

Comments to the author:

Dear Authors,

Thank you for submitting a revised version of your manuscript. You adequately addressed the issues raised by the reviewers. However, one question remains, which concerns the Monte Carlo simulations used to derive confidence intervals for the timing of seasonal maxima and minima (Section 2.4.4). Ideally, such a simulation would start from an error-free seasonal cycle, add pseudo-random noise, and do statistics about the timing of the extrema of the noisy cycle.

How did you do it? I guess you started from the GRACE-based seasonal cycle (which already contains noise) and then added additional pseudo-random noise.

Then the question would be how the actual GRACE noise affects the results. For example in Fig. 11b I would be concerned that the maximum at February is pronounced due to a little random spike and that this random spike causes the Monte-Carlo-based 50% confidence interval to be so narrow.

Typos:

p. 47 averge --> average

p. 20 line 20 ransect --> transect

Dear Dr. Eisen,

Thanks very much for your comments. We have corrected the typos that you have discovered. Regarding the calculation of uncertainty bounds on the timing of the GRACE cycle, we have indeed performed the Monte Carlo simulations as you have noted. For each mascon, we started from the GRACE seasonal cycle derived from GRACE-LM processing, and added pseudo-random noise for a normal distribution with a standard deviation equal to the stated error for that mascon from the GRACE dataset. Our approach is to add the errors to the GRACE-LM distribution to achieve an ensemble of possible “true” seasonal cycles. While none of these represents the true seasonal cycle, they represent a statistical distribution of possible cycles given the estimated error for each mascon. While we agree that noise from the original timeseries could potentially influence the distribution, with this approach, any noise that is within the error range from the GRACE timeseries should have a minimal effect. Any fluctuations that are larger than the estimated error will have an effect on the timing of the signal, but we consider these to be part of the “real” signal from GRACE, i.e. associated with variability in precipitation, runoff, or ice dynamics. We could smooth our initial GRACE seasonal cycle to remove these fluctuations, but we believe that doing this could possibly attenuate a real signal in the GRACE data. For this reason we prefer to adhere to the approach we have taken.

In the course of responding to your question, however, we discovered that one aspect of our procedure is perhaps incorrect. Namely, we added random error to the GRACE-LM timeseries for each day of the year. The stated error for each mascon is for the original GRACE-LM timeseries, which has a temporal resolution of roughly 10 days. Applying the 10-day error at a daily timescale is an underestimate, because the error on a shorter timescale should be higher than the error over longer periods, and therefore fluctuations within the error range in the original timeseries *do* influence the timing of peaks. We have found that applying the error at a temporal resolution of 10 days widens the distribution of seasonal cycle timing for the maximum and minimum of the cycle, in cases where the noise is relatively high in relation to the signal, for example, in the case of Fig. 11b. Therefore, we have decided to repeat our Monte Carlo simulations, applying error at a 10-day temporal resolution. As a result our sub-regions of the ice sheet are slightly different, and figures for the sub-regions have been updated. While the distribution on maximum and minimum peaks is widened, our model peaks are still outside the bounds of the distribution in areas of Northwest Greenland, and the conclusions have not changed.

Section 2.4.4 has been updated to indicate this change in the Monte Carlo procedure. Sections 3.3 and 3.4 have been modified in accordance with the changes to the figures. One sentence in the abstract has been slightly modified. Tables 1, 2, and S1, and Figures 5, 8-12, and S6-S8 have been modified.

Sincerely,  
Patrick Alexander  
On behalf of all Co-authors

# 1 **Greenland Ice Sheet seasonal and spatial mass variability** 2 **from model simulations and GRACE (2003-2012)**

3  
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## 23 24 **Abstract**

25 Improving the ability of regional climate models (RCMs) and ice sheet models (ISMs) to  
26 simulate spatiotemporal variations in the mass of the Greenland Ice Sheet (GrIS) is crucial for  
27 prediction of future sea level rise. While several studies have examined recent trends in GrIS  
28 mass loss, studies focusing on mass variations at sub-annual and sub-basin-wide scales are

1 still lacking. At these scales, processes responsible for mass change are less well understood  
2 and modeled, and could potentially play an important role in future GrIS mass change. Here,  
3 we examine spatiotemporal variations in mass over the GrIS derived from the Gravity  
4 Recovery and Climate Experiment (GRACE) satellites for the January 2003 – December 2012  
5 period using a “mascon” approach, with a nominal spatial resolution of 100 km, and a  
6 temporal resolution of 10 days. We compare GRACE-estimated mass variations against those  
7 simulated by the Modèle Atmosphérique Régionale (MAR) RCM and the Ice Sheet System  
8 Model (ISSM). In order to properly compare spatial and temporal variations in GrIS mass  
9 from GRACE with model outputs, we find it necessary to spatially and temporally filter  
10 model results to reproduce leakage of mass inherent in the GRACE solution. Both modeled  
11 and satellite-derived results point to a decline (of  $-178.9 \pm 4.4$  and  $-239.4 \pm 7.7$  Gt yr<sup>-1</sup>  
12 respectively) in GrIS mass over the period examined, but the models appear to underestimate  
13 the rate of mass loss, especially in areas below 2000 m in elevation, where the majority of  
14 recent GrIS mass loss is occurring. On an ice-sheet wide scale, the timing of the modeled  
15 seasonal cycle of cumulative mass (driven by summer mass loss) agrees with the GRACE-  
16 derived seasonal cycle, within limits of uncertainty from the GRACE solution. However, on  
17 sub-ice-sheet-wide scales, ~~there are some areas exhibit~~ significant differences in the timing of  
18 peaks in the annual cycle of mass change. At these scales, model biases, or unaccounted-for  
19 processes related to ice dynamics or hydrology may lead to the observed differences. This  
20 highlights the need for further evaluation of modelled processes at regional and seasonal  
21 scales, and further study of ice sheet processes not accounted for, such as the role of sub-  
22 glacial hydrology in variations in glacial flow.

23

## 24 **1 Introduction**

25 The Earth’s ice sheets represent substantial reservoirs of water stored in the form of ice,  
26 which contribute to fluctuations in global sea level. The Greenland Ice Sheet (GrIS) in  
27 particular is estimated to have lost mass at an average rate of  $-142 \pm 49$  Gt yr<sup>-1</sup> between 1992  
28 and 2011 (Shepherd et al., 2012). Roughly 50% of recent GrIS mass loss is associated with  
29 surface mass loss (Rignot et al., 2011; van den Broeke et al., 2009), characterized by multiple  
30 records in GrIS melt extent and duration over the past decade (Tedesco et al., 2008, 2011,  
31 2013a; Nghiem et al., 2012) which has led to increased meltwater runoff that exceeds small  
32 increases in ice-sheet-wide precipitation (van den Broeke et al., 2009; Ettema et al., 2009;

1 Fettweis et al., 2013a). The other portion of GrIS mass loss is associated with an acceleration  
2 of outlet glaciers (Rignot et al., 2011). The speedup of glaciers has been attributed to  
3 warming oceans (Rignot et al., 2012) and lubrication of the GrIS bed from meltwater  
4 generated at the surface, and channeled from the surface to the bed by vertical conduits,  
5 allowing glaciers to slide more easily (Zwally et al., 2002). This second factor has been  
6 shown to be more complex than initially thought, resulting in speed-ups or slow-downs that  
7 depend on the volume of meltwater reaching the bed and the time of year (e.g. Sundal et al.,  
8 2011).

9 Previous studies have generally focused on decadal trends in GrIS mass and the ability of  
10 models to capture these trends (e.g. Shepherd et al., 2012; Rignot et al., 2011), but seasonal  
11 variations in mass, and spatial variations at sub-basin-wide scales have not been explored  
12 extensively. At smaller temporal and spatial scales, poorly understood processes may play a  
13 particularly important role in mass variability. For example, numerous studies have identified  
14 seasonal variations in glacial flow (Bartholomew et al., 2010; Howat et al., 2010; Joughin et  
15 al., 2008, 2014; Moon et al., 2014), and local variations in flow associated with lake drainage  
16 events or summer melting (Das et al., 2008; Tedesco et al., 2013b; Hoffman et al., 2011). Ice  
17 sheet hydrology can also contribute to local variations in mass. For example, an observational  
18 study by Rennermalm et al. (2013) suggests that within one catchment along the GrIS coast,  
19 up to 50% of runoff generated at the surface may have been stored within the ice sheet over  
20 multiple seasons. Water can also be stored at or near the surface of the ice sheet, within  
21 supraglacial lakes, or within firn aquifers, which were recently discovered to persist during  
22 winter over large areas of the southwest and southeast GrIS margins (Forster et al., 2013;  
23 Koenig et al., 2014). While the amount of water stored within supraglacial lakes is likely  
24 small relative to the overall rate of GrIS mass change (Smith et al., 2015), the amount of  
25 water stored within the firn aquifers or englacially is currently unknown.

26 The overall GrIS mass balance (MB), the rate of ice sheet mass change, is generally  
27 considered to consist of two components, the Surface Mass Balance (SMB, i.e. the balance  
28 between accumulation and ablation at the ice sheet surface), and ice discharge (D), such that  
29  $MB = SMB - D$ . Simulations of SMB at high spatial and temporal resolutions (e.g. daily  
30 temporal resolution and <25 km spatial resolution) are conducted by Regional Climate  
31 Models (RCMs; e.g. Fettweis et al., 2013a; Ettema et al., 2009), and D can be simulated by  
32 Ice Sheet Models (ISMs) (e.g. Larour et al., 2012; Quiquet et al., 2012; Robinson et al., 2011;

1 Huybrechts et al., 2011), which simulate glacial flow subject to SMB forcing. At seasonal  
2 and sub-basin-wide scales other processes become more important, so that the full mass  
3 balance is expressed as follows (after Cuffey and Paterson, 2011):

$$4 \quad MB = SMB + EMB + BMB + DMB \quad (1)$$

5 where EMB is the englacial mass balance, BMB is the basal mass balance, and DMB is the  
6 mass balance associated with dynamic flow. Most processes related to EMB or BMB, as  
7 well as variations in DMB associated with ice-ocean interactions and meltwater lubrication  
8 are not accounted for by either RCMs or ISMs. In a warmer climate, more meltwater  
9 production and runoff is expected (Fettweis et al., 2013b), suggesting that unaccounted for  
10 processes will play an increasingly important role in future GrIS mass balance (Chu, 2014),  
11 and should be included in model projections of mass change.

12 Cognizant of the potential role of such processes in GrIS MB, and the need for evaluation of  
13 the combined results of ISMs and RCMs, we conducted a comparison between satellite-  
14 derived mass changes from the Gravity Recovery and Climate Experiment (GRACE; Luthcke  
15 et al., 2013), modeled DMB from the Ice Sheet System Model (ISSM; Larour et al., 2012),  
16 and SMB from the Modèle Atmosphérique Régionale (MAR; e.g. Fettweis et al., 2013a) for  
17 the period January 2003 through December 2012. The GRACE solution of Luthcke et al.  
18 (2013) (hereafter referred to as GRACE-LM) is provided at a high spatial and temporal  
19 resolution compared to other GRACE solutions (~100 km and 10 days respectively). We  
20 aggregate model results to the GRACE-LM grid and spatially and temporally filter the  
21 aggregated outputs in order to match inherent spatial and temporal attenuation of the  
22 GRACE-LM product (as discussed by Luthcke et al., 2013; Sabaka et al., 2010). After  
23 filtering model outputs, we compared spatial patterns of simulated and satellite-derived mean  
24 annual mass balance, the mean annual cycle of mass change, and the spatial distribution of the  
25 timing of the seasonal cycle. This analysis has two purposes: (1) to evaluate seasonal and  
26 spatial variations in mass from the combined results of an RCM and ISM applied over the  
27 GrIS, and (2) to reveal and analyse any discrepancy between GRACE-derived and modeled  
28 mass changes, while accounting for uncertainties associated with the GRACE-LM solution.

## 1 **2 Data and Methods**

### 2 **2.1 GRACE data**

3 We used the iterated global GRACE solution of Luthcke et al. (2013), which utilizes a mass  
4 concentration (mascon) approach to derive spatially and temporally distributed changes in the  
5 mass of land ice, at a 1 arc-degree (~100 km) spatial resolution and 10-day temporal  
6 resolution. The solution is available for the period January 2003 through June 2013, but we  
7 focus on the January 2003 through December 2012 period to avoid including an incomplete  
8 year in our analysis. GRACE-LM mass change estimates are provided for ~100x100 km<sup>2</sup>  
9 “mascon” regions, on what is essentially an equal area grid (shown in Fig. 1a). All GRACE  
10 solutions are ultimately derived from k-band range and range rate (KBRR) data for two co-  
11 orbiting satellites roughly 220 km apart (Tapley et al., 2004). The Luthcke et al. (2013)  
12 solution, used in this study, differs from other solutions in its approach: models of satellite  
13 motion are used to compute KBRR from forward-modeled mass changes, and through  
14 iteration, the residuals between the computed and observed KBRR are minimized. This  
15 contrasts with the spherical harmonic approach (e.g. Velicogna and Wahr, 2006) in which a  
16 set of Stokes coefficients or spherical harmonic fields provided by GRACE processing centers  
17 are spatially filtered and used to estimate spatial and temporal variations in mass. The  
18 mascon approach of Luthcke et al. (2013) attempts to minimize the loss of signal associated  
19 with processing GRACE data, and detailed error estimates, accounting for various steps in  
20 processing, are provided. The Luthcke et al. (2013) solution agrees within error estimates  
21 with other estimates for GrIS mass change derived from GRACE. As described by Luthcke et  
22 al. (2013), forward modelling is used during the processing of GRACE data to isolate the  
23 signal associated with land-ice changes. In particular, the static gravity field, orbital  
24 parameters, ocean and earth tides, terrestrial water storage, variations in mass associated with  
25 atmospheric and ocean circulation, and glacial isostatic adjustment are simulated by various  
26 models, and these simulated changes are used to correct GRACE-estimated mass change.  
27 The errors associated with each of these simulations are included in calculations of error for  
28 each GRACE-LM mascon. The GRACE-LM mascons are distributed at a resolution that is  
29 higher than the fundamental spatial resolution of GRACE (Luthcke et al., 2006), so that there  
30 is “leakage” of mass into and out of each mascon. This results in a spatial “smoothing” effect  
31 such that the change in mass for the area represented by a mascon is distributed over a radius  
32 of roughly 600 km from the mascon center (Luthcke et al., 2013). As a result, model outputs

1 need to be consistently spatially filtered to allow a fair comparison with the GRACE-LM  
2 data. The details of this process are described further in Section 2.4.

## 3 **2.2 The MAR RCM**

4 The MAR RCM (Gallée and Schayes, 1994; Gallée, 1997; Lefebvre et al., 2003) is a coupled  
5 surface-atmosphere RCM that has been applied over the GrIS to simulate current and future  
6 changes in SMB (e.g. Fettweis et al., 2013a; Franco et al., 2013). The atmospheric portion of  
7 MAR is described by Gallée and Schayes (1994), while the land surface model is the Soil Ice  
8 Snow Vegetation Atmosphere Transfer scheme (SISVAT), containing the Crocus snow model  
9 (Brun et al., 1992). We use model outputs from two versions of the MAR model, MAR v2.0  
10 (used by Fettweis et al., 2013a) for the period January 2003 – December 2010, with the model  
11 domain and setup described by Fettweis (2007), and MAR v3.5.2, the latest version of MAR  
12 (used by Colgan et al., 2015), for the period January 2003 – December 2012. For comparison  
13 with GRACE, we include MAR SMB for the entire island of Greenland, including the GrIS,  
14 peripheral ice-covered areas, and tundra areas, as Greenland mass changes related to snow  
15 and ice cover outside of the ice sheet boundaries are not removed in the GRACE solution. An  
16 overestimation of accumulation simulated by MAR v2.0 in the interior of the ice sheet  
17 (Vernon et al., 2012) was in part corrected in MAR v3.5.2 by slightly increasing the snowfall  
18 rate, producing more precipitation along the ice sheet margin and less inland. According to  
19 the recommendations of Alexander et al. (2014), MAR v3.5.2 features an updated bare ice  
20 albedo exponentially varying between 0.4 (dirty ice) and 0.575 (clean ice) as a function of the  
21 accumulated surface water height and slope. The bare ice albedo was fixed at 0.45 in MAR  
22 v2.0. Both MAR v3.5.2 and MAR v2.0 are forced every 6 hours at the lateral boundaries by  
23 the ERA-Interim reanalysis (Dee et al., 2011) beginning in January 1979, and are run at a 25  
24 km spatial resolution (as shown in Fig. 1b). This paper primarily focuses on results from  
25 MAR v3.5.2, which is used to force the ISSM ice sheet model. As will be discussed further in  
26 Section 2.4.2, the computationally intensive processing used in processing of GRACE-LM  
27 outputs had previously been applied to MAR v2.0 outputs, and we used these filtered MAR  
28 v2.0 outputs for the purpose of deriving a spatial filter to approximate spatial attenuation in  
29 the GRACE-LM solution.



## 1 **2.3 The ISSM Model**

2 ISSM (Larour et al., 2012) is a thermo-mechanical ice sheet model that simulates ice flow in  
3 response to forcing from surface mass balance. The model solves equations for conservation  
4 of mass, momentum, and energy, in conjunction with constitutive equations for ice properties  
5 and boundary conditions. It has the capability of incorporating multiple approximations to the  
6 full-Stokes (FS) ice flow equations in different regions. The model is implemented on a finite  
7 element mesh, which can be refined anisotropically to allow for a higher resolution in areas of  
8 high gradients in observed surface velocities. Inversion methods are used to derive  
9 constitutive properties such as ice rigidity and basal friction, by iteratively minimizing  
10 differences between radar-derived observed and modeled ice velocities (Morlighem et al.,  
11 2010; Larour et al., 2012).

12 In this study, ISSM has been run over the entire GrIS, following the model configuration of  
13 Schlegel et al. (2015), which uses a 2D Shelfy-Stream Approximation to the FS equations  
14 (MacAyeal, 1989) in order to increase computational efficiency (as described by Larour et al.,  
15 2012). Aside from the inversion methods used to perform initialization of parameters for ice  
16 properties and basal friction, the model is forced only by SMB at the surface, subject to the  
17 boundary conditions described by Larour et al. (2012). Bedrock topography is defined using  
18 the radar and mass-conservation-derived dataset of Morlighem et al. (2015) (described in  
19 Morlighem et al., 2014). The GrIS simulation consists of an anisotropic mesh, which ranges  
20 in spatial resolution from 1 km to 15 km, consisting of 91,490 elements. The MAR v3.5.2  
21 mean SMB for the period January 1979 – December 1988 is interpolated to a 5 km resolution  
22 using the method of Franco et al. (2012) to correct SMB with respect to subgrid topography  
23 as a function of the local vertical gradient of SMB, and is subsequently interpolated onto the  
24 ISSM mesh. Then it is used to spin up ISSM until the model reaches steady-state equilibrium,  
25 i.e. the change in GrIS mass over time is negligible (as described by Schlegel et al., 2013).  
26 Once the model reaches steady-state (after 30,000 years in this case), ISSM is forced monthly  
27 with SMB from the climate reconstruction of Box et al. (2013) and Box (2013) for the period  
28 January 1841 – December 1979, adjusted so that the mean SMB for this period is equal to the  
29 MAR mean SMB of January 1979 – December 1988. This ensures that ISSM responds to  
30 mean SMB from MAR, but incorporates anomalies from this mean beginning in 1841. MAR  
31 v3.5.2 SMB for the period January 1979 – December 2013 is then used to force ISSM at a  
32 daily temporal resolution with a model timestep of 12 hours. The cumulative mass change

1 from MAR v3.5.2 and ISSM are then combined for comparison with GRACE. ISSM mean  
2 DMB for the period January 2003 – December 2012 is shown in Fig. 1c.

