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sions as well as in their geographical distribution. Large uncertainties in past and current global inventories of anthropogenic and natural emissions limit the establishment of a reliable framework of emission inventories required for realistic emission scenarios. This is especially true for the Indian subcontinent for which reliable estimates of past and current emissions are crucially lacking. India is one of the two largest anthropogenic aerosol generating countries in the world (Lu et al., 2011). In the past decade, India has been identified as a hot spot in terms of high aerosol optical depth (AOD) observed from space (Prasad and Singh, 2007) with components such as sulfate, organic carbon (OC) or black carbon (BC) playing a very active role. India on its own contributes 10 to 20% of all current emissions globally (Bond et al., 2007) and has therefore received the greatest attention from compilers of emission inventories.

Recent works performed within the “Atmospheric Brown Cloud” and the “SHARE” projects (<http://www.evk2cnr.org>) indicated that the south-facing side of the Himalayas is particularly affected by emissions from the Indo-Gangetic plains. Unique atmospheric observations team at the global GAW (Global Atmosphere Watch) station of Nepal Climate Observatory-Pyramid (NCO-P) located at 5079 m a.s.l. on the southern foothills of Mt. Everest (Bonasoni et al., 2010), documented for the first time the efficient transport of short-lived climate pollutants (SLCPs) to the high altitude. Continuous observations over more than 7 yr revealed very high concentrations of SLCPs in particular BC (Marinoni et al., 2010) and ozone (Cristofanelli et al., 2010) especially in pre-monsoon season. The presence of light absorbing material affects the radiative balance of the atmosphere both at local (Marcq et al., 2010) and at a regional scale. Source areas of aerosols have been clearly identified for the high Himalaya by (Cristofanelli et al., 2010) and are located in the Indo-Gangetic plains for most part of the year. Source origin for aerosol species is likely linked to biofuel combustion (Decesari et al., 2010). Desert dust events are also regularly observed at NCO-P, either mixed with anthropogenic pollution from Pakistan and/or directly from Gobi desert (Bonasoni et al., 2010). Black carbon and dust are efficiently scavenged to the snow/ice surface by dry and wet deposition (Yasunari et al., 2010, 2013); these impurities can lead to surface albedo re-

laboratory and that the samples were going to melt during shipment, a few drops of formaldehyde solution were added to the samples in order to limit the bacteriological development in the vials. Temperature logger in the box recorded that samples were melted during 18 days of shipment before being refrozen in the laboratory.

2.3 Sample analyzes

Analyzes of major and organic chemical species were performed using a Dionex[®] ICS3000 dual ion chromatography system at LGGE. The chromatography system was setup for the analyzes of cations (Li^+ , Na^+ , NH_4^+ , K^+ , Mg^{2+} , Mn^{2+} , Ca^{2+} , Sr^{2+}) and anions (F^- , Cl^- , NO_2^- , Br^- , NO_3^- , $(\text{CH}_2)_2\text{C}_2\text{O}_4^{2-}$, SO_4^{2-} , $\text{C}_2\text{O}_4^{2-}$) down to sub-ppb level and high level accuracy (6 standards calibration, relative standard deviation < 2%). Other species determination was disturbed by the formaldehyde spiking. Insoluble dust measurements were performed using a microparticle counter (Coulter Counter[®] Multisizer III) for particles with equivalent diameter from 1.0 to 30 μm , divided in 300 equivalent size channels. The total mass of dust has been calculated from the volume size distribution assuming a density of 2.5 g cm^{-3} . Analytical systems are set up in the class-100 clean room of LGGE laboratory. Detailed analytical procedure and accuracy are described in (Delmonte et al., 2002).

The ^{18}O content of the ice was measured at the LAMA laboratory of HydroSciences Montpellier, on an Elementar Isoprime Mass Spectrometer coupled with an Aquaprep module, using the classical CO_2 equilibration technique. The results are expressed in $\delta^{18}\text{O}$ on the V-SMOW scale with an overall uncertainty of $\pm 0.06\%$. Half of the samples that were analyzed for chemical composition were measured for $\delta^{18}\text{O}$, which corresponds to a resolution of approximately 12 cm of snow.

The SP2 uses a laser-induced incandescence technique to measure the mass of individual refractory BC (rBC) independently from rBC particle morphology and light-scattering coating materials (Petzold et al., 2013; Cross et al., 2010; Moteki and Kondo, 2007, 2010; Schwarz et al., 2006). A single rBC particle passing through the laser

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beam intra-cavity absorbs light, reaches a vaporization temperature at which it incandesces, and emits visible thermal radiation. The incandescence signal is proportional to the mass of the individual rBC particle. Fullerene soot was used for calibration of incandescence signal. Use of SP2 for liquid samples requires nebulization which was performed by an APEX-Q, (EPOND, Switzerland) system. The efficiency of the APEX/SP2 system is accounted for applying a correction factor of 0.56 to rBC mass concentration in all samples. For rBC particles, we used the sum of individual rBC particle mass from ~ 50 to 600 nm mass equivalent diameter to obtain rBC mass concentration of each sample. Over the whole record, the highest number of particles appears between ~ 50 and 90 nm mass equivalent diameter (Lim et al., 2013).

The measurement of equivalent BC in the atmosphere, resulting from the aerosol absorption coefficient measure, was obtained at NCO-P by MAAP (Multi-Angle Absorption Photometer, © Thermo): it measures the transmission and the back scattering of a light beam (Petzold and Schönlinner, 2004) incident on a fibre filter where aerosol particles are deposited by sampling flow. The detection limit (3σ of blank measurements) was calculated as 11 ng m^{-3} , with an integration time of a 30 min basis (Marinoni et al., 2010).

