

# 3D surface properties of glacier penitentes over an ablation season, measured using a Microsoft Xbox Kinect.

Lindsey. I. Nicholson<sup>1</sup>, Michał. Pełlicki<sup>2,3</sup>, Ben. Partan<sup>4</sup>, and Shelley. MacDonell<sup>3</sup>

<sup>1</sup> *Institute of Atmospheric and Cryospheric Sciences, University of Innsbruck, Innsbruck, Austria*

<sup>2</sup> *Institute of Geophysics, Polish Academy of Sciences, ul. Księcia Janusza 64, 01-452 Warsaw, Poland*

<sup>3</sup> *Centro de Estudios Avanzados en Zonas Áridas (CEAZA), La Serena, Chile*

<sup>4</sup> *University of Maine, Orono, USA*

Correspondence to: L. I. Nicholson (lindsey.nicholson@uibk.ac.at)

**Abstract.** In this study, the first small-scale digital surface models (DSMs) of natural penitentes on a glacier surface were produced using a Microsoft Xbox Kinect sensor on Tapado Glacier, Chile (30°08'S; 69°55'W). The surfaces produced by the complete processing chain were within the uncertainty of standard terrestrial laser scanning techniques. The three-dimensional positional error of alignment between the digital surface and ground control points, was on average 0.08m, but in one case reached 0.3 m, due to poor overlap of individual scanned sections comprising the surface. Between November 2013 and January 2014 penitentes become fewer, wider, deeper, and the distribution of surface slope angles becomes more skewed to steep faces. Surface lowering during this core ablation season was in the order of 0.04m day<sup>-1</sup>. While morphological changes cannot be captured by manual point measurements, a key finding is that mean surface lowering is well captured by manual measurements of penitente surface height at a minimum density of 5 m<sup>-1</sup> over a 5 m transverse profile. Roughness was computed by applying two previously published geometrical formulae; one applied to the 3D surface and one to single profiles sampled from the surface. Morphometric analysis shows that skimming flow is persistent over penitentes, providing conditions conducive for the development of a distinct microclimate within the penitente troughs. Numerous options for representative roughness element height were used, and the calculations were done both with and without application of a zero displacement height offset to account for the likelihood of skimming air flow over the closely-spaced penitentes. Calculated roughness values are in the order of 0.01-0.10 m during the early part of the ablation season, increasing to 0.10-0.50 m after the end of December, in line with the largest previously published surface roughness values for glacier ice. Calculated surface roughness is strongly dependent on wind direction. For values calculated from 3D surfaces maximum roughness coincides with airflow across the penitente lineation while maximum roughness computed from sampled profiles coincides with airflow along the penitente lineation. These findings highlight the importance of determining directional roughness and wind direction for

37 strongly aligned surface features and also suggest more work is required to determine  
38 appropriate geometrical roughness formulae for linearized features.

## 39 **1. Introduction**

40 Penitentes are spikes of snow or ice, ranging from a few centimetres up to several metres  
41 in height that can form during the ablation season on snowfields and glaciers. They are a  
42 common feature of high elevation, low-latitude glaciers and snowfields (e.g. Hastenrath and  
43 Koci, 1981; Corripio and Purves, 2005; Winkler et al., 2009) where very low humidity,  
44 persistently cold temperatures and sustained high solar radiation favour their  
45 development (Lliboutry, 1954). As cryospheric water resources are relatively important to  
46 local dry season water supply in arid mountain ranges (Kaser et al., 2010), there is  
47 potential value in understanding how penitentes might influence both runoff and  
48 atmospheric humidity.

49 Penitentes form linearized, inclined fins of snow or ice on the surface. Both the latitudinal  
50 range (within  $55^\circ$  of the equator on horizontal surfaces) and geometry (aligned with the  
51 arc of the sun across the sky, and tilted toward the sun at local noon) of penitentes are  
52 governed by solar-to-surface geometry (Lliboutry, 1954; Hastenrath and Koci, 1981;  
53 Bergeron et al., 2006; Cathles et al., 2014). During the initial stages of penitente  
54 development, ablation is thought to proceed by sublimation alone, driven by low  
55 atmospheric humidity. Surface irregularities focus reflected solar radiation within  
56 depressions (Amstutz, 1958; Corripio and Purves, 2005; Lhermitte et al., 2014; Claudin et  
57 al., 2015) such that the energy receipts, and consequently ablation, are enhanced in the  
58 hollow and the surface irregularity becomes amplified. Subsequently, as the surface relief  
59 increases, a more humid microclimate can develop in the hollows between penitentes,  
60 suppressing sublimation and thereby allowing melting in the depressions. The penitentes  
61 tips continue to ablate by sublimation alone (Lliboutry, 1954; Drewry, 1970; Claudin et al.,  
62 2015) and, as melting requires approximately an eighth of the energy of sublimation to  
63 remove the same amount of ice, the spatial differentiation of ablation processes between  
64 penitente trough and tip amplifies the penitente surface relief.

65 The altered partitioning of ablation between sublimation and melting in penitente fields, as  
66 compared to surfaces without penitentes (e.g. Lliboutry, 1998; Winkler et al., 2009; Sinclair  
67 and MacDonell, 2016), is expected to alter the rate of mass loss and meltwater production  
68 of snow and icefields during the ablation season, but this has not yet been fully quantified  
69 (MacDonell et al., 2013). Previous studies, based on radiative modelling within idealized  
70 penitente surfaces, have investigated the impact of penitentes on the shortwave radiative  
71 balance (Corripio and Purves, 2005; Cathles et al., 2014; Lhermitte et al., 2014). The results  
72 suggest that penitentes reduce effective albedo by up to 40% compared to flat surfaces and

73 that both shape and penitente size impact the apparent albedo as measured by ground and  
74 satellite sensors (Lhermitte, et al., 2014). The development of penitentes also manifestly  
75 alters the surface roughness properties, but neither the impact of penitentes on surface  
76 roughness, nor the associated impact on turbulent energy fluxes has been investigated.  
77 While penitentes are a relatively rare form of linearized surface feature, linear crevasses  
78 are widespread, and penitentes offer a unique test bed for investigating the significance of  
79 linearized features on effective surface roughness for various wind directions.

80 Determining effective surface roughness on penitente-covered surfaces is complicated, as  
81 they present very closely spaced locally high relief surfaces. This means that calculating the  
82 aerodynamic roughness length ( $z_0$ ) is based not only on the absolute depth of the  
83 penitentes, but also on a zero-plane displacement ( $z_d$ ), which essentially means that the  
84 base of the eddy entering the penitente field is above the depth of the penitente trough. In  
85 addition, due to the irregularity of a penitente field, a roughness sub-layer can form and  
86 interact with the upper eddy system, therein creating a complex, chaotic flow. Closely  
87 packed roughness elements generally experience a wake interference regime, and in the  
88 most densely packed arrays of roughness elements skimming flow occurs (Grimmond and  
89 Oke, 1999). At the top of the roughness sublayer individual wakes caused by surface  
90 obstacles are smeared out and the flow is independent of horizontal position, and thus,  
91 observations at this level represent the integrated surface rather than individual surface  
92 obstacles. This level is known as the blending height ( $z_r$ ). All these properties are  
93 dependent on the size and arrangement of surface roughness elements.

94 Measurements of natural penitentes required to examine their morphometry and  
95 roughness are rare (e.g. Naruse and Leiva, 1997), and difficult to obtain because the  
96 complex, and partially overhanging, surface prevents the use of simplified automated tools  
97 such as photogrammetric determination of surface profile heights (e.g. Fassnacht et al.,  
98 2009; Manninen et al., 2012) or line-of-sight surveying from fixed positions. Recent  
99 advances in close-range, mobile, depth-of-field sensors and efficient feature tracking  
100 software used in interactive computer gaming offer potentially useful tools that can be  
101 applied to resolve such problems in earth science (e.g. Mankoff and Russo, 2013). In this  
102 study a Microsoft Xbox Kinect sensor is used as a close-range mobile distance ranger to  
103 produce a series of small-scale digital surface models (DSMs). The method of DSM  
104 generation is evaluated against standard terrestrial laser scanning, and the Kinect-derived  
105 DSMs of the penitentes are used to (i) perform the first detailed examination of the  
106 morphometry of natural penitentes over the course of an ablation season; (ii) compare the  
107 volume change computed from DSM differencing with estimates based on manual  
108 measurements of surface lowering and (iii) examine the geometrical roughness properties  
109 of the sampled penitente surfaces.

## 110 2. Methods

### 111 2.1 Description of field area and measurement setup

112 Tapado Glacier (30°08'S; 69°55'W), which is known to develop penitentes every summer,  
113 lies in the upper Elqui Valley of the semi-arid Andes of the Coquimbo Region of Chile (Fig  
114 1). Interannual climate variability is controlled by the El Niño Southern Oscillation (ENSO),  
115 such that during El Niño events, higher precipitation and warmer conditions are  
116 experienced (Escobar and Aceituno, 1998). Most precipitation is received during the winter  
117 (Vuille and Ammann, 1997), however convective storms can cause small precipitation  
118 events in the period from December to March (Schotterer et al., 2003). Although the glacier  
119 mass balance in the area is highly sensitive to precipitation, warming at elevation over the  
120 last 40 years has produced a rise of the glacier equilibrium line altitude of over 120 m  
121 (Carrasco et al., 2008). Annual mean temperature is below freezing and annual mean  
122 relative humidity is below 30% (Ginot et al., 1999). The glacier experiences year-round  
123 ablation by sublimation, however, melt is only produced during the summer (Sinclair and  
124 MacDonell, 2016).

125 Two measurement sites were analysed: the 'test site' and the 'glacier site' (Fig. 1). The 'test  
126 site' was established at a patch of snow penitentes (0.5 – 1.0 m height) within a dry stream  
127 bed at 4243 m a.s.l. in the glacier foreland (Fig 1). This site was used to (i) test instrumental  
128 setups to optimize the field operation of the Kinect sensor, and (ii) compare the  
129 performance of the Kinect sensor against a Terrestrial Laser Scanner (TLS). It was chosen  
130 to avoid the logistical difficulties of transporting the TLS to the glacier. Subsequently, two  
131 study plots were established at the 'glacier site' at an elevation of 4774 m a.s.l. in the glacier  
132 ablation zone (Fig 1). These sites were scanned repeatedly with the Xbox Kinect (see  
133 section 2.3) during the core ablation season between the end of November 2013 and the  
134 beginning of January 2014. The location and layout of the two glacier plots are shown in Fig  
135 1a. Site A (5 m by 2 m) was measured four times, on 25 November, 11 December, 20  
136 December and 3 January. Site B (2 m by 2 m) was only measured on the last three dates.  
137 The corners of the study sites were marked with 2 m lengths of plastic plumbing piping  
138 hammered vertically into the snow, or drilled into the ice (Fig 1c). The positions of these  
139 stakes were measured using a Trimble 5700 differential GPS with Zephyr antenna on the  
140 25th November to provide ground control points and a common reference for each survey.  
141 On each visit to the glacier, when possible, the stakes were hammered further into the  
142 snow and the resultant lowering of the stake top was noted. The maximum standard  
143 deviations of the GPS stake positions were < 1.0 cm, 1.1 cm and 1.7 cm in easting, northing  
144 and elevation respectively, with combined XYZ standard deviation < 2.0 cm for all stakes  
145 (Supplement A). Error on the manual measurements of height offsets of the marker stakes  
146 on subsequent survey dates is conservatively estimated to be 2.0 cm. This results in total  
147 positional errors of the ground control points at each scan date of between 2.3 and 2.7 cm

148 depending on the stake. Manual measurements of surface lowering were made along the  
149 eastern long side of site A. All surfaces heights were referenced to the elevation of the  
150 glacier surface at the upglacier end of this cross profile at the date of installation.