### 3 **2.4 Methods of Comparison**

4 In order to properly compare model results with the GRACE-LM solution, it was necessary to  
5 first spatially aggregate model data to the GRACE grid, to account for the different resolution  
6 of different products (Sect. 2.4.1). Second, in order to conduct a fair comparison with  
7 GRACE-LM at the spatial and temporal resolution of the GRACE-LM solution, model results  
8 must be spatially and temporally filtered to account for spatial and temporal attenuation of the  
9 GRACE signal, associated with the “leakage” of mass changes from each mascon into nearby  
10 mascons in space and time (Luthcke et al., 2013). The best means of filtering model data for  
11 comparison to GRACE-LM is to apply the equations used in GRACE-LM processing directly  
12 to the model data (Sect. 2.4.2). Because this process is computationally intensive, however,  
13 we approximated the effect of GRACE-LM processing using spatial and temporal Gaussian  
14 filters (Sect. 2.4.3). Although our approximation does not perfectly reproduce the effect of  
15 GRACE filtering in space and time (Sect. 2.4.5), we adopt a statistically conservative  
16 approach in our comparison between GRACE-LM and model outputs, to identify cases where  
17 differences are unlikely to be a result of filtering or errors in the GRACE-LM solution  
18 (discussed in Sect. 2.4.4).

#### 19 **2.4.1 Spatial aggregation**

20 MAR and ISSM daily outputs for the period January 2003 – December 2012 were spatially  
21 aggregated into GRACE-LM mascons (Figs 1d and 1e). In the case of ISSM data, ISSM  
22 dynamic thickness changes (ice thickness change associated only with dynamic motion of ice)  
23 on the anisotropic mesh were first interpolated onto a 10 km equal area grid, converted into  
24 mass changes using the density of ice ( $917 \text{ kg m}^{-3}$ ) and then aggregated to the nearest  
25 GRACE-LM mascon to produce timeseries of DMB for each mascon. In the case of MAR  
26 data, MAR SMB outputs at a 25 km resolution were aggregated to the nearest mascon. The  
27 sum of mass change simulated by each model was then calculated for each mascon. Over the  
28 oceans, all mass changes predicted by MAR (likely associated with accumulation over sea  
29 ice) were set to zero, as such accumulation does not result in changes in mass due to the  
30 presence of isostatic adjustment of sea ice over the oceans.

## 1 2.4.2 Spatial and temporal filtering using GRACE equations

2 The GRACE-LM solution uses a Gauss-Newton (GN) procedure to adjust an equivalent  
3 height of water within each mascon to produce perturbations in the GRACE spherical  
4 harmonic fields or Stokes coefficients. The partial derivatives of the Stokes coefficients with  
5 respect to the equivalent water height, and the partial derivatives of KBRR with respect to the  
6 Stokes coefficients are then used to determine the change in KBRR associated with a change  
7 in equivalent water height. The GN procedure iteratively adjusts equivalent water height  
8 within all mascons to minimize the residuals between computed KBRR and KBRR  
9 observations. The final GRACE-LM solution for a given mascon is not the “true” mascon  
10 state, but differs from it due to “leakage” between mascons and the presence of noise in the  
11 solution. The relationship between the true mascon state  $h_k$  and the updated mascon state  $\tilde{h}_k$   
12 is given by Equation 8 of Luthcke et al. (2013), expressed as :

$$13 \quad \tilde{h}_{k+1} = Rh_k + Ke \quad (2)$$

14 where  $e$  represents added noise, and  $R$  is referred to as the resolution operator, as it serves the  
15 function of “smoothing” the true mascon states  $h_k$  in space and time.  $K$  and  $R$  are in turn  
16 expressed by:

$$17 \quad K = (L^T A^T WAL + \mu P_{hh})^{-1} L^T A^T W \quad (3)$$

$$18 \quad R = KAL \quad (4)$$

19 where  $L$  represents the partial derivatives of the Stokes coefficients with respect to the mascon  
20 state,  $A$  represents the partial derivatives of the KBRR observations with respect to the Stokes  
21 coefficients, and  $W$  is a data weight matrix that accounts for orbital parameters and  
22 corrections for processes not related to ice sheet mass changes (e.g. isostatic adjustments,  
23 tides, etc.).  $P_{hh}$  is a regularization matrix, which constrains the solution so that differences in  
24 mass change between mascons closer together are minimized (Sabaka et al., 2010).  
25 Constraint regions for the GrIS are also defined (Fig. 1a) such that the constraint does not  
26 apply across the boundaries of the constraint region. Thus, for the GrIS, changes in mass  
27 above 2000 m in elevation, where the MB is generally positive, can occur independently of  
28 changes in mass below 2000 m in the GRACE-LM solution (Luthcke et al., 2013).

29 In order to spatially filter MAR data to match GRACE-LM, we applied the resolution matrix  
30 to the aggregated MAR v2.0 data, using Equation 2, taking the aggregated MAR v2.0 on

1 GRACE-LM mascons as the “true” mascon states  $h_k$ , and ignoring the added noise term  $e$ .  
 2 The resulting updated mascon states  $\tilde{h}_{k+1}$  are MAR v2.0 data spatially and temporally filtered  
 3 to match GRACE-LM. The effect of spatial smoothing on the MAR v2.0 aggregated outputs  
 4 (Fig. 2a), along with the impact of different constraint regions above and below 2000 m in  
 5 elevation, can be seen in Fig. 2b, which shows the mean January 2003 – December 2010  
 6 MAR v2.0 outputs filtered using the resolution matrix. As expected, the spatial filtering  
 7 decreases the magnitude of mass change for individual mascons by redistributing mass  
 8 change across other surrounding mascons.

9 Unfortunately, the methods discussed above (hereafter referred to as “GRACE-LM filtering”)  
 10 are computationally expensive and time consuming to perform. We only applied GRACE-  
 11 LM filtering to MAR v2.0 outputs as only these outputs were available when the GRACE-LM  
 12 filtering procedure was applied. To filter MAR v3.5.2 and ISSM data, we employed an  
 13 approximation to the GRACE-LM filtering procedure, which is described further below.

#### 14 2.4.3 A Gaussian approximation to GRACE-LM filtering

15 As discussed by Luthcke et al. (2013), the leakage associated with individual GRACE-LM  
 16 mascons is roughly equivalent to a spatial Gaussian filter with a radius of 300 km, with the  
 17 mascons within a 600 km radius accounting for almost 100% of the mass changes within a  
 18 mascon. To allow for spatial filtering of MAR v3.5.2 and ISSM outputs, we developed an  
 19 approximation to the GRACE-LM filtering using a Gaussian filter. The Gaussian function  
 20 can be expressed as a function of distance from the center of the distribution ( $x-\mu$ ), where  $x$  is  
 21 the  $x$ -coordinate and  $\mu$  is the mean of the distribution, and a standard deviation ( $\sigma$ ) as:

$$22 \quad g(x) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^2} \quad (5)$$

23 We used a gaussian function to weight the data for all surrounding mascons ( $j$ ) as a function  
 24 of radial distance from a central mascon ( $i$ ). In this case,  $x-\mu$  is replaced by the distance from  
 25 a central location to another mascon ( $r_{ij}$ ), and a discrete approximation to the Gaussian is  
 26 used, as follows:

$$27 \quad g(r_{ij}) = e^{-\frac{1}{2}\left(\frac{r_{ij}}{\sigma_i}\right)^2} \quad (6)$$

$$w_j = \frac{g(r_{ij})}{\sum_{j=1}^n g(r_{ij})} \quad (7)$$

The weight,  $w_j$ , assigned to a given mascon,  $j$ , at a distance  $r_{ij}$  from mascon  $i$ , is given by the value of the Gaussian function at the center of mascon  $j$  divided by the sum of all Gaussian values surrounding mascon  $i$ . A different  $\sigma_i$  value is chosen for each mascon, as will be explained further below.

We further modify Eq. (7) to account for the constraint regions discussed in the previous section, which for the GrIS, includes areas above and below 2000 m in elevation (Luthcke et al., 2013). For a given mascon within a constraint region (mascon  $i$ ), weights for mascons outside of the constraint region were multiplied by a leakage parameter,  $\lambda_{ij}$ , which represents the fraction of mass in mascon  $j$  outside the constraint region that influences the mass change in mascon  $i$ . Accounting for these constraints, Eq. (7) becomes:

$$w_j = \frac{g(r_{ij})\lambda_{ij}}{\sum_{j=1}^n g(r_{ij})\lambda_{ij}} \quad (8)$$

Where  $\lambda_{ij}$  for mascon  $i$  is set equal to 1 within the constraint region, and equal to a constant value between 0 and 1 for all mascons  $j$  outside of the constraint region.

The weights for mascons  $j$  surrounding a central mascon  $i$  are then used to create a weighted average of mass change for mascon  $i$  ( $\Delta m_{i,new}$ ) as a function of the modeled changes for mascon  $i$  ( $\Delta m_i$ ) and mascons  $j$  ( $\Delta m_j$ ):

$$\Delta m_{i,new} = \Delta m_i w_i + \sum_{j=1}^n \Delta m_j w_j \quad (9)$$

Finally, we added a time component to the filtering procedure, as the regularization matrix ( $P_{hh}$ ) discussed in Section 2.4.2 also includes a temporal component (Sabaka et al., 2010), and because GRACE-LM-filtering alters both the amplitude and timing of the seasonal cycle of mass change (as discussed in Section 2.4.5). After applying the spatial filter described by Eqs. (6) and (8), timeseries of cumulative mass from MAR v2.0 were interpolated onto GRACE-LM time-intervals. We then applied a temporal Gaussian filter to the cumulative mass timeseries for each mascon, using the temporal radius  $\Delta t_{t_0 t_k}$  where  $t_0$  is a point in time along the timeseries,  $t_k$  is a time before or after the time  $t_0$ , and  $\Delta t_{t_0 t_k} = |t_k - t_0|$ :

$$1 \quad g(\Delta t_{t_0 t_k}) = e^{-\frac{1}{2} \left( \frac{\Delta t_{t_0 t_k}}{\sigma_{time}} \right)^2} \quad (10)$$

$$2 \quad w_{tk} = \frac{g(\Delta t_{t_0 t_k})}{\sum_{k=n}^m g(\Delta t_{t_0 t_k})} \quad (11)$$

3 where  $n$  is the first value in the timeseries being filtered and  $m$  is the last value. The filtering  
4 was applied at each time  $t_0$  to produce a timeseries of filtered cumulative mass change.

5 We applied the spatial and temporal filters discussed above to the aggregated unfiltered MAR  
6 v2.0 data, and compared the resulting cumulative mass timeseries' from each mascon to the  
7 GRACE-LM filtered MAR v2.0 timeseries. Two filtering procedures were employed, one in  
8 which only spatial filtering was performed, and another in which both spatial and temporal  
9 filtering were performed to determine the impact of temporal filtering. We iteratively adjusted  
10 the values of  $\sigma_i$ ,  $\sigma_{time}$  (in the case of temporal filtering), and  $\lambda_{ij}$ , for each mascon  $i$ . Values of  
11  $\sigma_i$  were varied at 10 km increments over a range of 1 to 600 km, while values of  $\sigma_{time}$  ranged  
12 between 1 and 91 days at increments of 5 days, and  $\lambda_{ij}$  ranged between 0 and 1 at increments  
13 of 0.01. We tried all combinations of the three parameters over these ranges. The  
14 combination of parameters that yielded the minimum root mean squared error (RMSE)  
15 between the Gaussian-filtered and GRACE-LM-filtered cumulative mass timeseries were  
16 taken as the optimal set of parameters for a given mascon. We also tried applying the same  
17 values of  $\sigma_i$ ,  $\sigma_{time}$ , and  $\lambda_{ij}$  across all mascons  $i$  over the specified range of each parameter, but it  
18 was found that by spatially varying the values of these parameters the errors were reduced.  
19 We also set  $\lambda_{ij}$  equal to zero outside of the island of Greenland as defined by the GRACE-LM  
20 mascons, as this improved the agreement with the GRACE-LM-filtered results. We tried  
21 larger values of  $\sigma_i$  beyond the indicated range at larger increments, but did not find a  
22 reduction in RMSE for values larger than 600 km.

23 Average Gaussian-filtered MAR v2.0 SMB (with both spatial and temporal filtering applied)  
24 for the period 2003-2010 is shown in Fig. 2c. The Gaussian-filtered MAR v2.0 SMB is  
25 similar to GRACE-LM-filtered SMB. Differences between the GRACE-LM-filtered and  
26 Gaussian-filtered results (Fig. 2d) are an order of magnitude smaller than the average SMB  
27 values (ranging from -0.2 to 0.2 vs. -2 to 5 Gt), although in some regions where trends in  
28 SMB are small, the differences are a large percentage of the average SMB. Optimal values  
29 for  $\sigma_i$ ,  $\sigma_{time}$ , and  $\lambda_{ij}$  and the RMSE for the Gaussian vs. the GRACE-LM-filtered MAR v2.0

1 data are shown in Fig. 3. Further discussion of the impacts of filtering on model outputs is  
2 provided in Section 2.4.5.

### 3 2.4.4 Application of Gaussian filters, seasonal cycle analysis and trends

4 Following the choice of the optimal Gaussian filter using MAR v2.0, we applied the same  
5 chosen filter to MAR v3.5.2 and ISSM data forced by MAR v3.5.2, aggregated to the  
6 GRACE-LM grid. MAR v3.5.2 exhibits a less negative SMB along the Greenland coast and  
7 GrIS margins and a less positive SMB within the GrIS interior (Fig. S1) compared with MAR  
8 v2.0 (as there is more coastal accumulation and less interior accumulation in MAR v3.5.2).  
9 These differences do not affect our ability to filter MAR v3.5.2 outputs, as the Gaussian filter  
10 does not depend on mass changes, but approximates the GRACE-LM resolution operator,  
11 which serves to redistribute mass changes subject to specified constraint regions. Spatial  
12 filtering of MAR v3.5.2 and ISSM was conducted first at a daily temporal resolution. Filtered  
13 cumulative mass timeseries for each mascon were then interpolated onto GRACE-LM time  
14 steps, and temporal filtering was performed. We then summed the timeseries of cumulative  
15 mass change from MAR v3.5.2 and ISSM, to generate timeseries of integrated MB.

16 We examined differences between the modeled and GRACE-LM seasonal cycles of  
17 cumulative mass change by first linearly interpolating filtered cumulative model and  
18 GRACE-LM timeseries onto daily timesteps. This was necessary because the GRACE-LM  
19 timesteps are not evenly spaced, and do not occur at the same point in time every year. We  
20 then subtracted the long-term linear trend for the entire timeseries (2003-2012) obtained from  
21 least-squares regression, to remove the impact of differences in trends on the timing of the  
22 seasonal cycle. After removing trends, the cumulative mass value for a given day of the year  
23 was averaged across all years in the 2003-2012 period, to yield an average annual cycle for all  
24 years. The maximum and minimum peaks were computed from this average annual cycle.  
25 This was performed for the GrIS-wide timeseries, as well as for individual mascons and GrIS  
26 sub-regions.

27 GRACE data from Luthcke et al. (2013) include estimates of the error associated with the  
28 timeseries of cumulative mass change for each mascon. When examining aggregated data, we  
29 summed the error for all mascons. The error for a given day for the GRACE-LM seasonal  
30 cycle was determined to be the total GRACE-LM error for the cumulative timeseries divided  
31 by  $\sqrt{n}$ , where  $n$  was the number of years being averaged. Errors in the GRACE-LM

1 timeseries can lead to errors in the timing of the seasonal cycle because random errors can  
2 cause a shift in the timing of a local maximum or minimum point. To account for these  
3 errors, we performed 10,000 Monte Carlo simulations with the GRACE-LM seasonal data,  
4 assuming that the errors in the timeseries were normally distributed. For each of these  
5 simulations, a timeseries of randomly distributed noise with standard deviation equal to the  
6 error was added to the GRACE-LM timeseries at 10-day intervals beginning on 1 January  
7 (roughly corresponding to the temporal resolution of the GRACE-LM solution). The error  
8 was applied on 10-day intervals as the GRACE-LM error is defined for roughly this temporal  
9 resolution. For each simulation, we identified the local maximum and minimum peaks in the  
10 seasonal cycle, allowing us to generate a distribution of dates for maximum and minimum  
11 peaks. The temporal resolution of the GRACE-LM dataset, and the addition of errors at 10-  
12 day intervals can also lead to errors of roughly  $\pm 10$  days for the timing of any estimate.  
13 ~~Because the GRACE-LM timesteps are not regular, the uncertainty on the timing of peaks for~~  
14 ~~the average seasonal cycle due to temporal resolution is generally smaller than 10 days.~~  
15 Given that the error could be as large as 10 days, ~~however,~~ we calculated our error on the  
16 timing of seasonal cycle peaks as the 95% confidence interval from the Monte Carlo  
17 simulations,  $\pm 10$  days. If model peaks fell outside of this error range, the timing of the  
18 GRACE-LM and model peaks was deemed to differ.

19 We also compared modeled and GRACE-LM trends for the 2003-2012 period. To calculate  
20 the uncertainty in trends from GRACE-LM, we employed a similar procedure to estimate  
21 uncertainty in trends, conducting 10,000 Monte Carlo simulations and obtaining a distribution  
22 of trends and uncertainty values (from the 95% confidence interval for calculated each trend).  
23 The error on the trend was calculated as the average of the 2.5% and 97.5% deviations from  
24 the trend added to the 97.5% (upper) bound on the distribution of uncertainty values. For all  
25 model estimates, uncertainty on trends is reported as the 95% confidence interval obtained  
26 during linear regression.

