



# 1 Antarctic high-resolution ice flow mapping and increased mass loss 2 in Wilkes Land, East Antarctica during 2006–2015

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4 Qiang Shen<sup>1,3</sup>, Hansheng Wang<sup>1</sup>, Che-Kwan Shum<sup>2,1</sup>, Liming Jiang<sup>1,3</sup>, Hou Tse Hsu<sup>1</sup>,  
5 Jinglong Dong<sup>1,3</sup>  
6

7 <sup>1</sup>State Key Laboratory of Geodesy and Earth's Dynamics, Institute of Geodesy and Geophysics,  
8 Chinese Academy of Sciences, Wuhan 430077, China

9 <sup>2</sup>Division of Geodetic Science, School of Earth Sciences, The Ohio State University, Columbus, Ohio  
10 43210, USA

11 <sup>3</sup>University of Chinese Academy of Sciences, Beijing 100049, China  
12

13 *Correspondence to:* Qiang Shen (c1980606@asch.whigg.ac.cn)  
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15 **Abstract.** Substantial accelerated mass loss, extensive dynamic thinning on the periphery, and  
16 grounding line retreat in the Amundsen Sea Embayment, have amplified the long-standing concerns on  
17 the instability of the Antarctic ice sheet. However, the evolution of the ice sheet and the underlying  
18 causes of the changes remain poorly understood due in part to incomplete observations. Here, we  
19 constructed the ice flow maps for the years 2014 and 2015 at high resolution (100 m), inferred from  
20 Landsat 8 images using feature tracking method. These maps were assembled from 10,690 scenes of  
21 displacement vectors inferred from more than 10,000 optical images acquired from December 2013 to  
22 March 2016. We also estimated the mass discharges of the Antarctic ice sheet in 2006, 2014, and 2015  
23 using the high-resolution ice flow maps, InSAR-derived ice flow map, and the ice thickness data. An  
24 increased mass discharge ( $40 \pm 24 \text{ Gt yr}^{-1}$ ) from East Indian Ocean sector was found in the last decade,  
25 attributed to unexpected widespread accelerating glaciers in Wilkes Land, East Antarctica, while the  
26 other five oceanic sectors did not show any significant changes, contrary to the long-standing belief  
27 that present-day accelerated mass loss primarily originates from West Antarctica and Antarctic  
28 Peninsula. In addition, we compared the ice sheet mass discharge with the new surface mass balance  
29 (SMB) data to estimate the Antarctic mass balance. The most significant change of mass balance also  
30 occurred in East Indian Ocean during the last decade, reaching  $-40 \pm 50 \text{ Gt yr}^{-1}$ , the large uncertainty is  
31 caused mainly by error in the SMB data. The newly discovered significant accelerated mass loss and  
32 speedup of ice shelves in Wilkes Land suggest the potential risk of abrupt and irreversible  
33 destabilization, where the marine ice sheets on an inland-sloping bedrock, are adversely impacted by  
34 increasingly warmer temperature and warm ocean current intrusion, all of which may pose an  
35 unexpected threat of increased sea level rise.

## 36 37 1 Introduction

38 A large challenge for rigorous sea-level projection in the 21<sup>st</sup> century is that the dynamics of the  
39 Antarctic ice-sheet is not sufficiently understood under rapidly warming atmosphere and ocean (Church  
40 et al., 2013; Hanna et al., 2013; Joughin et al., 2012; Pritchard et al., 2012). Recent studies on Antarctic  
41 ice-sheet processes since the 1990s using satellite, airborne and *in situ* observations (McMillan et al.,  
42 2014; Rignot and Thomas, 2002; Shepherd et al., 2012; Vaughan et al., 2013), reported increasing



43 present-day ice-sheet changes, such as extensive dynamic thinning on the periphery (Pritchard et al.,  
44 2009), accelerated mass loss (McMillan et al., 2014; Shepherd et al., 2012), and grounding line retreat  
45 in the Amundsen Sea sector, West Antarctica (Rignot et al., 2014), all of which raised the long-standing  
46 concerns on ice-sheet instability (Joughin et al., 2014; Vaughan et al., 2013). Although the new  
47 observations have greatly improved our ability to quantify the changes to the Antarctic ice sheet, it  
48 remains unclear whether East Antarctic ice sheet is losing or gaining mass, especially in the large  
49 marine ice sheets of East Antarctica (Mengel and Levermann, 2014). It is also unclear whether the rate  
50 of Antarctic ice loss/gain has increased/decreased over the last two decades (Stocker et al., 2014).  
51 Furthermore, the underlying drivers of ice-sheet changes remain poorly understood (Alley et al., 2005).  
52 All these limitations make it difficult to determine the future behaviour of the ice sheet. The key to  
53 understanding the Antarctic ice-sheet dynamics is to more accurately determine its mass budget using  
54 extended observations to provide a longer and higher resolution observational record towards improved  
55 understanding of the ice-sheet evolutions, which is crucial for more reliable sea-level projections  
56 (Hanna et al., 2013; Vaughan et al., 2013).

57 Glacier ice flow or velocity, one of the critical ice dynamic parameters affecting the estimates of ice  
58 sheet mass balance and the corresponding sea level rise (Scheuchl et al., 2012), has been measured  
59 by traditional grounded-based measurements (e.g. GPS, electronic distance) since 1970s in the  
60 Antarctic ice sheet. However, the sporadic and discontinuous observations prohibit the studying of  
61 ice sheet mass balance as a whole. It was not until recently that the glaciologists began to present a  
62 complete picture of ice velocity in Antarctica by the use of multi-satellite interferometric synthetic  
63 aperture radar (InSAR) with a long data span (1996–2009) (Rignot et al., 2011). However, such a  
64 snapshot of ice motion of entire Antarctica is insufficient to provide a clear insight of the spatial  
65 and temporal characteristics of ice dynamics. Furthermore, the lack of higher-resolution ice  
66 velocity data limits a thorough investigation on localized ice dynamics (Favier and Pattyn, 2015;  
67 Nath and Vaughan, 2003), such as crevasse production, role of ice rises on the stability of ice sheet,  
68 etc. These limitations highlight the need for a new set of ice velocity observations over Antarctica.

69 Therefore, here we intend to construct two present-day ice flow maps covering the years of  
70 2014 and 2015 for all of the Antarctica inferred from Landsat 8 (L8) images acquired by the  
71 Operational Land Imager (OLI). The velocity data and the existing InSAR-derived ice velocity  
72 (Rignot et al., 2011) can be used to estimate the mass discharges in 2006, 2014, 2015 in  
73 combination with the Bedmap-2 ice thickness data (Fretwell et al., 2013) associated with IPR (Ice  
74 Penetrating Radar) track measurements from the IceBridge project (Allen, 2013, 2011;  
75 Blankenship et al., 2011, 2012). Furthermore, the mass balances of the Antarctic ice sheet can be  
76 estimated by comparing the mass discharges with the latest ice-sheet SMB data derived from a  
77 regional atmospheric climate model (RACMO2.3) (van Wessem et al., 2014), employing the  
78 input-output method (IOM) (Rignot et al., 2008), and the decadal changes can be easily found.

79

## 80 **2 Data and methods**

### 81 **2.1 Data**

82 We collected L8 orthorectified panchromatic bands in 15 m spatial resolution from December 2013  
83 to March 2016 to infer present-day ice velocities of Antarctic ice sheet. The images were acquired by  
84 the Operational Land Imager (OLI) on L8 and are managed by the U.S. Geological Survey (USGS)  
85 Earth Resources Observation and Science (EROS) Center. L8 is the eighth satellite in the Landsat



86 missions, launched on February 11, 2013, which provides a continuous series of land and ice surface  
87 observations with 16-day revisit cycle. The OLI has improved radiometric performance in 12-bit  
88 quantization, which can distinguish subtle contrast variations over bright targets (Fahnestock et al.,  
89 2016; Morfitt et al., 2015), such as Antarctica covered only by snow or ice with high reflectivity.  
90 Rigorous calibration and orbital control contribute to the resulting high-quality visible and infrared  
91 images. The OLI is calibrated to <5% uncertainty in absolute spectral radiance and ~8 m geodetic  
92 accuracies (circular error at 90% confidence (CE 90)) (Zanter, 2016).