151 An automatic weather station (AWS) on a free-standing tripod was installed beside the two  
152 glacier plots to provide meteorological context for the measurements, as well as an  
153 independent measure of local surface lowering (Fig 1). The sensors installed, and variables  
154 recorded, are detailed in Table 2. During the study period one significant snowfall event  
155 occurred on the 8<sup>th</sup> December 2013, when the sonic distance ranger recorded a surface  
156 height increase of 0.09 m over the course of the day (Table 2).

## 157 **2.2 Terrestrial laser scanning**

158 At the 'test site', surface scans produced by the Kinect sensor were compared with those  
159 produced by well-established TLS methods. An Optech ILRIS long-range terrestrial laser  
160 scanner was used as it is especially suitable for surveying snow and ice surfaces due to  
161 having a shorter wavelength laser beam (1064 nm) than other models. This equipment  
162 surveys surface topography based on time-of-flight measurement of a pulsed laser beam  
163 reflected to a given angle by a system of two rotating mirrors. It has a raw range accuracy  
164 of 4 mm at 100 m distance, raw angular accuracy of 80  $\mu$ rad, beam diameter of 27 mm at  
165 100 m distance and beam divergence of 250  $\mu$ rad. The instrument was placed in five  
166 locations around the surveyed snow patch and boulder, overlooking it from different  
167 directions. Positions of the TLS were measured with Trimble 5700 differential GPS with  
168 Zephyr antennae in static mode. Seventeen point clouds were obtained with nominal  
169 resolution of 0.11-0.75 cm. Resulting point clouds were corrected for atmospheric  
170 pressure, temperature and humidity using data from a weather station in the glacier  
171 forefield, and then trimmed using ILRIS Parser software, aligned with Polyworks IMAAlign  
172 software into a common local coordinate system and georeferenced with differential GPS  
173 measurements using Polyworks IMInspect software. The alignment error of the point  
174 clouds as estimated by this software is 0.36-0.87 cm and comparison with ground control  
175 points gives an error of 5.65 cm. The TLS scan of the snow penitentes is presented as an  
176 example of the nature of the DSM that can be obtained within a penitente field using TLS  
177 (Fig 2). Due to logistical constraints, the scans of snow penitentes could not be carried out  
178 with both the TLS and Kinect on the same day, so direct comparison of the TLS and Kinect  
179 scans is instead performed on a reference boulder within the test site, whose surface is  
180 assumed unchanged between different scan dates.

## 181 **2.3 Kinect surface scanning**

182 The Kinect sensor emits a repeated pattern of structured infra-red (IR) beams, and records  
183 the pattern distortion with an IR camera. The depth-of-field calculation is performed via a

184 proprietary algorithm and a distance map is the raw data output. Using the standard  
185 calibration the static raw depth field resolution of the Kinect is 1 mm and the distance error  
186 is < 1.0 cm at the distance range of the penitente scans (Mankoff and Russo, 2013).

187 For its original gaming usage, the Kinect is in a fixed position and proprietary software uses  
188 feature tracking to monitor movements of players within the field of view of the Kinect. The  
189 inverse of this workflow can also be applied whereby the Kinect sensor is moved  
190 interactively around a static surface or 3D body, using the same feature tracking to  
191 compute the position of the sensor relative to the object and thereby allowing a point cloud  
192 reconstruction of the object. In this work we apply the second work flow and sample Kinect  
193 data using ReconstructMe™ 2.0 software. In common with alternative reconstruction  
194 packages compatible with the Kinect, ReconstructMe™ performs bilateral filtering on the  
195 output depth map frame and converts the pixel version of each depth map frame to 3D  
196 coordinate maps of vertices and normals. An iterative closest point (ICP) alignment  
197 algorithm is then applied frame by frame at three scales to repeatedly rotate and translate  
198 the depth field to determine camera position and an aligned surface, giving weighted  
199 preference to portions of the surface that are perpendicular to the line of sight. The  
200 ReconstructMe™ software has the advantage of producing surface meshes in real-time, so  
201 that the operator can check the scan quality and coverage at the time of capture, but the  
202 disadvantage that the raw point cloud is not saved and if the real-time tracking is lost a new  
203 scan must be started.

204 The Xbox Kinect was connected via a 5m powered USB extension cord to an MSI GE60  
205 gaming laptop, powered using a 240V 600W inverter connected to the 160Ah 12V battery  
206 of the automatic weather station on the glacier. Scans were carried out by two people; one  
207 handling the Kinect and the other monitoring the quality of the surface being generated. In  
208 bright conditions, the return IR signal of the Kinect is swamped by natural radiation over  
209 snow and ice surfaces, which reflect a high proportion of incident shortwave radiation, and  
210 absorb or scatter much of the longwave radiation signal. Therefore, scanning was carried  
211 out at twilight or just after nightfall. Sudden movements caused by the operator slipping or  
212 the snow compacting underfoot resulted in loss of tracking of common reference points.  
213 Consequently, each study site was scanned in small sections and three to thirteen  
214 overlapping surface meshes were used to cover the area of each study site.

## 215 **2.4 Kinect surface mesh processing**

216 The full mesh processing procedure using the freely-available Meshlab software is  
217 presented in Supplement B, and briefly described here. Small surface components,  
218 unreferenced and duplicated vertices were removed from the meshes using inbuilt filters.  
219 The component meshes that cover each sampling date at a single site were aligned using an  
220 iterative closest point (ICP) algorithm which distributes the alignment error evenly across

221 the resultant mosaicked surface mesh. Alignment solutions consistently had mean  
222 distributed error < 4 mm (Supplement B). The aligned meshes were flattened into a single  
223 layer, remeshed using a Poisson filter and resampled to reduce the point density by setting  
224 a minimum vertex spacing of 2.5 mm.

225 The surface mesh for each scan date was georeferenced in Polyworks software using the  
226 known coordinates of the base of the marker stakes at the time of each scan because the  
227 upper portions of the symmetrical stakes are often poorly captured by the meshing  
228 software. The local elevation zero was set to be the north-east corner of site A. The  
229 mismatch evident in the georeferencing step (Table 1) is much larger than the mesh  
230 alignment error (Supplement B). This is most likely an artifact of a combination of (i)  
231 reduced mesh quality at the margins of the component scans and (ii) insufficient overlap  
232 between some scan sections producing distortion within the mesh alignment.

233 To eliminate the marker stakes and any data gaps near the margins of the study areas, each  
234 surface mesh was sub-sampled within the staked area. The sub-sampled area for site A is a  
235 2.0 by 3.5 m horizontal area (7.00 m<sup>2</sup>), and site B is a 1.5 m x 1.5 m horizontal area  
236 (2.25 m<sup>2</sup>) shown in the examples in Figure 3. Mesh vertices and an index file of the vertices  
237 comprising each face were exported from Meshlab for subsequent analysis in Matlab  
238 software.

## 239 **2.5 Calculations of surface geometrical properties**

240 The geo2d and geo3d toolboxes (available from the Matlab File Exchange) were used to  
241 compute the face areas and normals of the mesh, from which surface height distribution,  
242 aspect and dip of the sampled surface were calculated, weighted by the ratio of each face  
243 area to the total surface area of all faces. As the surfaces contain overhanging parts, DSM  
244 differencing cannot be performed by simple subtraction. Instead surface lowering was  
245 calculated in two ways: Firstly by differencing area-weighted mean surface elevations, and  
246 secondly by computing the volume change between scan dates. For the latter approach,  
247 volumes for all surfaces were computed relative to a baselevel horizontal reference.  
248 Volumes relative to this horizontal reference for upward-facing triangles were computed  
249 column-wise, by projecting the area of each triangular face onto the reference surface and  
250 using the height coordinate of the triangle centroid as the height dimension for each  
251 column. These were summed and volumes for overhanging triangles, calculated in the same  
252 way, were subtracted to derive the total volume between the reference surface and each  
253 scanned penitente surface. Successive volumes were then subtracted to obtain the volume  
254 change over each measurement interval. The volume-differencing approach is expected to  
255 be the more accurate of the two methods as it accounts for over-hanging surfaces.

## 256 **2.6 Manual measurements of surface change**

257 Traditional single-point stake measurements of glacier surface lowering are unreliable  
258 within the inhomogeneous surface of a penitente field. One alternative is to measure  
259 surface lowering at intervals along a profile perpendicular to the main axis of alignment of  
260 the penitentes. Such a reference was installed along the 5 m-long eastern margin of site A,  
261 between two longer corner stakes drilled 3 m into the ice using a Kovacs hand drill. The  
262 distance between a levelled string and the glacier surface was measured using a standard  
263 tape measure at 0.2 m intervals on 23 November. Subsequent measurements, on the 12 and  
264 21 December and on 4 January, were made at 0.1 m intervals. All measurements were  
265 recorded to the nearest centimetre, and the error on each measurement is estimated to be  
266 2.0 cm, which is assumed to capture the error associated with the horizontal position of the  
267 measurements along the reference frame and the vertical measurements of the distance to  
268 the surface beneath.

## 269 **2.7 Calculations of geometric surface roughness**

270 Morphometric determinations of surface roughness have been shown to be viable, and  
271 generally more easily implemented, alternatives to aerodynamic determinations of surface  
272 roughness (e.g. Kondo and Yamazawa, 1986; Munro, 1989; Grimmond and Oke, 1999;  
273 Fassnacht et al., 2009; Andreas, 2011). The surface meshes created from the Kinect  
274 measurements were used to calculate  $z_0$  using a widely-used relationship established by  
275 Lettau (1969), initially developed for isolated, regular obstacles distributed over a plane:

$$276 \quad z_0 = 0.5 h \left( \frac{s}{S} \right) \quad (1)$$

277 where  $h$  is the height of the obstacles,  $s$  is the upwind silhouette area of each obstacle and  $S$   
278 is the specific area occupied by each roughness element obstacle, also referred to as its lot  
279 area. Following Macdonald and others (1998),  $h$  was replaced by average obstacle height,  $s$   
280 with the sum of all the upwind silhouette areas, and  $S$  with the total area covered by the  
281 obstacles. While the upwind silhouette area, and indeed surface area in any direction, is  
282 relatively easily defined for each surface mesh area using trigonometry, it is difficult to  
283 define individual roughness elements and their representative heights, due to the lack of an  
284 apparent base level. Surfaces were first detrended to remove any general surface slope at  
285 the site, then roughness for the detrended 3D meshes is calculated assuming that the  
286 roughness elements cover the whole surface area (i.e  $S = \text{plot area}$ ), and for four possible  
287 representations of average obstacle height ( $h$ ) as follows: (i) the maximum range of the  
288 detrended mesh; (ii) twice the standard deviation of the detrended surface mesh; (iii)  
289 mean mesh height above the mesh minimum; and (iv) median mesh height above the  
290 minimum.