3 Ice core records

3.1 Dating

The firn core was annually dated between 2010 and 1999 based on seasonal cycles of stable isotopes and chemical proxies. In this region, the stable isotope composition of precipitation is mostly controlled by the rainout effect associated to the Indian monsoon activity (Araguás-Araguás et al., 1998; Zhang et al., 2001; Vuille et al., 2005). Precipitation data collected during one year at the high altitude GNIP Lhajung station ($27^{\circ}53'42''$ N; $86^{\circ}49'30''$ E; 4420 m a.s.l., Nepal) (International Atomic Energy Agency/World Meteorological Organization, 2006), showed depleted values

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layers (Marinoni et al., 2001; Shrestha et al., 2002). During winter, westerly circulation associated to distant or reduced sources may generate low NH_4^+ concentration in the snow layers; but these layers are usually scarce (very dry conditions in winter) and often not preserved in the snow column (snow remobilized into the atmosphere due to strong winds at high elevations in winter, Wagon et al., 2013). The ammonium core profile was therefore used for dating, with high concentration peaks corresponding to the annual dry season (Fig. 2).

Isotopic and chemical dating methods result in a seasonally resolved dated profile ranging from November 2010 to 1999 (Fig. 2). The lowest NH_4^+ concentration layer observed between the surface and 1.19 m water equivalent (w.e.) depth, also corresponding to the most depleted $\delta^{18}\text{O}$ over the whole profile, is atypical in the profile and was related to the temporary preservation of pre-early winter snow from September to November 2010. This layer was composed by cold small loose snow grain, in opposition to the deeper part of the core composed by more compact, larger and transformed snow crystals and ice layers. Such early-winter snow layer is not preserved elsewhere along the profile probably because of strong wind erosion and loose snow remobilization during winter as systematically observed on-site by Wagon et al. (2013) since 2007.

The profile was divided into three seasonal patterns. As said earlier only one single winter event was identified between surface and 1.19 m w.e. depth. The snow fraction related to monsoon and inter-monsoon conditions was attributed according to the $\delta^{18}\text{O}$ composition. Along the profile, 68 % of the core is attributed to “monsoon”, 24 % to “inter-monsoon” and 8 % to “winter” layers.

3.2 Observed concentration profiles, deposition processes and seasonal fluxes

The ice core record is characterized by a strong seasonality that can be sub-divided into 3 distinct patterns corresponding to winter, monsoon and inter-monsoon. Table 1 summarizes concentrations of major ions, dust and rBC for the whole core and for

(PM₁₀-PM₁) and equivalent BC in the atmosphere. This is an additional confirmation of the presence of a local dust source, formed with coarse particles not sampled in the PM₁₀ inlet at NCO-P.

5 The snow is clearly enriched in NO₃⁻ and NH₄⁺ with respect to the atmosphere, if we take SO₄²⁻ as a reference. While NO₃⁻/SO₄²⁻ and NH₄⁺/SO₄²⁻ rarely exceed 1 (average = 0.42 and 0.50 respectively) in the atmosphere (Decesari et al., 2010), they are much higher in the ice samples (average = 1.32 and 2.89, respectively). This is certainly due, on one side, to negative artifact on the filters leading to loss of NO₃⁻ (and NH₄⁺ to a lesser extent) but also to efficient scavenging (wet and dry) of highly soluble NH₃ and HNO₃ gases to snow surfaces.

3.4 Dust and rBC deposition fluxes

Annual deposition fluxes can be compared at both sites, with a few assumptions for the overlapping period 2006–2010. Here we have limited the calculation to BC from atmospheric measurements using deposition velocities derived by Yasunari et al. (2010, 2013). For the pre-monsoon season 2006, Yasunari et al. (2010) have obtained a total BC deposition amount of 0.266 mgm⁻² at NCO-P. The time evolution of the surface snow concentration is simulated using the firn core data and the dating procedure described previously and illustrated in Fig. 4. This reconstruction is used to simulate the impact of surface snow impurities impact on glacier energy balance and melting (Sect. 4.3).

20 Deposition fluxes for some specific aerosol like rBC and insoluble dust were calculated from the ice core data. Annual mean deposition fluxes (from 1 October to 30 September) over the last 10 yr are 3.2 ± 1.2 mgm⁻²yr⁻¹ for rBC and 10.1 ± 2.5 gm⁻²yr⁻¹ for dust. This value for rBC is in the upper range of fluxes modeled by (Yasunari et al., 2013) for this region of the Himalaya, and one order of magnitude higher than calculated from NCO-P station pre-monsoon measurements (Yasunari et al., 2010). Furthermore, these fluxes were calculated between monsoon and

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inter-monsoon regimes. If we apply the seasonal fluxes splitting between “monsoon” and “inter-monsoon”, we can observe that rBC and dust presents different behaviors. Dust mean concentration is still higher during inter-monsoon season (66.9 mg L^{-1} , vs. 26.5 mg L^{-1} during monsoon) but the deposition flux during monsoon represents 72 % of the annual flux. On the opposite, the rBC deposition flux is concentrated during inter-monsoon, with 75 % of the annual deposition, and with a higher mean concentration contrast between season ($9.25 \text{ } \mu\text{g L}^{-1}$ during inter-monsoon, $1.06 \text{ } \mu\text{g L}^{-1}$ during monsoon).