#### 27 2.4.5 Effect of filtering on seasonal variations in mass

28 As we were interested in examining seasonal variations in mass, we examined the impact of  
29 GRACE-LM filtering vs. Gaussian filtering (spatial only and spatial + temporal) on the  
30 cumulative timeseries of GrIS-wide mass changes, in relation to the seasonal cycle of mass  
31 change from GRACE-LM. While it is not possible to compare GRACE-derived mass changes  
32 directly to MAR, given that GRACE also records the effects of changes in DMB, a



1 comparison of de-trended timeseries of cumulative mass can be performed, if it is assumed  
2 that seasonal variations in ice discharge are small relative to those of SMB. A qualitative  
3 comparison of de-trended unfiltered and filtered MAR v2.0 and GRACE-LM cumulative  
4 mass timeseries for the GrIS over January 2003 – December 2010 (Fig. 4), suggests that this  
5 is a reasonable first-order assumption for the entire GrIS. (As will be seen in Section 3.2,  
6 modeled ice-sheet wide seasonal variability from ISSM is also less than 10% of variability  
7 from MAR.) Fluctuations in mass, coinciding with net loss of mass during summer months,  
8 and net gain of mass during winter months, are captured by both GRACE-LM and MAR v2.0.  
9 The average seasonal cycle of cumulative mass change in Fig. 5a indicates a larger amplitude  
10 of mass fluctuations for unfiltered and spatially Gaussian-filtered MAR v2.0 results (of 524  
11 and 500 Gt respectively) relative to GRACE-LM ( $287 \pm 30$  Gt), and a closer agreement  
12 between the amplitudes of GRACE-LM-filtered MAR v2.0 data (339 Gt) and GRACE-LM.  
13 On average, during periods of net ablation, GRACE-LM begins losing mass earlier (by 25  
14 days), and starts gaining mass later (by 8 days) as compared with MAR v2.0 unfiltered data  
15 (Table 1). GRACE-LM-filtering changes the timing of the start of mass loss such that the  
16 period of simulated mass loss begins 10 days sooner, extending the length of the mass loss  
17 period.

18 When both spatial and temporal Gaussian filtering are applied to the MAR v2.0 data, the  
19 amplitude of the seasonal cycle is reduced (to 351 Gt), resulting in a better agreement with  
20 GRACE-LM and with the GRACE-LM-filtered MAR v2.0 data (Fig. 5a). The timing of  
21 peaks in maximum and minimum mass are also changed, with the temporally Gaussian-  
22 filtered MAR v2.0 data exhibiting an extended period of mass loss (145 days) relative to that  
23 of the GRACE-LM-filtered MAR v2.0 data (123 days), resulting in the filtered seasonal cycle  
24 peaks occurring within 5 days of GRACE-LM peaks. In all cases, however, the timing of peak  
25 mass loss from MAR v2.0 falls within the 95% confidence bounds on maximum and  
26 minimum dates for GRACE-LM.

27 Adding temporal Gaussian-filtering improves the agreement between the Gaussian-filtered  
28 MAR v2.0 timeseries and the GRACE-LM-filtered timeseries in terms of amplitude, and  
29 lengthens the period of net ablation (perhaps too much relative to the period for GRACE-LM-  
30 filtered data). For both methods of Gaussian filtering, the timing of the peaks for filtered  
31 MAR v2.0 data fall within the 95% confidence bounds on the timing of the GRACE-LM  
32 seasonal cycle.

1 The comparison of GRACE and MAR timeseries and seasonal cycles in the case of MAR  
2 v3.5.2 (Figs. S2 and 5b respectively) is similar to that for MAR v2.0. MAR v3.5.2 features a  
3 seasonal cycle of smaller amplitude, likely as a result of snow falling more frequently along  
4 the coast, where it is more likely to be balanced by ablation during periods of net  
5 accumulation, and where it mitigates ablation during periods of net mass loss. The Gaussian  
6 filtering has a similar effect on the MAR v3.5.2 outputs, which are similar in timing to MAR  
7 v2.0 outputs (Table 1), by reducing the amplitude of seasonal variability and extending the  
8 length of the ablation season to be similar to that of GRACE-LM (Fig. 5b and Table 1). In our  
9 analysis of ISSM and MAR v3.5.2 outputs, we have chosen to focus on results obtained with  
10 temporal Gaussian filtering applied, as it results in reduced errors relative to the GRACE-LM  
11 filtering method. We consider this to be a statistically conservative approach. Because we  
12 are not able to fully capture the effect of filtering on the timing of the seasonal cycle, we  
13 choose the filter that brings the timing of the seasonal cycle closer to that of GRACE-LM.  
14 Thus, in locations where the timing of the Gaussian-filtered cycle falls outside of the range of  
15 dates from GRACE-LM, it is very likely that there is a difference between the modeled and  
16 GRACE-LM seasonal cycles that is not associated with filtering. In locations where the  
17 timing of the Gaussian-filtered cycle falls within the range of dates from GRACE-LM, we  
18 cannot confirm a difference.

### 19 **3 Results**

#### 20 **3.1 Trends and spatial differences in modelled vs. measured mean MB**

21 We first examine the timeseries of GrIS cumulative mass as simulated by MAR v3.5.2, ISSM,  
22 and GRACE-LM over the 2003-2012 period, as shown in Fig. 6. MAR v3.5.2 cumulative  
23 SMB shows a net accumulation of mass over Greenland (of  $246.9 \pm 4.7 \text{ Gt yr}^{-1}$ ), which varies  
24 seasonally in response to cycles of melting and accumulation. ISSM exhibits a net loss of  
25 mass ( $-425.8 \pm 0.3 \text{ Gt yr}^{-1}$  on average), with little seasonal variability relative to the long-term  
26 trend. There is a small seasonal cycle in ISSM dynamics driven by the SMB cycle (visible in  
27 the detrended timeseries shown in Fig. S3) which complements the mass changes from MAR  
28 (increased mass loss from MAR leads to decreased mass loss from ISSM, and vice versa),  
29 with an amplitude roughly an order of magnitude smaller than the SMB fluctuations.  
30 Together, ISSM and MAR v3.5.2 results produce a net loss of mass over 2003-2012, although  
31 the trend in simulated mass loss ( $-178.9 \pm 4.4 \text{ Gt yr}^{-1}$ ) is smaller in magnitude than that of  
32 GRACE-LM ( $-239.4 \pm 7.7 \text{ Gt yr}^{-1}$ ) by  $60.5 \pm 12.1 \text{ Gt yr}^{-1}$ .

1 The roughly complementary nature of modeled SMB and DMB is evident on a sub ice-sheet  
2 wide scale, as indicated by the unfiltered MAR v3.5.2 and ISSM 2003-2012 mean SMB and  
3 DMB (Fig. 1) as well as the Gaussian filtered data (Figs. 7a and 7b). Areas with a large  
4 positive SMB from MAR v3.5.2 show large dynamic mass loss from ISSM (e.g. areas higher  
5 than 2000 m in elevation), while areas with negative SMB from MAR v3.5.2 show smaller  
6 losses from ISSM. Summing SMB and DMB from MAR v3.5.2 and ISSM produces the  
7 pattern of MB shown in Fig. 7c, which indicates that the majority of modeled mass loss for  
8 the 2003-2012 period occurs below 2000 m in elevation. This is similar to pattern of MB  
9 from GRACE-LM (Fig. 7d). A map of the difference between modeled and GRACE-LM MB  
10 (Fig. 7e) indicates that the majority of the difference in trends observed in Fig. 6 results from  
11 an underestimation of mass loss from the filtered model results below 2000 m in elevation, in  
12 particular along the west and southeast coasts. Mass loss from GRACE-LM is larger in  
13 magnitude along the GrIS margins (by up to  $\sim 2.5 \text{ Gt yr}^{-1}$  per mascon). In areas below 2000  
14 m, overall mass loss is underestimated by  $92 \pm 10.3 \text{ Gt yr}^{-1}$  (with a trend of  $-149.4 \pm 3.5 \text{ Gt}$   
15  $\text{yr}^{-1}$  as compared with  $-242.4 \pm 6.8 \text{ Gt yr}^{-1}$  from GRACE). For areas above 2000 m, GRACE-  
16 LM suggests little change in mass ( $+3.0 \pm 4.2 \text{ Gt yr}^{-1}$ ), while the models suggest a loss of -  
17  $29.5 \pm 1.0 \text{ Gt yr}^{-1}$ . The differences between the models and GRACE-LM are larger than the  
18 uncertainties in trends from GRACE-LM and from the model outputs, suggesting that model  
19 errors or unaccounted-for processes contribute to the difference.

20 The differences between simulated and GRACE-LM MB may be due to a modeled MAR  
21 v3.5.2 SMB that is too high below 2000 m in elevation, or alternately, to simulated velocities  
22 that are underestimated in ISSM and vice versa at higher elevations. We evaluate the ISSM  
23 spin-up by comparing the ISSM ice thickness at steady-state to the ice thickness obtained  
24 from the mass conservation dataset of Morlighem et al. (2015), derived from radar data for  
25 1993-2014, interpolated onto the ISSM mesh. We also compare ISSM annual ice velocities  
26 with the radar-derived ice velocity data of Rignot and Mouginot (2012), derived from data  
27 spanning 2008-2009. A comparison of ice thicknesses from Morlighem et al. (2015) and  
28 velocities from Rignot and Mouginot (2012) with ISSM velocities and thicknesses for 1  
29 January 2003 is shown in Fig. S4. Some differences may result from the model outputs and  
30 observations not being coincident in time, but Fig. S3 indicates that temporal variability in  
31 ISSM is small ( $<10\%$ ) relative to long-term changes. (This relatively small variability is to be  
32 expected given that the ISSM simulation used here only considers forcing from SMB and not  
33 other factors that may lead to larger fluctuations in ice motion.) In particular, ice velocities

1 tend to be underestimated for glaciers along the northwest coast of the GrIS (Fig. S4b),  
2 possibly as a consequence of an upstream ice thickness that is also underestimated (Fig. S4a).  
3 This may contribute to underestimated mass loss along the northwest coast. In other areas,  
4 ice thickness is generally overestimated by ISSM, but some outlet glacier velocities are  
5 overestimated while others are underestimated, making it unclear how ISSM contributes to  
6 the observed discrepancies in these regions. It is difficult to determine if the observed  
7 differences are a result of errors in the MAR v3.5.2 SMB forcing (as the spinup relies  
8 primarily on the simulated SMB for forcing), simplifications to full-Stokes ice flow in ISSM,  
9 processes not considered in the ISSM simulation such as the role of hydrology in ice  
10 dynamics and ice-ocean interactions, or errors associated with the assumption that during the  
11 spin-up period, the ice sheet is in steady state. The assumption of 2D flow (the Shelfy-Stream  
12 Approximation) in the current ISSM simulation probably contributes to errors in dynamic  
13 mass balance, particularly at higher elevations, but it is not clear whether this would lead to  
14 faster or slower ice flow. A comparison between MAR v3.5.2 SMB and in situ measurements  
15 performed by Colgan et al. (2015) suggests that MAR v3.5.2 does not overestimate SMB  
16 below 2000 m, but this comparison is limited to one location in southwest Greenland: the  
17 Kangerlussuaq transect (van de Wal et al., 2005), and may not be representative of other  
18 areas. Another factor that may result in the discrepancy between model results and GRACE-  
19 LM is the 25 km resolution of the MAR outputs used in this study. ISSM is forced by a  
20 downscaled version of MAR v3.5.2. Using a higher resolution simulation or downscaled  
21 outputs could result in different SMB estimates (e.g. Franco et al., 2012). This could be  
22 particularly important along the borders of the GrIS, or for mountainous areas outside the ice  
23 sheet boundaries. In these areas, high spatial variability of topography can strongly influence  
24 SMB. To properly identify the source of the differences, further independent evaluations of  
25 MAR SMB and ISSM DMB are needed, for example, comparison between ISSM and remote-  
26 sensing derived discharge estimates (e.g. Rignot and Kanagaratnam, 2006), or comparison  
27 between MAR and additional in situ measurements in ablation areas.

### 28 **3.2 Seasonal mass changes from MAR, ISSM and GRACE**

29 The average seasonal cycle of filtered cumulative MAR v3.5.2 + ISSM for the 2003 through  
30 2012 period agrees well with that of GRACE-LM, as shown in Fig. 8a and Table 2. The  
31 amount of mass loss during the period of net ablation is similar for MAR v3.5.2 + ISSM (333  
32 Gt) and GRACE-LM (355±32 Gt). The dates of simulated maximum and minimum mass fall

1 within the range of uncertainty for these dates from GRACE-LM. It is possible, however,  
2 that differences in modeled and GRACE-LM trends alter the timing of the seasonal cycle,  
3 because changing the overall trend of a timeseries can alter the timing of local maxima and  
4 minima by altering local rates of change. Therefore, we also show the seasonal cycle for the  
5 de-trended timeseries in Fig. 8b and Table 2. For the de-trended seasonal cycle, the timing of  
6 the seasonal maximum occurs roughly 1 month before the maximum peak from the original  
7 seasonal cycle, and the timing of the seasonal minimum occurs roughly 1 week earlier (within  
8 the uncertainty associated with the ~10 day GRACE timesteps) for both MAR v3.5.2 + ISSM  
9 and GRACE-LM. The model timing for the filtered model data falls within the range of dates  
10 from GRACE-LM (with and without trends removed), and therefore we cannot confirm any  
11 difference between the modeled and GRACE-LM Greenland-wide seasonal cycles of mass  
12 change. The results are consistent with the comparison between the detrended MAR v3.5.2  
13 and MAR v2.0 seasonal cycles and GRACE. ISSM makes a small contribution to the  
14 simulated seasonal cycle; its seasonal cycle is the inverse of the MAR seasonal cycle with a  
15 lag of less than 1 month, indicating that ISSM responds to MAR SMB forcing by increasing  
16 discharge during periods of high SMB, and vice versa, with an amplitude roughly an order of  
17 magnitude smaller than that of MAR (Fig. S5). As noted earlier, this magnitude of simulated  
18 flow variability is the expected response to the SMB forcing applied to ISSM.

### 19 **3.3 Spatial variability in the seasonal cycle from MAR, ISSM and GRACE**

20 Maps of the timing of peaks in the seasonal cycle of de-trended cumulative mass change from  
21 GRACE-LM (Fig. 9) suggest that the timing of seasonal cycle peaks is spatially variable.  
22 Maps of the median GRACE-LM date for the maximum and minimum peaks (Figs. 9a and d)  
23 show that in some locations (e.g. northwest Greenland), GRACE-LM suggests that the peak  
24 in the seasonal cycle can occur as early as 1 November (i.e. mass loss begins during the fall),  
25 where in other areas it occurs as late as 1 July (for an area in north Greenland). The range of  
26 possible dates suggested by GRACE-LM, when taking into account GRACE-LM uncertainty,  
27 is fairly large, spanning ~~up to 5 months~~ more than 11 months in some locations in northern  
28 Greenland (Figs. 9b, c, e and f) where the magnitude of the seasonal cycle is relatively small,  
29 but spatial differences in seasonal timing are preserved even with ~~these large ranges~~ the large  
30 ranges in some areas. The GRACE-LM data suggest that the period of net mass loss begins in  
31 fall and ends in early summer in the northwestern portion of the ice sheet, while in most other  
32 parts of the ice sheet, mass loss begins in early or late spring, and mass begins to increase

1 again beginning in mid to late autumn. The period of summer mass loss over most of the ice  
2 sheet is consistent with what would be expected, given the cycle of climate forcing (warm  
3 conditions leading to increased melt), but the timing of the cycle in the northwest suggests  
4 that other processes may dominate seasonal variability in that region.

5 MAR v3.5.2 + ISSM suggest a more uniform pattern of timing in seasonal cycle peaks (Figs.  
6 10a, b), consistent with the SMB forcing. The filtered model results suggest that mass loss  
7 begins in late spring and early summer (between March and June) without much spatial  
8 variability across the ice sheet, and mass gain commences in late summer and early fall  
9 (between September and November), with the period of mass loss ending ~1 month later for  
10 mascons below 2000 m in elevation relative to those above 2000 m in elevation. These  
11 results are consistent with what would be expected given warmer temperatures at lower  
12 elevations and a longer period available for melting. (The differences are also likely to be  
13 larger without filtering.)

14 For many mascons in the Northwest, the modeled cycle maximum and minimum peaks can  
15 occur up to 3 months after and 2 months before the GRACE-LM peaks (Figs. 10c, d), with  
16 differences of ~1 month being quite common. Given the relatively large uncertainty in the  
17 timing of the GRACE-LM peaks, the model peaks often fall within the distribution of peaks  
18 from GRACE-LM. Along a portion of the northwest coast, however, the timing of the  
19 seasonal maximum occurs in May according to the models, roughly one or two months after  
20 the 95% confidence limit on the timing of maximum mass from GRACE-LM, and more than  
21 three months after the median peak from GRACE-LM. The clustering of the GRACE-LM  
22 peaks, despite the large uncertainty in the GRACE timing, suggests that the observed  
23 variations in timing are not associated with random deviations between mascons, but reflect  
24 seasonal variations in mass detected by GRACE-LM, that are not captured by the models. The  
25 timing of seasonal minimum falls within 95% of the distribution from GRACE-LM for all  
26 mascons., ~~with the exception of four mascons in the northwest GrIS region.~~

### 27 **3.4 The average seasonal cycle within ice sheet sub-regions**

28 In order to further examine discrepancies at regional scales, we created nine sub-regions of  
29 the GrIS based on the median timing of the maximum and minimum peaks of the average de-  
30 trended annual cycle from GRACE-LM (Figs. 9a and d). Mascons were grouped together if  
31 the timing of their maximum and minimum peaks were within 34 days of each other. (The

1 threshold was chosen to create a balance between different types of cyclical patterns and the  
2 ~~number of total~~ total number of regions.) The nine sub-regions are shown in Fig. 11a, along  
3 with the average seasonal cycle from four of these sub-regions (Figs. 11b-e). The average  
4 cycles for other regions are provided in Figs. S6a-e. The average GRACE-LM seasonal cycle  
5 for Region 2 significantly differs in its timing from the MAR v3.5.2 + ISSM cycle. GRACE-  
6 LM results suggest that the period of net mass loss begins no later than mid-~~February~~ April,  
7 while the models suggest that it begins in ~~late April~~ early May as a result of the SMB signal.  
8 For Region 7 (Fig. 11e), the difference in timing is larger. The maximum in the cycle of  
9 cumulative mass change occurs in mid May (although the ice sheet does not appear to start  
10 losing a substantial amount of mass until late June), while the GRACE-LM peak occurs in  
11 mid February. The entire modeled cycle appears to be offset by three months relative to  
12 GRACE-LM in this region, although seasonal mass changes are relatively small (on the order  
13 of 5 Gt). For Regions 8 and 9 (Figs. S6d and e) the timing of the GRACE-LM and modeled  
14 peaks is similar to that of Region 7, with the exception that in Region 9, the minimum  
15 GRACE-LM peak overlaps with a September minimum peak from the models. For Regions 5  
16 and 6 (Figs. 11c and d), the model maximum and minimum peaks fall within the distribution  
17 for GRACE-LM peaks. For Regions 1 ~~and~~, ~~3~~ ~~and~~ ~~8~~ (Figs. S6a ~~and~~, ~~b~~, ~~and~~ ~~d~~) the cycles are  
18 similar to those of ~~R~~Regions 2, with model peaks falling within the distribution from  
19 GRACE-LM for Region 3. ~~and~~ ~~7~~. For ~~R~~Region 4 (Fig. S6c), the cycle is similar to the cycle  
20 of Region 6, except for a sharp peak in mass in early July, which leads the GRACE-LM peak  
21 to occur after the peak from the models. ~~For Region 9 (Fig. S6e) the GRACE-LM maximum~~  
22 ~~peak is similar to that of other regions in northwest Greenland, preceding the model~~  
23 ~~maximum peak, but the minimum GRACE-LM peak overlaps with a September minimum~~  
24 ~~peak from the models.~~

25 Dividing the GrIS into high and low elevation areas (above and below 2000 m in elevation)  
26 also produces differences in the seasonal cycle (Fig. 12). For areas below 2000 m in elevation  
27 (Fig. 12a), there is a good agreement between the GRACE-LM and simulated seasonal cycles;  
28 the timing of MAR v3.5 + ISSM maximum and minimum peaks fall within the distribution of  
29 peaks for GRACE-LM as the signal is dominated by the summer surface mass loss. For areas  
30 higher than 2000 m in elevation (Fig. 12b), the period of simulated net mass loss is shortened  
31 relative to that of GRACE-LM. The good agreement between cycles at low elevations  
32 suggests that the timing of ablation and accumulation at low elevations on an ice sheet wide  
33 scale is well captured by MAR v3.5.2 + ISSM.