291 Penitente fields are expected to exceed the 20-30% roughness element density criterion  
292 below which Equation 1 applies (Macdonald et al., 1998). Above this limit, skimming flow is  
293 expected and consideration of an appropriate zero displacement height is needed. As  $z_d$  is  
294 also unknown in the case of penitente fields, sample calculations of three-dimensional  
295 roughness on the detrended surface meshes were made using three possible realizations of  
296  $z_d$ :  $z_d = h$ ;  $z_d = 2/3h$  (following Brutsaert, 1975);  $z_d = 1/3h$ . Each  $z_d$  case is computed for the  
297 four values of  $h$  previously outlined. Equation 1 is then applied to the roughness elements  
298 remaining above the plane of the general surface slope offset by a distance  $z_d$  above the  
299 minimum height of the surface mesh. The representative height  $h$  for this portion of the  
300 mesh exceeding the plane is taken to be the mean area-weighted height of all triangles  
301 above this plane,  $s$  is the summed frontal area of all mesh triangles above  $z_d$  that face into  
302 the chosen wind direction and  $S$  is the total horizontal area of the surface components  
303 above  $z_d$ .

304 Finally, roughness was calculated for cross-sections of length  $X$ , sampled perpendicularly to  
305 the wind direction following Munro (1989, 1990). Here,  $h$  is replaced with an effective  
306 height  $h^*$  expressed as twice the standard deviation from the standardized mean profile  
307 height;  $s$  is replaced with  $h^*X/2f$ , in which  $f$  is the number of profile sections that are above  
308 the mean elevation; and  $S$  is replaced with  $(X/f)^2$ . To investigate the nature of the  
309 roughness computed this way for north-south and east-west impinging wind directions,  
310 cross profiles longer than 1.5 m at 0.1 m intervals orientated E-W and N-S were extracted  
311 from each scanned surface. Cross-sections were detrended to remove the influence of any  
312 general surface slope at the site, and roughness was computed on each of these cross-  
313 sectional profiles following the modifications of Munro. Mean profile roughness for these  
314 two wind directions are presented for each sampled surface.

### 315 **3. Results**

#### 316 **3.1 Evaluation of the quality and suitability of penitente scans by TLS and Kinect**

317 The penitente surface produced by the TLS did capture some overhanging surfaces but only  
318 58% of the total horizontal surveyed area was captured as the deepest parts of the troughs  
319 were obscured from the view by the surrounding penitentes (Fig 2a). In comparison, the  
320 Kinect system scanned 100% of the survey area.

321 For the direct comparison of the two methods on a reference boulder, the Kinect-derived  
322 surface, produced from three mosaicked meshes was aligned to the surface produced from  
323 the TLS point clouds. The TLS scan was incomplete, with parts of the top and overhanging  
324 surfaces of the boulder missing due to being obscured from the TLS survey positions, while  
325 the Kinect scan achieved complete coverage of the boulder. The difference between the two

326 aligned meshes where overlapping data existed was always < 2 cm, which is well within the  
327 uncertainty of the georeferenced TLS surface model. Larger differences of up to 5 cm,  
328 evident in Figure 2b, occur only where there are data gaps in the TLS surface being  
329 compared.

330 It is difficult to formally assess the total error of the surfaces produced by the Kinect scans  
331 because the workflow involves several black box processing steps. The mean alignment  
332 errors of the mesh mosaicking step in Meshlab is < 0.4 cm and quantifiable errors  
333 associated with the GPS positions, subsequent measurement of the stake bottom positions  
334 relative to the GPS positions are all < 2.0 cm. However, the three-dimensional  
335 georeferencing error in this study is large (Table 1) compared to the other sources and is  
336 therefore taken as a reasonable value for the uncertainty of the total process chain. Errors  
337 given on the seasonal mass, volume and surface changes are based on summing the squares  
338 of the mean elevation difference between the marker stakes and ground control points  
339 (GPCs) at each site on the first and last survey dates.

### 340 **3.2 Morphometric changes and surface lowering**

341 The morphometry of the sampled penitentes changed visibly over the measured intervals  
342 (Figs 3 and 4). The strong east-west lineation and preferred north and south surface aspect  
343 predicted from theory developed early and were maintained throughout study period. Over  
344 time penitente troughs became fewer in number, but wider and deeper. This causes total  
345 surface area to increase; at site A the true surface is between 1.7 and 4.0 times the  
346 horizontal equivalent area, and at site B between 2.1 and 3.7 times the horizontal surface  
347 area equivalent (Fig 4 a & b). Snowfall during the first measurement interval decreases the  
348 surface area at site A over that interval. Surface relief, expressed by the vertical range of the  
349 mesh, also increases through time, except when snowfall partially filled the developing  
350 penitentes and reduces both the range of the surface and the general slope angle. The  
351 largest part of the surface is facing southwards, and the predominant angle generally  
352 steepens over time, though again this trend is reversed by snowfall (Fig 4 c & d). From the  
353 onset of measurements the surface aspect distribution is strongly dominated by north and  
354 south facing components and this becomes more pronounced in the latter measurements  
355 and the preferred orientation rotates slightly over the course of the season (Fig 4 e & f).

356 Surface lowering rates derived from calculated volume changes per unit area were 21, 41  
357 and 70 mm d<sup>-1</sup> over each interval at site A, and 57 and 61 mm d<sup>-1</sup> over the last two intervals  
358 at site B. In comparison, surface lowering rates calculated from area-weighted mean mesh  
359 elevation were within a few millimetres of those derived from volume changes: 22, 38 and  
360 69 mm d<sup>-1</sup> for the three measured intervals at site A, and 54 and 60 mm d<sup>-1</sup> for the last two  
361 intervals at site B. The increasing rate of surface lowering through time is associated with  
362 progressively increasing atmospheric energy supply and surface properties becoming

363 more conducive to melting. The warming atmosphere is expressed in the positive degree  
364 days of the three periods which are 3.7, 2.2 and 31.5 over the 16, 9 and 14 day-long periods  
365 respectively.

366 Total surface lowering over the whole available data period for each site computed by  
367 volume change (area-weighted mean height change) was  $1.68 (1.77) \pm 0.11$  m at site A and  
368  $1.37 (1.32) \pm 0.38$  m at site B. Over the common measurement period, surface lowering at  
369 site A was  $1.35 (1.31) \pm 0.21$  m, indicating that lowering rate is repeatable at both sites.  
370 Volume loss was converted to mass loss using the mean snow density of  $426 \text{ kg m}^{-3}$  (with  
371 an assumed uncertainty of  $\pm 5\%$ ) measured in a 1.10 m snow pit excavated on 22  
372 November 2013 beside the AWS. Mass loss at site A computed from mesh volume change  
373 (area-weighted height change) between 25 November and 3 January was  $716 \pm 58 (754 \pm$   
374  $59) \text{ kg m}^{-2}$ . Mass loss at site B from mesh volume changes (area-weighted height changes)  
375 between 11 December and 3 January was  $582 (562) \pm 166 \text{ kg m}^{-2}$ . Measurements at site A  
376 over the same period give mass loss of  $573 (558) \pm 95 \text{ kg m}^{-2}$ , so again, measurements at  
377 both sites are within the range of uncertainty.

### 378 **3.3 Manual measurements of reference cross-profile**

379 Intermittent measurements cross-cutting the predominant penitente alignment do not  
380 capture the complexity of the surface as revealed by the Kinect surface sampling (Fig 5).  
381 Over the 39 days of the study, the mean mass loss calculated from 26 points spaced at  
382 0.2 m intervals along a 5 m profile crosscutting the penitentes at site A was  $1.61 \pm 0.14$  m,  
383 which equates to a mass loss over the same period of  $688 \pm 70 \text{ kg m}^{-2}$ . This differs from the  
384 value calculated from volume change computed from surface meshes consisting of over 1.3  
385 million points and covering an area of  $7 \text{ m}^2$  by only  $28 \text{ kg m}^{-2}$ , which is within the  
386 uncertainty of the two measurement methods. Assuming that this difference holds true for  
387 the whole ablation season of 120 days, point measurements underestimate the seasonal  
388 mass loss obtained from the Kinect digital surface models by  $86 \text{ kg m}^{-2}$ .

389 To investigate the impact of sampling resolution on the manual measurements and how the  
390 derived surface lowering compares to the Kinect-derived lowering, maximum elevation  
391 range, mean surface height (compared to the horizontal reference) and mean surface  
392 lowering, were calculated from manual measurements at 0.1 ( $n = 52$ ), 0.2 ( $n = 26$ ), 0.4 ( $n =$   
393  $14$ ) and 1.0 m ( $n = 6$ ) intervals on the last three measurement dates. The highest resolution  
394 sample was taken as a reference against which to evaluate coarser sampling. Surface relief  
395 differed from that measured at 0.1 m by maxima of 0.13, 0.29 and 0.41 m for 0.2, 0.4 and  
396 1.0 m sampling intervals respectively. Mean measured surface height was within 0.03 m of  
397 the highest resolution measurements at 0.2 m and 0.4 m intervals, and within 0.12 m at  
398 1.0 m resolution. Mean lowering rates at 0.1, 0.2 and 0.4 m sampling intervals were all  
399 within  $3 \text{ mm d}^{-1}$ . This increased to a maximum of  $12 \text{ mm d}^{-1}$  when the sampling resolution

400 was decreased to 1.0 m. Decreasing the length of the sampled profile down to 2 m alters the  
401 mean lowering rate by less than 5 mm d<sup>-1</sup> at sampling resolutions of 0.1, 0.2 and 0.4 m.

402 Probing of the snow along the line of the horizontal reference on 25 November indicated  
403 mean snow depth of 1.83 m (standard deviation 0.56 m). The underlying ice surface does  
404 not appear to be influencing the structure of the overlying snow penitentes (Fig 5).  
405 However, it is difficult to draw a firm conclusion based on these measurements,  
406 particularly as, while the surface of the penitentes was still snow on the 3 January, in  
407 several instances the surface had lowered below the level of the ice interface suggested by  
408 the initial probing. This highlights the difficulty in identifying the underlying ice surface, or  
409 summer ablation surface, within a penitente field, suggesting that a single location must be  
410 sampled very densely to obtain a characteristic snow depth by this method.

### 411 **3.4 Surface roughness assessments**

412 The representative height,  $h$ , used in the calculations increases over time in all cases, and is  
413 bounded by the maximum ( $h$  as range of the detrended surfaces), and minimum ( $h$  as twice  
414 the standard deviation of the detrended surface) cases (Fig 6). Differences in  $h$  computed  
415 by the same method can reach as much as 0.2 m between the two sites, although the  
416 pattern of change over time is consistent.

