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Surface and snowdrift sublimation at Princess Elisabeth station, East Antarctica

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Abstract

In the near-coastal regions of Antarctica, a significant fraction of the snow precipitating onto the surface is removed again through sublimation – either directly from the surface or from drifting snow particles. Meteorological observations from an Automatic Weather Station (AWS) near the Belgian research station Princess Elisabeth in Dronning Maud Land, East-Antarctica, are used to study surface and snowdrift sublimation and to assess their impacts on both the surface mass balance and the surface energy balance. Comparison to three other AWSs in Dronning Maud Land shows that sublimation has a significant influence on the surface mass balance at katabatic locations by removing 10–23 % of their total precipitation, but at the same time reveals anomalously low surface and snowdrift sublimation rates at Princess Elisabeth (18 mm w.e. yr⁻¹ compared to 42 mm w.e. yr⁻¹ at Svea Cross and 52 mm w.e. yr⁻¹ at Wasa/Aboa). This anomaly is attributed to local topography, which shields the station from strong katabatic influence, and therefore on the one hand allows for a strong surface inversion to persist throughout most of the year and on the other hand causes a lower probability of occurrence of intermediately strong winds. These wind speed classes turn out to contribute most to the total snowdrift sublimation mass flux, given their ability to lift a high number of particles while still allowing for considerable undersaturation.

1 Introduction

The surface mass balance (SMB) includes the input and/or removal of mass at the surface and hence constitutes an essential part of the total mass balance. Consequently, its study is crucial for understanding the current and future evolution of the Antarctic ice sheet and associated sea-level changes. On a snow- or ice-covered surface, the SMB is given by

$$\text{SMB} = \text{PR} + \text{ME} + \text{ER}_{\text{ds}} + \text{SU}_{\text{s}} + \text{SU}_{\text{ds}} \quad (1)$$

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with the different terms of the SMB being solid precipitation (PR), melt (ME), erosion (deposition) due to divergence (convergence) of horizontal snowdrift transport (ER_{ds}), surface sublimation (deposition) (SU_s) and sublimation of drifting snow particles within a column extending from the surface to the top of the drifting snow layer (SU_{ds}) (Van den Broeke et al., 2004a).

As surface sublimation depends on near-surface temperature, the process is marked by a strong seasonality (Takahashi et al., 1994; Bintanja and Van den Broeke, 1995; Reijmer and Oerlemans, 2002; Frezzotti et al., 2004). In addition, the process is characterised by a strong spatial variability: in the extremely cold interior of the Antarctic ice sheet, SU_s is virtually absent (Déry and Yau, 2002), whereas blue ice areas are associated with high sublimation rates (e.g., at Seal rock in the Sør Rondane mountains, Takahashi et al. (1992) recorded SU_s rates from -200 to -280 mm w.e. yr^{-1}).

Whenever the friction velocity (u_*) becomes sufficiently high for the drag force to overcome inter-particle bonding forces and gravity, snow particles are lifted from the surface to form “drifting snow” (Bintanja, 1998). Snowdrift is a very common phenomenon in Antarctica: over certain parts of the ice sheet, snowdrift events have been observed on approximately one out of four days (Mann, 1998), and in the extremely windy Cape Denison, Terre Adélie (Wendler et al., 1997), snowdrift is even observed most of the year (Budd et al., 1966; Schwerdtfeger, 1984). During snowdrift events, continuous sublimation of the floating snow particles takes place, as the ambient air is usually undersaturated with respect to ice (Schmidt, 1982). Continuous ventilation of the floating particle thereby significantly enhances the transfer of moisture from the particle to the surrounding air (Bintanja, 1998). In addition, water vapour exchange takes place on all sides of each particle. Thus, snowdrift sublimation (SU_{ds}) – whenever it occurs – is more effective in ablating mass per unit of time than SU_s (Van den Broeke et al., 2004a; Bintanja, 1998). From a model study, Déry and Yau (2002) concluded that, on the scale of the Antarctic ice sheet, SU_{ds} removed 15.32 mm w.e. yr^{-1} on average from 1979 to 1993. Although this average appears small compared to continental scale values for the net SMB (149 – 171 mm w.e. yr^{-1} ; Vaughan et al., 1999; Van Lipzig et al.,

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2002b; Van de Berg et al., 2006), it masks a tremendous spatial variation, with virtually non-existent SU_{ds} in the interior of Antarctica to one of the major contributions near the relatively warm and windy coast (Bintanja, 1998; Déry and Yau, 2002).

In addition to the decrease in upward turbulent momentum flux, SU_{ds} constitutes a source of water vapour and a sink of sensible heat in the air (Déry and Taylor, 1996; Bintanja, 2001b). In fact, sublimation of floating snow particles tends to saturate the entire surface layer, starting in the lowest levels and therewith inhibiting further sublimation (Bintanja, 2001a). From the large number of numerical models for snowdrift sublimation and transport that have been developed over the last two decades (Pomeroy et al., 1993; Mobbs and Dover, 1993; Uematsu, 1993; Déry and Yau, 1999; Gallée, 1998; Liston and Sturm, 1998; Mann, 1998; Essery et al., 1999; Bintanja, 2000c; Liston and Elder, 2006; Lenaerts et al., 2010), several succeed in reproducing this self-limiting nature of snowdrift sublimation (Xiao et al., 2000).

In contrast, observing the different components of the SMB in Antarctica poses huge challenges. While precipitation measurements are disturbed by snowdrift, and SMB monitoring using for instance stakes requires reliable snow density measurements, it is nearly impossible to measure SU_s , SU_{ds} and ER_{ds} directly. The Automatic Weather Station (AWS) program in Dronning Maud Land coordinated by the Institute for Marine and Atmospheric Research of Utrecht University (IMAU), aims to tackle these problems by collecting high-quality meteorological observations which allow the calculation of the SMB components. In February 2009, this program was extended by installing AWS 16 near the Belgian Antarctic station Princess Elisabeth (<http://ees.kuleuven.be/hydrant>). Data collected at this new station are valuable, not only because they fill a vast observational gap spanning over 1070 km between the coastal stations Novolazarevskaya (70°46' S; 11°50' E) and Syowa (69°00' S; 39°35' E), but even more as the station is located in a specific microtopographical setting, i.e., close to a mountain range.

The two main goals of this study are (i) to compare sublimation rates calculated from AWS 16 to other, existing time series in Dronning Maud Land and (ii) to understand the observed differences from a physical perspective. In order to achieve this,

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first surface sublimation and snowdrift sublimation were calculated for four automatic weather stations, followed by an assessment of their relative contribution to the SMB. Next, the magnitude of the latent heat flux (LHF) in relation to other SEB components is investigated through application of a SEB model developed for snow-covered terrain (Van den Broeke et al., 2005). Finally, a sensitivity analysis was conducted to find the meteorological variables controlling SU_s and SU_{ds} . Combined with the study of the near-surface meteorology, these findings allow to attribute the sublimation anomaly at the Princess Elisabeth.

2 Data and methods

2.1 AWS data

Princess Elisabeth is erected on Utsteinen ridge ($71^{\circ}57' S$; $23^{\circ}21' E$) at the foot of the Sør Rondane mountains in the escarpment zone (Fig. 1). The Utsteinen ridge – oriented in a north-south direction and with an altitude of 1420 m a.s.l. – is 700 m long, a few meters wide and protrudes 20 m above the surrounding snow surface (Gorodetskaya et al., 2010; Pattyn et al., 2009). AWS 16 is located 300 m east of Utsteinen ridge, in the valley down Gunnestadbreen glacier. The area is influenced both by cyclones that bring in heat and moisture from the warmer regions surrounding Antarctica and katabatic winds advecting cold air from the interior of the Antarctic Ice Sheet (Gorodetskaya et al., 2012).