93 Compared to the satellite interferometric SAR data, the L8 panchromatic imagery is more suitable to  
94 estimate ice motion in fast-flowing regions for several reasons, (1) the nadir look results in similar  
95 viewing geometry between acquisitions. it can minimize the topographic artifacts, one of main error  
96 sources in SAR/InSAR processing (Mouginot et al., 2012); (2) despite a non-cloud free sensor as  
97 opposed to SAR, L8's 16-day revisit cycle and relative large swaths (185-kilometer), make it possible  
98 to obtain continuous snapshots of ice flow over entire Antarctica, (3) the optical imageries are almost  
99 free of atmospheric effect including ionosphere and tropospheric delays, which may introduce errors in  
100 the interferometric SAR imageries, for example, the ionosphere can produce large ice velocity error up  
101 to 17 m yr<sup>-1</sup> for L-band SAR imagery, (4) the feature tracking method can produce two-dimensional  
102 displacements with same accuracy, while SAR speckle-tracking method has lower accuracy in the  
103 azimuth direction, and differential radar interferometry method only measures one-dimensional  
104 line-of-sight (LOS) displacement.

105 The level 1 Terrain corrected (L1GT) products packaged as Geographic tagged image file formation  
106 (GeoTIFF) in 16-bit grayscale are used to produce the ice velocity of Antarctica. The L1GT products in  
107 Antarctica are terrain orthorectified data using Radarsat Antarctic Mapping Project Digital Elevation  
108 Model Version 2 (RAMP V2 DEM). The geometrically corrected products have minimal distortions  
109 related to the sensor (e.g., view angle effects), satellite (e.g., attitude deviations from nominal), and  
110 Earth (e.g., rotation, curvature, relief). Radiometric corrections were applied to remove relative  
111 detector differences, dark current bias, and some other artifacts. A complete L1GT product consists of  
112 13 files, i.e., the 11 band images, a product specific metadata file, and a Quality Assessment (QA)  
113 image. In our study, only the panchromatic band, specific metadata file and QA band are used. The  
114 specific metadata are used to obtain the cloud ratio as criteria (40%) to pick images for ice velocity  
115 extraction. The QA band is used to identify the spatial distributions of cloud and water, which are  
116 masked in velocity scenes. Based on visual interpretation and cloud cover ratio, a total of more than  
117 10,000 scenes were selected for producing ice velocities over Antarctica. The projection is polar  
118 stereographic with a true latitude of -71°. The reference ellipsoid used is the WGS84 model. In  
119 addition, for comparison of ice flux and mass balance at different periods, InSAR-derived ice velocity  
120 data (450 m resolution) inferred from multiple satellite InSAR data sets are used. The majority of  
121 InSAR data used are during 2007–2009, but with the data in the grounding lines acquired mostly in  
122 2006 (Rignot et al., 2008; Rignot et al., 2011).

123 In order to assess the accuracy of our ice velocity results, we also collected *in-situ* measurements



124 (Brecher, 1982; Frezzotti et al., 1998; Manson et al., 2000; Naruse, 1979; Rott et al., 1998; Skvarca et  
125 al., 1999; Zhang et al., 2008), compiled and managed by the National Snow & Ice Data Center  
126 (NSIDC). The *in-situ* measurements of ice velocity were obtained from a variety of methods such as  
127 differential GPS, electronic distance measurement and triangulation chain survey. The *in-situ* data in  
128 the Lambert-Amery basin were obtained mainly from 1988 to 2008, Siple Coast from 1984 to 1998,  
129 and Mizuho Plateau of Queen Maud Land from 1969 to 1978. Note that we only collected *in-situ*  
130 measurements in the slow-flow regions where ice velocities are less than  $100 \text{ m yr}^{-1}$  and thus assumed  
131 to have no significant secular changes.

## 132 2.2 Feature tracking method

133 To determine the horizontal displacement field of ice motion, we use feature tracking method  
134 (Bindschadler and Scambos, 1991; Leprince et al., 2007; Scambos et al., 1992), also called as the phase  
135 shift method. Since the input images are orthorectified, correlation can be directly implemented using  
136 the phase shift technique of low frequency calculated by Fourier-based frequency correlator (Leprince  
137 et al., 2007), which is produced within a specific sliding window (or patch) on the paired images  
138 repetitively. The result is given by a three-band file consisting the E/W displacement map (positive  
139 toward the East), the N/S displacement map (positive toward the North), and the SNR band as an  
140 indicator of the quality of the measurement. The technique enables us to resolve sub-pixel  
141 displacements of less than  $1/20$  of the pixel resolution at a high signal-to-noise ratio (SNR), which is  
142 generally greater than 0.9. All processes are performed using the COSI-Corr (Co-registration of  
143 Optically Sensed Images and Correlation) software package developed at the California Institute of  
144 Technology (Leprince et al., 2007).

145 The feature tracking is implemented in a two-step process. The first step (namely coarse correlation)  
146 is to roughly estimate the pixelwise displacement between two patches. In general, if noisy images or  
147 large displacements are expected, a larger initial sliding window should be used. In this study, the size  
148 of initial sliding window varies from 64 to 256 in pixels in both X and Y directions according to the  
149 prior knowledge of InSAR-derived Antarctic ice velocity, and the time interval between two paired  
150 images. Once the initial displacements are estimated, the final correlation (namely fine correlation) step  
151 is to retrieve the subpixel displacement by using smaller window. The new size of  $32 \times 32$  pixels is  
152 tentatively adopted in order to yield reliable estimates for the displacement at densely independent  
153 points. Other parameters of frequency correlator include the step sizes between sliding windows in both  
154 X and Y directions (in pixels), frequency masking threshold, the number of iterations for robustness,  
155 resampling and gridded output. The step size is set to be a constant value of 7 pixels in each dimension,  
156 which means that output product has 100-meter resolution. The frequency masking threshold of 0.9 is  
157 adopted as an optimum value as recommended in a previous study (Leprince et al., 2007).

## 158 2.3 Quality control for displacement vector

159 During the co-registration step, the Fourier frequency correlator is used in correlation estimates. The  
160 technique is more accurate compared with a statistical method; however, it is more sensitive to noise.  
161 High-performance L8 images can minimize the effect, but decorrelation still exists due to large ground  
162 motion, lack of measureable ground features (such as crevasses, or rise), sensor noise, and topographic  
163 artifact (thereby producing imprecise orthorectified data).



164 To overcome these problems, we devise three steps to enhance the signal and exclude unreliable  
165 measurements. First, we suppress the noise on each displacement scene by using an adaptive filter and  
166 a median filter respectively. The adaptive filter is the local sigma filter (Eliason and McEwen, 1990).  
167 The filter size is 9 pixels and sigma factor is 2. A median filter is further applied to remove “salt and  
168 pepper” noise in ice displacement scene. Second, the areas covered by cloud and water are excluded  
169 from the displacement scenes using the QA band (Zanter, 2016). In the QA band, each pixel contains  
170 16-bit integer that represents bit-packed combinations of surface, atmosphere, and sensor conditions at  
171 different confidence level. The pixels covered by cloud and water in paired images are unpacked from  
172 the QA band using our developed procedures, the pixels marked by cloud and water at high confidence  
173 (67–100%) are used to build a mask layer, and then they are masked from displacement scenes. It  
174 should be noted that the identification of cirrus is problematic in raw images based on our analysis;  
175 snow and ice are easily considered as cirrus. Here, we only use cloud to build mask layer. Third, since  
176 the frequency correlation easily gives rise to errors at the edges of displacement image, the pixels are  
177 also masked.

#### 178 **2.4 Ice velocity extraction**

179 The cloud contamination is a main challenge in ice flow generation using optical images. In order to  
180 overcome the problem, we process all image pairs using a one-year time intervals as time baseline with  
181 the minimum repeat cycle of 16 days in Worldwide Reference System (WRS-2). Some adjacent paths  
182 in WRS-2 are also paired to produce ice velocity for some void areas where there are no valid scenes  
183 for pairing in same path and row. The one-year time interval is derived from our correlation  
184 experiments. When the time interval is more than one year, the decorrelation may appear due to large  
185 surface motion or geomorphic change. Finally, 10,690 image pairs are selected from more than 10,000  
186 scenes of L8 panchromatic images, and processed for the production of ice velocity.