417 The application of Lettau's (1969) formula is considered to be invalid if the ratio of the  
418 frontal area to the planar area of the obstacles exceeds 20-30%, implying skimming airflow.  
419 .This ratio is greater than 20% for all of the penitente surfaces, and after the 20<sup>th</sup> December  
420 is always greater than 30%. If this issue is ignored, calculated  $z_0$  values increase with time  
421 and show a strong dependence on the impinging wind direction, with values peaking for  
422 wind directions perpendicular to the alignment of the penitentes (Fig 7). Calculated  $z_0$   
423 ranges from 0.01 – 0.90 m, depending on the way in which the representative height is  
424 expressed, the date and the wind direction (Fig 8). Given the close spacing of the penitentes  
425 it is likely to be more valid to explore what  $z_0$  would be when a zero displacement height  
426 offset is applied. Introducing the zero displacement height reduces the maximum  
427 calculated roughness by about half, and also reduces the variability between different  
428 representative heights (Fig 8), as a smaller  $h$  value translates into a smaller  $z_d$  so that the  
429 calculation is performed on a larger portion of the mesh.

430 Surface roughness assessments on the basis of calculations following Munro's modification  
431 for single profile measurements were applied to cross profiles longer than 1.5 m at site A  
432 (B) yielding 20 (6) profiles orientated N-S and 33 (7) E-W. Surface amplitude increases  
433 over time, and the amplitude of the N-S running cross profiles is generally larger than the  
434 E-W running cross profiles, as illustrated in the example of site B (Fig 9). Table 3 shows the  
435 calculated roughness values at each survey date, revealing that while profile-computed

436 roughness length increases monotonically over time at site B, it reduces over the first  
437 period at site A, associated with snowfall during this period. Both the range and relative  
438 increase in roughness over time is larger for the N-S running profiles. The computed  
439 roughness at both sites is 4.3 to 6.8 times larger for airflow impinging on the penitente field  
440 in an E-W direction than for airflow in the N-S direction. This is contrary to the results  
441 computed on the full 3D mesh surface, but is understandable because this formulation  
442 relies on the amplitude of the surface, which is generally larger in the N-S orientated cross  
443 profiles than the E-W running cross profiles.

444 Prevailing wind direction differs only slightly in each period with an increasing  
445 northwesterly component in the second two periods compared to the first. This may be  
446 related to the occurrence of snow during the first period, which is expected to alter  
447 thermally-driven valley wind systems. Over the whole study period wind direction is  
448 predominantly from the south-easterly and north-westerly sectors, and swings through  
449 both extreme wind angles used in the roughness calculations here (Fig 10). This indicates  
450 that the effective roughness at this site can be expected to differ significantly over time  
451 depending on the wind direction.

## 452 **4. Discussion**

### 453 **4.1 Methods of measuring change of rough glacier surface elements**

454 The test site for scanning penitentes with a TLS was chosen as scanning positions could be  
455 established on the surrounding higher ground overlooking the penitente field, thereby  
456 offering the best viewing angles possible. Nevertheless, the terrestrial laser scanning could  
457 only capture the upper portions of the penitentes. As ablation is at its maximum in the  
458 troughs, TLS data cannot capture the true volume change of penitentes. In contrast, the  
459 Kinect sensor can be moved across the complex relief of the penitente field to make a  
460 complete surface model. Although it is in principle possible to capture a large area with the  
461 ReconstructMe software used here, and it offers the advantage of providing real time  
462 feedback on the mesh coverage, it proved difficult to capture the study sites in a single scan  
463 given (i) the reduced signal range of the sensor over snow and ice (Mankoff and Russo,  
464 2013), and (ii) the difficulty of moving around the penitente field. As a result, partial scans  
465 were obtained, with the disadvantage that subsequently combining these introduces a  
466 substantial degree of additional error associated with alignment if the component scans  
467 were not of high quality at the margins, or did not overlap adjacent scan areas sufficiently.  
468 The practical utility of the Kinect on glacier surfaces is limited to small study areas, but  
469 integrating local findings with glacier wide TLS or photogrammetric information of surface  
470 conditions may offer a means to usefully extrapolate small scale findings to the glacier  
471 scale. Surface scanning technology and software is an area of rapid development, and

472 ongoing development of new sensors and airborne platforms may eliminate the challenges  
473 of producing high quality depth maps over larger areas using similar technology to the  
474 Kinect.

475 Despite not visually capturing the complex morphology of the penitentes, manual  
476 measurements of surface height change in a penitente field along a profile cross-cutting the  
477 penitentes are found to be robust for determining mean surface lowering rates, and show  
478 good agreement to the volume changes computed from differencing the digital surface  
479 models scanned in detail using a Kinect. Comparison of the manual sampling at different  
480 intervals suggests that five samples per meter is adequate to characterize surface change of  
481 penitentes, but that data will be unreliable if the cross-profile is too short.

## 482 **4.2 Penitente morphology**

483 Although the penitentes sampled here are more convoluted than the parallel rows of  
484 penitentes used in model representations (Corripio and Purves, 2005; Lhermitte et al.,  
485 2014), the morphometric properties of the meshes are similar to the morphometric  
486 properties of simplified surfaces. The penitente surface represents a much larger total  
487 surface area than the equivalent non-penitente surface and the control of solar radiation on  
488 penitente morphology (Cathles et al., 2014) means that the vast majority of the surface  
489 consistently dips steeply to the north and south at all stages of development.

490 Unless a snowfall event occurs to partially fill the troughs, surface relief, slope angle,  
491 penitente spacing and total surface area all increase over time as the penitentes develop.  
492 Thus the impact of penitentes on surface properties will also change along with the  
493 morphological changes. At Tapado Glacier, penitentes are initially overhanging to the  
494 north, and the southfacing sides are convex compared to the northfacing overhanging faces.  
495 Over the season the penitentes become more upright as the noon solar angle gets higher,  
496 and these changes in morphology may in turn alter the effect of the penitentes on surface  
497 albedo (Lhermitte, et al., 2014). In the context of the numerical theory of Claudin and  
498 others (2015), penitente spacing controls the atmospheric level at which water vapor  
499 content is representative of the bulk surface properties. Simultaneous field or laboratory  
500 measurements of penitente spacing evolution and vapor fluxes above the surface would be  
501 required to confirm this, but the spacing from the field measurements provided here can be  
502 used as an indication of the level at which measurements would need to be made in order  
503 to capture the bulk surface fluxes rather than fluctuations governed by the small-scale  
504 surface terrain.

### 505 4.3 Surface roughness

506 Given that aerodynamic measurements to determine the most suitable representative  
507 height and zero displacement level for penitentes are thus far unavailable, the approach  
508 taken here was to do an exploratory study and compute geometric surface roughness  
509 values using various ways of expressing  $h$  and  $z_d$ . As a consequence the results are purely  
510 illustrative and while patterns can be drawn from them that have meaning for  
511 understanding the nature of the computation, the applicability of these values in turbulent  
512 exchange calculations remains to be established.

513 The ratio of frontal to planar area of the penitentes implies that skimming flow prevails,  
514 such that turbulent airflow in the overlying atmosphere does not penetrate penitente  
515 troughs. This is in agreement with the theory of formation and growth of penitentes, in  
516 which the development and preservation of a humid microclimate within the penitente  
517 troughs is required to facilitate differential ablation between the trough and tip of the  
518 penitente. Although the data here shows that penitentes become less densely packed over  
519 time, skimming flow regime persists over the study period, and available data is insufficient  
520 to determine if this holds true to the end of the ablation season.

521 The changing morphometry of the penitentes alters the geometrical surface roughness as  
522 they develop over the ablation season. Roughness calculated using a range of possible  
523 representations of  $h$  and  $z_d$  give roughness values in the order of 0.01-0.10 m during the  
524 early part of the ablation season and 0.10-0.50 m after the end of December. These values  
525 are in line with values previously published for rough glacier ice (Smeets et al., 1999;  
526 Oblitner, 2000). This roughness increase is related to the deepening of the penitentes over  
527 time and an increase of the surface amplitude. Lettau's (1969) formula, which does not  
528 account for  $z_d$ , overestimates roughness for densely packed obstacles, but this does not  
529 compensate sufficiently to reproduce values of  $z_d + z_0$  for densely packed obstacles from  
530 formulations that include  $z_d$  in the computation of  $z_0$ . Thus, Lettau's formula is expected to  
531 estimate the zero velocity point of a logarithmic wind profile to be lower than formulations  
532 that include  $z_d$  in the computation of  $z_0$ . The pattern of the computed roughness properties  
533 is consistent between the two neighbouring sites, but individual values can differ,  
534 suggesting that relief varies substantially over short distances and sampling a large area is  
535 necessary to capture mean properties.

536 The strong alignment of penitentes means that calculated roughness is strongly dependent  
537 on wind direction. Roughness calculated from 3D surface meshes are higher for wind  
538 impinging in a north-south direction, as the large faces of the penitentes form the frontal  
539 area in this case. In contrast, roughness calculated for individual profiles extracted from the  
540 mesh to mimic manual transect measurements in the field, is between 3 and 6 times larger  
541 for air flow impinging in an east-west direction, than in a north-south direction. As neither

542 approach has been evaluated against independent surface roughness derived from  
543 atmospheric profile measurements over penitentes, the available data is insufficient to  
544 distinguish if maximum effective aerodynamic roughness is associated with wind flowing  
545 across or along the penitente lineation, and the appropriate relationship between wind  
546 direction and surface roughness for calculating turbulent fluxes over penitentes remains  
547 elusive. In principle it sounds reasonable to expect airflow across the penitente lineation to  
548 maximize turbulence as the penitentes present a large surface area to the wind, yet, if  
549 skimming flow is established, with the result that only the tips of the penitentes are  
550 determining the structure of the turbulence then effective roughness in this direction  
551 would be strongly reduced, and perhaps even be less than for air flow along the penitente  
552 lineation, for which the smaller frontal area reduces the likelihood of skimming flow.  
553 Further investigation of this in order to quantify the impact of penitentes on turbulent  
554 fluxes for various airflow patterns would require high resolution turbulence modelling or  
555 direct measurement of aerodynamic roughness and turbulent fluxes over penitentes in all  
556 wind directions.

557 In this study we did not explicitly compute the blending height as available formulae are  
558 dependent upon  $z_0$  and  $z_d$ . Estimates of the blending height independently from  $z_0$  and  $z_d$   
559 have been suggested to be 2.5 - 4.5 times  $h$ , as twice the mean element spacing, or as  
560 combination of the height and spacing (see examples within Grimmond and Oke, 1999).  
561 Given that only atmospheric measurements above the blending height give representations  
562 of integrated surface fluxes and conditions, the first approach would imply that  
563 aerodynamical or flux measurements over penitentes would have to be carried out at  
564 considerable height above the surface to capture mean surface properties rather than the  
565 effects of individual roughness elements. The mathematical model of Claudin and others  
566 (2015) gives a characteristic length scale for the level at which the vapour flux is constant  
567 in horizontal space that is related to the spacing of the penitentes. Interpreting this level as  
568 the blending height implies that the blending height might be determined on the basis of  
569 spacing of penitentes alone, and that this in turn might contain useful data for  
570 understanding the structure and efficiency of turbulence above penitentes. Exploring these  
571 ideas requires information from detailed meteorological measurements as well as the  
572 geometrical information offered in this paper.