AWS 5 is situated near the Antarctic stations Wasa and Aboa, on the coastal slopes of the ice sheet, just inland of the grounding line. AWS 6 at Svea Cross stands at the foot of the Heimefront Mountains and is strongly influenced by katabatic winds. Finally, AWS 9 is located a few kilometres east of the German Antarctic station Kohnen, on Amundsenisen on the East Antarctic plateau (Van den Broeke et al., 2004a). AWS characteristics are listed in Table 1.

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Each AWS records air temperature (T), pressure (p), wind speed (u) and direction (WD), relative humidity with respect to water (RH) and down- and upward short- and long-wave radiation fluxes (SW_{in} , SW_{out} , LW_{in} , LW_{out}) at a single level, initially about 4 m above the surface (Gorodetskaya et al., 2010; Van den Broeke et al., 2004a). In addition, the distance between the surface and the T and RH measurement level (H) is monitored by a sonic height ranger. Sensor specifications are shown in Table 2. Where possible, the analysis is conducted on hourly resolution; else 2-hourly values are used as model input (Table 1). Finally, observed RH values are converted to relative humidity with respect to ice (RH_i) similar to Anderson (1994).

While 13-yr time series are used for AWS 5 and 9, and 11 yr for AWS 6, currently only 2 yr of measurements are available for AWS 16. Obviously, this period is insufficient to determine a reliable climatology or seasonal cycle of the station's mass and energy fluxes. Moreover, data gaps in December 2009 and November 2010 might cause a bias towards winter values; especially for the summer phenomenon surface sublimation this might be an issue. However, although analysis of longer time series might be useful to reduce the uncertainties of estimates presented below, results indicate features which cannot be ascribed to data or model uncertainty.

2.2 Sublimation and SMB

The in-situ measurements are used to dissect the net accumulation signal at each station into its components following the method described by Van den Broeke et al. (2004a). The core of this SMB model builds on the findings from Monin-Obukhov similarity theory to calculate surface sublimation, while snowdrift sublimation is retrieved through averaging three different parameterisations for snowdrift sublimation. Details of the method are given below.

Surface sublimation SU_s ($\text{kg m}^{-2} \text{s}^{-1}$) can be obtained from the surface latent heat flux LHF (W m^{-2}) using the relationship $SU_s = \text{LHF} / L_{\text{sub}}$, where $L_{\text{sub}} = 2.84 \times 10^6 \text{ J kg}^{-1}$

is the constant latent heat release associated with the sublimation process (assumed constant). The surface latent heat flux LHF can be expressed as:

$$\text{LHF} = \rho_{\text{air}} L_{\text{sub}} u_* q_* \quad (2)$$

With u_* (ms^{-1}) and q_* (kg kg^{-1}) the turbulent scaling parameters of momentum and moisture and ρ_{air} (kg m^{-3}) the density of the near-surface air (the latter retrieved using the ideal gas law). Together with the turbulent scaling parameter of heat θ_* (K), u_* and q_* can be determined based on the “bulk” method (Denby and Greuell, 2000), i.e. through the flux-profile relationships:

$$u_* \cong \frac{\kappa[u(z_V) - u(z_{0,V})]}{\ln\left(\frac{z_V}{z_{0,V}}\right) - \Psi_m\left(\frac{z_V}{L_{\text{MO}}}\right)} \quad (3)$$

$$\theta_* \cong \frac{\kappa[\theta(z_T) - \theta(z_{0,T})]}{\ln\left(\frac{z_T}{z_{0,T}}\right) - \Psi_h\left(\frac{z_T}{L_{\text{MO}}}\right)} \quad (4)$$

$$q_* \cong \frac{\kappa[q(z_q) - q(z_{0,q})]}{\ln\left(\frac{z_q}{z_{0,q}}\right) - \Psi_h\left(\frac{z_q}{L_{\text{MO}}}\right)} \quad (5)$$

With $\kappa = 0.4$ the von Kármán constant, $u(z_V)$ (ms^{-1}) the wind speed measured by the AWS at a height z_V above the surface and $u(z_{0,V})$ the velocity at the surface. Mutatis mutandis, Eqs. (4) and (5) require values for the potential temperature θ (K) and the specific humidity q (kg kg^{-1}), each at a level z_T (z_q) above the surface and at the height of the temperature (moisture) roughness lengths $z_{0,T}$ ($z_{0,q}$). Ψ_m and Ψ_h are the momentum, respectively heat/moisture stability functions. RH_i and T are measured at the same level, so $z_T = z_q = H$. Wind speed is measured about 0.75 m higher ($z_V = H + 0.75 \text{ m}$). In our SMB calculations we adopted a value of 0.1 mm for the momentum surface roughness length $z_{0,V}$ (Bintanja, 2000d). Following the method

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developed by Andreas (1987), $z_{0,V}$ is related to $z_{0,T}$ and $z_{0,q}$ through two empirical expressions depending on the roughness Reynolds number $Re_* = u_* z_{0,V} / \nu_{\text{air}}$, with ν_{air} the kinematic viscosity of air ($\text{m}^2 \text{s}^{-1}$). By definition, $u(z_{0,V}) = 0$. Furthermore, T_s can be obtained from the Stephan-Boltzmann law: $T(z_{0,T}) = T_s = (LW_{\text{out}} / \varepsilon \sigma)^{1/4}$, with $\sigma = 5.67 \times 10^{-8} \text{W m}^{-2} \text{K}^{-4}$ the Stephan-Boltzmann constant and the longwave emissivity $\varepsilon = 1$. Subsequently, θ_s can be derived. The value for $q(z_{0,q})$ is calculated from the saturation vapour pressure with respect to ice $e_{s,i}$ at the surface (Curry and Webster, 1999), assuming the snow surface to be saturated with respect to ice. Note that, there- with, the near-surface RH_i-gradient is assumed negative or zero at all times. Finally, the momentum and heat/moisture stability functions Ψ_m and Ψ_h are calculated using the expressions proposed by Holtslag and de Bruin (1988) for stable and very stable conditions, while equations put forward by Dyer (1974) are applied for the unstable regime. The ratio z/L_{MO} is used to discern between the three regimes: if $z/L_{\text{MO}} > 0.5$ conditions are very stable, $0.5 > z/L_{\text{MO}} > 0$ stands for stable conditions and $z/L_{\text{MO}} < 0$ represents unstable conditions. In each case, L_{MO} (m) stands for the Obukhov length scale, which can be calculated following (Van den Broeke et al., 2004a):

$$L_{\text{MO}} = \frac{u_*^2}{\kappa \frac{g}{\theta} (\theta_* + 0.62 \theta q_*)} \quad (6)$$

Thus the problem is implicit, and requires iterative solving. Convergence (< 0.001% difference) is usually reached after 10–15 iterations.

Although the threshold friction velocity for snow transport depends on numerous environmental conditions such as temperature (Li and Pomeroy, 1997) and snow-pack properties such as snow particle size and density (Gallée et al., 2001), snow transport was set to occur only when $u_* > 0.3 \text{m s}^{-1}$. In that case, SU_s vanishes and column-integrated snowdrift sublimation SU_{ds} is estimated as the average of three parameterisations: Bintanja and Reijmer (2001), Déry and Yau (2001) and Bintanja (1998) (hereafter referred to as BR01, DY01 and B98, respectively).

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The expressions put forward by BR01 account for the increase in $z_{0,V}$, $z_{0,T}$ and $z_{0,q}$ due to enhanced momentum dissipation, lower temperature and increased saturation levels, respectively, during snowdrift (Schmidt, 1986; Wamser and Lykossov, 1995; Mann et al., 2000; Bintanja, 2000b). BR01 suggest three empirical expressions depending on u_* to account for the respective roughness length dependency upon wind speed, subsequently yielding SU_{ds} in a way similar as described above. As the parameterisation is found to be no longer valid at wind speeds with corresponding friction velocities above 1 ms^{-1} , a cap of 0.1 m on the momentum, heat and moisture roughness lengths was introduced.

On the basis of multiple integrations with the numerical blowing snow model PIEKTUK-D, DY01 derived another explicit parameterisation for snowdrift sublimation requiring the input of u at 10 m above the surface, and T and RH_i at 2 m. For the latter two, observed values were used without applying any correction, whereas the 10 m wind speed was retrieved from the flux profile relationships assuming a roughness length of 1 mm and neglecting stability effects.