1 As ~~for~~in the case of Greenland-wide fluctuations in mass, most of the simulated seasonal  
2 variability within sub-regions of the ice sheet is dominated by MAR, as expected given that  
3 the only forcing applied to ISSM is the SMB from MAR. ISSM exhibits a seasonal cycle that  
4 is a lagged inverse of the MAR cycle with less than 10% of the amplitude of MAR v3.5.2 in  
5 all sub-regions of the ice sheet (Figs. S7, S8, and S9), and the seasonal response is consistent  
6 across all areas of the ice sheet. The timing of the seasonal cycle for GRACE, MAR v3.5.2,  
7 ISSM, and MAR v3.5.2+ISSM for all sub-regions is provided in Table S1.

8 The differences in the GRACE-LM and modeled seasonal cycles within individual regions  
9 seem unlikely to be caused by errors in the simulated timing of surface ablation, as they occur  
10 either during times of the year when melting does not occur at the surface (i.e. the ‘early’ start  
11 to the period of net mass loss in the northwest from November through February; Figs. 9a, b,  
12 c). The results therefore suggest that the observed changes could be associated with errors in  
13 seasonal accumulation from MAR v3.5.2, or processes not currently incorporated into ISSM,  
14 which induce seasonal fluctuations in ice discharge or liquid water. These processes are  
15 difficult to validate, and therefore it is difficult to determine which processes are most  
16 responsible for the observed differences. As discussed in the following section, significant  
17 seasonal variations in glacier velocities have been observed and could contribute to the  
18 observed discrepancies. Accumulation or ice flow errors could also affect differences at  
19 higher elevations, where the net ablation due to melting is small (i.e. above 2000 m in  
20 elevation). Such discrepancies could also be influenced by differences below 2000 m due to  
21 leakage between constraint regions, but the amount of leakage in terms of amplitude is small  
22 and is comparable to the GRACE-LM uncertainty (Luthcke et al., 2013). Additionally, it is  
23 possible that although the GRACE-LM solution includes error estimates associated with the  
24 forward models used in GRACE processing, unaccounted for errors or processes, such as  
25 errors in model simulations used to correct for variability in atmospheric or ocean circulation  
26 (for which observations for validation are limited) may contribute to the differences.  
27 However, we cannot envision any obvious reason for the discrepancy other than the potential  
28 errors in ISSM or MAR v3.5.2 that have been noted.

#### 29 30 **4 Concluding remarks**

31 The above results show several areas of agreement as well as areas of disagreement between  
32 modeled and GRACE-derived Greenland mass balance. We have shown that in order to



1 compare spatial and temporal variations in GrIS mass from RCM, ISM results and the  
2 GRACE-LM solution, it is necessary to spatially and temporally filter the model outputs. We  
3 have developed a Gaussian approximation to the GRACE-LM resolution operator, which  
4 accurately captures the effect of the GRACE-LM solution on spatial variations in mean MB.  
5 We also find that applying temporal filtering reduces differences between the modeled and  
6 GRACE-LM seasonal cycles. We have therefore also applied a temporal Gaussian filter to the  
7 model outputs to reproduce the attenuation inherent to GRACE-LM processing. The  
8 Gaussian temporal filtering does not completely capture the seasonal cycle of mass changes  
9 obtained using the GRACE-LM resolution operator in that it extends the period of mass loss  
10 simulated by the models further than the period obtained from GRACE-LM filtering. As the  
11 filter extends the length of the modeled period of mass loss, and tends to bring the timing of  
12 modeled seasonal cycle peaks closer to those from GRACE-LM (which exhibits a longer  
13 period of mass loss relative to the unfiltered model results), our approach is conservative: in  
14 cases where the cycles disagree, there is likely a difference between the GRACE-LM and  
15 modeled seasonal cycles.

16 A comparison between Gaussian-filtered MAR v3.5.2 + ISSM and GRACE-LM Greenland  
17 mass trends for 2003-2012 indicates that the models tend to underestimate the magnitude of  
18 this mass loss, as a result of underestimated mass loss below 2000 m in elevation. This  
19 difference is either due to an overestimation of SMB from MAR v3.5.2 in low elevation areas,  
20 or to intrinsic errors in ice flow from ISSM. MAR v3.5.2 SMB for low elevation areas is  
21 higher than that of MAR v2.0, in part due to a relatively high bare ice albedo (as described by  
22 Alexander et al., 2014; a higher albedo persists despite modifications made to MAR v3.5.2  
23 albedo), and in part due to a shift in precipitation from high to low elevation areas. A  
24 comparison at eight in situ stations at the Kangerlussuaq transect on the southwest GrIS  
25 suggests that MAR v3.5.2 SMB is closer to in situ measurements (Colgan et al., 2015), but  
26 such measurements are limited to this transect with the exception of a comparable number of  
27 ablation stake measurements from the Programme for Monitoring of the Greenland Ice Sheet  
28 (PROMICE; [www.promice.org](http://www.promice.org)). The only means of determining the relative contribution of  
29 ISSM and MAR v3.5.2 to underestimated mass loss would be to conduct an independent  
30 evaluation of each model against DMB (e.g. using the methods of Rignot and Kanagaratnam,  
31 2006) and SMB estimates over large portions of the GrIS. These analyses are beyond the  
32 scope of this study, and the evaluation of SMB is limited by sparsely available data, although  
33 radar measurements of snow accumulation may help to fill this gap. A preliminary

1 comparison by Koenig et al. (2015) suggests that there is a good agreement between MAR  
2 v3.5.2 and radar-derived accumulation estimates over the interior of the GrIS during the years  
3 2009-2012, but that MAR tends to overestimate accumulation along the southeastern coast.

4 We examined the mean seasonal cycles of de-trended cumulative mass change from GRACE-  
5 LM and MAR v3.5.2 + ISSM as a means of examining the ability of the models to capture  
6 mass changes at a relatively high spatial and temporal resolution. We have shown that on a  
7 Greenland-wide scale, the timing of modeled and GRACE seasonal cycles agree, within the  
8 limits of GRACE uncertainty, but on sub-ice-sheet-wide scales, there are significant  
9 differences in the timing of annual cycle peak. On the scale of individual mascons, there is  
10 considerable variability in the timing of the seasonal cycle from GRACE-LM, while model  
11 outputs suggest a more uniform timing across Greenland driven mainly by summer surface  
12 mass loss and mass gain simulated by MAR. While some of this variability is likely due to  
13 GRACE errors, other variations likely reflect real differences in the seasonal variability  
14 within different regions, particularly as the differences are not random, but spatially clustered.  
15 In particular, in northwestern Greenland, the simulated period of mass loss is shorter than that  
16 of GRACE-LM, and the timing of the simulated maximum in the seasonal cycle occurs up to  
17 three months after the GRACE-LM peak in some areas.

18 Spatial and seasonal differences in the seasonal cycle may result from various factors  
19 including (1) underestimation or overestimation of accumulation and ablation by MAR  
20 v3.5.2, (2) cycles of ice sheet motion associated with processes not incorporated into ISSM,  
21 (3) cycles of water storage and release, and (4) errors in the GRACE-LM solution. We have  
22 attempted to account for the last factor by considering the impact of errors of the GRACE-LM  
23 solution estimated by Luthcke et al. (2013) on the timing of the seasonal cycle, and by  
24 filtering our model results to match GRACE-LM. However, as GRACE does not provide  
25 direct observations of mass changes, and different methods of processing can produce  
26 somewhat different mass change solutions (Shepherd et al., 2012), it is possible that some of  
27 the observed discrepancies may be due to errors not considered in this solution. With regard  
28 to MAR v3.5.2 accumulation, the Colgan et al. (2015) study suggests that MAR v3.5.2  
29 effectively captures spatial variations in SMB, but few observations of SMB are available in  
30 areas of net ablation. The seasonal cycle of MAR v3.5.2 has not been evaluated against  
31 observations, as few sub-annual estimates of accumulation are available. With regard to GrIS  
32 discharge, an analysis of ISSM annual discharge has not been conducted, although

1 comparison with satellite-derived ice velocities suggests that ISSM velocities may be  
2 underestimated in some areas at the ice sheet margins. Data on seasonal velocities are not  
3 available for the entire GrIS, but various studies have indicated seasonal variations in the flow  
4 of GrIS glaciers occur, particularly in association with meltwater that reaches the ice sheet  
5 bed (e.g. Joughin et al., 2008), as well as interactions between ocean circulation and ice at  
6 calving fronts (Howat et al., 2010). Using GPS measurements for west coast GrIS glaciers,  
7 Ahlstrøm et al. (2013) showed that the glaciers examined underwent seasonal cycles in  
8 velocity, with several glaciers showing a decline in velocity in late summer associated with  
9 increased efficiency of subglacial drainage systems. Moon et al. (2014) present the most  
10 comprehensive evaluation of seasonal velocity cycles to date, identifying three types of  
11 seasonal cycles in velocity near the terminus of marine-terminating glaciers, one in which  
12 meltwater production produces acceleration during summer months (“type 2”), another in  
13 which deceleration occurs late in the melt season, followed by acceleration peaking in the  
14 early melt season (“type 3”), and a third in which fluctuations are more likely associated with  
15 ice-ocean interactions (“type 1”). Different glaciers exhibit different patterns of flow  
16 variability, and a single glacier may exhibit different patterns of flow in different years.  
17 Seasonal variations in flow generally represent ~10-20% of mean annual velocities. The type  
18 1 and especially type 3 seasonal patterns could potentially lead to the patterns of mass change  
19 from GRACE-LM in northwest Greenland, but it is unclear how variations in flow of  
20 different glaciers contribute to seasonal fluctuations in ice sheet discharge, and a study  
21 examining this would be useful for evaluating ISMs such as ISSM, which do not currently  
22 take into account the influence of these processes.

23 The general agreement between modeled and GRACE-LM MB below 2000 m in elevation,  
24 where ice sheet hydrology might be expected to play a role, suggests that factors such as  
25 water storage and release as indicated by Rennermalm et al. (2013), and observed on glaciers  
26 in locations outside of Greenland (Jansson et al., 2003) do not play a large role in the timing  
27 of seasonal variations in mass on the ice-sheet-wide scale. It is possible that these processes  
28 influence the amplitude of mass variations, or lead to changes in mass on shorter timescales  
29 that we cannot observe given the uncertainties in GRACE-LM results and filtering, and that  
30 they play a role in longer-term variations in mass. Long-term water storage could contribute  
31 to underestimated trends in mass loss below 2000 m in elevation, which could also result from  
32 underestimated SMB from MAR v3.5.2, or an underestimation of D from ISSM.

1 Further studies are also needed to understand the impact of temporal variations in mass on the  
2 observations presented here, i.e. whether they are associated with processes that reoccur from  
3 year-to-year, or whether isolated events influence the timing of the seasonal cycle. In  
4 addition, future studies are needed to validate seasonal variations of RCM accumulation and  
5 simulated SMB in ablation areas. The spatial and temporal resolution of this analysis was  
6 limited by the fundamental spatial and temporal resolutions of GRACE to seasonal-scale  
7 variability and spatial scales of ~600 km. It is possible that if seasonal variations in GrIS  
8 mass are examined at higher spatial and temporal resolutions, with reduced errors, further  
9 discrepancies between modeled and measured cycles will be observed. As the ice sheet  
10 changes in the future, such processes could potentially become more important to GrIS-wide  
11 changes in mass, and therefore they need to be better understood and their impact quantified.

12

### 13 **Author contribution**

14 P. M. A. and M. T. devised the study. P. M. A. carried out the analysis. N-J. S. and E. L.  
15 performed simulations with and developed the ISSM model. S. L. produced the GRACE  
16 solution. X. F. performed simulations and development of the MAR model. P. M. A.  
17 prepared the manuscript. All co-authors revised and contributed to the editing of the  
18 manuscript.

19

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7

1 Table 1. Timing of maximum and minimum peaks in the seasonal cycle of GrIS-wide  
 2 | detrended cumulative mass change for GRACE-LM and MAR ~~v2.0~~ for the January 2003 –  
 3 December 2010 period. For GRACE-LM, the median value and bounds for the 95%  
 4 confidence interval of the distribution after accounting for uncertainty in GRACE-LM are  
 5 listed. The uncertainty bounds have been extended by 10 days to account for potential errors  
 6 associated with temporal resolution.

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	<u>GRACE</u>	<u>MAR v2.0</u>	<u>MAR</u>	<u>MAR v2.0</u>	<u>MAR v2.0</u>	<u>MAR v2.0</u>	<u>MAR v3.5.2</u>
		<u>Unfiltered</u>	<u>v3.5.2</u>	<u>GRACE-</u>	<u>Gaussian-</u>	<u>Gaussian-</u>	<u>Gaussian-</u>
			<u>Unfiltered</u>	<u>Filtered</u>	<u>(Spatial)</u>	<u>(Space, Time)</u>	<u>(Space, Time)</u>
<u>Maximum</u> (2.5% Bound)	<u>22 Mar</u>						
<u>Maximum</u>	<u>1 May</u>	<u>18 May</u>	<u>21 May</u>	<u>8 May</u>	<u>18 May</u>	<u>28 Apr</u>	<u>30 Apr</u>
<u>Maximum</u> (97.5% Bound)	<u>10 Jun</u>						
<u>Minimum</u> (2.5 % Bound)	<u>19 Aug</u>						
<u>Minimum</u>	<u>18 Sep</u>	<u>7 Sep</u>	<u>31 Aug</u>	<u>8 Sep</u>	<u>7 Sep</u>	<u>20 Sep</u>	<u>20 Sep</u>
<u>Minimum</u> (97.5% Bound)	<u>28 Oct</u>						

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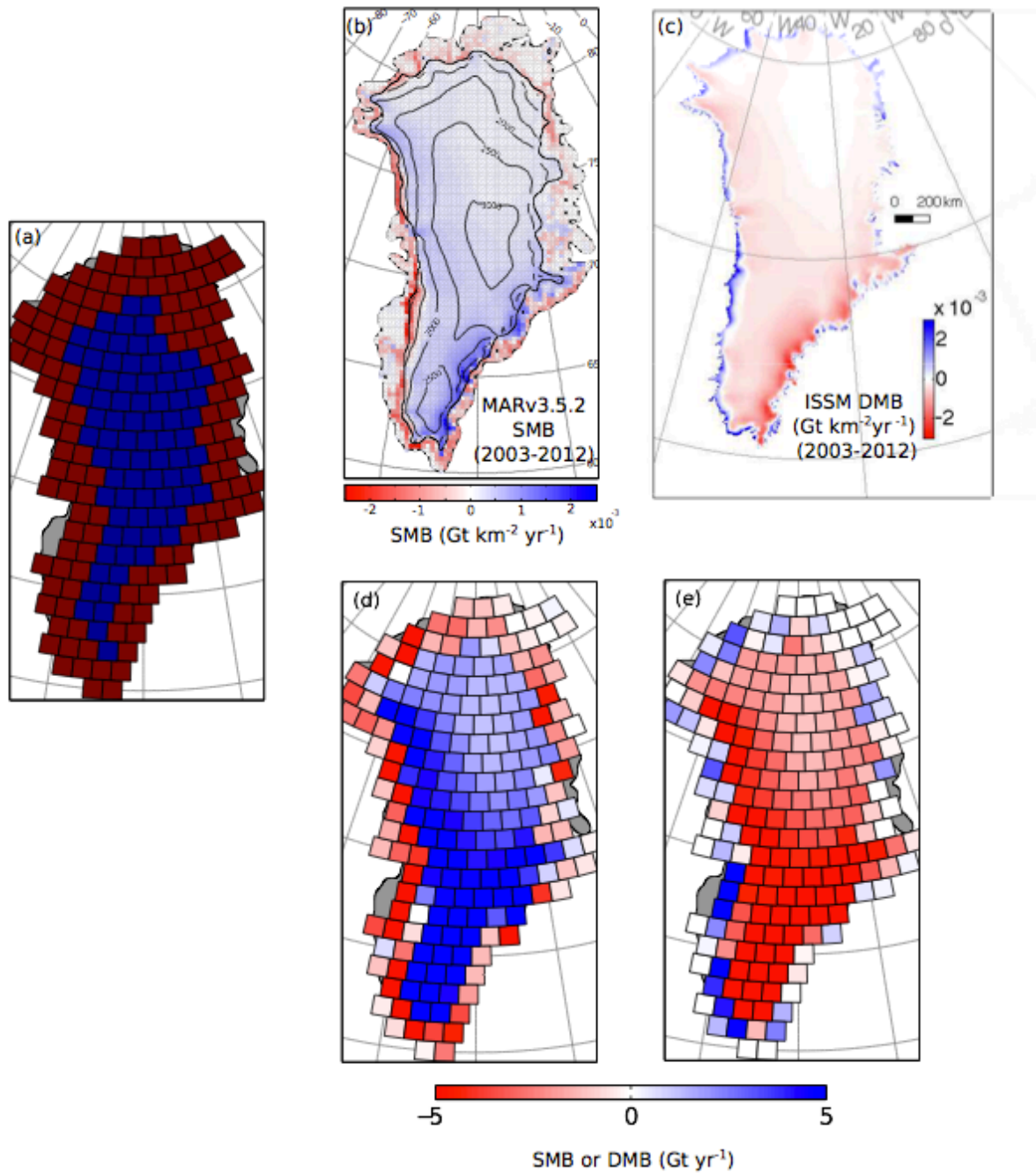
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1 Table 2. Same as Table 1, but for GRACE-LM and MAR v3.5.2 + ISSM (with Gaussian  
 2 spatial and temporal filtering), for the period January 2003 – December 2012.