## 573 **5. Conclusion**

574 This study demonstrates that the Microsoft Kinect sensor be used successfully at close  
575 range over rough snow and ice surfaces under low light conditions, to generate small-scale  
576 digital surface models useful for assessing morphometry and surface roughness properties  
577 of complex terrain, as well as detailed assessments of spatial variability of surface ablation.  
578 The data collected in this study offers the first detailed study of how the geometry of



579 penitentes evolve through time, highlighting the rate of change of surface properties over  
580 an ablation season that can serve as a guideline for parameterizing surface properties  
581 required for energy and mass balance modelling of penitente surfaces. The method  
582 demonstrated here could be useful for investigating glacier surface features such as  
583 sastrugi, crevasses or meltwater streams and determining the patterns of surface change  
584 associated with such features.

585 Relatively crude manual measurements of penitente surface lowering are shown to be  
586 adequate for quantifying the seasonal mass loss, which is good news for the validity of  
587 existing measurements of surface change on glaciers with penitentes. However, further  
588 measurements and/or modelling studies are required to determine if the mass loss from  
589 the expanded and convoluted surface of penitentes is enhanced or inhibited compared to  
590 mass loss in the absence of penitentes.

591 Aerodynamical roughness properties and related metrics over very rough surfaces remain  
592 poorly quantified and both geometric and meteorological determinations of these values  
593 show a wide spread; consequently it remains unclear what the best methods to use are or  
594 what values modellers would be best to use (Grimmond and Oke, 1999). In this context  
595 further study of penitentes offers a useful opportunity as (a) their morphometric evolution  
596 over time allows various geometries to be evaluated by monitoring a single site, and (b)  
597 they offer a bridge between wind tunnel and urban field experimentation of turbulence and  
598 roughness over extreme terrain. Although validity of surface roughness calculations based  
599 on surface geometry remains to be established for penitentes, this study highlights that (i)  
600 skimming flow is expected to persist over penitentes field, but is more likely under wind  
601 directions perpendicular to the penitente alignment; (ii)  $z_d$  is certainly greater than zero,  
602 and while the depth of penetration of surface layer turbulence into a penitente field is not  
603 clearly established it is likely to evolve with the developing penitentes, and values of  $z_d$   
604  $\sim 2/3h$  give results that are theoretically reasonable in the framework outlined by  
605 Grimmond and Oke (1999); (iii) the two methods of geometric computation of surface  
606 roughness applied here give conflicting results as to whether the effective surface  
607 roughness of penitentes is greater for airflow along or across the penitente lineation and  
608 (iv) more complete understanding of the impact of penitentes on the turbulent structure,  
609 its evolution in time, and its directional dependency, would require atmospheric  
610 measurements with no directional bias concurrent with measurements of penitentes  
611 morphology.

612 Potential future applications and analyses of the surfaces generated in this study include (i)  
613 using surface properties and roughness values as a guide for input into surface energy  
614 balance models; (ii) assessing the performance of models against the measured volume  
615 loss over time and (iii) evaluating how well simplified representations of penitente  
616 surfaces used in small scale radiation models and turbulence models capture the real-

617 world complexity. Such studies would help establish the nature of the likely micro-climatic  
618 distribution of the surface energy balance within a real penitente field, and as a result the  
619 impact of penitentes on runoff and exchange of water vapour with the atmosphere.

620 **Author contributions.** LN designed the study. Fieldwork was carried out by LN and BP with MP providing  
621 the TLS data. TLS and AWS equipment was provided by SM through collaboration with CEAZA. The data was  
622 analysed by LN and MP. Preparation of the manuscript and figures was led by LN with contributions from all  
623 co-authors.

624 **Acknowledgements.** Fieldwork for this study was funded by a National Geographic Waitt Grant awarded to  
625 L N and S M. LN was supported by an Austrian Science Fund Elise Richter Grant (V309). MP was supported  
626 within statutory activities No 3841/E-41/S/2016 of the Ministry of Science and Higher Education of Poland.  
627 International cooperation was supported by the Centre for Polar Studies from the funds of the Polish Leading  
628 National Research Centre (KNOW) in Earth Sciences (2014–18). Thanks are also due to Mathias Rotach for  
629 reading the paper prior to submission.

## 630 **References**

- 631 Amstutz, G. C. (1958) On the formation of snow penitentes. *Journal of Glaciology*, 3 (24), 304-311.
- 632 Andreas, E. L. (2011). A relationship between the aerodynamic and physical roughness of winter sea ice.  
633 *Quarterly Journal of the Royal Meteorological Society*, 137(659), 1581–1588. doi:10.1002/qj.842
- 634 Bergeron, V., Berger, C., & Betterton, M. D. (2006). Controlled irradiative formation of penitentes. *Physical*  
635 *Review Letters*, 96(9), 098502, doi:10.1103/PhysRevLett.96.098502
- 636 Brutsaert, W. (1975). A theory for local evaporation (or heat transfer) from rough and smooth surfaces at  
637 ground level. *Water Resources Research*, 11(4), 543–550.
- 638 Carrasco JF, Osorio R, Casassa G. 2008. Secular trend of the equilibrium-line altitude on the western side of  
639 the southern Andes, derived from radiosonde and surface observations. *Journal of Glaciology*, 54(186): 538-  
640 550, doi: 10.3189/002214308785837002.
- 641 Cathles, L. M., Abbot, D. S., & MacAyeal, D. R. (2014). Intra-surface radiative transfer limits the geographic  
642 extent of snow penitents on horizontal snowfields. *Journal of Glaciology*, 60(219), 147–154.  
643 doi:10.3189/2014JG13J124
- 644 Claudin, P., Jarry, H., Vignoles, G., Plapp, M., & Andreotti, B. (2015). Physical processes causing the formation  
645 of penitentes. *Physical Review E*, 92(3), 033015. doi:10.1103/PhysRevE.92.033015
- 646 Corripio, J. G., & Purves, R. S. (2005). Surface energy balance of high altitude glaciers in the Central Andes: the  
647 effect of snow penitentes. In: De Jong C., Collins D.N. and Ranzi, R. (Eds) *Climate and Hydrology in Mountain*  
648 *Areas*. Wiley and Sons, Chichester, 15-27. Drewry, D. J. (1970). Snow penitents. *Weather*, 25(12), 556.
- 649 Drewry, D. J. (1970). Snow penitents. *Weather*, 25(12), 556.
- 650 Escobar F., & Aceituno P. 1998. The ENSO phenomenon influence on snowfall in the Andean sector of central  
651 Chile during winter. *Bull. Inst. Fr. Etud. Andines* 27(3): 753-759.

652 Fassnacht, S. R., Stednick, J. D., Deems, J. S., & Corrao, M. V. (2009). Metrics for assessing snow surface  
653 roughness from Digital imagery. *Water Resources Research*, 45, W00D31 doi:10.1029/2008WR006986

654 Ginot, P., Kull, C. Schotterer, U., Schwikowski, M. and Gäggeler, H. W. (1999) Glacier mass balance  
655 reconstruction by sublimation of chemical species on Cerro tapado (Chilean Andes). *Climate of the Past*, 2, 21-  
656 30.

657 Grimmond, C. S. B., & Oke, T. R. (1999). Aerodynamic Properties of Urban Areas Derived from Analysis of  
658 Surface Form. *Journal of Applied Meteorology*, 38(9), 1262–1292.

659 Hastenrath, S., & Koci, B. (1981). Micro-morphology of the snow surface at the Quelccaya ice cap, Peru. *Journal*  
660 *of Glaciology*, 27(97), 423–428.

661 Kaser, G., Großhauser, M., & Marzeion, B. (2010). Contribution potential of glaciers to water availability in  
662 different climate regimes. *Proceedings of the National Academy of Sciences*, 107(47), 20223–20227.  
663 doi:10.1073/pnas.1008162107

664 Kondo, J., & Yamazawa, H. (1986). Aerodynamic roughness over an inhomogeneous ground surface.  
665 *Boundary-Layer Meteorology*, 35(1983), 331–348.

666 Lettau, H. (1969). Note on Aerodynamic Roughness-Parameter Estimation on the Basis of Roughness-Element  
667 Description. *Journal of Applied Meteorology*, 8(5), 828-832.

668 Lhermitte, S., Abermann, J., & Kinnard, C. (2014). Albedo over rough snow and ice surfaces. *The Cryosphere*,  
669 8(3), 1069–1086. doi:10.5194/tc-8-1069-2014

670 Lliboutry, L. (1954). The origin of penitents. *Journal of Glaciology*, 2, 331–338.

671 Lliboutry, L. (1998). Glaciers of Chile and Argentina. In, R. S. Williams and J. G. Ferrigno (Ed). Satellite image  
672 atlas of glaciers of the world: South America. USGS Professional Paper 1386-I.

673 Macdonald, R. W., Griffiths, R. F. F., & Hall, D. J. J. (1998). An improved method for the estimation of surface  
674 roughness of obstacle arrays. *Atmospheric Environment*, 32(11), 1857–1864. doi:10.1016/S1352-  
675 2310(97)00403-2

676 MacDonell, S., Kinnard, C., Mölg, T., Nicholson, L. I., & Abermann, J. (2013). Meteorological drivers of ablation  
677 processes on a cold glacier in the semi-arid Andes of Chile. *The Cryosphere*, 7(5), 1513–1526. doi:10.5194/tc-  
678 7-1513-2013

679 Mankoff, K. D., & Russo, T. A. (2013). The Kinect: a low-cost, high-resolution, short-range 3D camera. *Earth*  
680 *Surface Processes and Landforms*, 38(9), 926–936. doi:10.1002/esp.3332

681 Manninen, T., Anttila, K., Karjalainen, T., & Lahtinen, P. (2012). Automatic snow surface roughness estimation  
682 using digital photos. *Journal of Glaciology*, 58(211), 993–1007. doi:10.3189/2012JoG11J144

683 Munro, D. S. (1989). Surface roughness and bulk heat transfer on a glacier: comparison to eddy correlation.  
684 *Journal of Glaciology*, 35(121), 343–348.

685 Munro, D. S. (1990). Comparison of Melt Energy Computations and Ablatometer Measurements on Melting Ice  
686 and Snow. *Arctic, Antarctic, and Alpine Research*, 22(2), 153–162. doi:10.2307/1551300

687 Naruse, R. and Leiva, J. C. (1997) Preliminary study on the shape of snow penitentes at Piloto Glacier, the  
688 central Andes. *Bulletin of Glacier Research*, 15, 99-104.

689 Obleitner, F. (2000). The energy budget of snow and ice at Breidamerkurjökull, Vatnajökull, Iceland.  
690 *Boundary-Layer Meteorology*, 97(3), 385–410.