Finally, B98 derived an expression for vertically integrated snowdrift sublimation from a set of simulations with the numerical blowing snow model SNOWSTORM forced with 3 m u and T . Since the B98 parameterisation was developed for application to AWS data for which no humidity measurements were available, RH_i was assumed constant and set to 70%. However, judging from the RH_i values observed at AWS 5, 6, 9 and 16, this would lead to an overestimation of column snowdrift sublimation. Therefore a linear rescaling of the form

$$SU_{ds} = SU_{ds} \left(\frac{1 - RH_i}{1 - 0.7} \right) \quad (7)$$

was applied in order to account for the variations in humidity.

Since the surface mass balance can be obtained by multiplying snow height changes with the snow density ρ (Table 1), the “residual term” containing PR and ER_{ds} can be obtained from Eq. (1). Note that first, the raw snow height data had to be corrected

for false reflections and mast extensions, the former using a histogram approach, the latter by imposing an upper limit to the accumulation (Table 1).

2.3 SEB model

Considering a skin layer, the surface energy balance (SEB) of a snow covered surface can be written as (Van den Broeke et al., 2005):

$$M = SW_{in} + SW_{out} + LW_{in} + LW_{out} + SHF + LHF + G = R_n + SHF + LHF + G \quad (8)$$

with M the amount of melt energy, SW_{in} and SW_{out} the incoming and outgoing short-wave radiative fluxes, LW_{in} and LW_{out} the incoming and outgoing longwave radiative fluxes, R_n the net radiation, SHF and LHF the sensible, respectively latent heat flux and finally G the subsurface conductive heat flux. All terms are in Wm^{-2} and defined positive when directed towards the surface.

With the goal of calculating the different components of the SEB, we use the SEB model described in detail by Van den Broeke et al. (2005) (for applications see Kuipers Munneke et al., 2009; van As et al., 2005; Giesen et al., 2008; Van den Broeke et al., 2011). Here, we only provide a very short description of the model's main steps and assumptions. Essentially, AWS measurements (SW_{in} , SW_{out} , LW_{in} , T , RH, u , ρ and H) serve as input to the model that iterates for a surface temperature T_s for which closure of all energy terms is achieved. An important advantage of this approach is that the observed T_s can be used to evaluate the model. Turbulent fluxes are determined analogous to the SMB model, whereas the subsurface conductive heat flux G is obtained by upwards extrapolating the one-dimensional heat-transfer equation at 2 and 6 cm depth. Melting energy is determined from the excess energy whenever modelled T_s reaches the melting point. The initial T_s was chosen to minimise spin-up effects (see also Van Lipzig et al., 2002a). The module calculating the subsurface SW radiation penetration (Brandt and Warren, 1993) is not used in our simulations, since it has been shown that its impact upon T_s both in Greenland and Antarctica is virtually inexistent (Kuipers Munneke et al., 2009; Brandt and Warren, 1993).

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3 Results

3.1 Model testing and evaluation

Since the vertical RH_i -gradient is assumed negative or zero (Sect. 2.2), one might expect similar gradients for q assuming an isothermal stratification. However, even given this restriction, an important T -inversion might lead to a positive q -gradient or “ q -inversion”, as can be inferred from the Clausius-Clapeyron relationship. From Eq. (5), it can then be inferred that a q -inversion is always associated with a downward turbulent moisture flux. At low wind speeds, this leads to the prediction of surface deposition. If this happens during snowdrift however, BR01 will predict the deposition of mass rather than sublimation. But it is physically impossible that the suspended particle accrete mass as long as the ambient air is not supersaturated. Moreover, DY01 and B98 do predict snowdrift sublimation rather than deposition during these events, as they do not depend upon the q -gradient but on a single-level RH_i -value. Whereas at AWS 5, 6 and 9, this situation is very rare, the micro-topography at Utsteinen ridge strongly favours the build-up of a q -inversion (Sect. 3.4.), forcing us to reject BR01 at AWS 16.

The SMB calculations are tested by applying the method to the first four years of available data for AWS 5, 6 and 9. Comparison with Van den Broeke et al. (2004a) shows that the model is successful in reproducing the SMB signal and the different components SU_s and SU_{ds} , as differences are small and non-systematic. Hence it is appropriate to apply the SMB routine.

With the actual T_s directly derived from the observed LW_{out} , the comparison of the modelled and observed T_s presents a powerful tool to evaluate the SEB model (Fig. 2). The good agreement between both quantities indicates that the model is doing well, with a mean difference ranging from 0.16°C (AWS 16) to 1.13°C (AWS 5) and a RMSE from 0.76°C (AWS 6) to 1.72°C (AWS 16).

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3.2 Surface and snowdrift sublimation

Table 3 depicts the SMB characteristics for the respective measurement periods. During the 2-yr measurement period, SU_s removed 14 mm w.e. at AWS 16. The signal shows a marked seasonality (Fig. 3a) with most of the sublimation taking place during summer. Results show a continuous snow deposition during the winter of 2009. SU_s attains its highest values in katabatic wind zone, whereas SU_s is nearly absent at AWS 9 (Table 3). Surprisingly, surface and snowdrift sublimation rates at AWS 5 and 6 are two to three times larger than at AWS 16. In Sect. 3.4, we investigate reasons for this large difference.

At AWS 16, a marked difference between the first and second measurement year (hereafter referred to as 2009, respectively 2010) can be observed: in 2010, SU_s removed 11 mm w.e., compared to only 3 mm w.e. in 2009. The second period coincides with lower saturation levels (61 % in 2009 compared to 52 % in 2010), so one could state that the dryer air opens a larger potential to sublimation during 2010. However, although the Spearman rank correlation ($\rho = -0.24$) is significant at the 1 % level, RH_i only explains 6 % of the total variance in SU_s . Moreover, since surface deposition occurs at all values of RH_i , this variable does not even allow for discrimination between sublimation and deposition. Instead, both the sign and magnitude of SU_s predominantly depend upon the near-surface specific humidity gradient, represented by q_* (Eq. 5), as demonstrated by a Spearman rank correlation of 0.88, significant at the 1 % level. For the near-surface q decreasing with altitude surface sublimation takes place, while for a q -inversion surface deposition is to be expected. In Dronning Maud Land, q -inversions are formed when intrusions of warm, moist air encounter the cold Antarctic snow surface (Van den Broeke et al., 2004a).

The annual SU_{ds} rate at AWS 16 is similar to SU_s (Table 3), and removed 20 mm w.e. snow during the entire measurement period. Given the important uncertainties related to SU_{ds} predictions, annual rates predicted by B98 and DY01 agree surprisingly well (Fig. 3c). A clear mark of seasonality is absent: we rather observe that an important

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fraction of the total SU_{ds} mass flux is removed during circa 5 major events, coinciding with a decaying or developing cyclone (Gorodetskaya et al., 2012). Also here, SU_{ds} is two to three times larger at AWS 5 and 6 and virtually inexistent at AWS 9. Even though, at AWS 9, snowdrift occurs about as often as at the other stations, air temperatures there are too low to generate sublimation from the suspended particles.

The difference in sublimation rates between AWS 16 and the other katabatic stations remains significant, even when considering the short measurement period and the presence of data gaps. In fact, calculation of the sublimation mass fluxes for the 2009–2010 period at AWS 5 and 9 shows no strong change compared to their entire measurement period, suggesting that both years were not exceptional in terms of sublimation (Table 4). Furthermore, estimating the mass loss during both data gaps at AWS 16 (by interpolating the mean SU_s and SU_{ds} rates in the month before and after each data gap) indicates that SU_s at AWS 16 might be 3–4 mm w.e. yr⁻¹ stronger, while SU_{ds} would increase with 1–2 mm w.e. yr⁻¹. Again this effect is important, but not high enough to close the gap with the other katabatic stations.