As dust shows high deposition fluxes for all seasons, we used the available data to track source changes. Soluble species analyzed by ion chromatography as calcium, magnesium, manganese and strontium were used to calculate ratios of soluble ionic species vs. dust. The negative ionic balance is attributed to the carbonate load. These ratios and carbonate are used to identify some changes in the dust sources. The ionic balance is negative all over the record ($-6.7 \text{ } \mu\text{Eq L}^{-1}$ mean value), but with a weak higher value for winter snow ($-4.5 \text{ } \mu\text{Eq L}^{-1}$) and some spikes ranging down to -20 to $-40 \text{ } \mu\text{Eq L}^{-1}$. The calcium/dust ratio is quite stable over the record. The two dustiest layers located during inter-monsoon, at 4.1 and 8.9 m w.e. depth are atypical (Fig. 2), with highest carbonate concentration and high ratios of manganese and magnesium, which may characterize a different dust source that was not identified. A third source, with high lithium ratio associated with sea salt, and with or without carbonate is reported in the most recent part of the record (winter and monsoon 2010), and could represent inputs of aerosols coming from salt flat or saline lakes located in the Tibetan plateau.

As we can see in Fig. 2, rBC concentration peaks build up between each annual monsoon event, during the drier period. According to the observation at NCO-P station (Marinoni et al., 2010), the highest atmospheric BC concentration occurs during pre-monsoon season (February–May) and is transported both by valley breeze form Indo-Gangetic Plain (India, Nepal) and by longer range from Middle East and Europe sources air masses. The deposition occurrence between BC and other anthropogenic

activities tracers, like ammonium and fluoride, confirm the regional source associated to the “brown cloud”.

4 Dust and rBC impact on surface albedo and glacier melting

4.1 Albedo changes with BC and dust deposition

5 For this study, we used the ice core based aerosol fluxes reconstruction to investigate the impact of dust and rBC contents of the snow surface on its energy balance by changing the surface albedo. Our approach was to simulate the resulting albedo change produced by different impurity concentrations, dust and rBC, measured from the ice core. For this purpose, we used the radiative transfer model DISORT (Stamnes
10 et al., 1988) with several simplistic assumptions since we aim at providing order of magnitude of the impurity effects on the snow radiative balance.

The snow grains are considered as spherical and we used Mie theory to compute their single scattering properties. The refractive index of ice is taken from (Warren and Brandt, 2008). The refractive index of BC and its density is taken from (Flanner
15 et al., 2012) whereas the refractive indices and density of dust with different contents of hematite are taken from (Balkanski et al., 2007). DISORT model was used to simulated the spectral albedo of an external or internal mixture of impurities and snow as detailed in (Flanner et al., 2012). In the absence of any measurement, the snow grain size was
20 chosen as constant over the profile (optical radius (r_{opt}) of 0.16 mm or 0.65 mm) and the impurities were equally distributed into the snowpack (internal or external).

Dust and BC contents were computed from the seasonally dated core record. Dust is considered as being external to the snow grains given its size (mass equivalent diameter ranging from 6 μm when deposited with precipitation up to 10 μm when transported dry). BC is assumed as being external during the dry season (dry deposition) and internal during the monsoon (wet deposition). BC mean diameter is taken as 90 nm for
25 all seasons as shown by SP2 the measurements.

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The simulated spectral albedo using DISORT for the 1 June 2003 and the 1 June 2005 and for diffuse irradiance is illustrated for examples in Fig. 5. These two dates were chosen as they represent extreme values of dust and rBC content in the record at the end of the simulated post-monsoon season. For June 2003, the rBC content is 26.64 ppb and the dust content is 66.9 ppm whereas for June 2005, the rBC content is 47.9 ppb and the dust content is 13.13 ppm. The blue solid line corresponds to the albedo of pure fine snow ($r_{\text{opt}} = 0.16$ mm) while the black solid line is the albedo of fine snow containing both rBC and dust. The two red lines correspond to the spectral albedo of pure and contaminated snow for coarser grains ($r_{\text{opt}} = 0.65$ mm). For these simulations we choose a hematite content of 1.5 % for the dust refraction index which corresponds to the average content determined by (Balkanski et al., 2007). In the 350–500 nm range for fine grains, impurities caused a decrease of about 0.1 of the value of the albedo for June 2003 whereas only dust implies a decrease of 0.05 for June 2005 (around 0.1 for 2003) and only rBC implies a decrease of around 0.05 for the two years. This figure enhances the fact that the impurities have a non linear effect on the value of the albedo. Indeed rBC and dust combined together have less effect than the sum of the effects of only rBC and only dust. This can be easily explained by the fact that a mixture of impurities and snow has a much higher absorption coefficient (imaginary part of the refractive index) than pure snow. Thus adding one type of impurity to the snow diminishes the impact of the other impurity compared to the impact of the same amount of impurity on pure snow. In other words, the higher the absorption coefficient is, the less the impact of supplementary impurities will be. The effect of snow grain size is also noticeable on this figure while comparing the simulations. The same amount of impurities has more effect on coarse grain snow than on fine grain snow. Indeed, the albedo decrease due to impurity in the visible range is at least two times higher for coarse grains than for small grains since photons generally travel a greater distance in the snow pack for coarser grains (Warren and Wiscombe, 1980). Different content of hematite were tested, minimal (0.9 %) and maximal (2.7 %) according to (Balkanski et al., 2007), and illustrated in Fig. 5. It shows the impact on the albedo of the refractive

index of dust chosen in the simulation. In the following, we will only use the average value of 1.5 % for the hematite content.

4.2 Additional absorbed energy

To evaluate the amount of additional solar energy absorbed by the snowpack due to the presence of impurities, it was necessary to estimate a mean solar incident radiation incident and its characteristics. The solar irradiance was therefore computed using the radiative transfer model SBDART (Ricchiuzzi et al., 1998). SBDART is based on the same computation rules than DISORT. It considers an atmosphere made of different plane parallel layers. The characteristic of the atmospheric profiles (aerosols, water vapor, ozone ...) are set according to the site location. The solar zenith angle was selected according to the location and the hour of the day. The sky is assumed to be clear for all days in the clear sky simulation and cloudy for alldays in the cloudy simulation. One simulation was done at the beginning of each month during the 11 yr of the study. For each day, the diurnal cycle of the energy absorbed by the snowpack was reconstructed, and interpolated on a daily basis over the whole period.