691 Schotterer U, Grosjean M, Stichler W, Ginot P, Kull C, Bonnaveira H, Francou B, Gäggeler HW, Gallaire R,  
692 Hoffman G, Pouyaud B, Ramirez E, Schwikowski M & Taupin JD. 2003. Glaciers and climate in the Andes  
693 between the equator and 30°S: What is recorded under extreme environmental conditions? *Climatic Change*,  
694 59(1-2), 157-175, doi:10.1023/A:1024423719288.

695 Sinclair, K. & MacDonell, S. (2016) Seasonal evolution of penitente geochemistry at Tapado Glacier, northern  
696 Chile. *Hydrological Processes*, 30(2), 176-186, doi: 10.1002/hyp.10531.

697 Smeets, C. J. P. P., Duynkerke, P., & Vugts, H. (1999). Observed wind profiles and turbulence fluxes over an ice  
698 surface with changing surface roughness. *Boundary-Layer Meteorology*, 92(1), 99–121.

699 Winkler, M., Juen, I., Mölg, T., Wagnon, P., Gomez, J., & Kaser, G. (2009). Measured and modelled sublimation  
700 on the tropical Glaciar Artesonraju, Peru. *The Cryosphere*, 3(1), 21–30.

701 **Supplementary material**

- 702
- 703 • A: GPS position of ground control points at each glacier site
  - 704 • B: Mesh surface components and processing steps used for Kinect surface scans
  - 705 • C: Kinect surface meshes for both sites on all dates as .PLY files [sX\_DDMM.PLY]
  - 706 • D: 3D viewer files of surfaces at site B can be seen at:  
707 <https://sketchfab.com/LindseyNicholson/folders/penitentes-on-glaciar-tapado-chile>  
708
- 709

710 *Table 1: Maximum absolute georeferencing error at each marker stake for site A and B, relative to the standard*  
 711 *deviation of the differential GPS measurement.*

	$\Delta X$ [mm]	$\Delta Y$ [mm]	$\Delta Z$ [mm]	$\Delta XY$ [mm]	$\Delta XYZ$ [mm]	<i>dGPS XYZ standard deviation [mm]</i>
<b>A-1</b>	63	25	38	68	77	17
<b>A-2</b>	214	118	259	233	312	15
<b>A-3</b>	14	57	53	57	62	14
<b>A-4</b>	23	29	61	33	69	16
<b>A-5</b>	54	32	128	56	139	18
<b>B-1</b>	59	46	19	75	77	16
<b>B-2</b>	121	11	102	164	193	17
<b>B-3</b>	11	48	2	49	49	12
<b>B-4</b>	85	37	34	85	92	12

712

713

714 *Table 2: Mean meteorological conditions during the measurement intervals: incoming shortwave (SW in), albedo*  
 715 *( $\alpha$ ), incoming longwave (LW in), windspeed (u), wind direction (dir), surface temperature computed from*  
 716 *measured outgoing longwave radiation (T surface), air temperature (T air), relative humidity (RH), air pressure*  
 717 *(P) and the distance between the sonic ranger and the glacier surface (dist).*

	<b>SW in</b>	<b><math>\alpha</math></b>	<b>LW in</b>	<b>u</b>	<b>dir</b>	<b>T surface</b>	<b>T air</b>	<b>RH</b>	<b>P</b>	<b>dist</b>
	[W m <sup>-2</sup> ]	[-]	[W m <sup>-2</sup> ]	[m s <sup>-1</sup> ]	[°]	[°C]	[°C]	[%]	[hPa]	[m]
<b>sensor</b>	<i>Kipp and Zonen CNR1</i>			<i>Young 05103</i>		CNR1	<i>Vaisala HMP45</i>		<i>Setra 278</i>	<i>SR50</i>
<b>26/12 - 11/12</b>	413	0.54	205	3.0	170	-5.3	-2.7	32.5	442	1.62
<b>12/12 - 20/12</b>	441	0.48	212	2.8	214	-2.9	-0.8	41.4	448	1.96
<b>21/12 - 03/01</b>	426	0.41	224	3.1	217	-1.4	1.9	39.5	456	2.56

718

719

720 *Table 3: Surface roughness ( $z_0$ ) computed according to Munro (1989) on detrended profiles longer than 1.5 m,*  
 721 *extracted at 0.10 m intervals from the Kinect surface meshes at site A and B for E-W impinging wind and N-S*  
 722 *impinging wind. The number of profiles used for each wind direction is given in parenthesis. The likely*  
 723 *displacement of the zero velocity plane ( $d_{top} \pm$  standard deviation), was computed as the mean of  $2/3h$  for all*  
 724 *profiles and expressed as a distance from the top of the penitentes. The range of the detrended 3D mesh (3D*  
 725 *range) provides a reference for the penetration depth of turbulence.*

	site A						site B					
	$z_0$ E-W (20)			$z_0$ N-S (33)			$z_0$ E-W (6)			$z_0$ N-S (7)		
	mean	max	min	mean	max	min	mean	max	min	mean	max	min
<b>25-Nov</b>	45	111	11	8	19	3						
<b>11-Dec</b>	33	68	12	6	13	2	28	41	22	6	9	1
<b>20-Dec</b>	70	146	57	25	67	7	122	156	84	22	47	14
<b>03-Jan</b>	136	211	71	45	136	11	133	186	101	21	30	12
	3D range [m]	$d_{top} \pm$ std [m]		3D range [m]	$d_{top} \pm$ std [m]		3D range [m]	$d_{top} \pm$ std [m]		3D range [m]	$d_{top} \pm$ std [m]	
<b>25-Nov</b>	0.41	0.27	0.06	0.41	0.34	0.02						
<b>11-Dec</b>	0.48	0.33	0.05	0.48	0.41	0.01	0.58	0.45	0.02	0.58	0.51	0.02
<b>20-Dec</b>	0.76	0.58	0.03	0.76	0.61	0.04	0.98	0.76	0.02	0.98	0.84	0.04
<b>03-Jan</b>	1.07	0.79	0.03	1.07	0.86	0.05	1.14	0.86	0.03	1.14	0.98	0.02

726

727

728

729

730 Figure 1: Map of Tapado Glacier in the Elqui catchment of the Coquimbo Region of Chile, showing  
731 the location of the measured sites and insets of (a) the glacier site layout, showing the location of  
732 the horizontal reference (black line) and (b) the test site, indicating the boulder (red star) at which  
733 the Kinect scans were compared against TLS

734 Figure 2: (a) Oblique view of the TLS-derived DSM of the test site highlights the patchy coverage of  
735 the penitentes obtained by this method. (b) Absolute differences between DSMs of the sample  
736 boulder produced using TLS and Kinect.

737 Figure 3: Shaded DSM meshes of N-S orientated DSMs for the 1.5 m x 1.5 m glacier site B on (a)  
738 12.12.2013 (b) 20.12.2013 and (c) 03.01.2013 obtained using the Kinect.

739 Figure 4: Summary of the DSM properties through time at site A (left) and B (right). (a,b) Surface  
740 height distribution as a percentage of total surface area, in local coordinates [m] relative to the  
741 position of the northern end of ablation frame. Inset tables show weighted mean mesh elevation,  
742 range, surface area and surface area as a function of the horizontal area of the sampled site. (c,d)  
743 Distribution of surface angles as a percentage of total surface area. (e,f) Aspect distribution as a  
744 percentage of total surface area.

745 Figure 5: Comparison of surface height through time from manual measurements (points) and  
746 extracted from the Kinect scans (solid lines  $\pm$  vertical error) along the horizontal reference (site A,  
747 Figure 1). Triangles indicate original snow depth compared to the surface measured on 25/11/13  
748 and solid black triangles indicate locations where snowdepth exceeded the length of the 3 m probe

749 Figure 6: Representative surface heights computed on detrended surface meshes for site A (solid)  
750 and site B (open) over time where  $h_1$ - $h_4$  refer to representative surface heights computed as range  
751 ( $h_1$ ), twice the standard deviation ( $h_2$ ), area weighted mean height above the minimum ( $h_3$ ), and  
752 area weighted median above the minimum mesh height ( $h_4$ ).

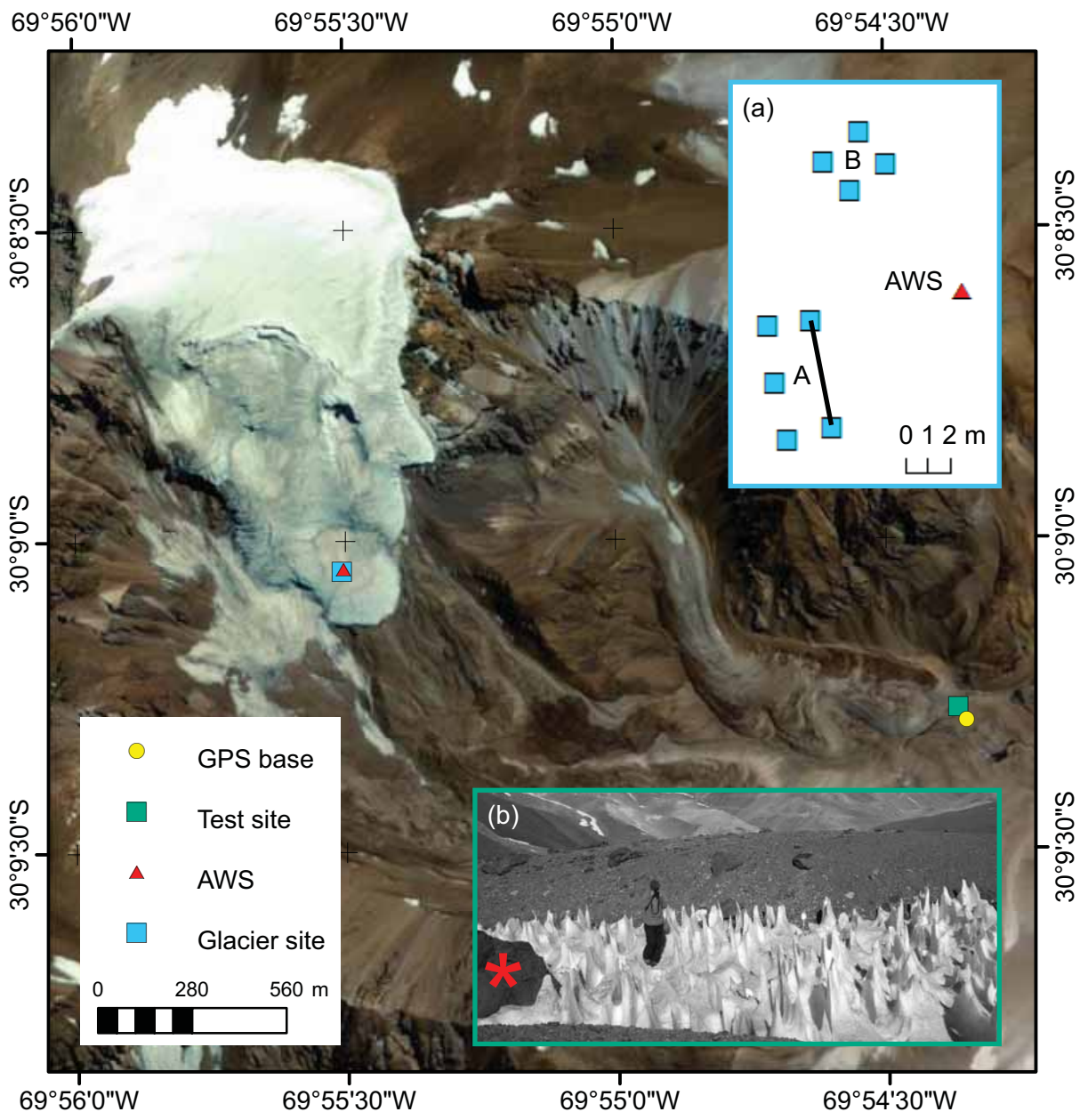
753 Figure 7: 3D  $z_0$  computed for  $10^\circ$  aspect intervals for all detrended DSMs highlighting peak  
754 roughness occurs in N-S airflow. Maximum values take  $h$  to be the detrended mesh elevation range,  
755 and minimum values take  $h$  to be twice the standard deviation of the detrended mesh.