187 Despite of the improvement in geometric accuracy, the residual geolocation errors with L8  
188 panchromatic band still exists (~8m) in CE90, the errors could cause offsets between the displacement  
189 scenes which should be removed (Fahnestock et al., 2016). In fact, the offset tuning is often called  
190 absolute calibration of the ice velocity data. In Antarctica, absolute calibration is a challenging issue  
191 because the ice is active almost in everywhere and available rock outcrops are extremely scarce. Here,  
192 we use the InSAR-derived Antarctic velocity map to determine the relatively stagnant areas (the  
193 magnitude of ice velocity  $<5 \text{ m yr}^{-1}$ ) for absolute calibration of our ice velocity estimates.

194 There are three steps for the velocity calibration. First, the differences of the displacements between  
195 InSAR-derived velocity map and our displacements are calculated in the stagnant areas. Second, to  
196 eliminate outliers, a three-sigma filter is applied recursively on these differences in which the  
197 differences will be omitted if the magnitudes of the values are larger than three times the standard  
198 deviation ( $3\sigma$ ). Third, the mean of the rest differences is considered as the offset of displacement  
199 scenes. Furthermore, the offsets for the displacement scenes outside of stagnant areas (such as in the  
200 Ross and Ronne ice shelves) are estimated by overlapped neighboring scenes at nearly the same time  
201 periods. The two velocity components are independently estimated and rectified.

202 The mosaicked velocity map is produced on the basis of processed displacement scenes as above. To



203 increase the accuracy of mosaicked velocity map, we stack all displacement scenes after the pixels with  
204 SNR < 0.9 are masked. In general, the velocity map contains 8–10 scenes in a specific location. For a  
205 specific pixel denoted by  $i$ , all displacement scenes ( $m=1, 2, \dots, n$ ) are stacked to give the estimate of  
206 the ice velocity ( $V_i$ ) as follows,

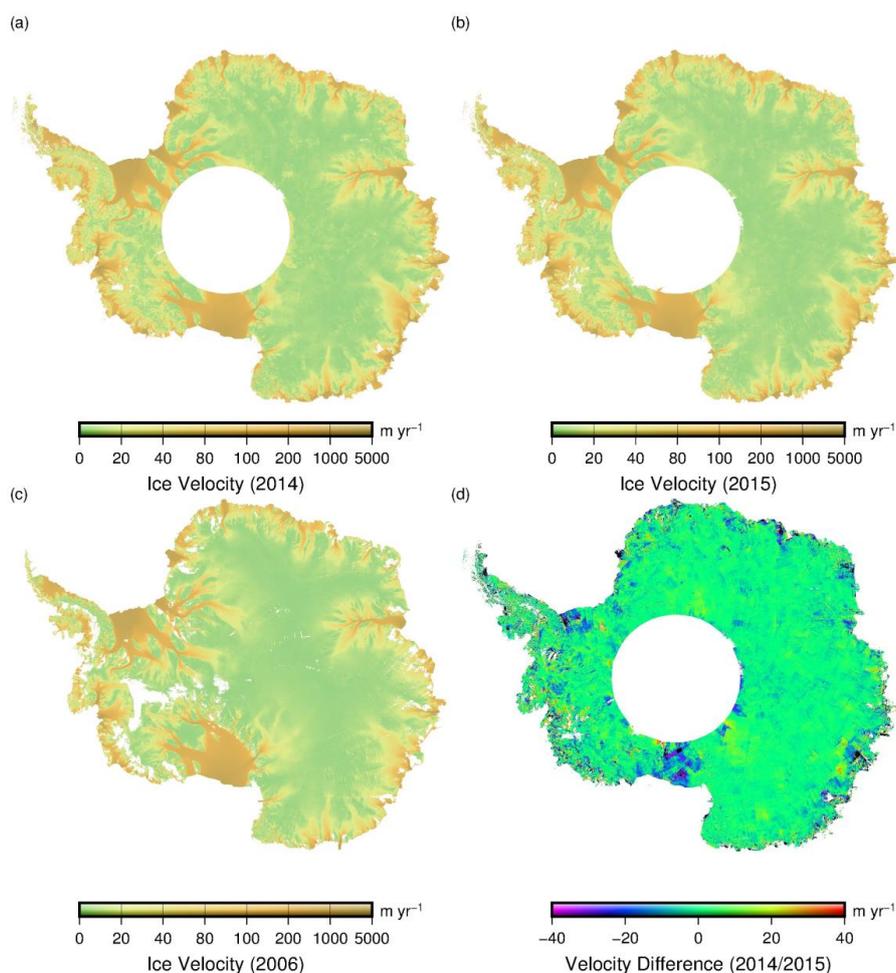
$$V_i = \frac{\sum_{m=1}^n \Delta d_m^i}{\sum_{m=1}^n \Delta t_m^i} \quad (1)$$

209 where  $\Delta d_m^i$  denotes the generated displacement during the time interval  $\Delta t_m^i$ .

### 210 3 Results of ice velocity

#### 211 3.1 Ice velocity maps

212 In Antarctica, the valuable L8 images are available just in summer (November, December, January,  
213 February and March). Due to the short observational span at the end of 2013 and at the beginning of  
214 2016, it is difficult to produce the individual mosaic for the entire Antarctica, thus the images acquired  
215 in the two years are used to produce 2014 mosaic and 2015 mosaic respectively. In Figure 1, we show  
216 two mosaicked ice velocity maps for 2014 (Fig. 1a) and 2015 (Fig. 1b), respectively for Antarctica. Ice  
217 velocity differences between the two maps are usually very small relative to the magnitudes of the  
218 velocities since the mean and standard deviation are  $0.17 \text{ m yr}^{-1}$  and  $7.6 \text{ m yr}^{-1}$  (Fig. 1d). The  
219 InSAR-derived ice velocity data are also shown (Fig. 1c), in which the data at grounding lines used for  
220 the analysis of glacier discharge changes were derived from the SAR images in about 2006 (Rignot et  
221 al., 2011). Our results exhibit a similar pattern in the ice flow field compared with a previous  
222 InSAR-based study over a long time span (Rignot et al., 2011). The spatial resolution of our velocity  
223 maps is 100-m, which is 4 times higher than the InSAR-derived ice flow map. Our two ice velocity  
224 maps thus provide an opportunity to investigate localized ice dynamics, such as crevasse production,  
225 and the roles of ice rise and rumples on ice-sheet dynamics and evolution. They also have a better  
226 coverage over Antarctica except for the south of 82.5S. The two mosaicked ice velocity maps cover the  
227 majority of the ice sheet and nearly 99% of fast-flow glaciers and ice shelves, and fast ice, except for a  
228 few ice streams of Ronne ice shelf (e.g. Academy, Foundation glaciers) and Ross ice shelf (e.g.  
229 Whillans glacier in Siple Coast). In addition, in order to be computationally efficient, the entire  
230 Antarctica is divided into 11 sub-regions, and data stacking is processed independently, then 11 sub  
231 regions (Fig. 7) are mosaicked to generate an ice velocity map for the entire Antarctica.



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**Figure 1.** L8-derived '2014' ice flow in a) from December, 2013 to December, 2014, and L8-derived '2015' ice flow in b) from January, 2015 to March, 2016, and InSAR-derived ice velocity in c) from 1996 to 2009; the difference of ice flow between '2015' and '2014' in d). L8-derived ice velocity maps are drawn with a resolution of 500m.

### 3.2 Uncertainty analysis

The uncertainty of ice velocity maps derived from the L8 data primarily resulted from mis-registration, the time interval of the pairs used to extract displacement, and the amount of stacking data. The mis-registration is mainly caused by three error sources: (1) decorrelation due to severe ground change, lack of measureable features between the scenes due to long time interval or single land cover (e.g. snow or ice); (2) low image quality caused by sensor noise, pixel oversaturation, aliasing and cloud contamination; (3) topographic artifacts primarily due to shadowing differences and imprecise orthorectification of satellite attitudes. The co-registration accuracy is conservatively set to be 1/25 of the pixel size in E/W and N/S displacement components, which is larger than 1/50 of the pixel size



247 prompted by Leprince et al. (2007) (Leprince et al., 2007). Using the co-registration error together with  
248 the total amount of stacking data, and time interval between two acquisitions, the ice velocity error per  
249 year can be calculated on the basis of error propagation law.