Figure 6 presents for the year 2003 the daily mean surface forcing due to impurities. These daily values correspond to the difference between the solar energy absorbed by a pure snowpack and the energy absorbed by a contaminated snowpack. The simulation shows that the radiative forcing of impurity is maximal for the pre-monsoon season, up to 40 W m^{-2} for clear sky and coarse grains snow, as also observed by MODIS investigation (Gautam et al., 2013). For 2003, 75 % of the radiative forcing can be attributed to the presence of dust in the snowpack. The Fig. 6 shows also that the radiative forcing is twice greater for coarse snow ($r_{\text{opt}}=0.65 \text{ mm}$) than for fine thin snow ($r_{\text{opt}}=0.16 \text{ mm}$). These results are comparable with those provided by (Sterle et al., 2013) in the Sierra Nevada snow. Moreover cloudy sky induces half radiative forcing compared to clear sky simulation (maximum radiative forcing of 24 W m^{-2} for coarse snow grains).

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The whole period covered by the ice core, 2000–2010 was finally simulated (Fig. 7). It appears clearly that the high dust content is mainly responsible of most for the additional absorbed energy. The effect of rBC is indeed, except for some high rBC concentration peaks (e.g. June 2006), much lower than the dust effect. Over the whole period, the total impurities are responsible for a 6.3 % increase of the absorbed energy in case of fine snow, with 5.3 % related to dust and 1.3 % to BC compared to pure snow. In case of coarse snow consideration, the impact is more important reaching 12 % increase of absorbed energy.

4.3 Potential melting evaluation

The impurities deposited on the glacier surface increase the absorbed energy and have an impact on the glacier melting. Here we only estimate a potential excess melt rate attributed to impurities over years for each season (dry or monsoon) over the experiment period. The melt rate has been calculated using the mean value for clear sky and cloudy sky simulations and for the average of coarse snow and fine snow radiative forcing. In order to represent the melt caused by the impurities any given day the following assumptions have to be done, (a) all other terms of the surface energy budget corresponding to turbulent fluxes and long wave net radiation, were remained unchanged and, (b) the snowpack was at the fusion temperature. The resulting values are illustrated in Fig. 8. It is important at this point to noteworthy that almost no melt occurs in the highest part of the glacier and at the location of the ice core drilling, because the firn temperature always remains below the melting point except for some exceptional events, as confirmed by the core analyzes that show only sign of very limited refreezing layers. Nevertheless, this study can help evaluating the additional melt due to the presence of the same quantity of impurities in the lower part of the glacier where melt does occur in summer assuming that dust and BC contents are equally distributed over the glacier. This later assumption needs to be further investigated since e.g. in the ablation area of the glacier, the melt water might “wash” the snow surface by dragging the impurities, or impurities could concentrate on ablation surface (Sterle et al., 2013).

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If it is well demonstrated that the impact of dust is much higher than BC on the potential melt for Mera Glacier, it is interesting to observe that this impact does not change much over seasons, with moderately higher values during the dry season (with higher impurity content) than during the monsoon season (Fig. 8). This is mainly related to lower solar irradiance during part of the dry season (winter) compared to the summer irradiance. These results might thus vary while considering the effect of clouds during the monsoon only and according to localization on the glacier depending on the aspect and shading.

4.4 Comparison with glacier energy and mass balance

Dust and rBC particles snow content increases the glacier surface melting in some areas of the glacier. Their impact, as illustrated in Fig. 8, is estimated at an annual mean melting of $+1.8 \text{ kg m}^{-2} \text{ d}^{-1}$ for dust and rBC, with the fraction related to rBC only of $+0.78 \text{ kg m}^{-2} \text{ d}^{-1}$ and $+0.19 \text{ kg m}^{-2} \text{ d}^{-1}$ during inter-monsoon and monsoon respectively. These rates correspond to about $+662 \text{ kg m}^{-2}$ annual melting associated with dust and rBC, and $+214 \text{ kg m}^{-2}$ for rBC only.

In order to evaluate this impact on Mera glacier, we compare these values with mass-balance and energy-balance measurements taken on the glacier since 2007. The melting rate is calculated from the energy balance using as input data measurements from the free standing automatic weather station (AWS) located in the ablation zone at the elevation of 5360 m a.s.l. The specific mass balance is calculated annually from a stakes network and topographical data derived from Pleiade satellite image of 2012 (Wagnon et al., 2013). For the year 2009–2010 at 5400 m a.s.l., the mass balance measurement gives an annual point-mass balance of $-2280 \text{ kg m}^{-2} \text{ yr}^{-1}$ (corresponding to annual snow accumulation minus annual ablation), whereas energy balance approach results in an annual melting of $4300 \text{ kg m}^{-2} \text{ yr}^{-1}$ melting without taking into account sub-layers refreezing. Considering that in 2009–2010, the net accumulation measured at the highest part (6330 m a.s.l.) of Mera Glacier is 720 kg m^{-2} (Wagnon et al., 2013), annual accumulation at 5400 m a.s.l. is at least 720 kg m^{-2} making the

effective melting at 5400 m a.s.l. higher than 3000 kg m^{-2} ($= 2280 + 720$). The correct melting value is in turn comprised between these previous 3000 and 4300 kg m^{-2} . The melting related to impurities calculated in this study for the year 2009–2010 correspond to $+543 \text{ kg m}^{-2} \text{ yr}^{-1}$ annual melting associated to dust and rBC, and $+215 \text{ kg m}^{-2} \text{ yr}^{-1}$ for rBC only.