3.3 Contribution to SMB

The surface mass balance at AWS 16 shows a net accumulation of 299 mm w.e. for the 2-yr measurement period, corresponding to 161 mm w.e. yr⁻¹ (Fig. 4a, Table 3). Note that even though this value lies within the net SMB range of 149–171 mm w.e. yr⁻¹ computed for various time periods and averaged over the entire Antarctic ice sheet (Vaughan et al., 1999; Van Lipzig et al., 2002b; Van de Berg et al., 2006), it is lower than one would expect given its location relatively close to the coast. Of this increase of the snow height by about 0.89 m, more than 80 % (0.75 m) was already achieved after the first year.

While the “residual processes” term obviously is the dominant component of the SMB (Table 3), estimation of the relative weight of PR and ER_{ds} within this term is not straightforward. A precise estimate of the wind contribution of ER_{ds} at this particular

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5 site requires fluid dynamic modelling including realistic topography and is beyond the scope of this study. Nonetheless, an attempt can be made to quantitatively constrain the magnitude of ER_{ds} . Application of the parameterisation for snowdrift transport by Mann et al. (2000), see also Van Lipzig et al. (2004) to AWS 16 yields an estimate of 192 mm w.e. snow transported over the station during the measurement period (104 mm w.e. yr^{-1}). This transport term provides an upper bound for ER_{ds} , being the snow transport divergence, which cannot be derived from single-point observations. There are good reasons to assume that the cumulative ER_{ds} is much lower: (i) suspended particles are known to be transported over large distances in Antarctica (Déry et al., 1998), (ii) most probably, both erosion and deposition occur at the station during the course of time, therewith partially cancelling out each other, and (iii) a modelling study for Antarctica by Déry and Yau (2002) found much lower values for the cumulative mass flux (0.05 mm w.e. yr^{-1} for the latitudinal band 70° – 80° S). Hence, although ER_{ds} can be significant on short time scales, setting “residual processes” equal to PR is most likely a valid assumption when considering time scales of one or more years.

10 If this assumption is applied, it is found that PR is responsible for an accumulation of 332 mm w.e. at AWS 16 during the measurement period. With AWS 6 and 16 both situated at the foot of the steep transition towards the Antarctic Plateau (“the escarpment zone”), orographic forcing is a likely source for PR (Fig. 1). In agreement with observations in more inland stations such as South Pole (Braaten, 2000), most of the accumulation at all AWSs occurs during sporadic precipitation events (Reijmer and Van den Broeke, 2003; Gorodetskaya et al., 2012).

15 Considering the entire measurement period, SU_s and SU_{ds} together removed 10 % of all PR at AWS 16. The fractional removal increased from 4 % in 2009 to 31 % in 2010, indicating both limited accumulation and enhanced (summer) sublimation during the second year. At AWS 5, 6 and 9, sublimation removed up to 23 %, 14 % and 6 %, respectively, of total precipitation during the measurement periods.

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3.4 Contribution to SEB

Both SW_{in} and SW_{out} show a clear seasonal cycle (Fig. 5a). From around mid-May till mid-August, the sun does not rise anymore above Princess Elisabeth and both SW components drop to zero. Atmospheric scattering and cloud amount determine the final fraction reaching the surface, hence the small-scale variations in SW_{in} . The high correlation between SW_{in} and SW_{out} (Spearman correlation coefficient $\rho = 0.9975$, significant at the 1 % level) underlines the high and relatively constant albedo α of the snow surface ($0.8 < \alpha < 0.95$). As for the LW radiation components, a seasonal cycle is much less pronounced, and small scale fluctuations are stronger for LW_{in} compared to LW_{out} (Fig. 5a) given the stronger variations in the atmospheric emissivity caused by cloud cover variability. Overall, AWS 16 is characterised by a radiation deficit during nearly the entire measurement period, except for the high summer.

At high latitudes and especially during winter, this radiative cooling of the surface may last the entire day and tends to render the near-surface air warmer than the underlying surface. As a response, the SHF transfers heat from near-surface air towards the surface (Fig. 5b) (Van Lipzig, 2007). It has been noted that this process causes increasing densities in near-surface air, and therewith leads to the formation of the katabatic wind system in the presence of a surface slope (Reijmer and Oerlemans, 2002). The LHF, representing the total amount of energy available for sublimation, is clearly less important in the SEB, except for high summer, when a clearly negative LHF consumes the R_n excess. The impact of G on the energy budget is rather small, while M is even completely absent at the station during the measurement period, consistent with observed T_s (Figs. 2a and 5b).

In Fig. 6a–d, monthly mean values of the different SEB components are depicted for the four AWSs. Compared to AWS 16, radiative losses are much smaller at AWS 9 but similar at AWS 5 and 6, with a maximum loss during wintertime when SW_{in} is absent and LW_{in} equilibrates towards the cloud base temperature (Van den Broeke et al., 2004b). Consequently, annual average positive SHF are larger at AWS 5, 6 and

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16 and the seasonal cycle largely mirrors seasonal variations in R_n (Table 3, Fig. 6b). Again, compared to this balance, the LHF calculated by the SEB model are found to be much weaker except during high summer (Table 3, Fig. 6c).

3.5 Attributing the AWS 16 sublimation anomaly

5 In order to account for the remarkably low sublimation rates at AWS 16 compared to other katabatic AWSs 5 and 6, one first needs to identify the main controlling variables for SU_s and SU_{ds} . This was achieved through a number of sensitivity experiments, where the values for T , RH_i , u and the magnitude of the near-surface temperature inversion (ΔT) were varied pairwise between bounds realistic for katabatic stations, 10 while mean values for AWS 16 during snowdrift-free, respectively snowdrift conditions were taken for the parameters remaining constant. As these four variables together constrain q and Δq , and since the dependency of SU_s upon Δq is directly evident from Eqs. (2) and (5), the latter variables need not to be tested. Figures 7–8 depict the most relevant of these experiments, together with the location of each station based 15 on the mean meteorological conditions during snowdrift free, respectively snowdrift conditions.

From Fig. 7a, it can be inferred that SU_s predominantly depends upon ΔT : a strong inversion dampens thermal turbulence and therewith SU_s . Also RH_i is an important determinant. For instance, taking the AWS 16's mean wind velocity and temperature 20 inversion, it is found that for RH_i values below $\sim 50\%$, sublimation takes place, while deposition occurs at higher values (Fig. 7b). T and u are equally important, but only represent second order effects by allowing for a higher moisture content and inducing a more efficient mixing of the moisture input in the surface layer (Fig. 7c).

Positioning of the different AWSs relative to each other on the basis of average 25 near-surface meteorological conditions during snowdrift-free conditions allows us to attribute the negative SU_s anomaly at AWS 16 (Fig. 7). Indeed the average ΔT is much stronger at AWS 16 than at the other stations (8.2 K at AWS 16 versus 2.4 K, 2.7 K and 1.3 K at, respectively AWS 5, 6 and 9). Consequently, as strong surface winds tend to

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destroy the near-surface T -inversion by enhancing vertical mixing (Ohata, 1985; Jons-son, 1995; Van den Broeke et al., 1999; Reijmer and Oerlemans, 2002), the low SU_s at AWS 16 can be related to reduced katabatic influence at this station.

As for snowdrift sublimation, an extensive sensitivity analysis reveals that SU_{ds} is pre- dominantly limited by RH_i (Fig. 8d–i), in agreement with findings from King et al. (1996) and Déry et al. (1998), but only when the ambient moisture content is high. Below a RH of about 70–80%, T and u take over as the main controlling variables for SU_{ds} (Fig. 8a–c). Higher T allow for larger q -gradients to build across the drifting snow layer, while the impact of wind speed peaks at values just above the threshold for snow trans- port, when enhanced saltation drift density (Pomeroy and Gray, 1990) and stronger ventilation (Bintanja, 1998) are most effective. Hence, suppressed seasonality in SU_{ds} (Fig. 3c,d) is probably due to higher u balancing for lower T during winter. Although minor differences exist, all three parameterisations show consistent behaviour.