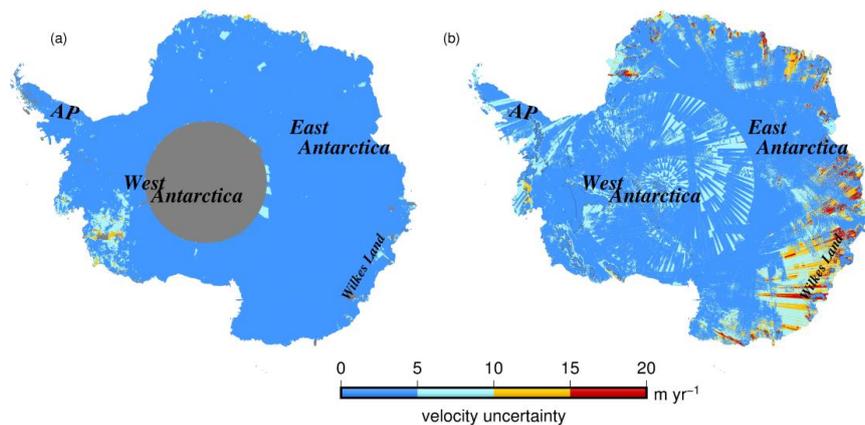
250 According to the mosaicking method as mentioned above (Eq. 1), the uncertainty of one mosaicked  
251 velocity component at i-th pixel denoted by  $\sigma_{v_i}$  can be estimated using the following error  
252 propagation formula under the assumption that the errors of different sources are independent:

$$\sigma_{v_i} = \pm \frac{\sqrt{\sum_{m=1}^n (\sigma_m^i)^2}}{\left(\sum_{m=1}^n \Delta t_m^i\right)^2} \quad (2)$$

253  
254 Where  $\sigma_m^i$  is the co-registration error, i.e., the standard deviation of m-th displacement observation  
255 during time interval of  $\Delta t_m^i$ . Since the co-registration errors are constant in space (the whole scene)  
256 and time domain (all stacked displacements), if the  $\sigma_m^i$  is denoted by a constant of  $\sigma$ , Eq. 2 can be  
257 simplified as follows,

$$\sigma_{v_i} = \pm \sqrt{n} \frac{\sigma}{\sum_{m=1}^n \Delta t_m^i} \quad (3)$$

258  
259 The uncertainty of a mosaicked velocity map is dependent on the amount of stacking data and the time  
260 intervals during the velocity stacking. That means that the larger the time span, the higher resulting ice  
261 velocity accuracy. Since the E/W and N/S components at i-th pixel have the same uncertainty, the  
262 uncertainty as calculated with Eq. 3 is actually valid for the magnitude of the velocity vector. The error  
263 of magnitude of mosaicked velocity vector with magnitudes of 0-20 m yr<sup>-1</sup> is shown in Figure 2a. For  
264 comparison, the uncertainty of InSAR ice velocity is also shown in Figure 2b.

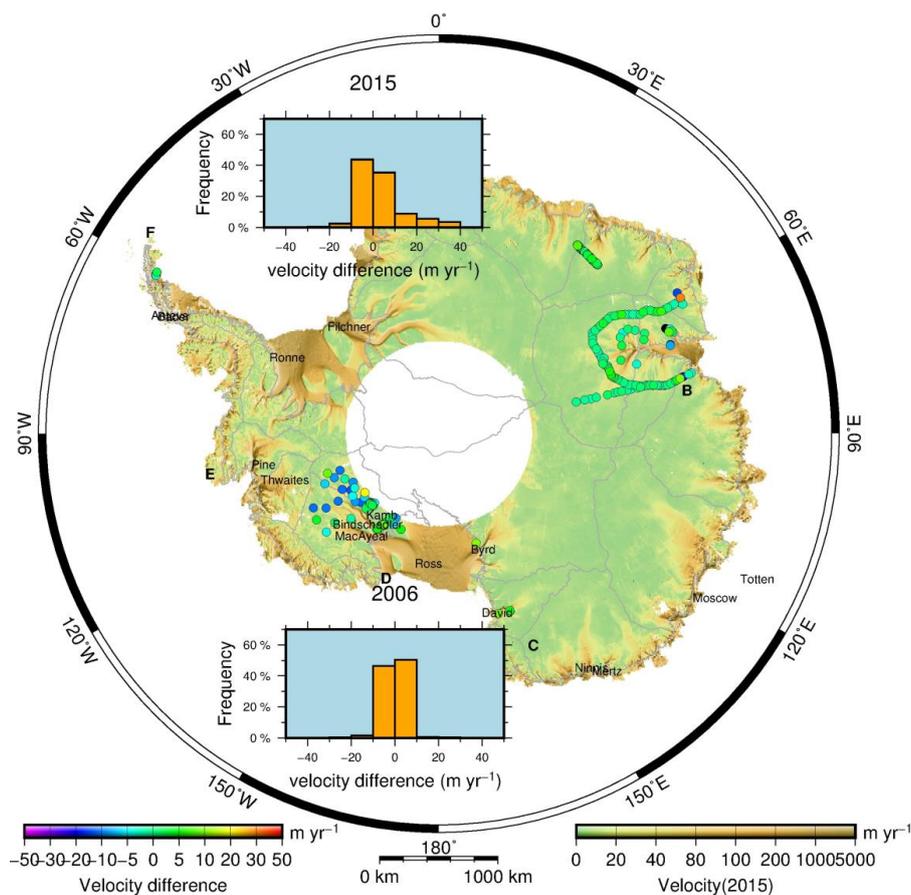


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**Figure 2.** The uncertainty maps of L8-derived Antarctic ice velocity in 2015 (a) and InSAR-derived ice velocity (b).

### 3.3 Comparison with *in-situ* measurements

Our ice velocity results are only compared with the *in-situ* measurements in the slow-flow areas ( $<100$   $\text{m yr}^{-1}$ ). The 538 sites chosen for the comparison are shown by the dots of Figure 3, and the differences are shown by the color dots. From the upper inset, the differences are usually  $<10$   $\text{m yr}^{-1}$  and the average of the difference is  $3$   $\text{m yr}^{-1}$  with a standard deviation of  $10$   $\text{m yr}^{-1}$ . For comparison, the differences between InSAR velocity and field surveying data are also shown in lower inset in Figure 3. The average of the difference is  $0.3$   $\text{m yr}^{-1}$  with a standard deviation of  $4.2$   $\text{m yr}^{-1}$ . The differences in accuracy performance may result from the measurement errors and different time spans of surveys.



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278 **Figure 3.** The comparison between L8-derived ice velocities in 2015 and data from *in-situ*  
 279 measurements. The colored dots show the differences between L8 ice velocity in 2015 and *in-situ*  
 280 survey data. Upper inset shows the histogram of differences between L8-derived ice velocity and field  
 281 data, and lower inset also shows the same but for InSAR-derived ice velocity.