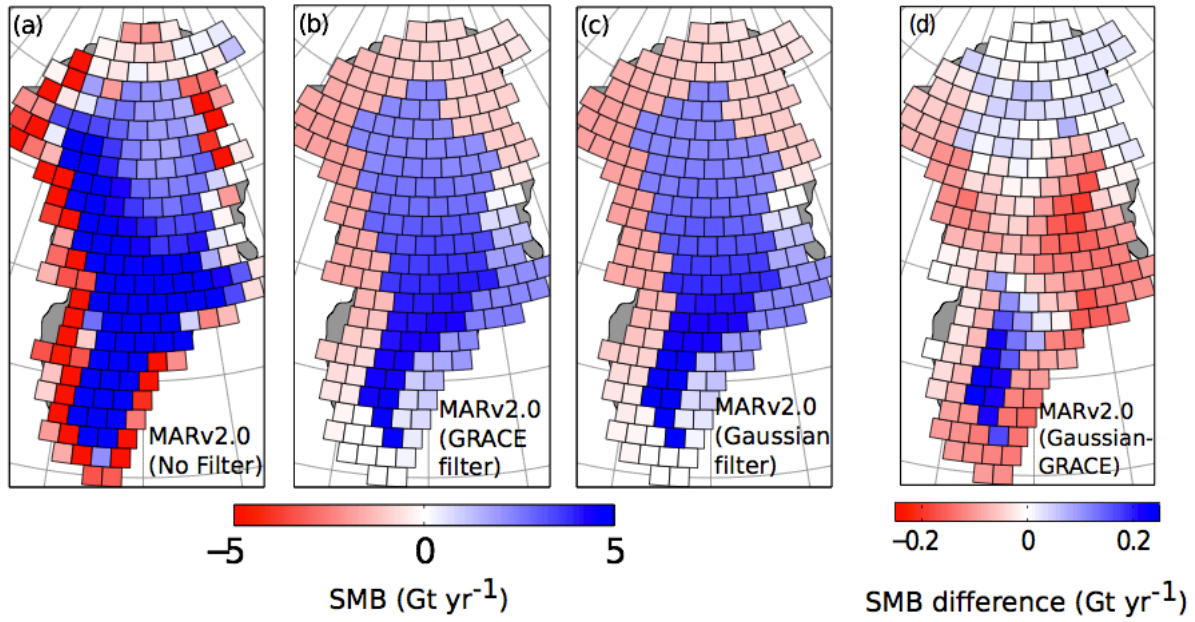
	<u>MAR v3.5.2 + ISSM</u>	<u>GRACE (median and 95% CI)</u>	<u>GRACE (median and 95% CI, detrended)</u>	<u>MAR v3.5.2 (detrended)</u>
<u>Maximum (2.5% Bound)</u>		<u>20 Feb</u>	<u>12 Mar</u>	
<u>Maximum</u>	<u>20 Apr</u>	<u>22 Mar</u>	<u>21 Apr</u>	<u>30 Apr</u>
<u>Maximum (97.5% Bound)</u>		<u>11 May</u>	<u>21 May</u>	
<u>Minimum (2.5 % Bound)</u>		<u>8 Sep</u>	<u>29 Aug</u>	
<u>Minimum</u>	<u>1 Oct</u>	<u>8 Oct</u>	<u>18 Sep</u>	<u>20 Sep</u>
<u>Minimum (97.5% Bound)</u>		<u>17 Dec</u>	<u>28 Oct</u>	

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 2 Figure 1. (a) Grid of mascons over the GrIS and constraint regions for the GRACE solution of  
 3 Luthcke et al. (2013). Areas below 2000 m in elevation are red, while areas above 2000 m are  
 4 blue. (b) MAR v3.5.2 average specific Surface Mass Balance (SMB,  $\text{Gt km}^{-2} \text{yr}^{-1}$ ) for January  
 5 2003 – December 2012 on the MAR grid, for the GrIS and periphery (contours show  
 6 elevation above sea level). (c) ISSM average specific Dynamic Mass Balance (DMB,  $\text{Gt km}^{-2}$   
 7  $\text{yr}^{-1}$ ) for the same period on the ISSM mesh. SMB and DMB for the 2003-2012 period  
 8 aggregated to the GRACE-LM grid (without filtering) are also shown for MAR v3.5.2 (d),  
 9 and ISSM (e).

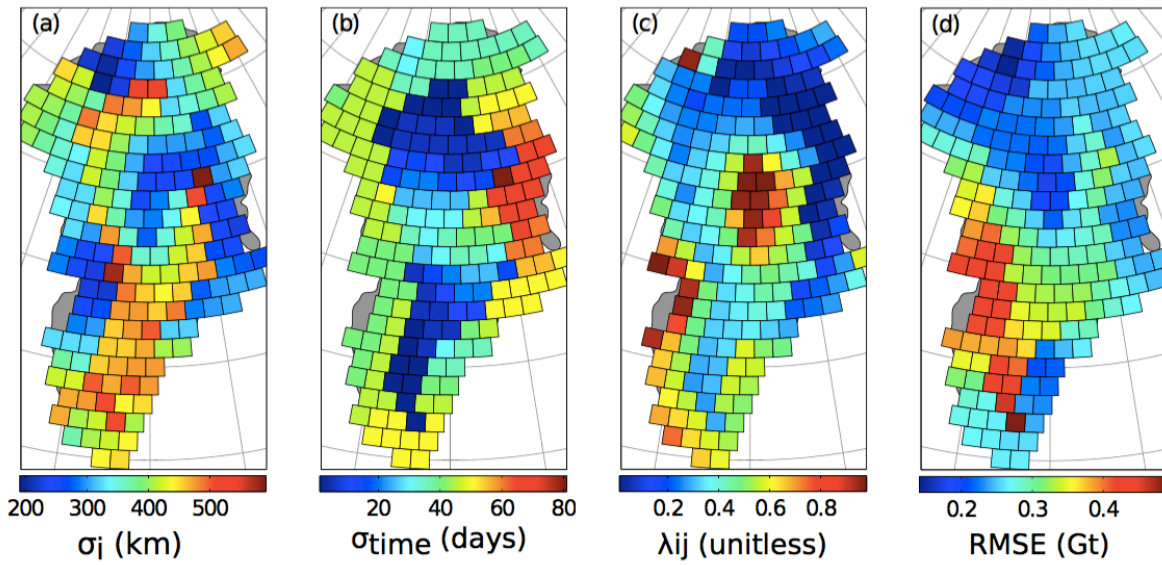
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3 Figure 2. Average MAR v2.0 SMB ( $\text{Gt yr}^{-1}$ ) for the period January 2003 – December 2010:  
4 (a) averaged onto GRACE-LM mascons with no filtering, (b) filtered using the resolution  
5 operator from GRACE-LM processing, (c) filtered using a Gaussian approximation to  
6 GRACE-LM filtering in space and time. (d) The difference between (a) and (b). Note the  
7 difference in color scale for (d).

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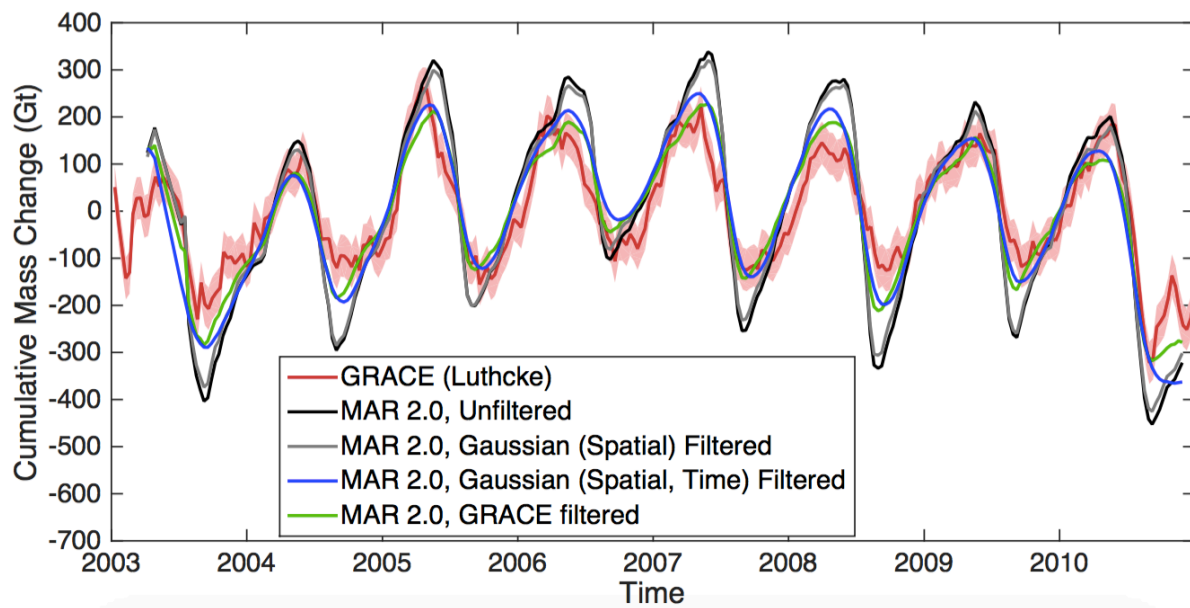


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2 Figure 3. Optimal values of parameters used in spatial and temporal Gaussian filtering of  
 3 MAR v3.5.2 and ISSM data: (a) the spatial Gaussian radius, (b) the temporal gaussian radius,  
 4 and (c) the leakage parameter. (d) RSME (Gt) for GRACE-LM-filtered vs. Gaussian-filtered  
 5 MAR v2.0 (January 2003 – December 2010).

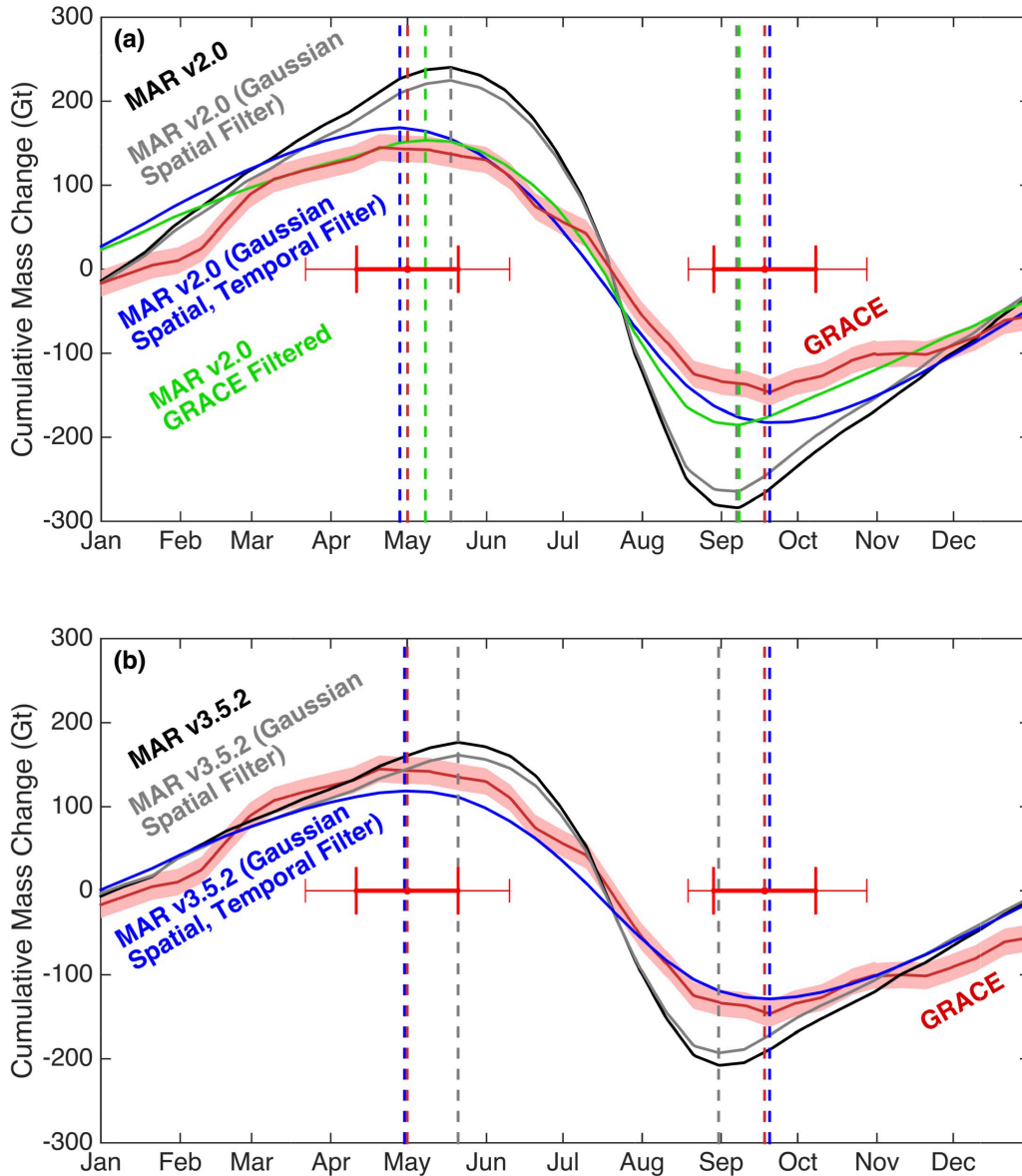
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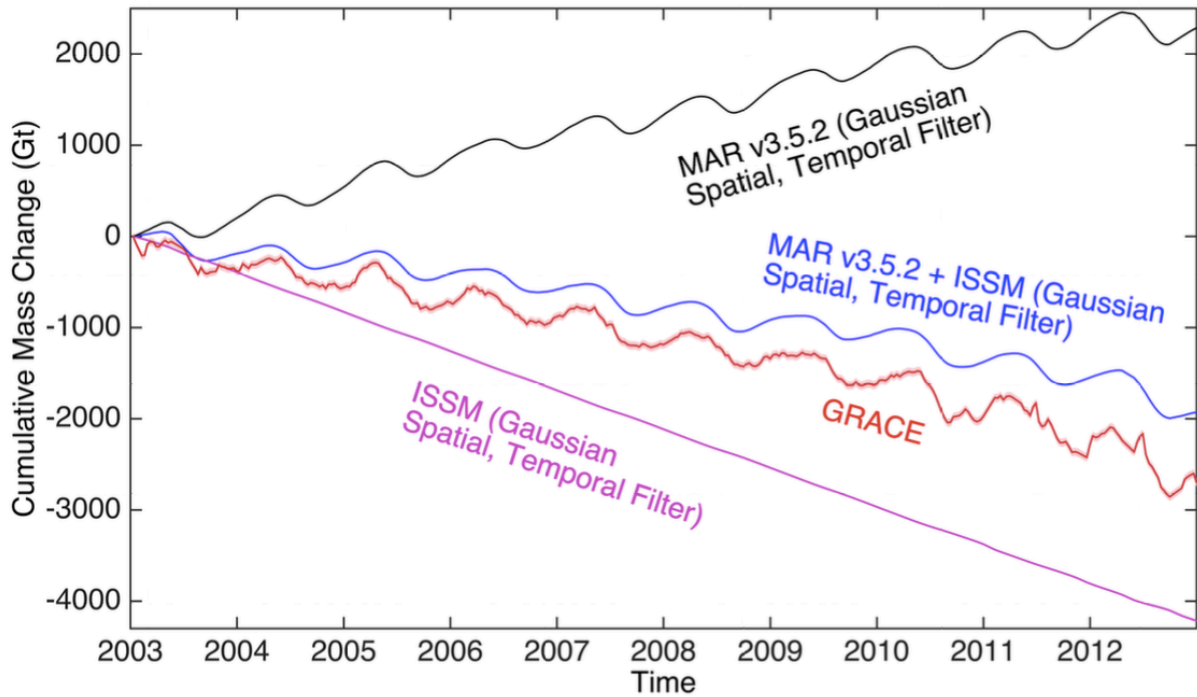
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 2 Figure 4. Detrended timeseries of cumulative GrIS-wide mass change, for GRACE-LM,  
 3 MAR v2.0 (unfiltered), MAR v2.0 (GRACE-LM-filtered), and MAR v2.0 (Gaussian-filtered)  
 4 for January 2003 – December 2010. Timeseries are shown for Gaussian filtered MAR v2.0  
 5 outputs subject to only spatial filtering (gray curve) and both spatial and temporal filtering  
 6 (blue curve). The pink shading indicates the range of error for the GRACE-LM timeseries.





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 2 Figure 5. Average seasonal cycle of January 2003 – December 2010 GrIS-wide de-trended  
 3 cumulative mass change (a) for GRACE-LM, unfiltered MAR v2.0 data, GRACE-LM-  
 4 filtered MAR v2.0 data, and Gaussian-filtered MAR v2.0 data (with spatial filtering and  
 5 spatial+temporal filtering). (b) The same as (a), for MAR v3.5.2 (for which GRACE-LM-  
 6 filtered data are not available). Vertical dashed lines indicate the timing of peaks of  
 7 maximum and minimum mass in the cycle. Red horizontal error bars indicate the error in the  
 8 timing of the GRACE cycle. Bold error bars indicate 50% of the GRACE-LM distribution,  
 9 and thin red lines indicate 95% of the distribution. The error bars have been extended for 10  
 10 days in either direction to account for errors associated with temporal resolution.