Contrary to SU_s , visualisation of the mean atmospheric conditions during snowdrift (Fig. 8) does not lead to an understanding why SU_{ds} rates at AWS 16 are two, re- spectively three times lower compared to AWS 5 and 6. Further investigation of the near-surface meteorology, described below, can however explain this difference.

Due to the specific geographic location of the station, wind fields observed at AWS 16 show a marked behaviour. Predominantly two wind speed regimes reign at AWS 16 (Fig. 9a). Most often, a S-SSE katabatic wind is blowing at low speed ($0\text{--}10\text{ ms}^{-1}$). Strong katabatics are absent at the station, as they are blocked by the surrounding Sør Rondane mountains (Fig. 1b). On the other hand, during approximately 10–20% of the observation period, a strong E synoptic wind prevails at AWS 16. The highest wind speeds ($20\text{--}30\text{ ms}^{-1}$) are all situated within this wind regime, and low wind speeds ($< 5\text{ ms}^{-1}$) blowing from this direction are very rare. Both wind regimes also coincide with a clear distinction in T : while katabatics are associated with cold temperatures ($225\text{--}255\text{ K}$) the synoptic winds advect warmer air masses ($250\text{--}267\text{ K}$) (Fig. 10a). At high u , the difference is even more pronounced, as the data cloud forks into a synoptic ($\sim 260\text{ K}$) and a katabatic (247 K) peak. Clearly the katabatic flow generally remains

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colder than the E maritime winds, even after adiabatic heating upon descend from the Antarctic Plateau (Fig. 1b). Finally, as the katabatic downflow also induces a drying of the near-surface air ($10\% < RH_i < 60\%$), the contrast in RH_i with the nearly saturated maritime air masses ($80\% < RH_i < 100\%$) is evident from the AWS 16 data (Fig. 10b).

5 Comparison to the wind regimes at AWS 5 and 6 allows to attribute the AWS 16 SU_{ds} anomaly. At AWS 5, the katabatic winds (ENE) dominate, with occasional high wind speeds coming from the synoptic regime (NE; Fig. 9b). The wind regime at AWS 6 appears comparable to AWS 16, taking into account that wind directions have backed by about 45° (katabatic: SE; synoptic: NE; Fig. 9c), probably due to local orography and a stronger Coriolis effect (Bintanja, 2000a). However, comparative histograms of the 2-hourly mean u , accompanied by a generalised extreme value (GEV) distribution fit to each dataset (Fig. 11a), reveal a marked difference between the different stations. Clearly, for u ranging between 5 ms^{-1} and 15 ms^{-1} , the probability of occurrence is much larger at AWS 5 and 6 compared to AWS 16. Subsequent binning of SU_{ds} on the basis of the wind speed data (1 ms^{-1} bin width, Fig. 11b), shows that the wind speed classes with the largest contribution to the total SU_{ds} mass flux are $8\text{--}15 \text{ ms}^{-1}$. While wind velocities below 8 ms^{-1} usually do not succeed in lifting particles from the surface, also strong winds (exceeding 15 ms^{-1}) contribute only little to the total SU_{ds} , not only because they occur less often, but especially as they are usually associated with storm depressions (synoptic regime). In fact, given the maritime origin of these air masses and the precipitation which is regularly observed during such events (Gorodetskaya et al., 2012), associated RH_i values will be close to saturation and will therefore inhibit significant sublimation rates (Figs. 8d–i and 10). Binning the measured RH_i values on the basis of u demonstrates that this saturation effect is present at each station (Fig. 11c).

25 On the contrary, moderate wind velocities ($8 \text{ ms}^{-1} < u < 15 \text{ ms}^{-1}$) are strong enough to lift particles but still allow for significant ambient undersaturation. This wind regime, generally created by medium strength katabatic winds which dry upon descend from the plateau, provides the ideal conditions for strong SU_{ds} rates at all stations

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(Fig. 11a–c). The anomalously low SU_{ds} rates at AWS 16 can now be understood: in essence, moderate wind velocities occur less often at AWS 16 compared to AWS 5 and 6, this due to orographic shielding of the former station (Fig. 11a). In fact, this shielding reduced the probability of snowdrift occurrence to 12 % during the measurement period at AWS 16, whereas snowdrift occurred during 29 %, respectively 23 % of the time at AWS 5 and 6.

4 Conclusions

The goal of this study is to compare sublimation rates calculated from a newly installed automatic weather station (AWS) to existing time series in Dronning Maud Land and to explain the observed differences from a physical perspective. Remarkable differences were found between AWS 16 and two other katabatic stations in Dronning Maud Land. From February 2009 till February 2011, annual surface (SU_s) and snowdrift sublimation (SU_{ds}) rates at AWS 16 were found to amount up to $-7 \text{ mm w.e. yr}^{-1}$, respectively $-11 \text{ mm w.e. yr}^{-1}$. Both processes together have a significant influence on the surface mass balance at this station, by removing 10 % of all solid precipitation (PR) during the measurement period. However, sublimation rates at AWS 16 are 2 to 3 times lower compared to AWS 5 and 6, where sublimation annually removes 22 %, respectively 14 % of all PR.

In order to account for this sublimation anomaly at AWS 16, a process study was undertaken in order to determine the controlling variables for surface and snowdrift sublimation. The specific humidity gradient, i.e. the combined effect of ambient RH_i and ΔT , is predominant for both the sign and magnitude of SU_s in Antarctica, whereas T and u only represent second order effects. Subsequently, it was found that the strong surface inversion, persisting throughout most of the year at AWS 16, strongly dampens SU_s compared to AWS 5 and 6. In contrast, the process study for sublimation during snowdrift events showed that SU_{ds} is primarily limited by RH_i but only if the near-surface air is close to saturation. Below a RH_i of $\sim 70\text{--}80\%$, T and u take over as

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the main controlling variables for SU_{ds} , with higher temperatures and stronger winds both allowing for enhanced SU_{ds} rates. Surprisingly, application of these findings to the mean T , RH_i and u conditions during snowdrift suggests comparable sublimation rates at all three AWSs. Hence, further investigation of the near-surface meteorology at AWS 16 was undertaken. This showed that not the mean conditions during snowdrift, but rather the markedly lower probability of occurrence of moderate, katabatic winds ($8\text{ms}^{-1} < u < 15\text{ms}^{-1}$) is responsible for the lower SU_{ds} rates at AWS 16. Precisely these wind speed classes show to contribute most to the total SU_{ds} amounts, given their ability to lift particles while at the same time allowing for significant ambient under-saturation. To conclude, the sublimation anomaly at AWS 16 is the result of the twofold effect of the local orography, which protects the station from medium-strength katabatic winds and therewith (i) allows for a strong, dampening surface inversion to persist throughout most of the year, and (ii) reduces the occurrence of snowdrift by 50–70 % compared to nearby katabatic stations.

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Table 1. AWS characteristics. The snow density at AWS 16 was retrieved from the first 85 cm of an ice core drilled in February 2009 near the station. Start and end of observation mark the first, respectively last entry used in this study; all stations except AWS 6 are still operational. Also, note that AWS 16 has no lower bound for the mass balance since no mast extension has been conducted yet.

	AWS 16	AWS 5	AWS 6	AWS 9
Location	Princess Elisabeth	Wasa/Aboa	Svea Cross	Kohnen
Latitude	71°57' S	73°06' 19" S	74°28'53" S	75°00'41" S
Longitude	23°21' E	13°09'53" W	11°31'06" W	00°00'44" E
Altitude (m a.s.l.)	1420	363	1160	2892
Start of observation	3 Feb 2009	3 Feb 1998	3 Feb 1998	3 Feb 1998
Start of hourly observation	12 Jan 2010	2 Feb 2003	9 Jan 2004	13 Jan 2001
End of observation	2 Feb 2011	2 Feb 2011	2 Feb 2009	2 Feb 2011
Snow density (kg m ⁻³)	336	383*	396*	307*
Mass balance cap (mm w.e.)	/	-150	-100	-200

* values from Van den Broeke et al. (2004a) for 1998–2001.