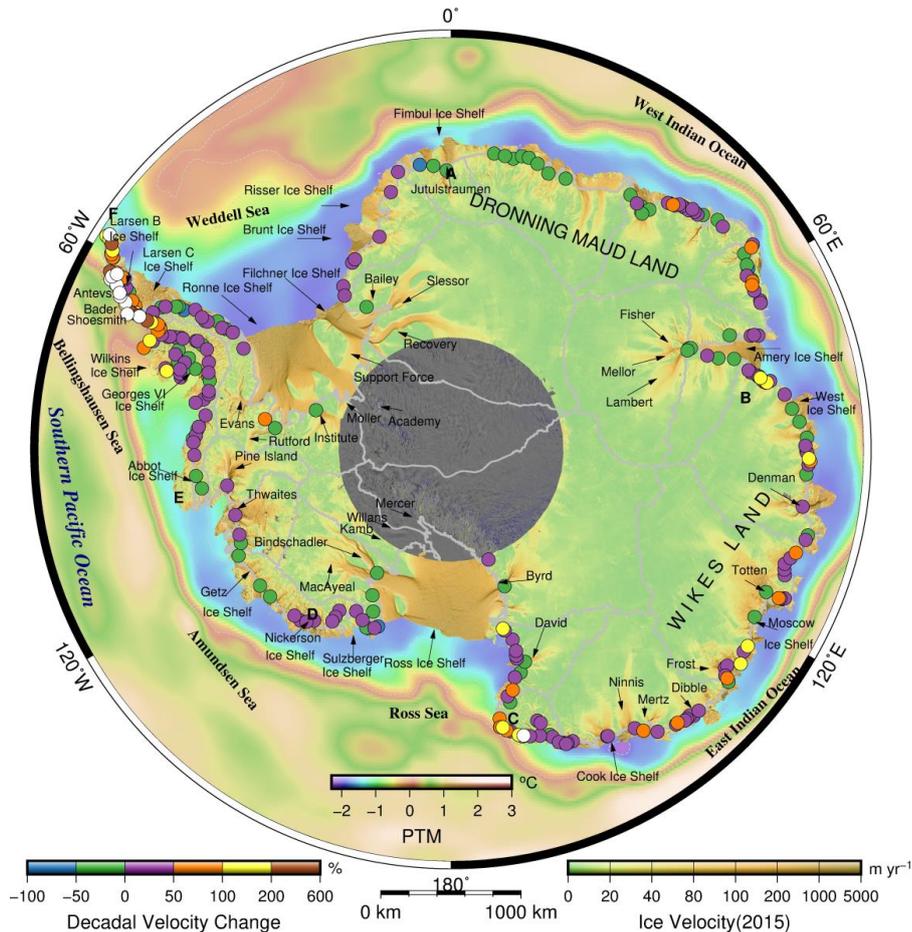
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#### 283 4 Decadal glacier dynamics

284 We investigated the decadal evolution of ice dynamics of 465 glaciers, nearly all of the glaciers in  
 285 Antarctica, based on our estimate of high-resolution ice velocity maps in 2015 and an InSAR-derived  
 286 map in 2006 (Rignot et al., 2011) (Table S1, supplementary materials). 218 glaciers were found to be  
 287 accelerating, and only 82 glaciers underwent decelerations at a high confidence level ( $2\sigma$ ) (Fig. 4).  
 288 We found significant outlet glacier accelerations ( $>50\%$  in velocity change, same hereafter) over much  
 289 of the Antarctic Peninsula (AP), nearly 30% in Ellsworth Land in West Antarctica, and approximately  
 290 25% in the Victoria Land and the Wilkes Land in East Antarctica. In contrast, glacier decelerations  
 291 were found at rates of 20–40% in the Dronning Maud Land, and at 3–20% for the glaciers in the three  
 292 largest ice-shelf systems (Filchner-Ronne, Ross, and Amery). In particular, majority of the glaciers  
 293 accelerated by more than 200% in the northern part of the Western AP (WAP) along the Bellingshausen  
 294 Sea coast, resulting from the intrusion of the warmer Circumpolar Deep Water (CDW) (Martinson et al.,



295 2008; Smith et al., 1999) and increasing air temperature (Vaughan et al., 2003). The acceleration in  
 296 WAP is more significant than that over the period 1993-2003 (Pritchard and Vaughan, 2007). In the  
 297 Weddell Sea sector, velocities of most of the glaciers in the Eastern AP (EAP) and Coats Land in East  
 298 Antarctica evidently accelerated by 5–50%, whereas the glaciers draining into the Filchner and Ronne  
 299 ice shelves exhibited deceleration. We found complicated decadal variations of glacier dynamics in  
 300 East Antarctica. In the West Indian Ocean sector, Dronning Maud Land glacier velocities exhibited an  
 301 overall deceleration, whereas its adjacent region, Enderby Land, exhibited acceleration by 5–30%.  
 302 However, in the East Indian Ocean sector, most glaciers accelerated by ~25%. In the Ross Sea sector,  
 303 the glaciers in Victoria Land accelerated widely by ~20%, whereas much of the glaciers draining into  
 304 the Ross ice shelf decelerated, especially for the fast-flowing large Byrd and Mulock glaciers in the  
 305 Transantarctic Mountains, Bindschadler and other unnamed glaciers in the Siple Coast. In the  
 306 Amundsen Sea sector, although the Pine Island and Thwaites glaciers evidently accelerated by ~15%,  
 307 many of the remaining glaciers draining into the Getz ice shelf decelerated. The details of ice dynamics  
 308 on individual glaciers can be found in the supplementary materials.



309  
 310 **Figure 4.** Ice velocity in 2015 and decadal velocity change in Antarctica. The mosaic of the Antarctic  
 311 ice velocity (2015) derived from L8 panchromatic images from January 2015 to March 2016 is shown



312 here overlaid on a MODIS mosaic of Antarctica (MOA)(Bohlander and Scambos, 2007). The  
313 magnitude of ice velocity is coloured on a logarithmic scale and overlaid on gridded potential  
314 temperature data of seawater (PTM) at a depth of 200 m from the World Ocean Circulation Experiment  
315 (WOCE). The white-filled dots show that the velocity changes are larger than 600%. The velocity  
316 changes on grounding lines are calculated for 465 glaciers between 2015 and 2006, and shown for the  
317 300 glaciers with high confidence levels ( $>2\sigma$ ) coloured on a logarithmic scale. The names of  
318 selected glaciers and ice shelves are labelled. ‘A’ through ‘F’ delimits the six oceanic sectors. The  
319 details of ice velocity changes on grounding lines are presented in Table S1. The solid grey lines  
320 delineate major ice divides.

### 321 5 Decadal variations of mass discharge and mass balance

322 We use ice flow measurements for 2014 and 2015 and the existing InSAR results for 2006 to infer  
323 the corresponding Antarctic ice sheet losses at the drainage basin scale (Zwally et al., 2012) in  
324 combination with Bedmap-2 ice thickness data (Fretwell et al., 2013) and ice penetrating radar (IPR)  
325 thickness from multiple campaigns from 2002 to 2014 from the IceBridge project (Allen, 2013, 2011;  
326 Blankenship et al., 2011, 2012) (see supplementary materials). We compare the ice sheet discharge with  
327 the new surface mass balance (SMB) data (1979–2014) (van Wessem et al., 2014) to estimate the  
328 Antarctic mass balance using input-output method (Rignot et al., 2008). The mass discharges across the  
329 Antarctic grounding lines (Depoorter et al., 2013) are derived from the flux gate method (Rignot et al.,  
330 2013) using a developed procedure (see supplementary materials). Here, we calculate the ice-sheet  
331 inflow mass using the new SMB data at a horizontal resolution of 27.5 km resulting from the updated  
332 regional Atmospheric Climate Model RACMO2.3 on the 27 glacier drainage basins (Zwally et al.,  
333 2012) (Table S3). Figure 5 shows the mass discharge, mass balance and their changes between 2015  
334 and 2006 covering the entire Antarctic ice sheet. The total mass balance estimates of the Antarctic ice  
335 sheet under the constant accumulation rate (Monaghan, 2006) during the survey period were  $-181\pm 68$   
336  $\text{Gt yr}^{-1}$ ,  $-232\pm 60 \text{ Gt yr}^{-1}$ , and  $-230\pm 60 \text{ Gt yr}^{-1}$  in 2006, 2014 and 2015, respectively (Table 1, S2). These  
337 results are comparable with the latest results inferred from GRACE (Williams et al., 2014) and  
338 Cryosat-2 (McMillan et al., 2014) data, and consistent with recent InSAR mass balance estimates in  
339 2006 (Rignot, 2008). However, our estimated rates are larger than the previous results obtained using  
340 ICESat altimetry data (Shepherd et al., 2012). Table S4 shows detailed estimates of mass balance using  
341 altimetry, gravimetry, and IOM method in the last several decades. The Amundsen Sea sector had the  
342 largest imbalance of  $-212\pm 24 \text{ Gt yr}^{-1}$  in 2015 (similar to previous studies (Vaughan et al., 2013)),  
343 accounting for nearly the total imbalance ( $-230\pm 60 \text{ Gt yr}^{-1}$ ) of the entire Antarctic ice sheet. Besides  
344 the Amundsen Sea sector, another significant negative imbalance ( $-78\pm 32 \text{ Gt yr}^{-1}$ ) was observed in the  
345 East Indian Ocean sector of East Antarctica, whereas the West Indian Ocean sector exhibited an  
346 obvious positive mass balance ( $64\pm 29 \text{ Gt yr}^{-1}$ ). The Ross Sea sector exhibited slight mass gain, whereas  
347 the Weddell and the Bellingshausen Sea sectors exhibited no significant mass changes. However, the  
348 mass balance estimates in the Bellingshausen Sea sectors are most likely underestimated owing to the  
349 summer meltwater not being considered (see supplement materials). Ice-sheet mass budgets are also  
350 shown in Table 2 for East Antarctica, West Antarctica and the Antarctic Peninsula. On the East  
351 Antarctic ice sheet, no significant mass change occurred ( $11\pm 86 \text{ Gt yr}^{-1}$ ), similar to recent estimates  
352 (Shepherd et al., 2012), because the net mass deficit in the East Indian Ocean sector was compensated  
353 by the mass gain in the West Indian Ocean sector. In West Antarctica, the total mass balance was  $-$   
354  $274\pm 41 \text{ Gt yr}^{-1}$  in 2015, which is larger than recent altimetry estimates ( $-134 \text{ Gt yr}^{-1}$ ) from 2010 to  
355 2013 (McMillan et al., 2014) and much larger than recently reconciled estimates ( $-65\pm 26 \text{ Gt yr}^{-1}$ ) from