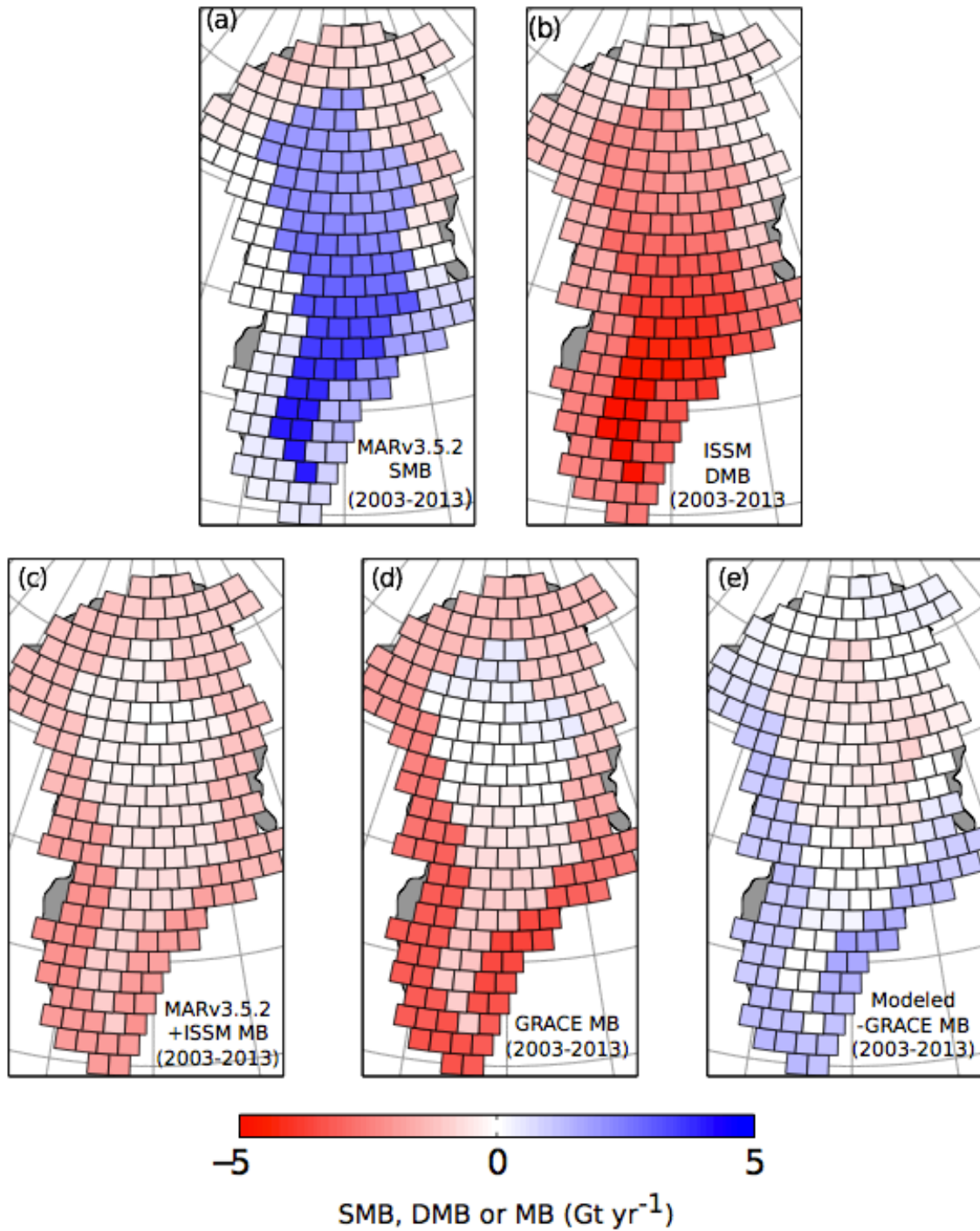
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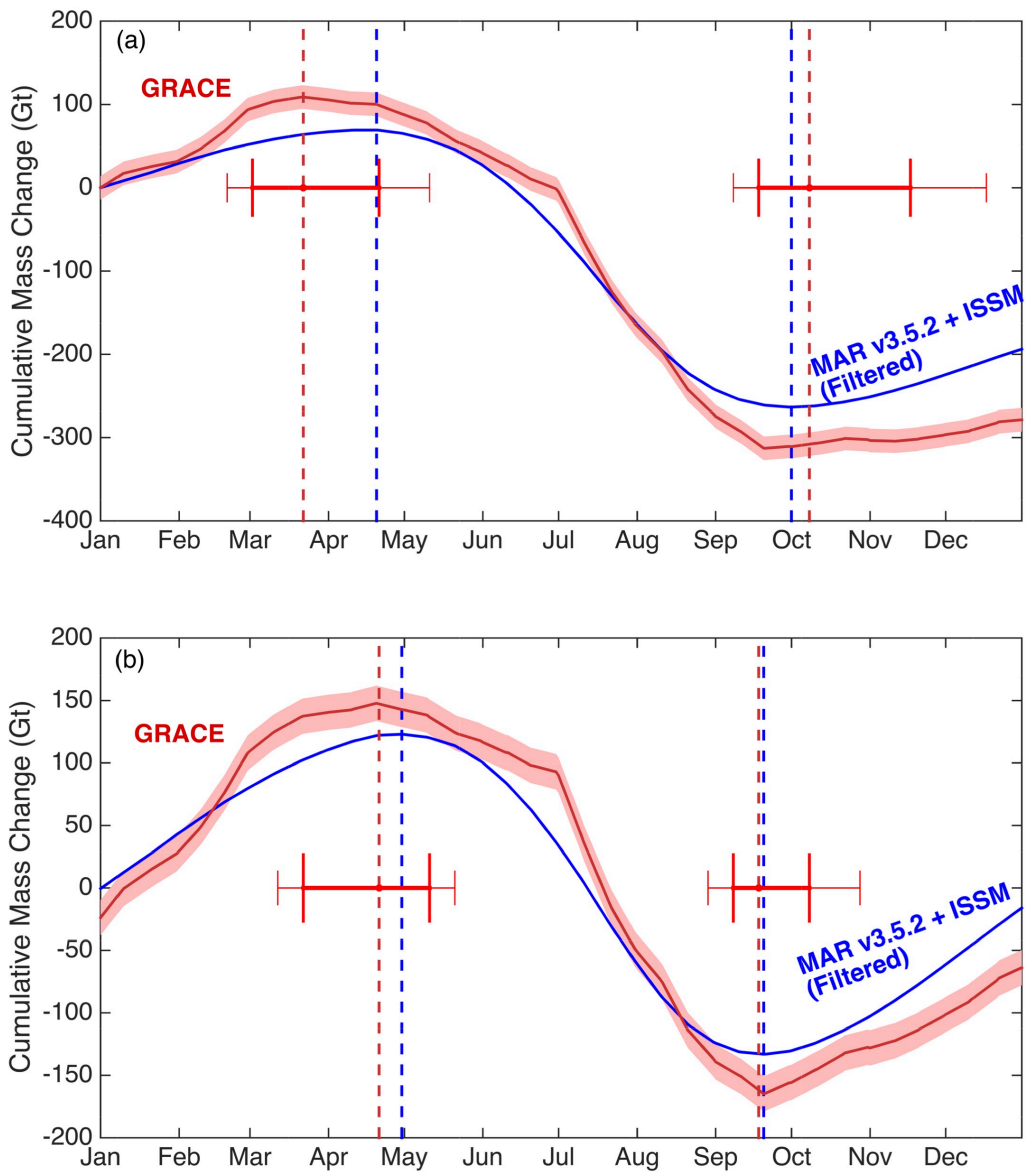
3 Figure 6. Cumulative GrIS mass change for the January 2003 – December 2012 period from  
4 GRACE-LM, MAR SMB, ISSM DMB, and the combined results of MAR v3.5.2 + ISSM.  
5 Spatial and temporal Gaussian filtering is applied to model outputs. The estimated error  
6 associated with the GRACE-LM solution is small relative to the overall trend, and is denoted  
7 with pink shading surrounding the GRACE-LM timeseries.

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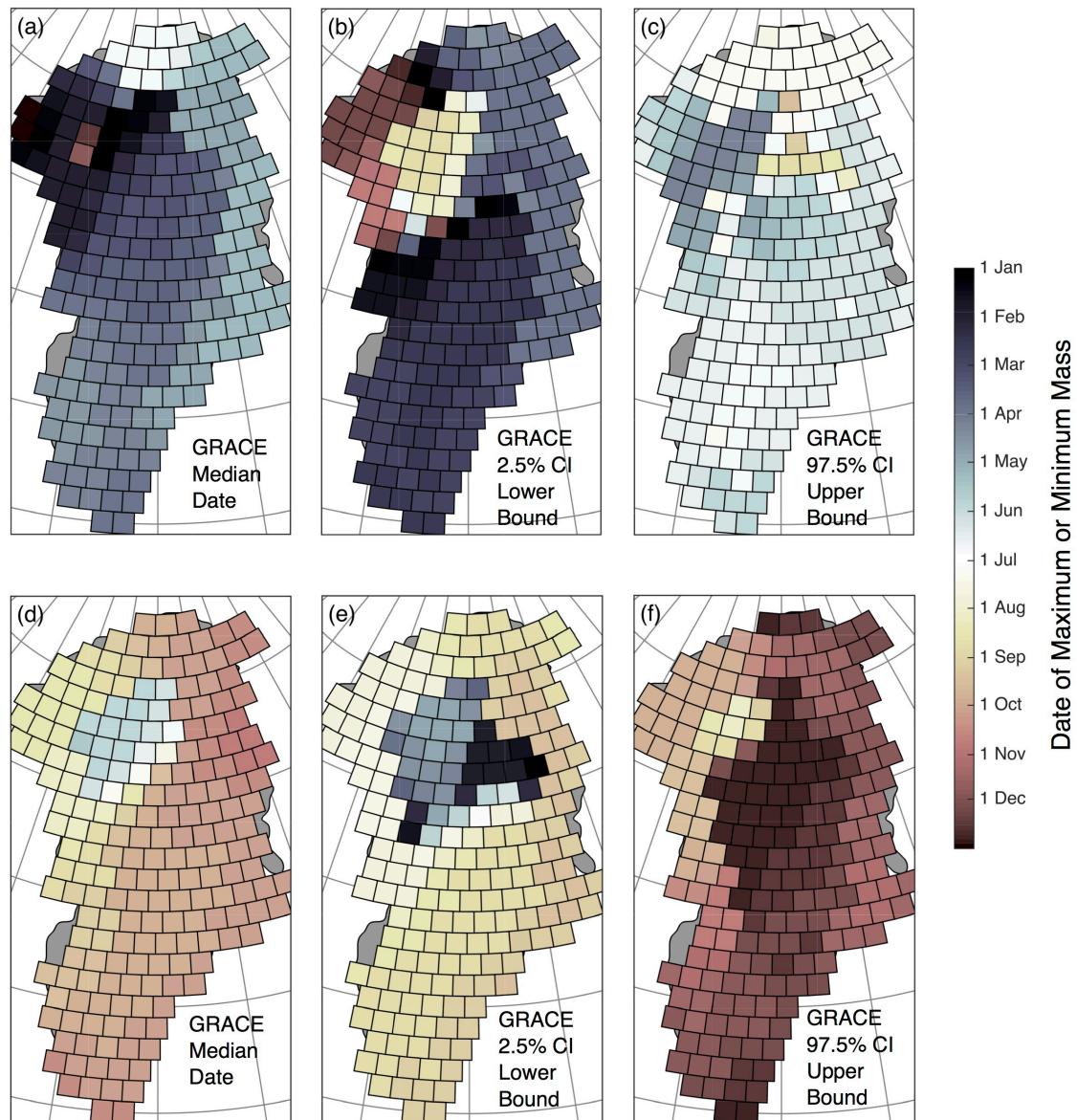
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 3 Figure 7. Average January 2003 – December 2012 mass balance (Gt) for MAR, ISSM and  
 4 GRACE-LM. Gaussian spatial and temporal filtering have been applied to MAR and ISSM  
 5 outputs. (a) SMB from MAR, (b) Dynamic mass change from ISSM, (c) The sum of (a) and  
 6 (b), giving the mean MB. GRACE-LM mean MB is shown in (d), and (e) depicts the  
 7 difference between modeled MB (c) and GRACE-LM MB (d).

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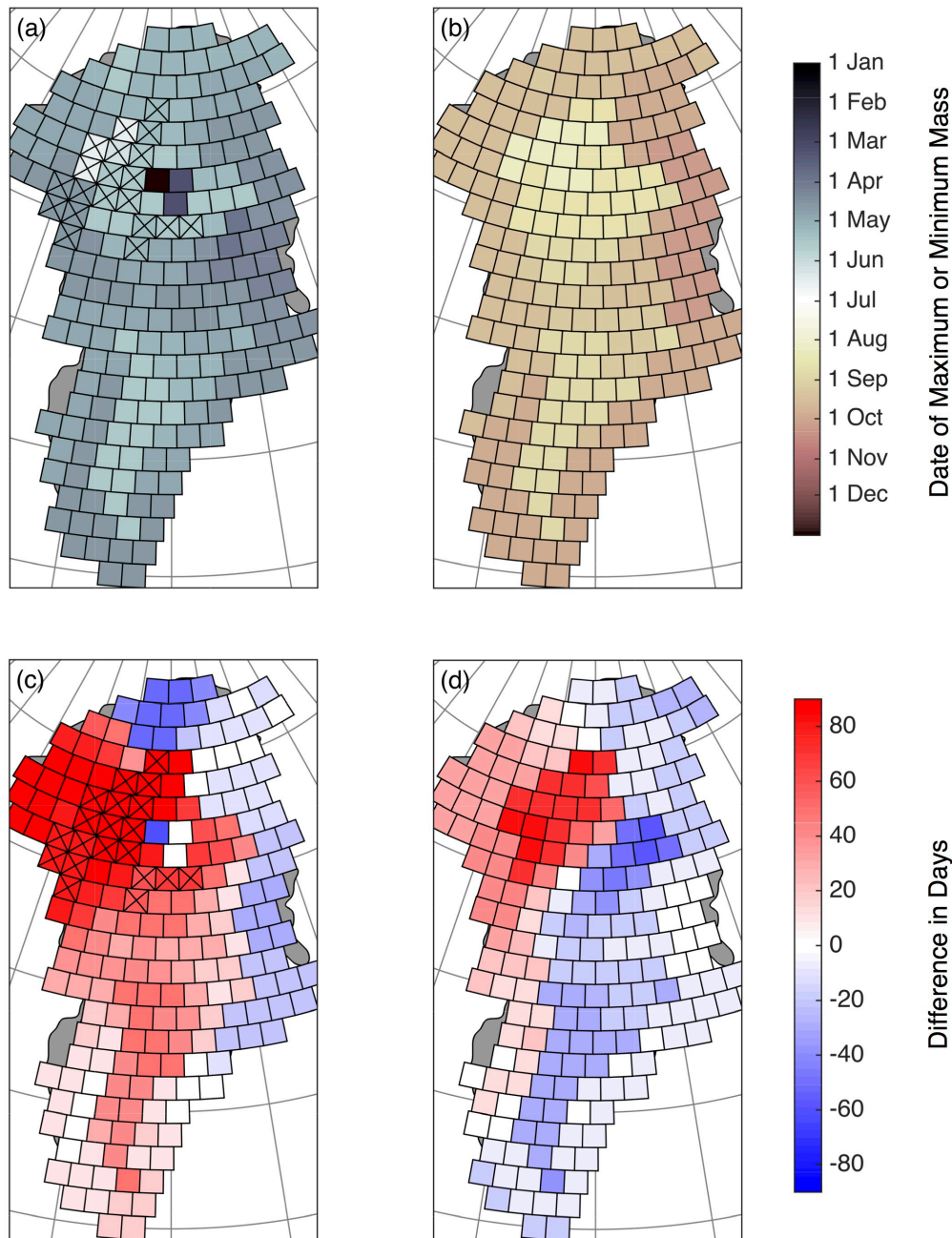
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 2 Figure 8. (a) The mean January 2003 – December 2012 seasonal cycle of GrIS cumulative  
 3 mass change from GRACE-LM and MAR v3.5.2 + ISSM. (b) The same as (a) for the case  
 4 when the timeseries from all mascons are detrended.

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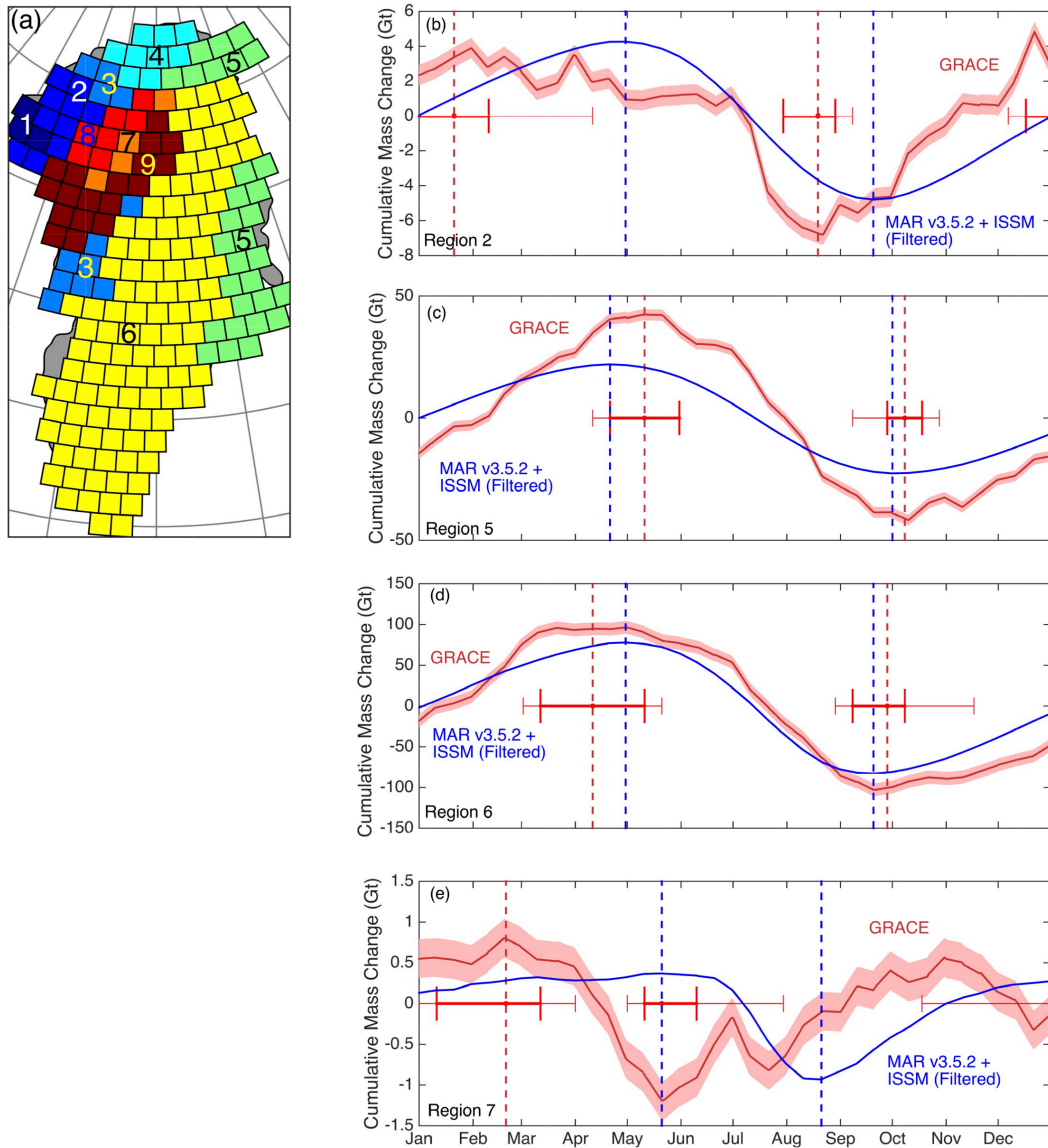
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2 Figure 9. Timing of peaks in the average January 2003 – December 2012 seasonal cycle of  
3 detrended cumulative mass from GRACE-LM. (a) Dates of maximum mass for each mascon  
4 from the median of the distribution from GRACE-LM (b) the 2.5% limit on the distribution of  
5 maximum mass dates, and (c) the 97.5% limit. (d) Dates of minimum mass for each mascon,  
6 and (e) the 2.5% and (f) 97.5% limits. 10 days have been subtracted (added) to the lower  
7 (upper) bound on the distributions to account for potential errors associated with temporal  
8 resolution.