If we compare these values, it appears clearly that, with all our assumptions and considering a site where firn is close to 0°C , total impurities control a maximum of about 18% of the melting, when rBC alone accounts for a maximum of about 7% only at 5400 m a.s.l. Some intuitive consideration for a spatial distribution over the whole glacier would conclude to much less impact for the impurities on glacier melting. Mera Glacier specific mass balance measured between 2007 and 2012 results in $-100 \pm 400 \text{ kg m}^{-2} \text{ yr}^{-1}$ (Wagnon et al., 2013). The impact of rBC on the mass balance is estimated to be lower than its inter-annual variability, while the impact of dust and BC together can be of the same order of magnitude.

4.5 Limits of the simulations

The simulations presented in this section rely on a range of simplistic assumptions. These assumptions have been used to circumvent some difficulties: the exact profile of impurity content in the snowpack was not translated into surface concentration evolution, meteorological conditions (clouds, precipitations . . .) were not taken into account and surface conditions over the whole glacier area have not been taken into account either. Snow metamorphism was not taken into account either and in the simulation ice grains are supposed to be spherical (Libois et al., 2013). Nevertheless, the results presented here are deduced from the difference between the solar radiation budget for pure snow and contaminated snow, for which these assumptions are not critical. At glacier scale, our modeling approach based on a comparison between various impurity contents of surface snow and based on theoretical considerations regarding incident solar radiation, or surface conditions (melting or not) cannot provide accurate results

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on the effect of impurities on glacier melting but it is still useful to roughly quantify it, at least for the debris-free ablation zone.

Further investigations should be conducted to test these assumptions. Simulations using the detailed snow model Crocus (Vionnet et al., 2012) and reanalysis for past meteorological forcing can be done to investigate the effect of snow metamorphism and local meteorological conditions. Remote sensing data such as images from the imager MODIS can be used to infer the daily radiative forcing due to impurities (Dumont et al., 2012; Painter et al., 2012) and its spatial distribution over the Mera Glacier.

5 Conclusions

In this study, a 20 m firn core extracted from Mera Glacier in Nepal was used to reconstruct 10 yr of aerosol deposition fluxes at 6376 m on the southern flank of the Himalaya. Some species, like water stable isotopes, rBC, ammonium or fluoride deposition are consistent with the precipitation patterns related to the monsoon, with higher or spiking concentration during the dry season. rBC concentration were compared with BC atmospheric measurements from NCO-P station, both agrees for the timing but at different concentration levels in accord with the altitude difference between the sites. Other proxy, like dust revealed high and constant fluxes all over the year. These preliminary glaciochemical profiles quality and the in glacial temperature divulge that this site is a good candidate for environmental and climatic reconstruction based on ice core studies.

Dust and rBC aerosols present on the snow surface increase the potential snow melting by increasing the surface albedo. With this study using aerosol fluxes reconstructed from Mera Glacier firn core analyzes integrated by atmospheric measurements at the Nepal Climate Observatory – Pyramid, we estimate their impact on glacier melting.

The analyzes reveal that mass fluxes are a few orders of magnitude higher for dust ($10.2 \pm 2.5 \text{ gm}^{-2} \text{ yr}^{-1}$) than for rBC ($3.2 \pm 1.2 \text{ mgm}^{-2} \text{ yr}^{-1}$), and that their deposition is distributed evenly over the year for dust but centered around pre-monsoon for rBC

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Table 1. Data summary: EOF analyzes parameters and aerosol concentration and fluxes.

	All/Annual	Inter-monsoon	Monsoon	Winter	
Samples	276	67 (24 %)	188 (68 %)	21 (8 %)	
Ice core	19.8 m	4.81 m	12.62 m	2.37 m	
	12.47 m weq	3.19 m weq	8.09 m weq	1.19 m weq	
PC1 Variance	34 %	32 %	32 %	48 %	
PC2 Variance	15 % (49 %)	18 % (50 %)	16 % (47 %)	18 % (66 %)	
PC3 Variance	11 % (61 %)	14 % (65 %)	12 % (59 %)	12 % (79 %)	
$\delta^{18}\text{O}$ Mean composition	-17.21 ± 2.95	-13.92 ± 2.56	-17.75 ± 1.80	-21.98 ± 1.11	(‰)
NH_4^+ Mean concentration	1.69 ± 0.84	2.49 ± 0.96	1.55 ± 0.49	0.32 ± 0.15	($\mu\text{Eq L}^{-1}$)
F^- Mean concentration	0.070 ± 0.060	0.120 ± 0.094	0.056 ± 0.026	0.026 ± 0.015	($\mu\text{Eq L}^{-1}$)
Ca^{2+} Mean concentration	4.55 ± 2.80	5.81 ± 3.43	4.11 ± 2.34	4.36 ± 3.20	($\mu\text{Eq L}^{-1}$)
SO_4^{2-} Mean concentration	1.05 ± 1.30	1.53 ± 1.60	0.87 ± 1.12	1.08 ± 1.37	($\mu\text{Eq L}^{-1}$)
NO_3^- Mean concentration	1.08 ± 0.97	1.96 ± 1.30	0.80 ± 0.62	0.66 ± 0.37	($\mu\text{Eq L}^{-1}$)
rBC Max. concentration	47.90	47.90	8.22	1.20	($\mu\text{g L}^{-1}$)
Mean concentration	3.00	9.25	1.06	0.35	($\mu\text{g L}^{-1}$)
Deposition flux	3.2 ± 1.2	(75% \pm 12%)	(25% \pm 12%)		($\text{mg m}^{-2} \text{ yr}^{-1}$)
Dust Max. concentration	66.9	66.9	26.5	15.4	(mg L^{-1})
Mean concentration	10.1	11.1	10.1	6.5	(mg L^{-1})
Deposition flux	10.1 ± 2.5	(28% \pm 11%)	(72% \pm 11%)		($\text{g m}^{-2} \text{ yr}^{-1}$)