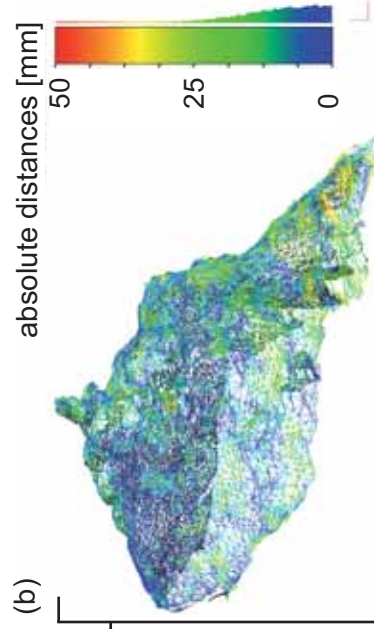
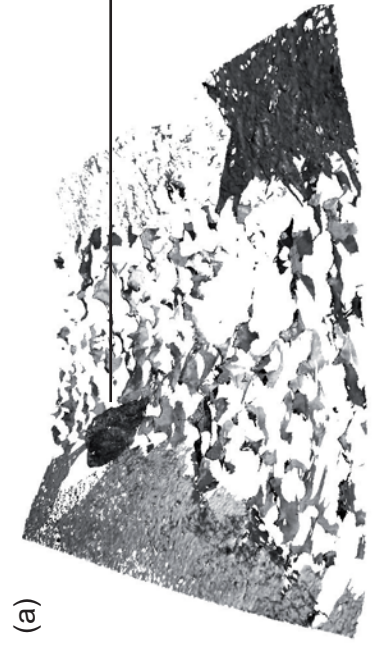
756 Figure 8: Comparison of three-dimensional surface roughness through time, indicating the range of  
757  $z_0$  computed for all incident wind angles (at  $10^\circ$  intervals). Upper panels show the roughness with  
758 no zero level displacement and lower panels show values with a zero displacement offset  $d_1 = h$ ;  $d_2$   
759  $= 2/3h$  and  $d_3 = 1/3h$ . As before,  $h_1$ -  $h_4$  refer to representative surface heights computed as range,  
760 twice the standard deviation, area weighted mean height above the minimum, and area weighted  
761 median above the minimum mesh height respectively.

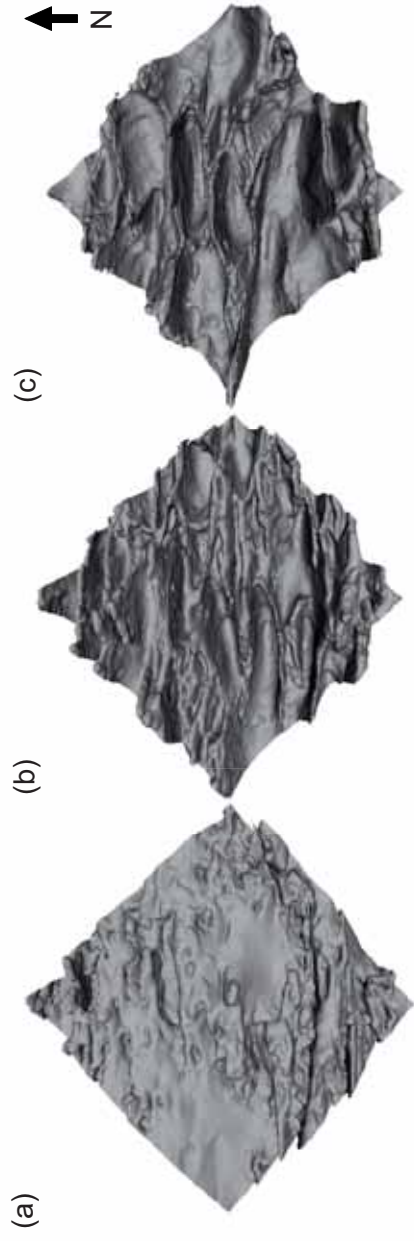
762 Figure 9: Examples of (a) N-S, and (b) E-W orientated cross sections longer than 1.5 m, sampled at  
763 0.1 m intervals from which effective surface roughness properties were computed using the  
764 methods of Munro (1989, 1999). The local coordinates are relative to the NE corner marker of site A  
765 (Fig 1).

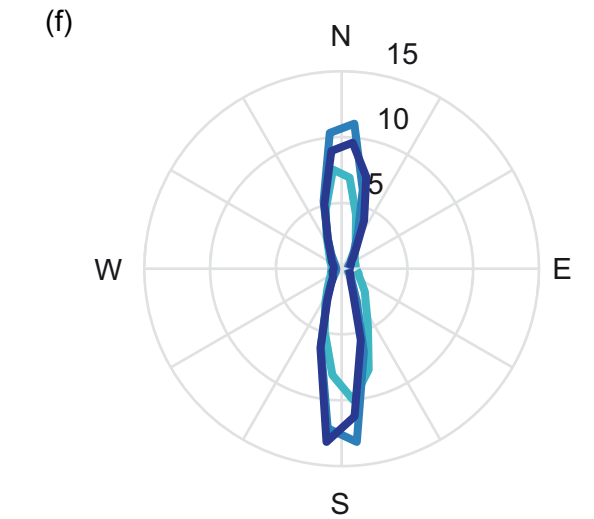
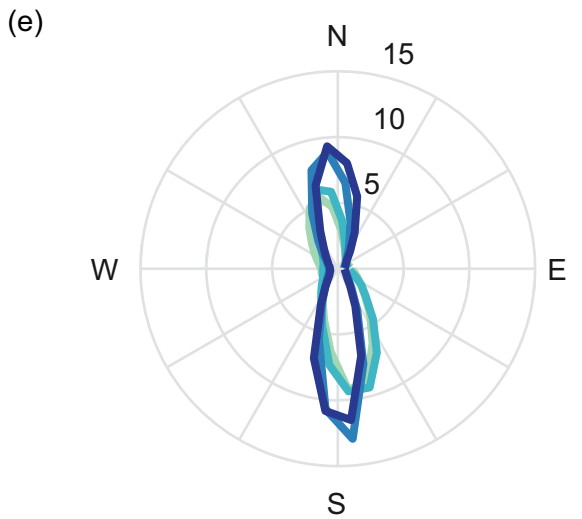
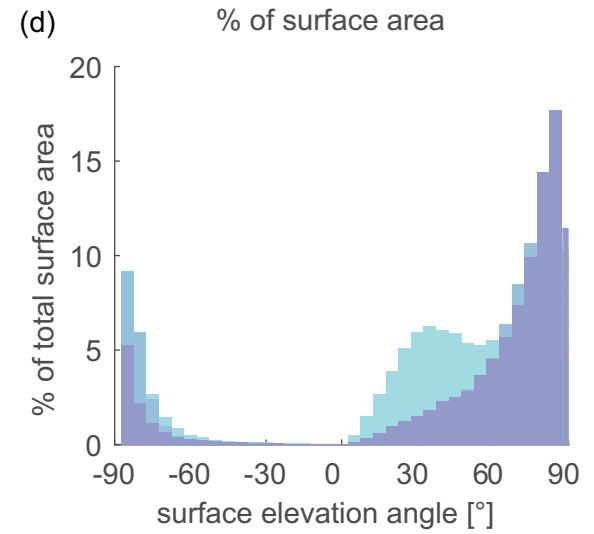
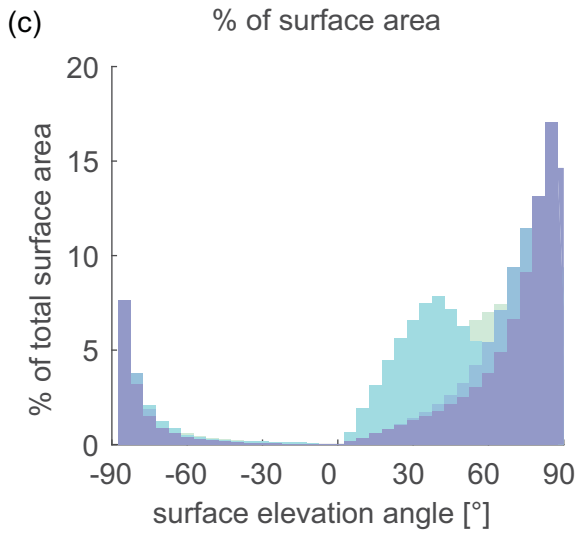
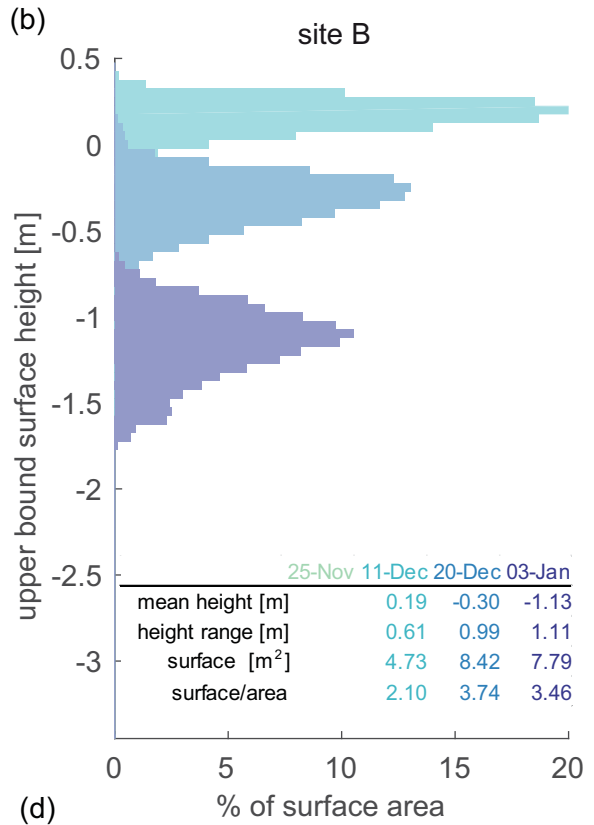
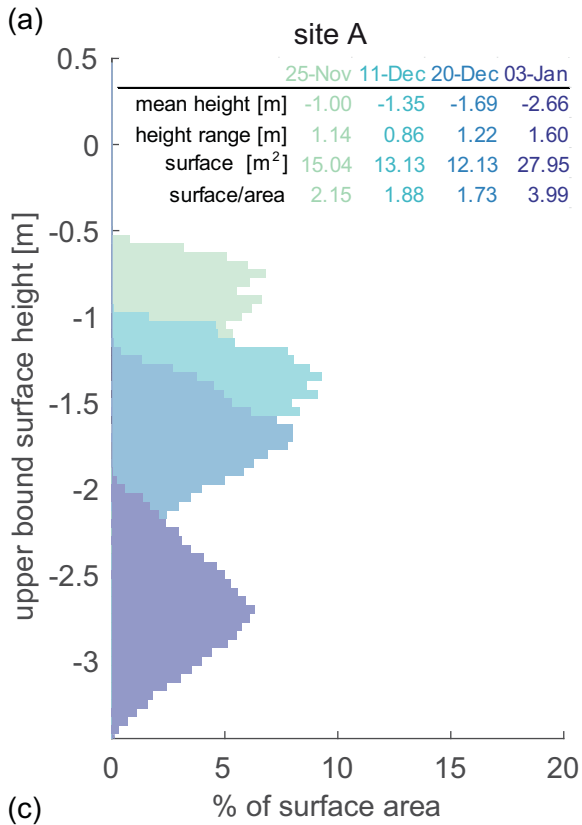
766 Figure 10: Wind rose for the whole study period (26 Nov 2013 – 3 Jan 2014).