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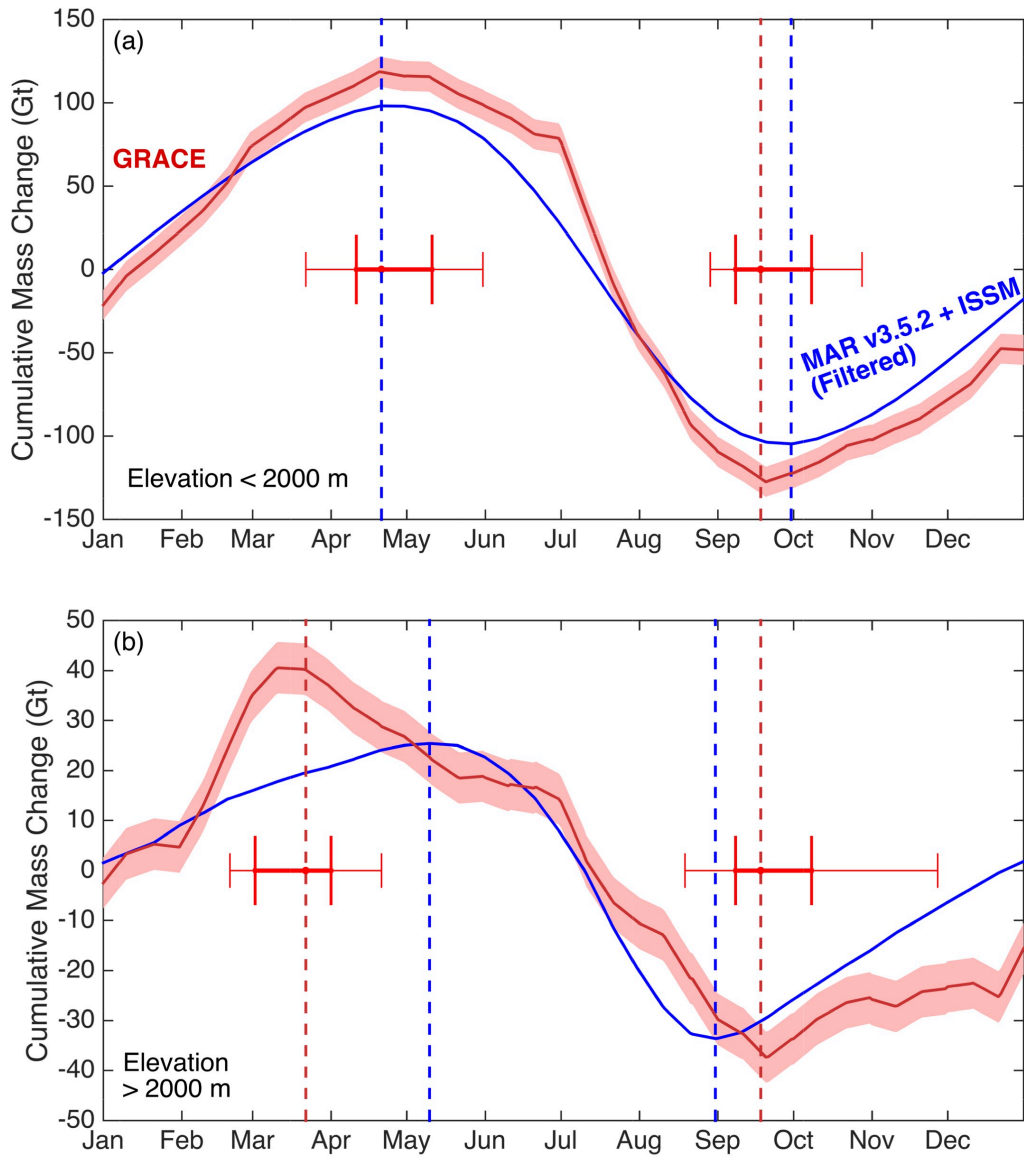
2 Figure 10. Timing in peaks of the seasonal cycle of de-trended cumulative MB simulated by  
 3 GRACE-LM-filtered MAR v3.5.2+ISSM outputs. (a) The timing of the maximum peak for  
 4 each mascon, and (b) the timing of the minimum peak for each mascon. The number of days  
 5 between MAR v3.5.2 + ISSM and the GRACE-LM median dates for the cycle maximum and  
 6 minimum are shown in (c) and (d) respectively. For (c) and (d) red colors indicate that the  
 7 model date occurs later than that of GRACE-LM, and blue colors indicate an earlier date. 'x'  
 8 marks indicate where the modeled peak falls outside of the error range of dates for the  
 9 GRACE-LM peak shown in Fig. 9.



1  
 2 Figure 11. (a) GrIS regions defined based on the timing of peaks in the average cycle of  
 3 detrended cumulative mass change from GRACE-LM. Also shown is the average January  
 4 2003 to December 2012 seasonal cycle from MAR v3.5.2 + ISSM and GRACE-LM for  
 5 selected regions: (b) Region 2, (c) Region 5, (d) Region 6, and (e) Region 7.

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3 Figure 12. Same as Fig. 8b, but for (a) regions below 2000 m in elevation, and (b) above 2000  
4 m in elevation.



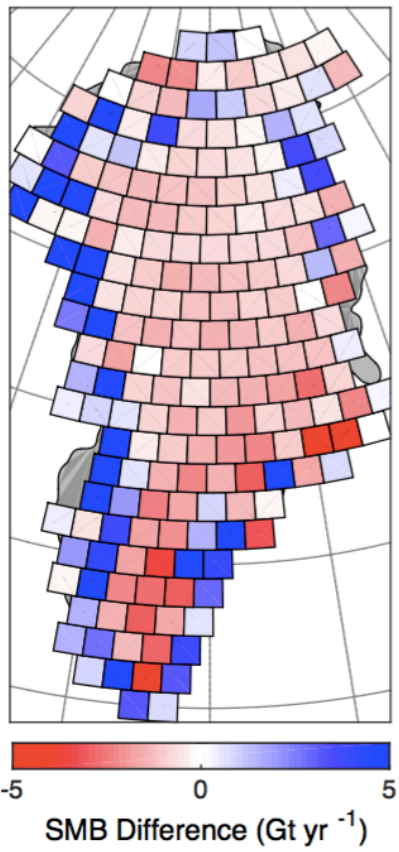
1 Table S1. Timing of the seasonal cycle maximum and minimum for cumulative mass change  
 2 (January 2003 – December 2012) for GRACE-LM, MAR v3.5.2, ISSM forced by MAR  
 3 v3.5.2, and MAR v3.5.2 + ISSM, within the different regions of the ice sheet defined in  
 4 Figure 12a, as well as the high (above 2000 m) and low (below 2000 m) elevation regions  
 5 defined in Figure 1a. The 2.5 % and 97.5% bounds and median value for GRACE-LM are  
 6 provided.

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	<u>&lt; 2000 m</u>	<u>&gt;2000 m</u>	<u>Reg 1</u>	<u>Reg 2</u>	<u>Reg 3</u>	<u>Reg 4</u>	<u>Reg 5</u>	<u>Reg 6</u>	<u>Reg 7</u>	<u>Reg 8</u>	<u>Reg 9</u>
<u>Maximum:</u>											
<u>GRACE (2.5 %)</u>	<u>22 Mar</u>	<u>20 Feb</u>	<u>7 Dec</u>	<u>7 Dec</u>	<u>21 Jan</u>	<u>20 Jun</u>	<u>11 Apr</u>	<u>2 Mar</u>	<u>18 Oct</u>	<u>28 Sep</u>	<u>11 Jan</u>
<u>GRACE median</u>	<u>21 Apr</u>	<u>22 Mar</u>	<u>27 Dec</u>	<u>21 Jan</u>	<u>1 Apr</u>	<u>30 Jun</u>	<u>11 May</u>	<u>11 Apr</u>	<u>20 Feb</u>	<u>11 Jan</u>	<u>10 Feb</u>
<u>GRACE (97.5%)</u>	<u>31 May</u>	<u>21 Apr</u>	<u>11 Jan</u>	<u>11 Apr</u>	<u>1 May</u>	<u>10 Jul</u>	<u>31 May</u>	<u>21 May</u>	<u>1 Apr</u>	<u>2 Mar</u>	<u>11 Apr</u>
<u>MAR v3.5.2</u>	<u>30 Apr</u>	<u>11 May</u>	<u>30 Apr</u>	<u>30 Apr</u>	<u>30 Apr</u>	<u>10 May</u>	<u>21 Apr</u>	<u>30 Apr</u>	<u>21 May</u>	<u>21 May</u>	<u>30 Apr</u>
<u>MAR v3.5.2 + ISSM</u>	<u>21 Apr</u>	<u>10 May</u>	<u>30 Apr</u>	<u>30 Apr</u>	<u>30 Apr</u>	<u>10 May</u>	<u>21 Apr</u>	<u>30 Apr</u>	<u>21 May</u>	<u>11 May</u>	<u>30 Apr</u>
<u>ISSM</u>	<u>21 Oct</u>	<u>1 Oct</u>	<u>10 Oct</u>	<u>10 Oct</u>	<u>1 Oct</u>	<u>1 Oct</u>	<u>21 Oct</u>	<u>10 Oct</u>	<u>1 Oct</u>	<u>10 Sep</u>	<u>1 Oct</u>
<u>Minimum:</u>											
<u>GRACE (2.5 %)</u>	<u>29 Aug</u>	<u>19 Aug</u>	<u>30 Jul</u>	<u>30 Jul</u>	<u>20 Jul</u>	<u>19 Aug</u>	<u>8 Sep</u>	<u>29 Aug</u>	<u>1 May</u>	<u>21 Apr</u>	<u>10 Jul</u>
<u>GRACE median</u>	<u>18 Sep</u>	<u>18 Sep</u>	<u>19 Aug</u>	<u>19 Aug</u>	<u>29 Aug</u>	<u>18 Sep</u>	<u>8 Oct</u>	<u>28 Sep</u>	<u>21 May</u>	<u>21 May</u>	<u>30 Jul</u>
<u>GRACE (97.5%)</u>	<u>28 Oct</u>	<u>27 Nov</u>	<u>8 Oct</u>	<u>8 Sep</u>	<u>28 Sep</u>	<u>18 Oct</u>	<u>28 Oct</u>	<u>17 Nov</u>	<u>30 Jul</u>	<u>9 Aug</u>	<u>18 Sep</u>
<u>MAR v3.5.2</u>	<u>30 Sep</u>	<u>31 Aug</u>	<u>20 Sep</u>	<u>21 Sep</u>	<u>20 Sep</u>	<u>20 Sep</u>	<u>1 Oct</u>	<u>20 Sep</u>	<u>21 Aug</u>	<u>11 Aug</u>	<u>10 Sep</u>
<u>MAR v3.5.2 + ISSM</u>	<u>30 Sep</u>	<u>31 Aug</u>	<u>20 Sep</u>	<u>20 Sep</u>	<u>20 Sep</u>	<u>20 Sep</u>	<u>1 Oct</u>	<u>20 Sep</u>	<u>21 Aug</u>	<u>11 Aug</u>	<u>10 Sep</u>
<u>ISSM</u>	<u>1 Jun</u>	<u>20 Jun</u>	<u>21 May</u>	<u>21 May</u>	<u>1 Jun</u>	<u>22 May</u>	<u>31 May</u>	<u>10 Jun</u>	<u>1 Jun</u>	<u>1 Jun</u>	<u>1 Jun</u>

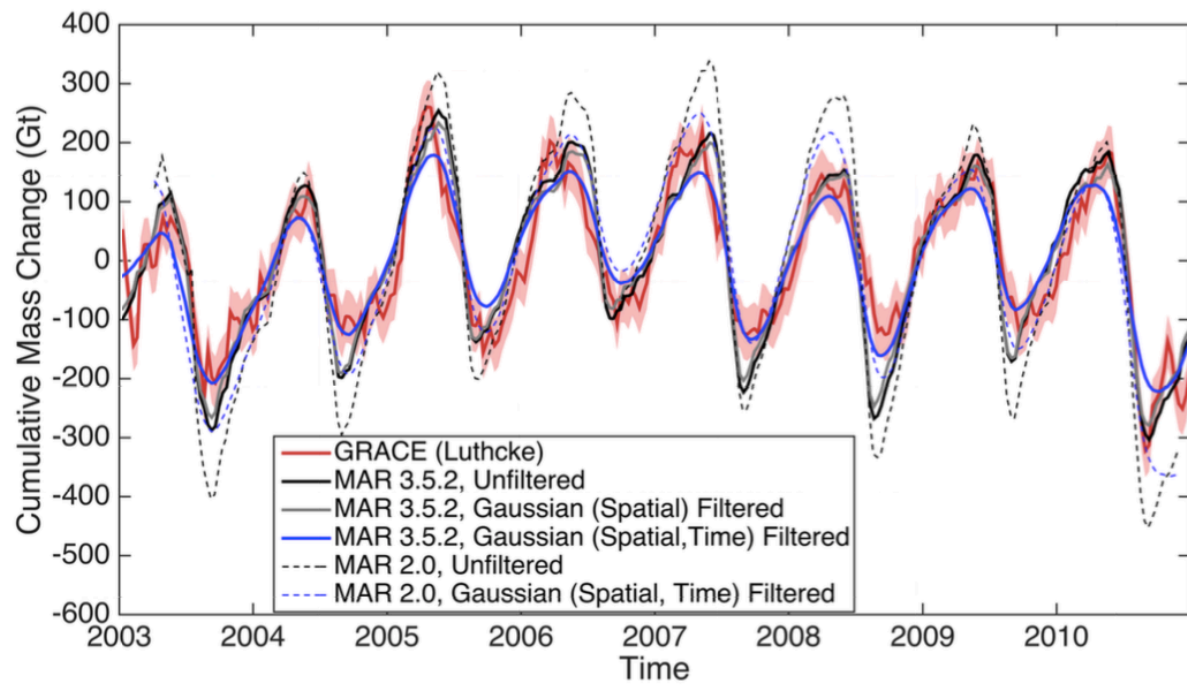
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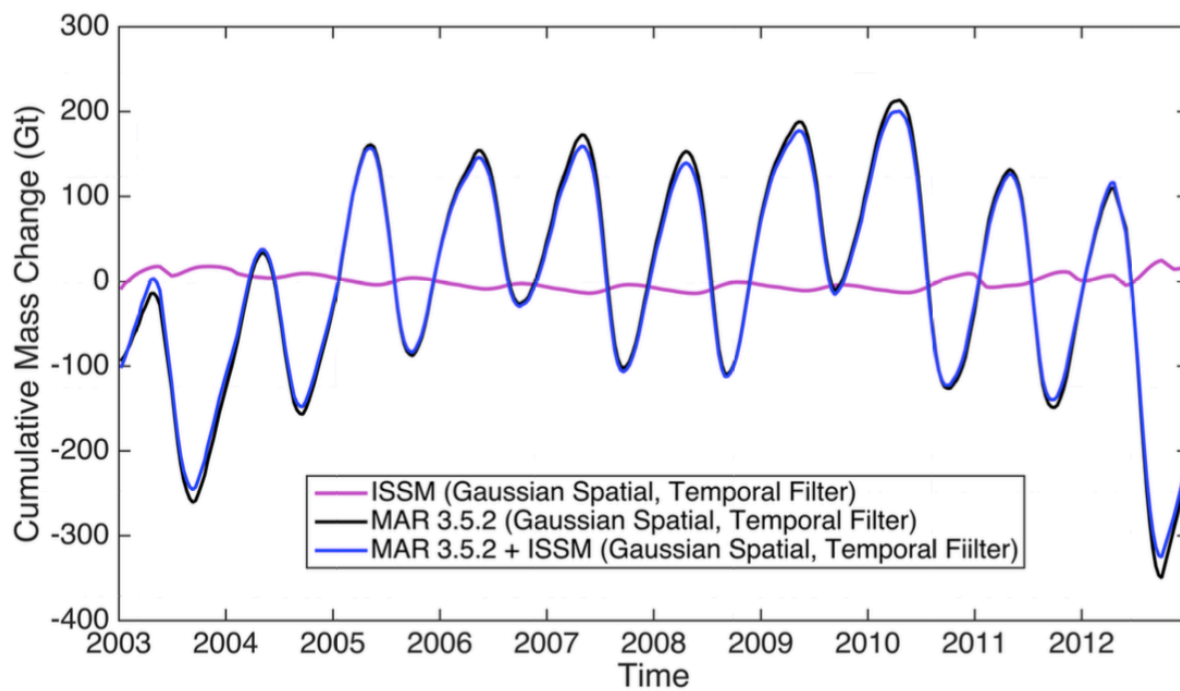
Figure S1. MAR v3.5.2 – MAR v2.0 average SMB, aggregated to the GRACE-LM grid, with no filtering applied, for the period January 2003 – December 2010.



