 (f_{vol}), and 0.055 (f_{geo}).

Various factors can contribute to the observed differences: (1) data quality, (2) short-term changes in snow properties, (3) footprint size differences, (4) inherent challenges comparing the camera and satellite data.

First, a thorough quality assessment for both datasets formed the prerequisite for the comparison. Only airborne data from cloudless conditions (at solar zenith angles below 70°) with a RMSE lower than 0.1 were compared with full inversion MODIS retrievals (quality flag values 0 or 1, $\text{RMSE} < 0.1$). However, possible artifacts such as the latitudinal band around 76° S with values of exactly zero cannot be ruled out entirely.

5 Secondly, the case studies demonstrated the strong influence of the solar zenith angle, the surface roughness, and the optical-equivalent snow grain size on the anisotropy of the reflection. Precipitation events or drifting snow altering the optical-equivalent snow grain size and the surface roughness can change the surface properties on short time scales. These short-term changes in snow properties are not captured well by the 16-day BRDF product from MODIS. Even though the 434 observations stem from 7 different research flights, most of the data originates from the two flights on 29 December 2013 and 2 January
10 2014 (morning flight). For most of the remaining flights, no full inversion retrieval was possible for MODIS. Being restricted to mainly two different days, short-term changes in snow properties can indeed explain the low correlation between the airborne and MODIS retrievals. Also, the solar zenith angle at the time of the aircraft measurements varies from the different zenith angles during the satellite overpasses used for the MODIS retrieval.

Thirdly, the airborne measurements still included varying footprint sizes (mostly between 1 and 3 km) compared to the
15 500 m resolution of the MODIS product. This was the result of trading off comparability with MODIS and data coverage. Also, no correlation between the f_k retrieved from the camera measurements and the footprint size was found (not shown here).

Lastly, comparing the airborne camera data with the MODIS product poses challenges that are inherent to the measurement setup. The observation times of the MODIS product do not coincide with the times of the research flights. MODIS spectral
20 band 4 was chosen as its RSR function agrees best with the green camera channel. As the HDRF is wavelength dependent, differences naturally occur due to different spectral coverage but should be small.

Thus, despite the thorough quality assessment analogously applied to both airborne and satellite data, a quantitative comparison between the two retrievals seems challenging. Nevertheless, the comparison highlights the difficulties to accurately measure the snow BRDF from MODIS as well as how quickly snow properties can change leading, in this study, to an under-
25 estimation of the anisotropic reflection. For the analyzed cases, MODIS underestimates the volumetric model weight f_{vol} on average by at least a factor of 10 compared to the airborne measurements (neglecting MODIS values of exactly zero). Many satellite retrievals of surface albedo and atmospheric parameters over snow surfaces rely on a correct angular modeling of the surface reflectance. Thus, inaccuracies in the surface BRDF can propagate and influence other retrieval results.

5 Conclusions

30 The HDRF of snow surfaces in the wavelength band 490-585 nm was ~~investigated with respect to the influence of solar zenith angle, surface roughness, and optical-equivalent snow grain size utilizing an airborne digital 180° fish-eye camera~~ derived from airborne measurements in Antarctica during austral summer in 2013/14. An airborne digital 180° fish-eye camera installed on a research aircraft was used for this purpose. The surface roughness was retrieved using an airborne laser scanner, while the

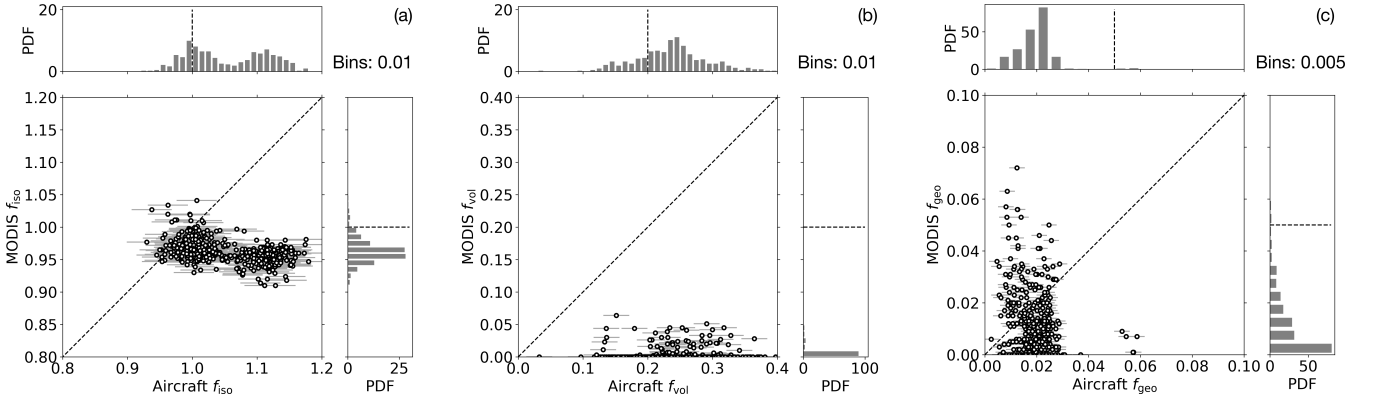


Figure 10. (a) Scatter plot between airborne (490-585 nm wavelength) and satellite-MODIS (resampled on flight track, 545-565 nm) retrieval for f_{iso} (490-585 wavelength)-including all 7 research flights used for the comparison. Error bars correspond to the mean standard deviation of the airborne retrieval. Dashed: 1:1 line. In addition, the histograms for both airborne and satellite retrievals are given. Dashed: center of axis. (b) As in (a), but for f_{vol} . (c) As in (a), but for f_{geo} .

optical-equivalent snow grain size was calculated from spectral surface albedo measurements. While based at Kohlen station (75°0' S, 0°4' E) at the outer part of the East Antarctic Plateau, the airborne measurements with the Polar 6 research aircraft covered an area of around 1000×1000 km² in Dronning Maud Land. Thus, the snow HDRF was obtained for a variety of conditions with different solar zenith angle (θ_0), surface roughness (l_{rough}), and optical-equivalent snow grain size (R_{opt}).