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Table 2. AWS sensor specifications (EADT: estimated accuracy of daily totals) (<http://ees.kuleuven.be/hydrant/>). All four AWSs used in this study are similar in design.

Sensor	Type	Range	Accuracy
Air pressure	Vaisala PTB101B	600–1060 hPa	4 hPa
Air temperature	Vaisala HMP35AC	−80 to +56 °C	0.3 °C
Relative humidity	Vaisala HMP35AC	0–100 %	2–3 %
Wind speed	Young 05103	0–60 m s ^{−1}	0.3 ms ^{−1}
Wind direction	Young 05103	0–360 °	3 °
Pyranometer	Kipp and Zonen CNR1	305–2800 nm	10 % EADT
Pyrradiometer	Kipp and Zonen CNR1	5000–50 000 nm	10 % EADT
Sonic height ranger	Campbell SR50	0.5–10 m	0.01 m or 0.4 %

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Table 3. Comparison of specific surface mass balance and its components (mm w.e. yr^{-1}), and mean surface energy balance components (W m^{-2}) at each station. The values at each station are valid for the respective measurement period at this station. As explained in the text, BR01 was not applied to AWS 16.

	AWS 16	AWS 5	AWS 6	AWS 9
Surface mass balance (mm w.e. yr^{-1})	161	174	256	71
Surface sublimation (mm w.e. yr^{-1})	-7	-20	-17	-3
Melt (mm w.e. yr^{-1})	0	-2	0	0
Snowdrift sublimation (mm w.e. yr^{-1})	-11	-31	-25	-1
B98 (mm w.e. yr^{-1})	-10	-25	-19	0
BR01 (mm w.e. yr^{-1})	/	-31	-24	-1
DY01 (mm w.e. yr^{-1})	-11	-37	-31	-1
Residual processes (mm w.e. yr^{-1})	179	227	298	75
Net radiation (W m^{-2})	-31	-17	-23	-2
Sensible heat flux (W m^{-2})	33	20	26	2
Latent heat flux (W m^{-2})	-2	-3	-3	-1
Ground flux (W m^{-2})	-0.4	0.05	-0.1	-0.1
Melting energy (W m^{-2})	0	0.02	0	0

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Table 4. Surface mass balance and its components (all units mm w.e. yr⁻¹) at AWS 5 and 9 for 2009–2010.

	AWS 5	AWS 9
Surface mass balance	177	81
Surface sublimation	-17	-1
Melt	0	0
Snowdrift sublimation	-30	-1
Residual processes	225	83

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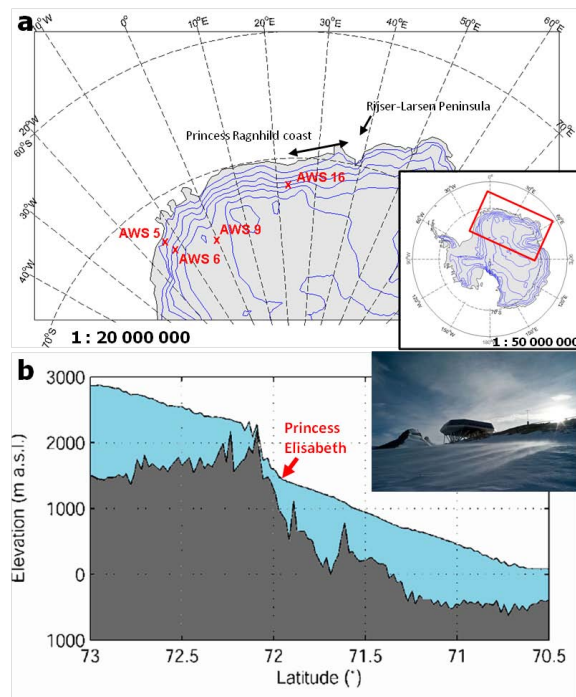


Fig. 1. (a) Map with focus on Dronning Maud Land, indicating the location of the 4 automatic weather stations (AWSs) under consideration. 500 m height contours are depicted in blue (<http://www.eos.ubc.ca/~rich/map.html>). (b) Surface (Bamber et al., 2009) and bedrock (Lythe and Vaughan, 2001) topographic transect along the 23° E meridian (adapted from Pattyn et al., 2009). The outcropping bedrock at $\sim 72.1^\circ$ S form the Sør Rondane mountains. Inset: snowdrift event at Princess Elisabeth, where AWS 16 is located.

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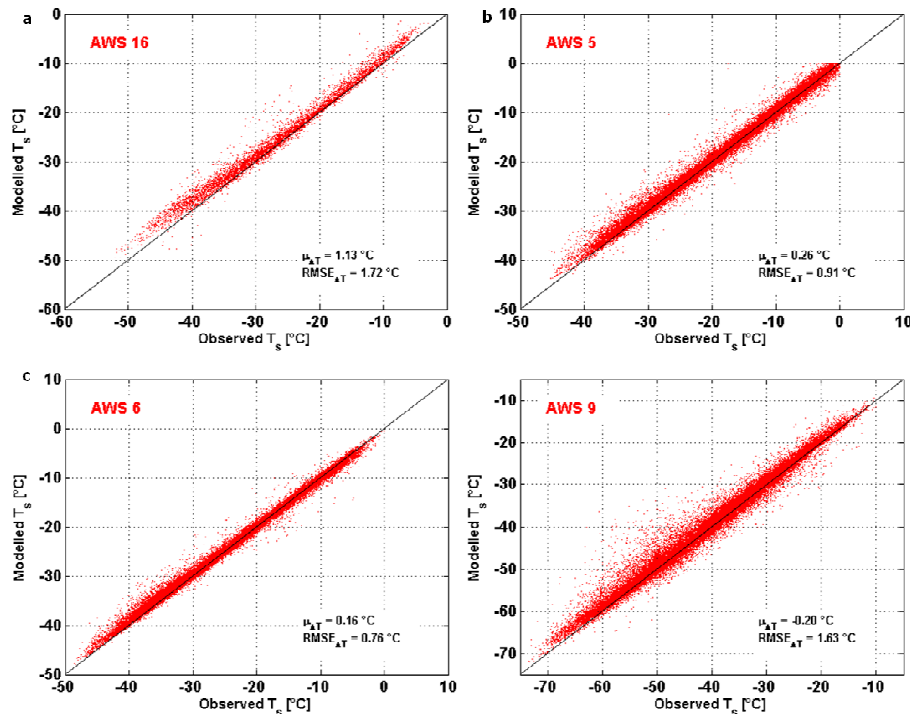


Fig. 2. Observed versus modelled surface temperature T_s (2 h averages), varying time periods, at **(a)** AWS 16 – Princess Elisabeth, **(b)** AWS 5 – Wasa/Aboa, **(c)** AWS 6 – Svea Cross, and **(d)** AWS 9 – Kohlen. Note the different axes' ranges. $\mu_{\Delta T}$ presents the mean surface temperature difference and $\text{RMSE}_{\Delta T}$ the Root Mean Squared Error of the surface temperature difference.