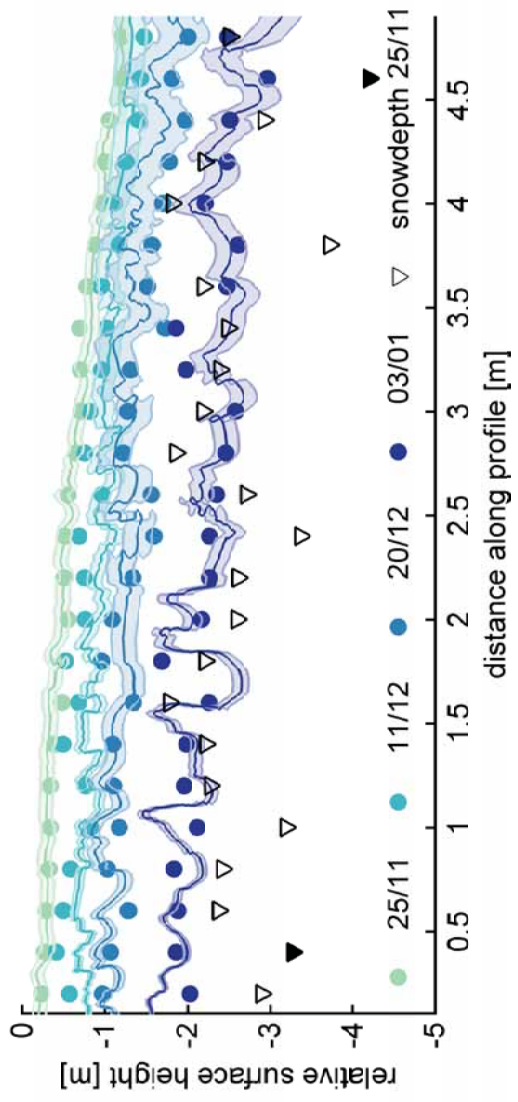


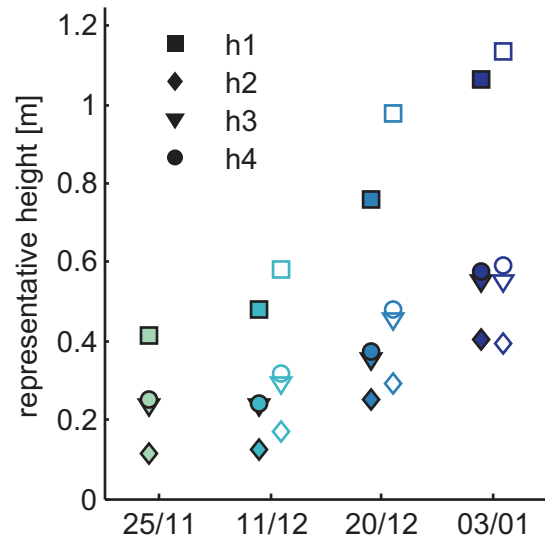


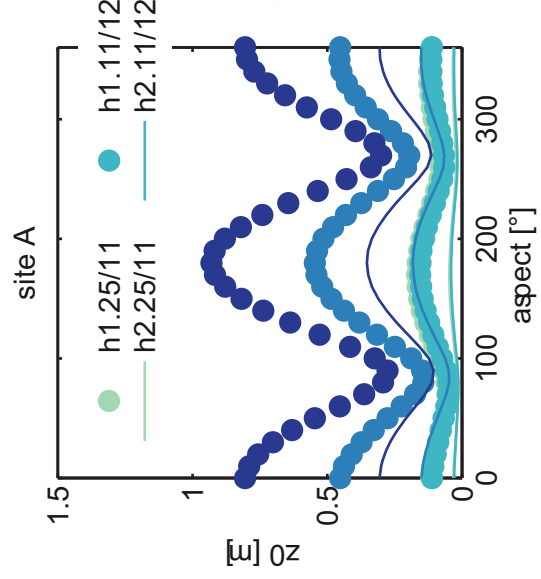
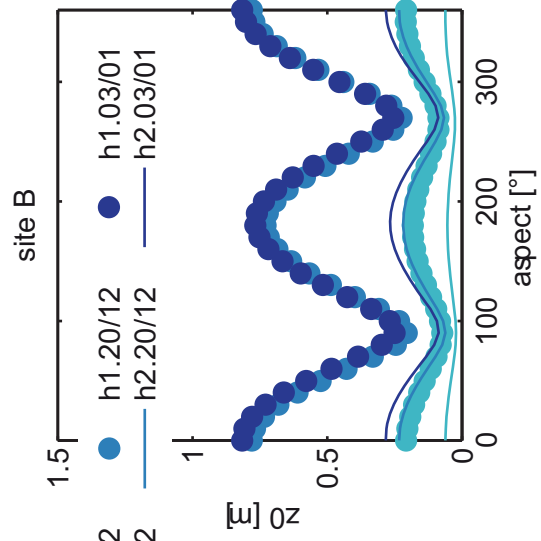


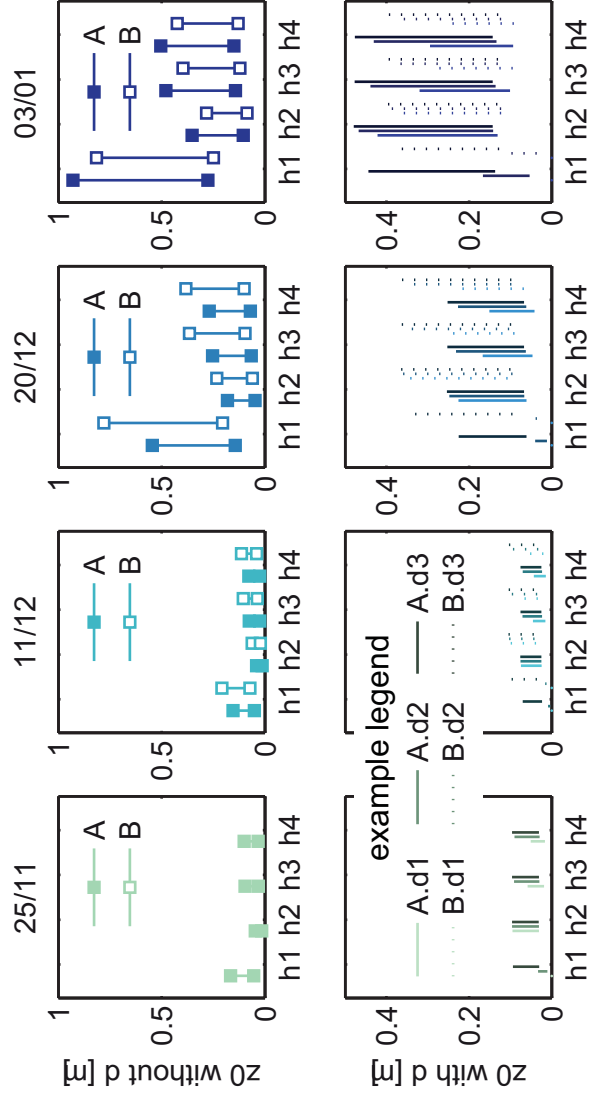




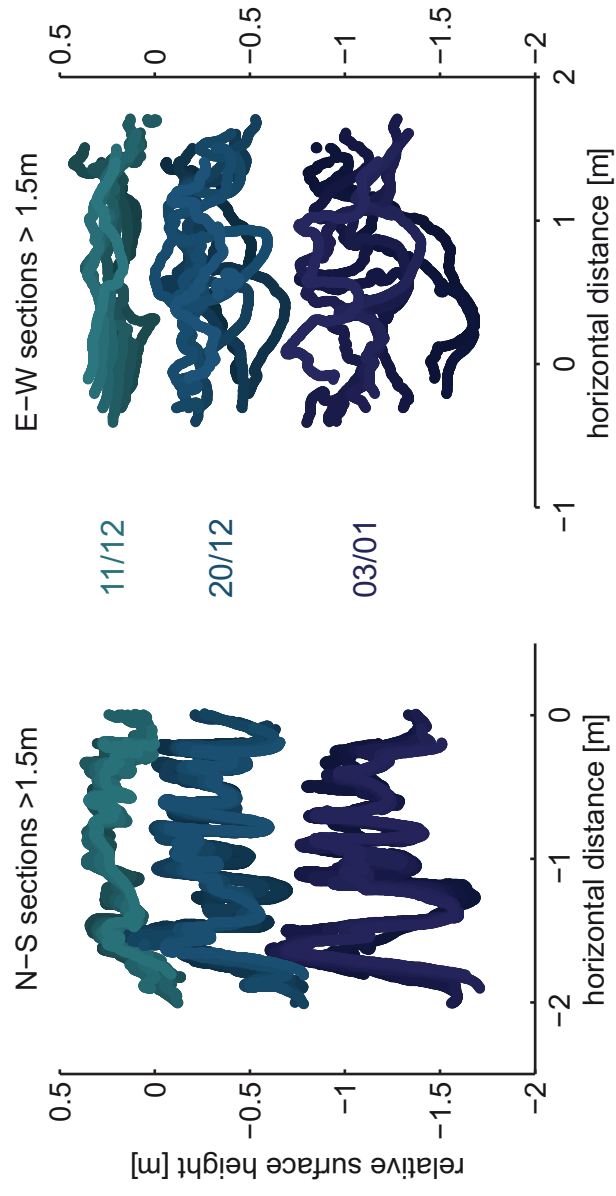












26.11.13 - 03.01.2014