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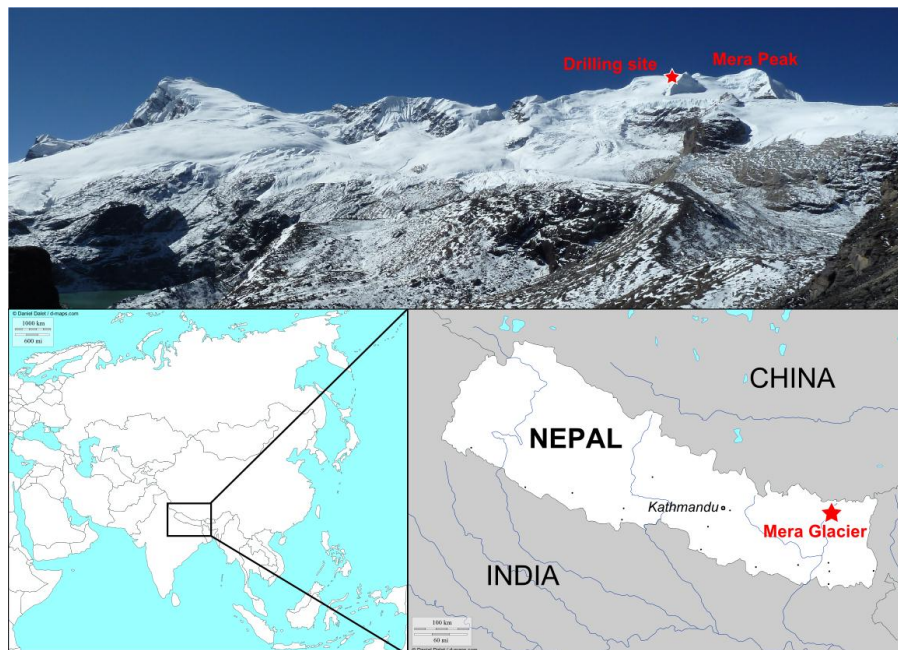


Fig. 1. Location of Mera Glacier in Nepal. The picture shows Mera Glacier view from the north and the drilling site position.

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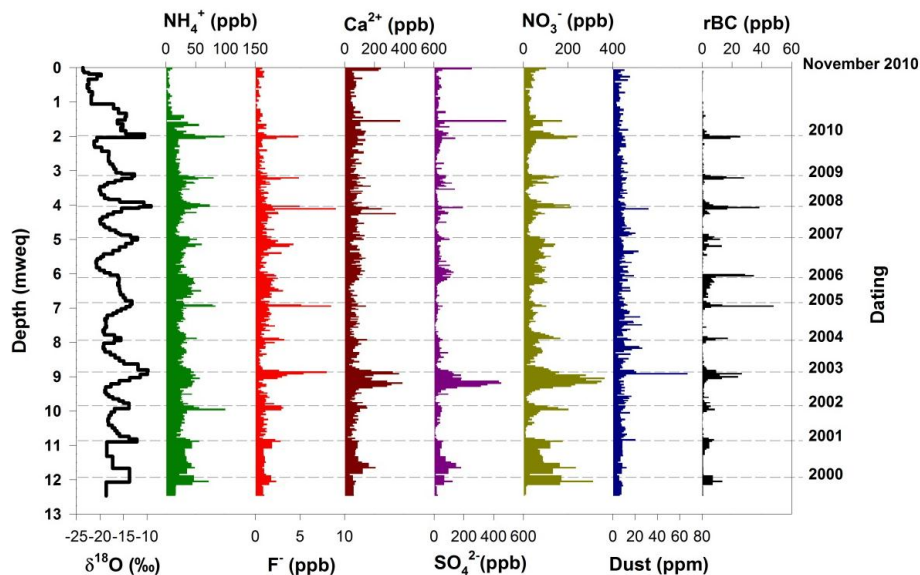


Fig. 2. Mera Glacier firn core profiles with depth for water stable isotope ($\delta^{18}\text{O}$), ammonium, fluoride, calcium, sulfate, nitrate, dust and refractory black carbon (rBC). The right vertical scale years corresponds to the annual dry season.