356 2003 to 2009 (Shepherd et al., 2012). In the Antarctic Peninsula, there was a positive mass balance  
 357 ( $33\pm 21$  Gt yr<sup>-1</sup>) in 2015, contrary to previously studies (Rignot et al., 2008; Shepherd et al., 2012),  
 358 probably due to a larger estimate of snow accumulation rate from the new high-resolution (5.5 km)  
 359 SMB data (van Wessem et al., 2016), and the summer meltwater not included (see supplement  
 360 materials).

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**Table 1.** Mass budgets for the six oceanic sectors of the Antarctic ice sheet

Oceanic Sector	Area	SMB	GLF(2006)	GLF(2014)	GLF(2015)	Net(2006)	Net(2014)	Net(2015)	GLL
	km <sup>2</sup>	Gt yr <sup>-1</sup>	km						
Ross Sea (ROS)	2763447	191±12	171±11	171±6	171±6	20±16	20±13	20±13	8334
Amundsen Sea (AMU)	590119	319±24	524±5	532±5	531±5	-205±24	-213±24	-212±24	4481
Bellingshausen Sea (BEL)	206768	221±11	226±4	232±3	229±4	-5±11	-11±11	-8±11	5295
Weddell Sea (WED)	3240372	393±26	396±23	406±11	407±11	-3±34	-13±28	-14±28	12138
West Indian Ocean (WIS)	2544605	267±29	216±13	210±7	203±6	51±31	57±29	64±29	7978
East Indian Ocean (EIS)	2549133	511±32	549±23	581±6	589±4	-38±39	-70±32	-78±32	7213
<b>Total in Antarctica</b>	<b>11894445</b>	<b>1901±58</b>	<b>2082±37</b>	<b>2133±16</b>	<b>2131±16</b>	<b>-181±68</b>	<b>-232±60</b>	<b>-230±60</b>	<b>45439</b>

363 The glacier mass discharge or grounding-line flux is denoted by ‘GLF’, the mass balance by ‘Net’ is  
 364 SMB minus GLF, and grounding line length by ‘GLL’. The results for 2014 are given for the period  
 365 from December 2013 to December 2014, and 2015 from January 2015 to March 2016. The ice-sheet  
 366 area (Area) excludes ice rises and islands, which isolate the main ice sheet. The details about the  
 367 glacier’s affiliation to the six oceanic sectors can be found in the supplementary materials.

368  
 369

**Table 2.** Ice-sheet mass budgets of East Antarctica, West Antarctica and the Antarctic Peninsula

Sector	Area	SMB	GLF(2006)	GLF(2014)	GLF(2015)	Net(2006)	Net(2014)	Net(2015)	GLL
	km <sup>2</sup>	Gt yr <sup>-1</sup>	km						
East Antarctica	9915811	1106±84	1066±52	1097±21	1095±19	40±98	9±86	11±86	25697
West Antarctica	1747718	578±39	846±21	856±13	852±15	-268±44	-278±41	-274±41	12130
Antarctic Peninsula	230916	217±12	170±6	179±3	184±3	47±13	38±12	33±12	7612
<b>Total in Antarctica</b>	<b>11894445</b>	<b>1901±93</b>	<b>2082±56</b>	<b>2133±25</b>	<b>2131±24</b>	<b>-181±108</b>	<b>-232±96</b>	<b>-230±96</b>	<b>45439</b>