- 5 The digital camera provides airborne radiance measurements with high angular resolution. The technical characterization of the sensor-digital camera in the laboratory revealed excellent linearity as well as negligible dark current and read-out noise. The From the measured angular radiances, the HDRF was calculated from the angular radiance measurements applying simulations of the global irradiance with the library for radiative transfer libRadtran using the discrete ordinate radiative transfer solver DISORT. The relative uncertainty of the HDRF measurements is estimated with 4.5%. The footprint of the snow HDRF
- 10 measurements analyzed within this study varies between 1 km to 4.5 km.

- Three case studies were investigated to separate the effects of θ_0 , l_{rough} , and R_{opt} on the snow HDRF. With increasing solar zenith angle, the HDRF maximum in the forward scattering direction becomes more pronounced. Conversely, roughness structures enhance the backscatter the backscatter is enhanced over rougher surfaces by changing the effective angle of incidence and by casting shadows. This was quantified by inverting the semi-empirical kernel-driven RTLSR model and parameterizing the snow HDRF with respect to θ_0 , l_{rough} , and R_{opt} (Eqs. 20-22). The uncertainty of the inversion is estimated with 10%. The increased anisotropy (with increasing θ_0) is mainly shown by the increase in the model parameter f_{vol} . Vice versa, f_{vol} decreases with increasing l_{rough} . The snow grain size reveals a similar effect as surface roughness structures in terms of a decrease in of the anisotropy with increasing R_{opt} , which is a different result than found by earlier studies differs from earlier findings (e.g., Steffen, 1987). Possible reasons have been identified are discussed (footprint size, ice crystal shape, contamina-

tion with dust particles). However, to yield stronger dependence on the optical-equivalent snow grain size, a camera sensitive to radiation in the near-infrared part of the EM spectrum needs to be employed.

Continuous satellite observations of the snow surface albedo with global coverage ~~can be achieved by satellite observations.~~ ~~However, critically depend on~~ the angular modeling of the surface BRDF, which may introduce large errors especially in polar regions due to low sun elevations and the anisotropic scattering phase function of ice-snow crystals. The RTLSR model was chosen for the inversion as it forms the basis of the MODIS 16-day BRDF/albedo product. Thus, the retrieved model parameters f_k from the HDRF measurements allow for direct comparison with BRDF products from satellite remote sensing. Despite the similar quality assessment applied to airborne and satellite data, large deviations were found especially for the volumetric model weight f_{vol} . The airborne values for f_{vol} are larger than the corresponding MODIS retrievals (by at least a factor of 10). Short-term changes in snow properties (precipitation, drifting snow), which are not captured by the 16-day MODIS retrieval, and different solar zenith angles at the time of ~~measurement are suspected to be the main measurements are discussed as likely~~ reasons for the observed differences. Although the effects are presumably small, influences of a varying footprint size for the airborne observations and the different RSR functions of the green camera channel and MODIS spectral band 4 can further contribute to this discrepancies.

Generally, MODIS observations underestimated the anisotropy of the reflection ~~at by~~ the snow surfaces. This has possible implications for satellite retrievals of surface albedo ~~as well as and~~ atmospheric parameters over snow surfaces (e.g., AOD, cloud properties), which strongly depend on a correct angular modeling of the surface reflectance (e.g., Qu et al., 2015).

The analysis of the snow HDRF measurements can be readily applied to other airborne campaigns. For example, the same digital camera with a similar setup of the Polar 6 research aircraft was used (a) during the Arctic CLOUD Observations Using airborne measurements during polar Day (ACLOUD) between 22 May and 28 June 2017 based at Longyearbyen, Svalbard (Wendisch et al., 2019; Ehrlich et al., 2019; Jäkel and Ehrlich, 2019), and (b) during the Polar Airborne Measurements and Arctic Regional Climate Model Simulation Project (PAMARCMiP) between 12 March and 6 April 2018 at Station North, Greenland (Herber et al., 2012). ~~Thereby, the database of the parameterizations could be extended to wider ranges of R_{opt} , l_{rough} , and θ_0 .~~

Data availability. The tabulated data including the retrieved f_{iso} , f_{vol} , and f_{geo} as well as θ_0 , l_{rough} , R_{opt} , and auxiliary information (time and position of the measurements) is available for all research flights in the Publishing Network for Geoscientific & Environmental Data (PANGAEA): <https://doi.pangaea.de/10.1594/PANGAEA.921707>. The MODIS BRDF/Albedo Model Parameters product (MCD43A1, Version 6) and the quality dataset (MCD43A2) were downloaded from NASA's Land Processes Distributed Active Archive Center (LP DAAC) website (<https://doi.org/10.5067/MODIS/MCD43A1.006> and <https://doi.org/10.5067/MODIS/MCD43A2.006>).

Author contributions. TC performed the post-processing and data analysis of the camera and surface albedo measurements, produced the figures, and wrote the paper with help from all authors. GB and MS conducted the measurements at Kohnen station and with the Polar 6 research aircraft and helped write the paper. VH analyzed the laser scanner data for the surface roughness measurements and helped write

the paper. EJ gave substantial input with regard to the data analysis and helped write the paper. AE and MW planned the measurements, designed the study and helped write the paper.

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