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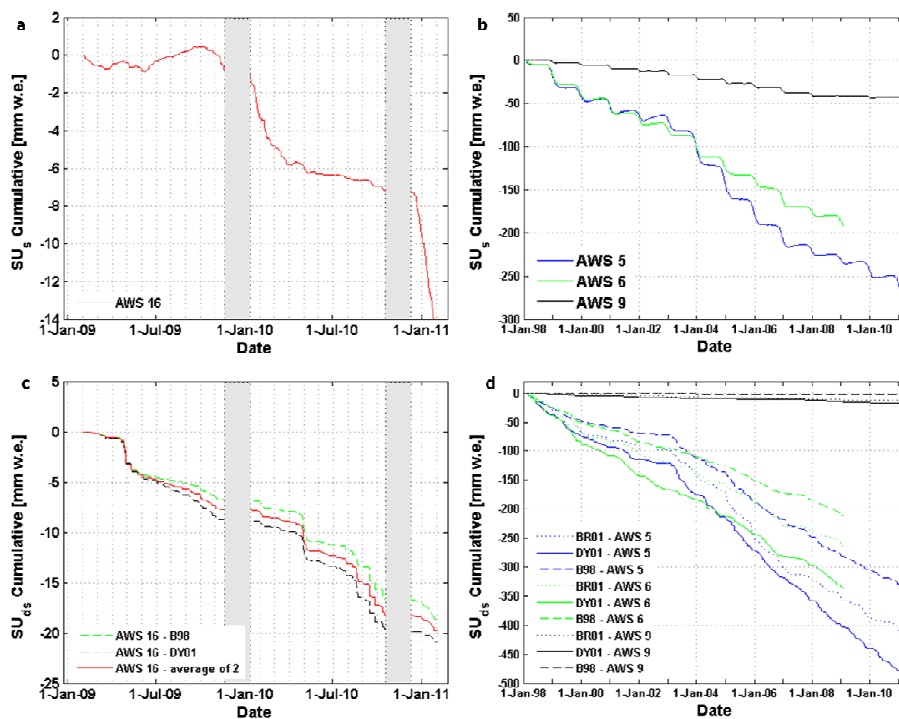


Fig. 3. Cumulative sublimation mass fluxes (mm w.e.): Surface sublimation SU_s at (a) AWS 16 and (b) AWS 5, 6 and 9; Snowdrift sublimation SU_{ds} at (c) AWS 16 and (d) AWS 5, 6 and 9.

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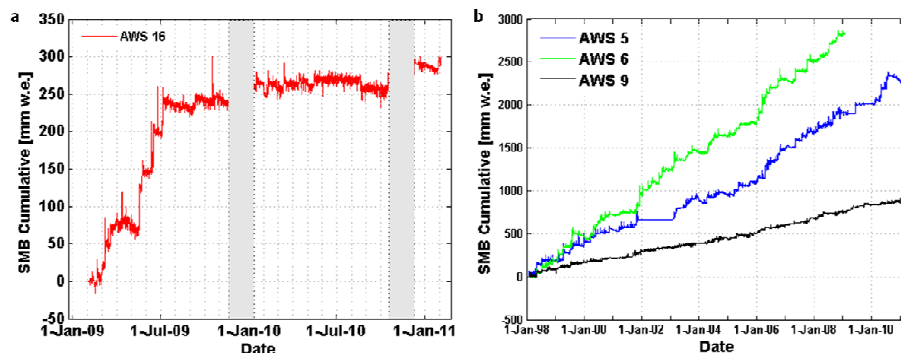


Fig. 4. Cumulative surface mass balance SMB (mm w.e.) for the respective measurement periods at (a) AWS 16 and (b) AWS 5, 6 and 9. Note that at AWS 16, accumulation during the December 2009 and November 2010 data gaps is captured by a jump in the snow height observations.

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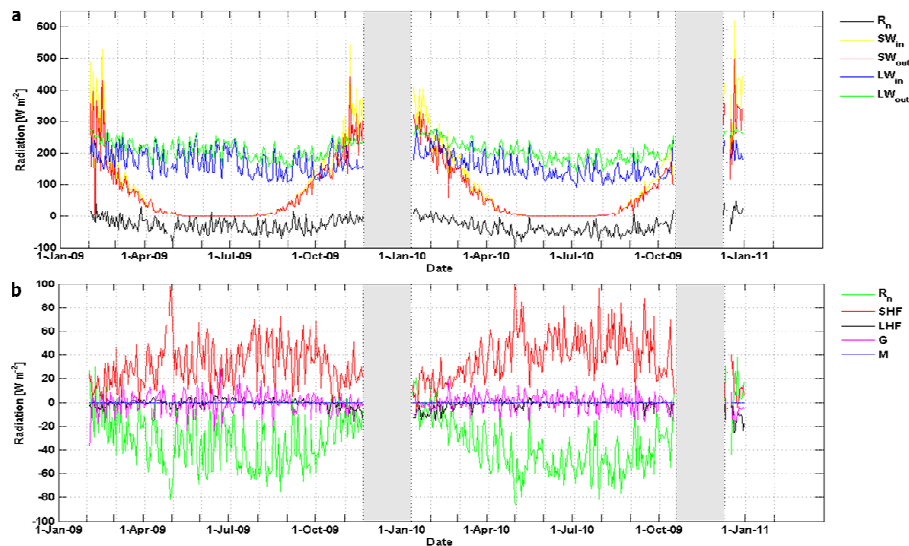


Fig. 5. Daily mean **(a)** net radiation R_n and its components, and **(b)** SEB components, 2009–2010, at AWS 16. Abbreviations stand for incoming (outgoing) shortwave (longwave) radiation (SW_{in} , SW_{out} , LW_{in} , LW_{out} , respectively), sensible heat flux (SHF), latent heat flux (LHF), sub-surface conductive flux (G) and melting energy (M).

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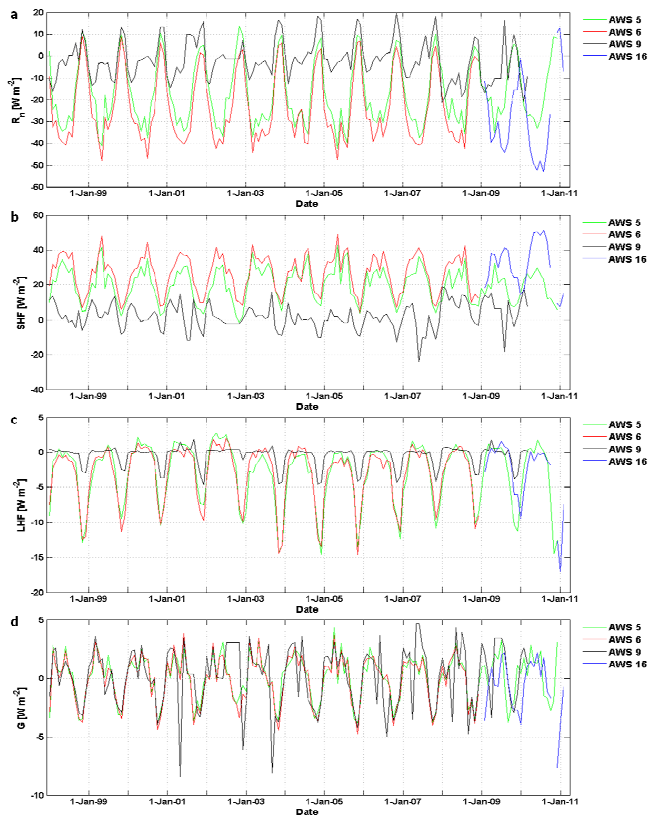



Fig. 6. Monthly mean SEB components (Wm^{-2}) at AWS 5, 6, 9 and 16, respective measurement periods: **(a)** net radiation (R_n), **(b)** sensible heat flux (SHF), **(c)** latent heat flux (LHF), **(d)** subsurface conductive heat flux (G). Not shown is melting energy (M), as this energy flux is absent at AWS 6, 9 and 16 and negligible at AWS 5.