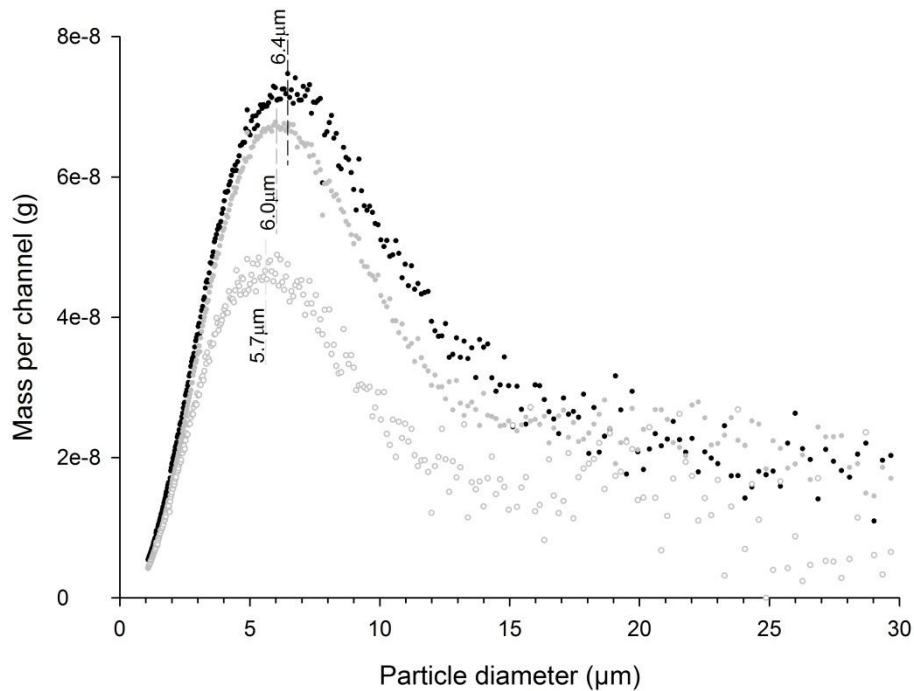


Fig. 3. Dust size distribution for the different season: inter-monsoon (black), monsoon (dark grey) and winter (white).

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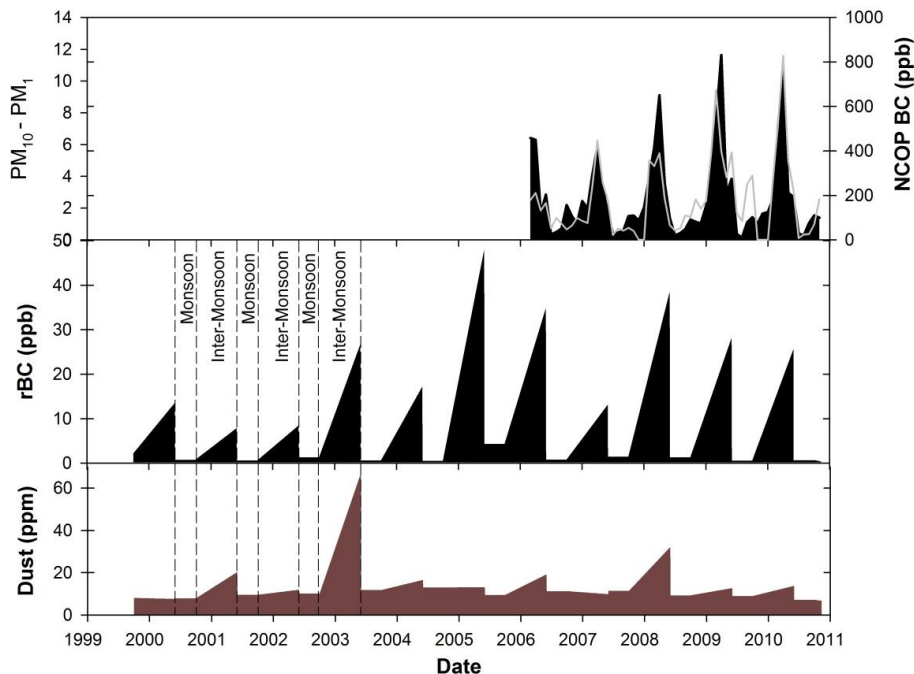


Fig. 4. Atmospheric BC (black) and dust (grey line) measured monthly at NCO-P station (top), and surface snow concentration evolution for rBC (black bar) and dust (brown bar) simulated from firn core data.

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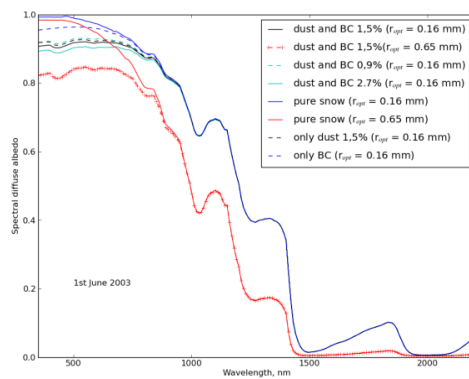
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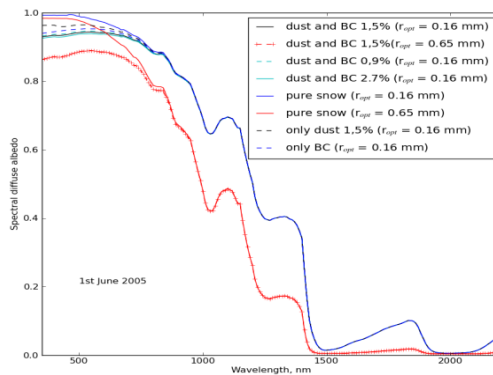
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a)



b)

Fig. 5. Simulated spectral albedo for the 1 June 2003 **(a)** and the 1 June 2005 **(b)**.

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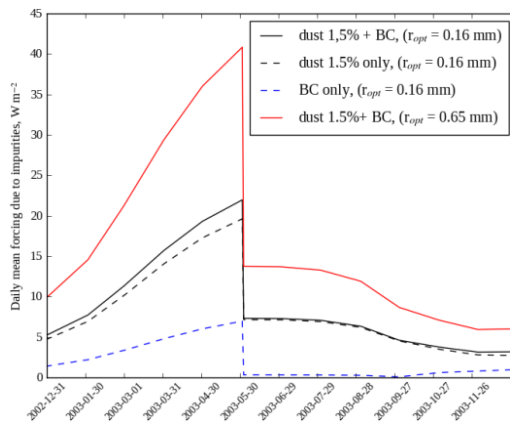
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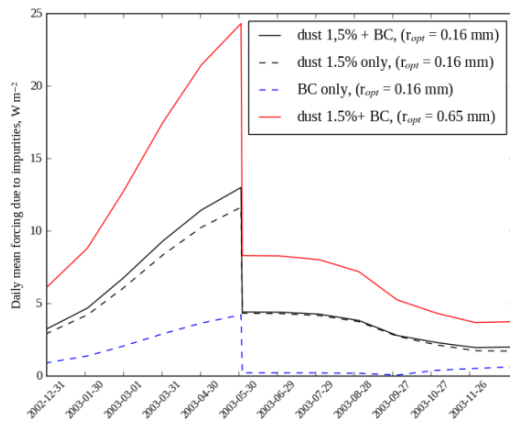
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a)



b)

Fig. 6. Daily mean forcing due to impurities **(a)** for clear sky **(b)** for cloudy sky for year 2003.