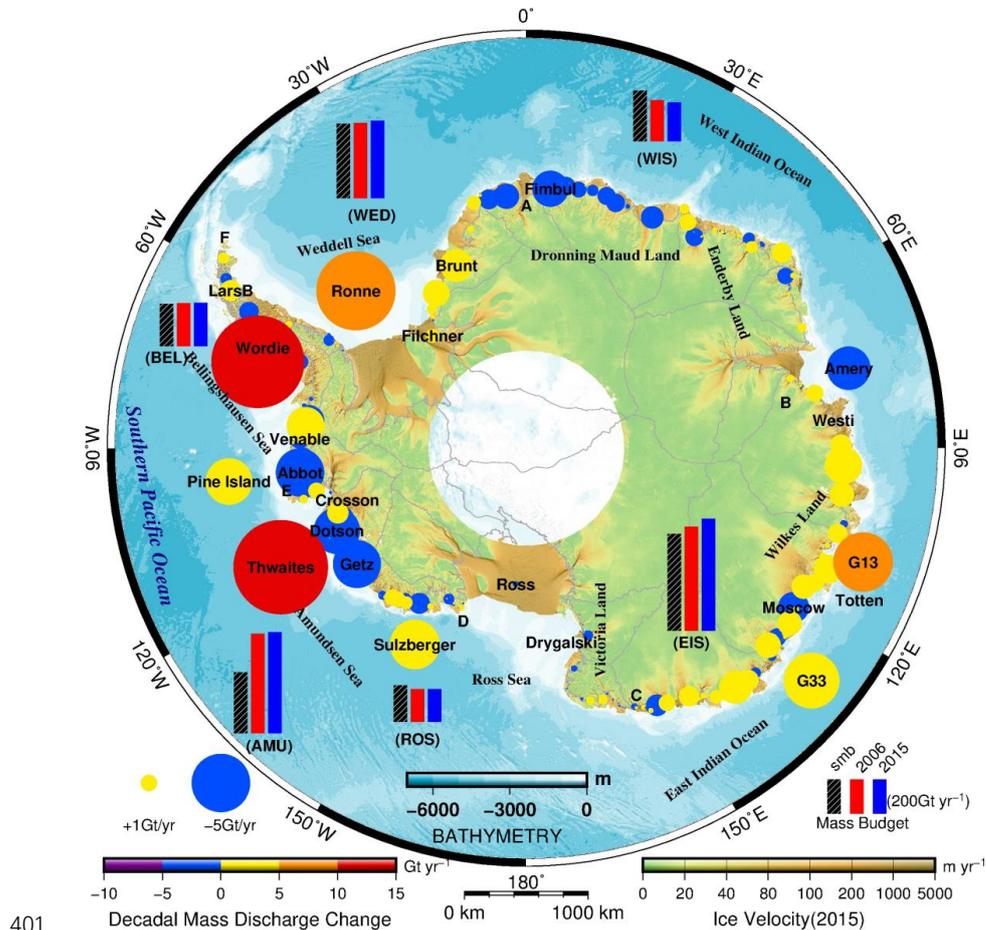
370  
 371

We further analyzed the decadal change of mass balance in the Antarctic ice sheet from 2006 to 2015  
 (Fig. 5). The mass balance decreased by 27% during the last decade to reach a rate of  $-230\pm 60$  Gt yr<sup>-1</sup>  
 in 2015, compared with  $-181\pm 68$  in 2006. The change of mass balance ( $-49\pm 90$  Gt yr<sup>-1</sup>) is not  
 significant in comparison to its large uncertainty caused mainly by SMB. The most significant change  
 of mass balance occurred in East Indian Ocean, reaching  $-40\pm 50$  Gt yr<sup>-1</sup>. We found an increased mass  
 discharge from East Indian Ocean sector in the last decade by up to  $40\pm 24$  Gt yr<sup>-1</sup>, attributed to  
 unexpected widespread accelerating glaciers in Wilkes Land, East Antarctica. The underlying cause for  
 this accelerated mass discharge is most likely linked to the incursion of warm CDW towards glacier  
 termini and a reduction in sea ice (Miles et al., 2016). In Wilkes Land, the large accelerated mass  
 discharge, together with anomalous glacier retreat (Miles et al., 2016), the contemporary thinning along  
 its margins (Pritchard et al., 2012) and unstable inland-sloping bedrock topography, suggests potential  
 instability of the marine ice sheet in a warmer temperature and warm ocean current environments (Gille,  
 2002; Vaughan et al., 2013). These results are inconsistent with the previously documented persistent  
 state for the last 14 Myr (Aitken et al., 2016). The Aurora Subglacial Basin (ASB) in the western

384



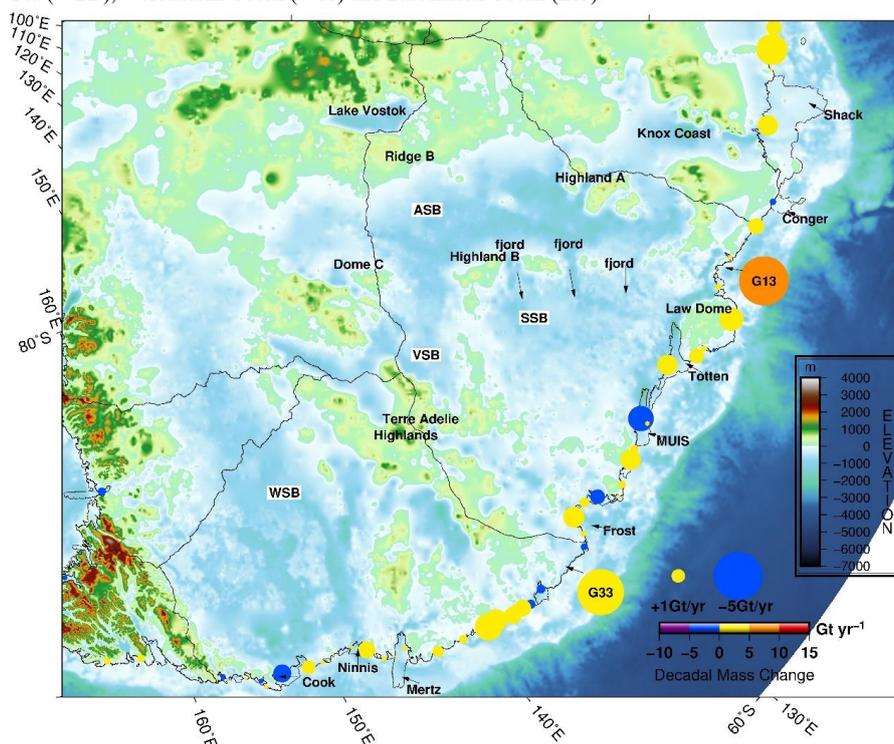
385 Wilkes Land is located to the northeast of elevated Dome A and Ridge B on the Antarctic ice sheet (Fig.  
 386 6). The ASB is overlain by 2–4.5 km of ice, which holds an ice mass equivalent to 9 m of sea level rise.  
 387 IPR data have identified a series of deep topographic troughs (more than 1 km below sea-level) within  
 388 a mountain block landscape oriented nearly orthogonal to the modern margins (Young et al., 2011). The  
 389 accelerated mass discharge at the margins of the ice sheet may trigger instability of the upstream ice  
 390 sheet (e.g., ASB), which has happened many times throughout the paleo-climate era and has  
 391 significantly contributed to sea level changes (Young et al., 2011). In the Wilkes Subglacial Basin  
 392 (WSB) of the Eastern Wilkes Land, which holds marine ice equivalent to 19 m of sea-level rise  
 393 (Mengel and Levermann, 2014), the marginal glaciers (e.g., Cook, Ninnis) which function as an ice  
 394 plug supporting the marine ice sheet of WSB also exhibit obviously accelerated mass discharge. In  
 395 contrast, the other five sectors exhibit no significant mass discharge changes. Interestingly, in Pine  
 396 Island and the Thwaites catchment (basins 21, 22), West Antarctica, and the Antarctic Peninsula, the  
 397 accelerated mass discharges are observed to be  $13\pm 1$  Gt yr<sup>-1</sup>,  $10\pm 25$  Gt yr<sup>-1</sup> and  $14\pm 7$  Gt yr<sup>-1</sup>,  
 398 respectively, in the past 10 years, which are obviously less than the previous estimates of  $46\pm 5$  Gt yr<sup>-1</sup>,  
 399  $46\pm 23$  Gt yr<sup>-1</sup> and  $29\pm 13$  Gt yr<sup>-1</sup> in 1996-2006 (Rignot et al., 2008). However, the underlying causes are  
 400 unclear in these regions.



401



402 **Figure 5.** Changes of mass discharges and mass balances over Antarctic ice sheet between 2006 and  
 403 2015. The colour and size of the circles denote the magnitudes of the decadal mass discharge changes  
 404 for individual glaciers with no ice-shelf linked and for the combinations of the glaciers linked the same  
 405 ice shelves. Note that the circles are drawn in variable size scales for clarity. The details about the  
 406 glaciers can be found in Table S2. In addition, the SMB, the mass discharges in six oceanic sectors in  
 407 2015 and 2006 are denoted by black-hatched and coloured bars in six oceanic sectors. The mosaic of  
 408 ice velocity in 2015 and ice divides, as in Figure 4, and an overlain bathymetry map are shown. The six  
 409 oceanic sectors include Ross Sea (ROS), Amundsen Sea (AMU), Bellingshausen Sea (BEL), Weddell  
 410 Sea (WED), West Indian Ocean (WIS) and East Indian Ocean (EIS).



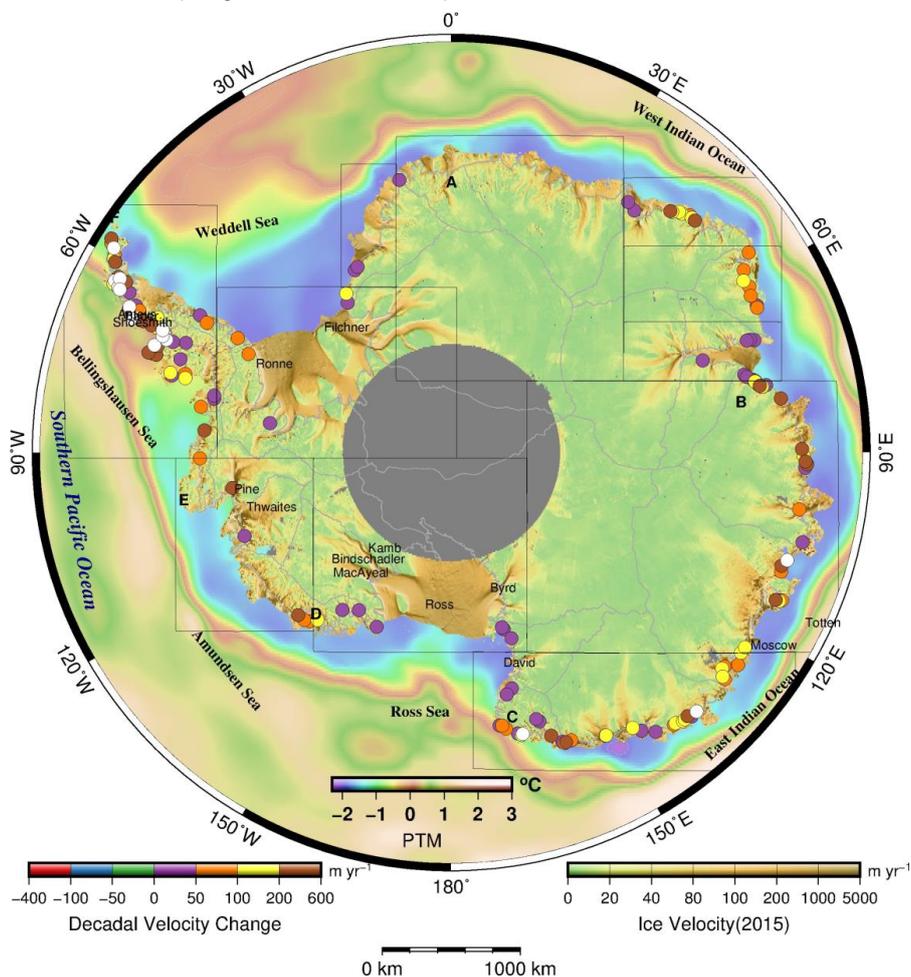
411 **Figure 6.** Bed topography of the Wilkes Land, East Antarctica. Color circles show the mass balance  
 412 changes between 2015 and 2006 for individual glaciers with no ice-shelf linked and for glacier  
 413 combinations in view of the same linked ice shelves. ASB: Aurora Subglacial Basin, VSB: Vincennes  
 414 Subglacial Basin, VST: Vanderford Subglacial Trench, SSB: Sabrina Subglacial Basin, WSB: Wilkes  
 415 Subglacial Basin.