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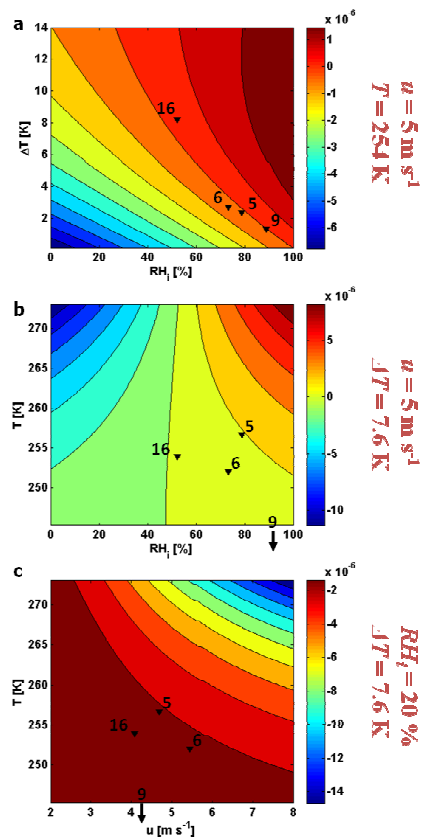


Fig. 7. SU_s rates ($\text{kg m}^{-2} \text{s}^{-1}$) from 3 sensitivity experiments, with **(a)** wind speed $u = 5 \text{ m s}^{-1}$ and air temperature $T = 254 \text{ K}$, **(b)** $u = 5 \text{ m s}^{-1}$ and surface temperature inversion $\Delta T = 7.6 \text{ K}$, and **(c)** relative humidity $RH_i = 20 \%$ and $\Delta T = 7.6 \text{ K}$. Mean values for RH_i , u and ΔT were calculated for AWS 5, 6, 9 and 16 for snowdrift-free conditions and used to plot each station.

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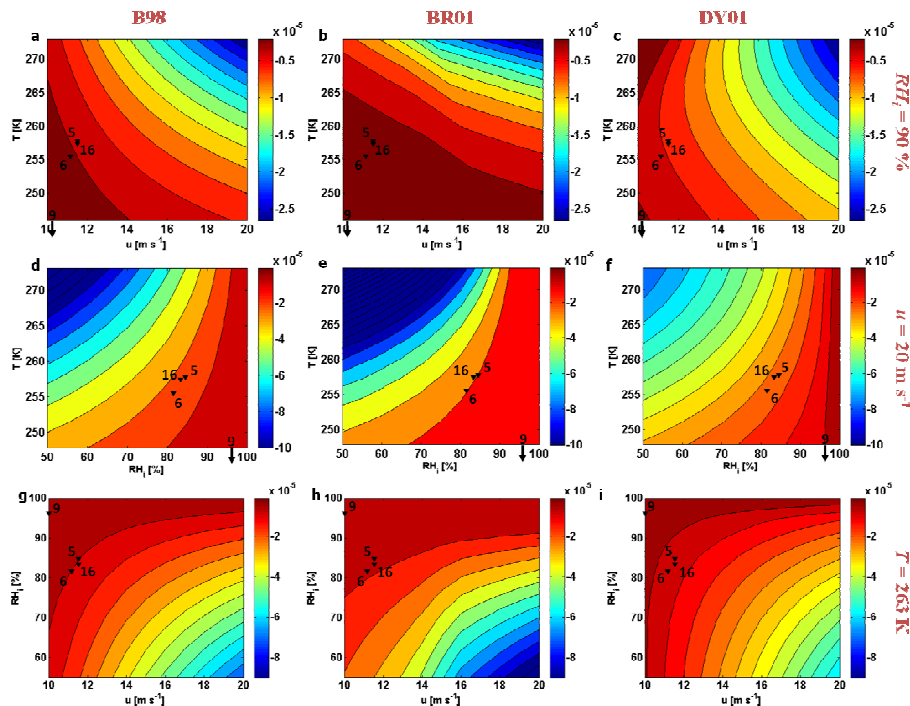


Fig. 8. SU_{ds} rates ($\text{kg m}^{-2} \text{s}^{-1}$) from 3 sensitivity experiments, with **(a–c)** $RH_i = 90\%$, **(d–f)** $u = 20 \text{ m s}^{-1}$, and **(g–i)** $T = 263 \text{ K}$, each one conducted for B98 (left panels), BR01 (central panels), and DY01 (right panels), respectively. Mean values for RH_i , u and T were calculated for AWS 5, 6, 9 and 16 for snowdrift conditions and used to plot each station.

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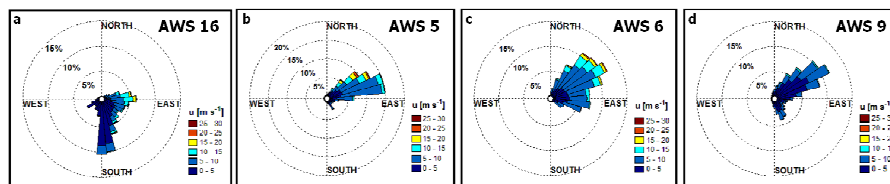


Fig. 9. Dominant wind directions measured at **(a)** AWS 16 – Princess Elisabeth, **(b)** AWS 5 – Wasa/Aboa, **(c)** AWS 6 – Svea Cross and **(d)** AWS 9 – Kohnen. The contribution of each wind speed class to a given wind direction is shown in the colour scale.

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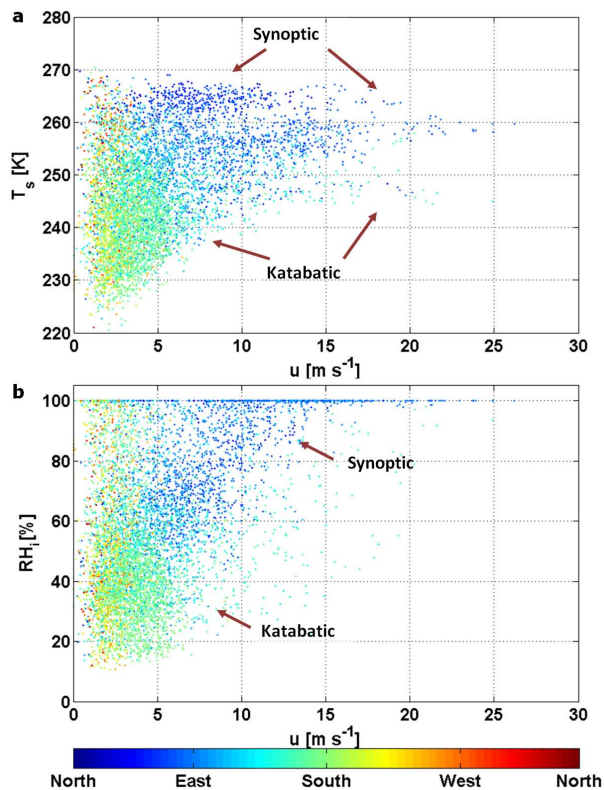


Fig. 10. Scatter plot of (a) surface temperature T_s and (b) relative humidity with respect to ice RH_i versus wind speed u . The colour of the scatter points represents the wind direction.

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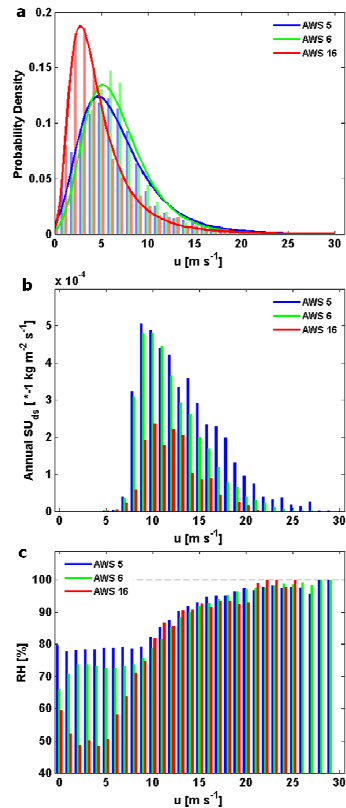


Fig. 11. (a) Histogram of 2-hourly mean wind velocity (u), with the Generalised Extreme Values (GEV) probability density function fitted to the data. From $u > 25 \text{ m s}^{-1}$ onwards, the three GEV curves overlay each other. Note that here the GEV distribution is preferred over the Weibull distribution, since R^2 is higher for the GEV fit ($0.97 < R^2 < 0.99$) than the Weibull fit ($0.91 < R^2 < 0.95$) at each location. **(b)** Total annual snowdrift sublimation (SU_{ds}) for each wind speed bin (width = 1 m s^{-1}). **(c)** mean RH_i for each wind speed bin (width = 1 m s^{-1}).