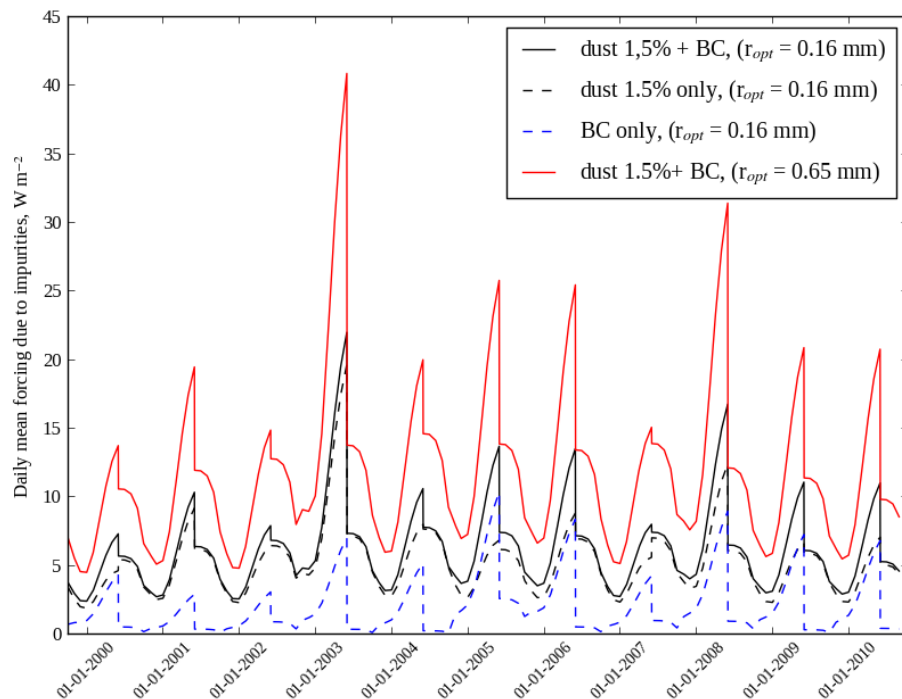


Fig. 7. Daily mean radiative forcing of impurity for clear sky simulation. The solid red and black lines are the additional energy due to the presence of dust and BC for coarse and fine snow. The black dotted line is for dust only and the blue dotted line is for BC only. The two solid lines are simulations for fine snow. For the results, the hematite content of dust is 1.5% as Balkanski et al. (2007), shows that the central content is around this value.

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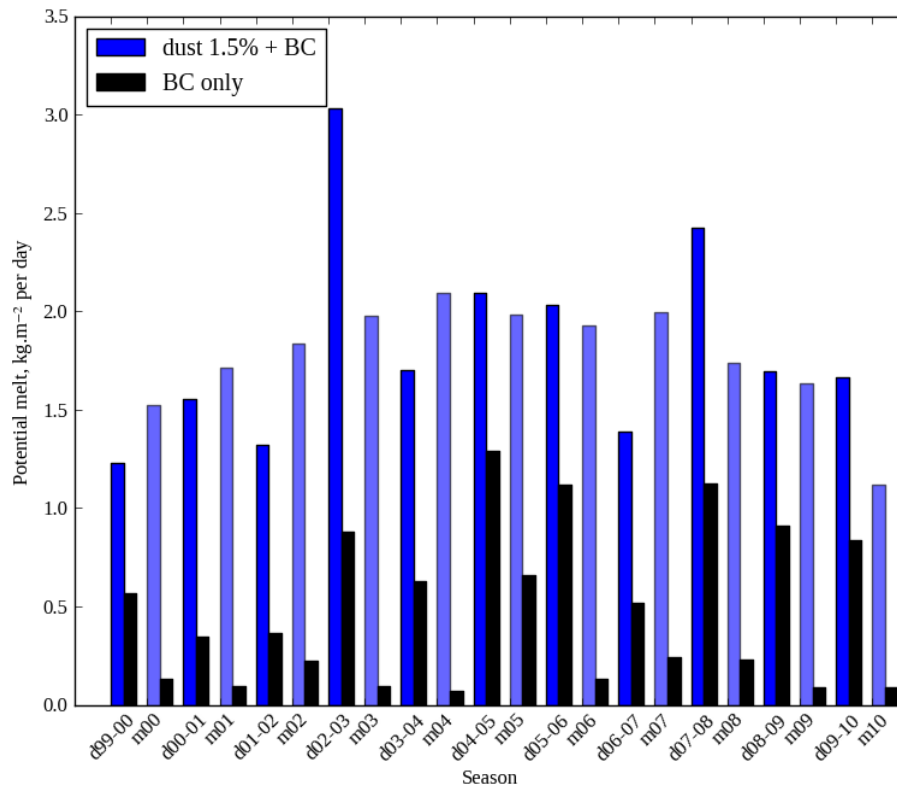


Fig. 8. Potential melt per day due to the presence of impurity in the snowpack for each dry (d) and monsoon (m) season of the study. The hematite content used is 1.5%. For each season, the first bar (blue) corresponds to the presence of both dust and BC, the second (black) to BC only.

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