## 417 6 Simultaneous acceleration of ice shelves and glaciers

418 The velocity changes of ice shelves are also investigated to reveal their underlying relationships with  
 419 linked glaciers, since the Antarctic ice shelves can be seen as the ice plugs of their bounded ice-sheets  
 420 and tributary ice-streams, effectively deterring their retreat or abrupt disintegration (Mengel and  
 421 Levermann, 2014). 204 of the surveyed ice shelves (Table S1, Fig. 7) were found to accelerate mainly  
 422 in Wilkes Land in the East Indian Ocean sector, Enderby Land in the West Indian Ocean sector, and  
 423 WAP in the Bellingshausen Sea sector. This result suggests that acceleration of ice shelves is a possible



424 cause of the fast flow of glaciers. Especially in Wilkes Land, the glaciers and corresponding ice shelves  
 425 exhibited nearly simultaneous acceleration. This acceleration further enhances the concerns of the  
 426 instability of the marine ice sheets in Wilkes Land, East Antarctica. The marine ice sheets in Wilkes  
 427 Land hold ice equivalent to more than 28-m of global sea-level rise, which is more than six times that  
 428 of West Antarctica (Mengel and Levermann, 2014).



429  
 430 **Figure 7.** The velocity change of the Antarctic ice shelves between 2015 and 2006. The color dots  
 431 show the velocity changes of Antarctic ice shelves. The white dots show the changes of ice-shelf  
 432 velocity are larger than 600 m yr<sup>-1</sup>. However, in western Antarctic Peninsula, the ice velocity changes  
 433 are shown mostly for glaciers. The mosaic of present ice velocity for 2015 and a gridded potential  
 434 temperature data of seawater (PTM) at 200 m depth are also shown as background. The boxes show the  
 435 11 mosaicked sub-regions for ice velocity.

436  
 437 **7 Conclusions**

438 In this contribution, we constructed two high-resolution ice flow maps covering the years of 2014



439 and 2015 for the entire Antarctica, which can accurately describe the current ice dynamics in the area.  
440 We also found a significantly increased mass discharge in East Indian Ocean sector by  $40\pm 24$  Gt yr<sup>-1</sup>  
441 over the last decade, attributable to the widespread accelerating glaciers in Wilkes Land, East  
442 Antarctica, while the other five oceanic sectors did not show obvious changes in mass discharge,  
443 contrary to the long-standing belief that present-day accelerated mass loss primarily originates from  
444 West Antarctica and Antarctic Peninsula. In the entire Antarctic ice sheet, total mass balance decreased  
445 by  $49\pm 90$  Gt yr<sup>-1</sup> during the last decade, a decline of 27% from 2006 ( $-181\pm 68$  Gt yr<sup>-1</sup>). The most  
446 significant change of mass balance was found in East Indian Ocean during the last decade, reaching  
447  $-40\pm 50$  Gt yr<sup>-1</sup>. The large uncertainty of mass balance change is mainly due to error in the SMB data.  
448 The significant increased mass discharge together with synchronized speedup of the linked ice shelves  
449 in Wilkes Land suggests a potential risk of destabilization of the marine ice sheet in the region overlain  
450 by the large subglacial basins with inland-sloping bedrock and deep troughs, an instable bedrock  
451 configuration like West Antarctica. Our new high-resolution ice flow maps together with existing  
452 InSAR ice velocity allow the first continent-wide assessment of ice flow and ice discharge changes  
453 during the last decade, which will contribute to our understanding of the entire Antarctic ice dynamics,  
454 and to potentially improving ice-sheet modelling and sea-level projections in the 21<sup>st</sup> century.

455  
456 Supplement materials include:  
457 Table S1, S2, S3, and S4 in Excel format  
458 Supplementary materials in pdf format

#### 459 **Data availability**

460 The data used in this paper include ice velocity data, ice thickness data, optical satellite images and  
461 grounding line products. The ice velocity data are from  
462 [http://nsidc.org/data/docs/measures/nsidc0484\\_rignot/](http://nsidc.org/data/docs/measures/nsidc0484_rignot/), <http://nsidc.org/data/nsidc-0545/index.html>, and  
463 <http://nsidc.org/data/velmap/>; to obtain our velocity products, please contact the author. The ice  
464 thickness products are provided by the Bedmap program from  
465 [http://www.antarctica.ac.uk/bas\\_research/data\\_access/bedmap/download/](http://www.antarctica.ac.uk/bas_research/data_access/bedmap/download/), and by IceBridge (IPR data)  
466 from <http://nsidc.org/data/docs/daac/icebridge/irmcr3/>,  
467 <https://nsidc.org/data/docs/daac/icebridge/brmcr2/>, <http://nsidc.org/data/ir1hi2>,  
468 <https://nsidc.org/data/ir2hi2>. The optical satellite images are from <http://glovis.usgs.gov> and  
469 <http://earthexplorer.usgs.gov/> for Landsat. Other data are from  
470 [http://icesat4.gsfc.nasa.gov/cryo\\_data/ant\\_grn\\_drainage\\_systems.php](http://icesat4.gsfc.nasa.gov/cryo_data/ant_grn_drainage_systems.php) for Antarctic drainage basins,  
471 [http://nsidc.org/data/atlas/news/antarctic\\_coastlines.html](http://nsidc.org/data/atlas/news/antarctic_coastlines.html) for the MOA grounding line and coastline,  
472 <http://nsidc.org/data/nsidc-0489> for the ASAIID project grounding line,  
473 [http://nsidc.org/data/docs/measures/nsidc0498\\_rignot/](http://nsidc.org/data/docs/measures/nsidc0498_rignot/) for the DInSAR grounding line,  
474 <http://pangaea.de/> for the grounding line provided by Depoorter et al.(2013),  
475 <http://nsidc.org/data/nsidc-0082> for the Antarctica DEM, and <http://www.staff.science.uu.nl/> for the  
476 FDM and SMB data, <http://nsidc.org/data/docs/agdc/nsidc0280/index.html> for MODIS Mosaic of  
477 Antarctica(Bohlander and Scambos, 2007), and <http://woceatlas.tamu.edu/> for potential temperature of  
478 seawater and bathymetry.

#### 480 **Author contribution**

481 Q.S. conceived the study, analysis and wrote the article. H.W. contributed to the research framework,  
482



483 helped develop the methodology and helped interpret the results and edited the manuscript. C.K.S.  
484 contributed to the interpretation of the results and edited the manuscript. L.J. contributed to the data  
485 analysis. H.T.H. contributed to the result analysis. J.D. contributed to the data processing. All authors  
486 commented on the manuscript.

487

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644