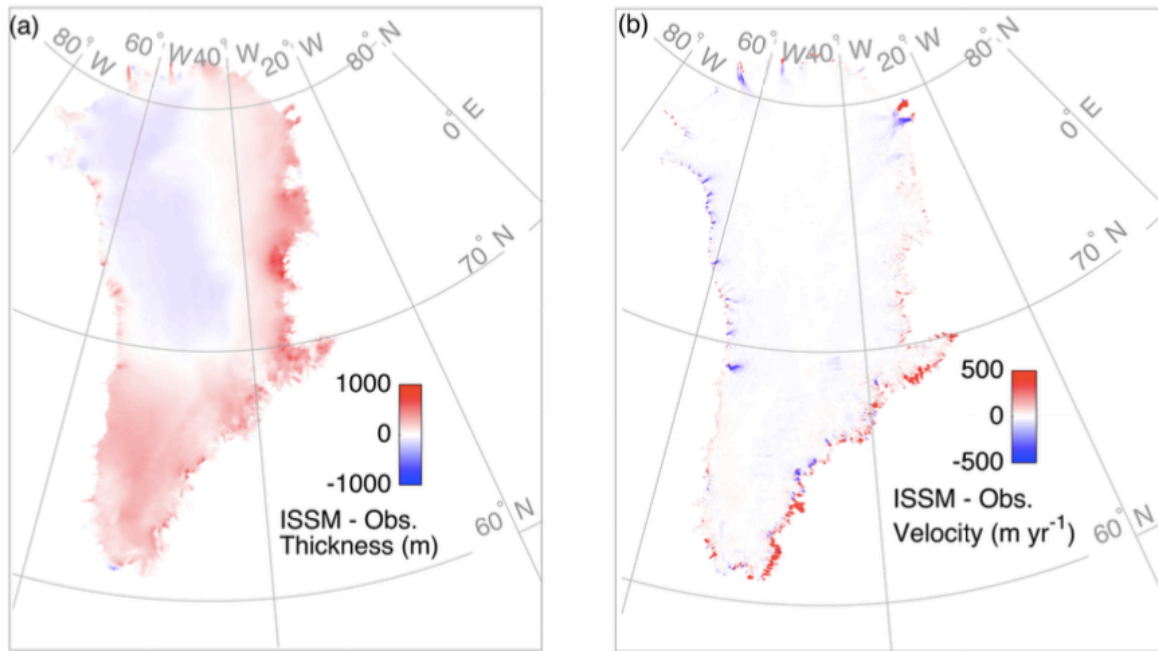
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2 Figure S2: Same as Figure 4, but for MAR v3.5.2. Dashed lines show the MAR v2.0  
3 detrended unfiltered and Gaussian-filtered timeseries for comparison with MAR v3.5.2.

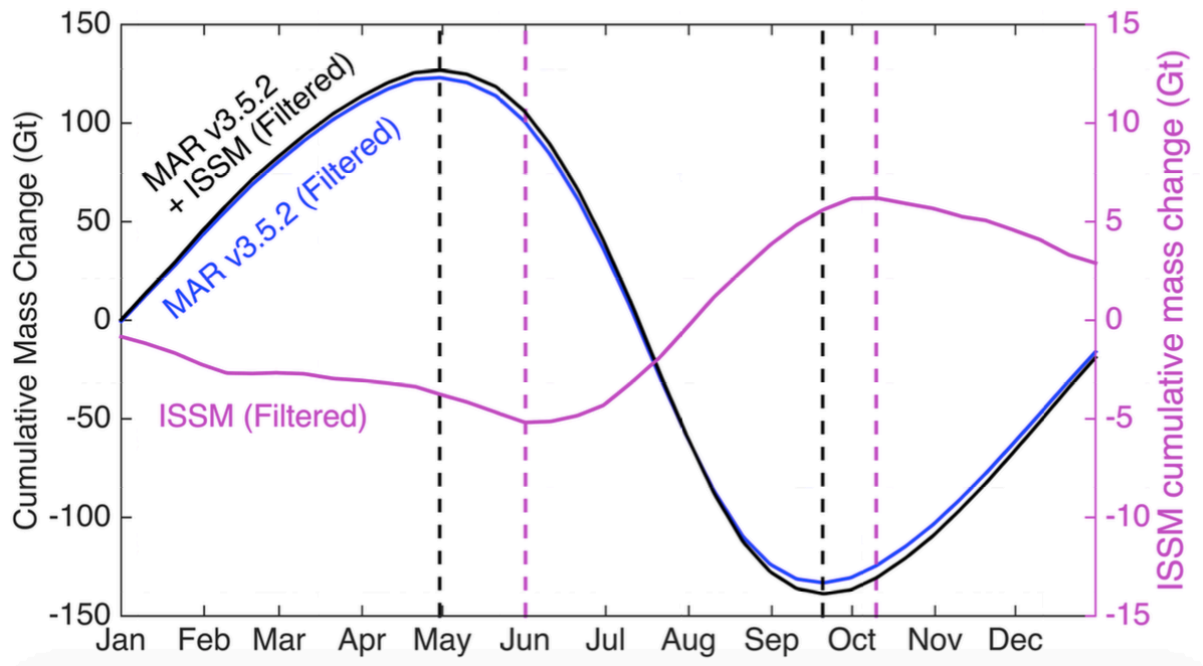
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 2 Figure S3: Detrended cumulative timeseries of Greenland mass change for MAR v3.5.2,  
 3 ISSM, and MAR v3.5.2 + ISSM. Spatial and temporal Gaussian filtering has been applied to  
 4 the model outputs.  
 5

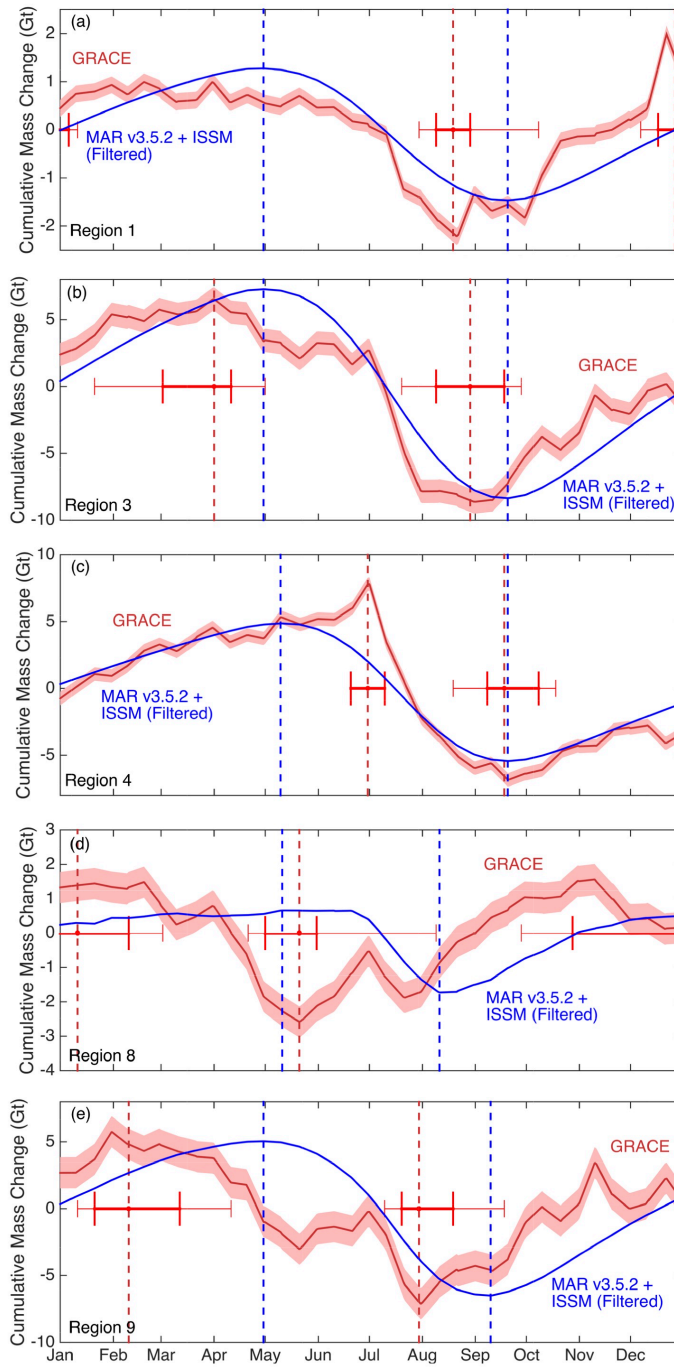


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 2 Figure S4: Difference between ISSM and observations for (a) Greenland Ice Sheet thickness  
 3 (m), and (b) Annual ice velocities, for the year 2003. Thicknesses are obtained from  
 4 Morlighem et al. (2015), and velocities are those of Rignot et al. (2012).  
 5



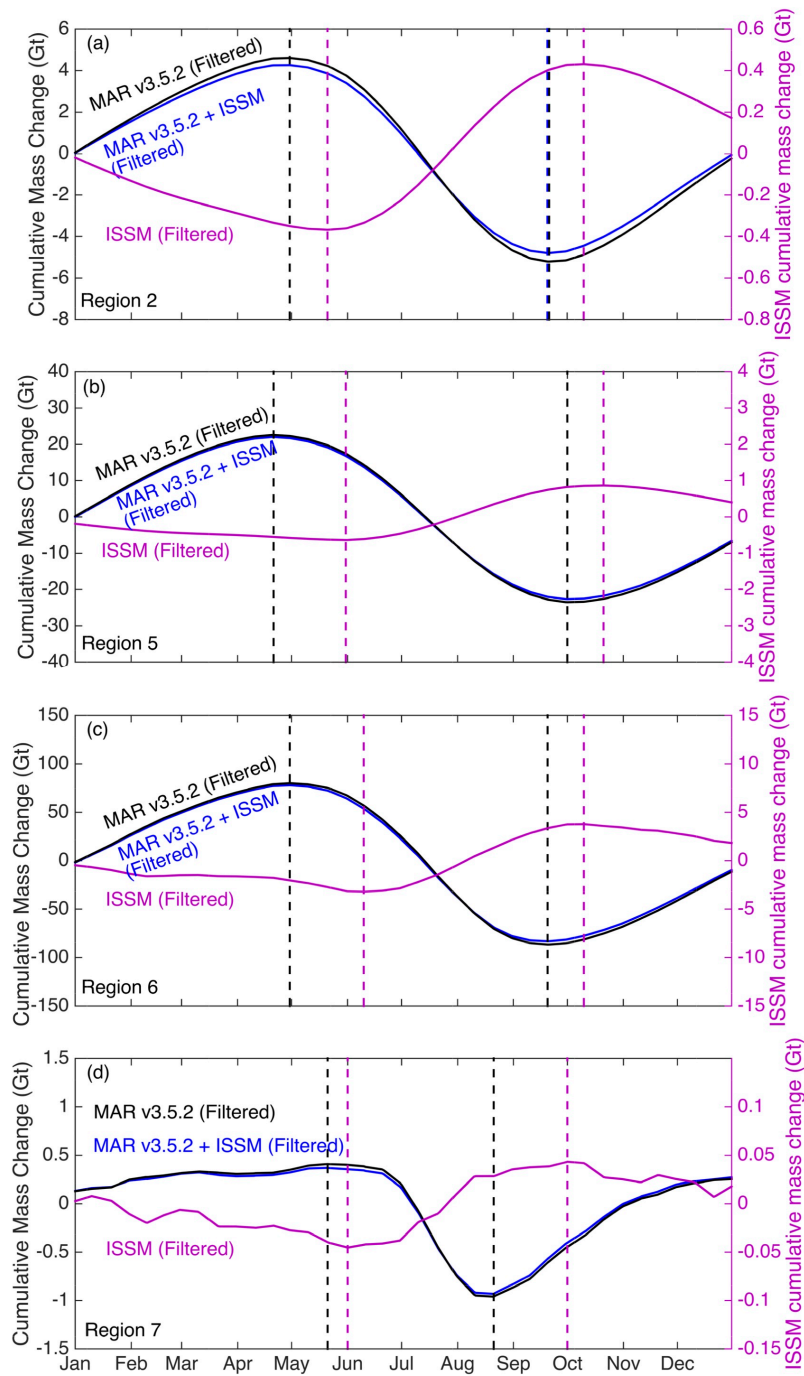
1  
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Figure S5: Same as Fig. 8b for MARv 3.5.2 + ISSM, MAR v3.5.2 and ISSM. Note a different axis is used for ISSM data.



1  
 2 Figure S6: Same as Fig. 11b-e, but for additional regions shown in Fig. 11a: (a) Region 1, (b)  
 3 Region 3, (c) Region 4, (d) Region 8, and (e) Region 9

4

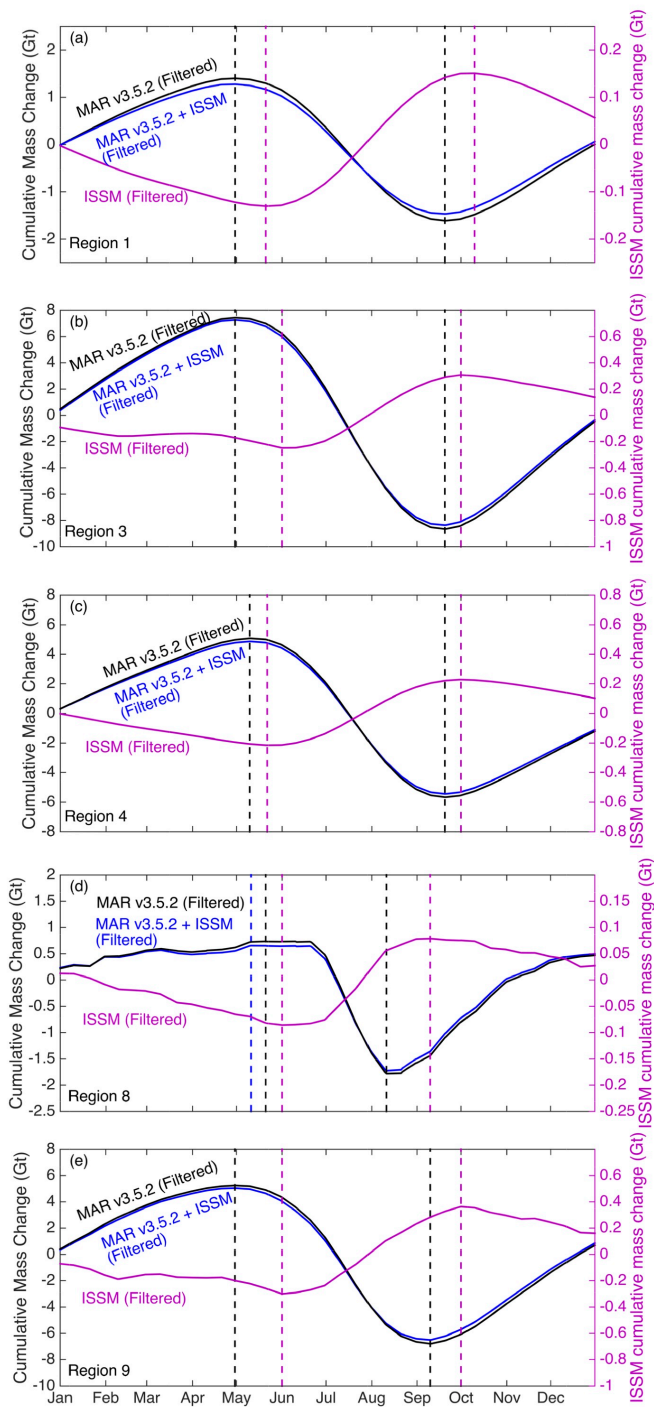


1

2 Figure S7: Same as Fig. 11b-e, but showing seasonal cycles for MAR v3.5.2, ISSM and  
 3 MAR v3.5.2 + ISSM. Note the different axis for ISSM.

4

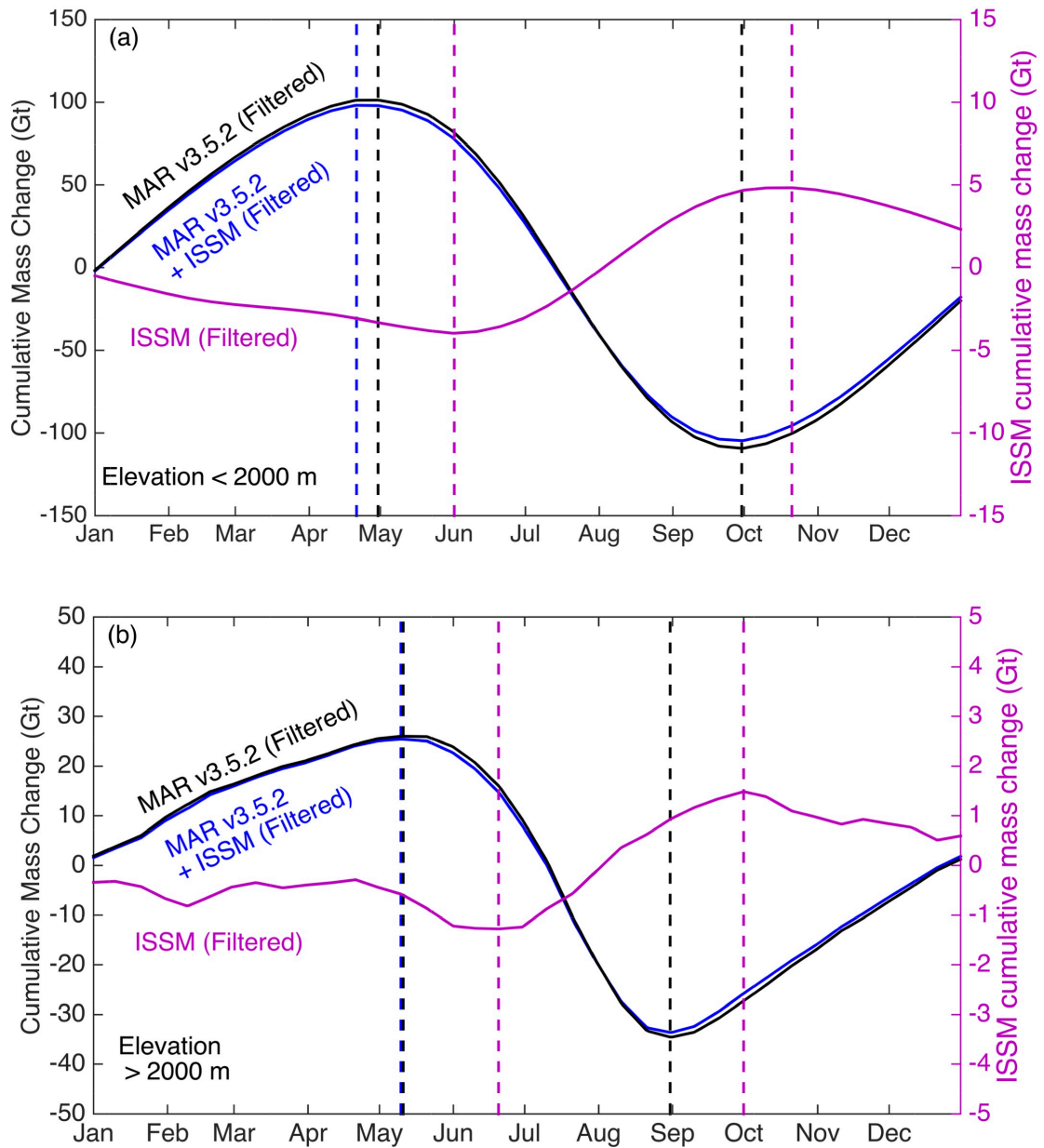




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3 Figure S8: Same as Fig. S6, but showing seasonal cycles for MAR v3.5.2, ISSM and MAR  
4 v3.5.2 + ISSM. Note the different axis for ISSM.

5



1  
 2 Figure S9: Same as Fig. 12, but showing seasonal cycles for MAR v3.5.2, ISSM and MAR  
 3 v3.5.2 + ISSM. Note the different axis for ISSM.

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